


## Lecture 6 : The extragalactic universe I

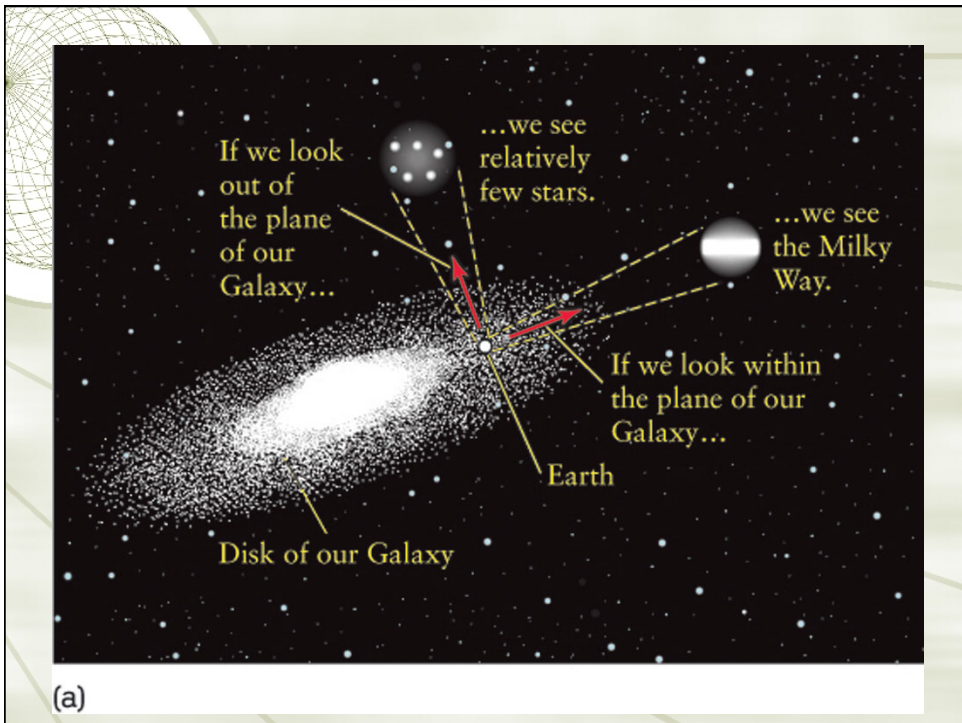
- ★ Our place in the Galaxy
- ★ The Great Debate
- ★ Measuring distance in astronomy
  - ✦ Parallax and Cepheids
- ★ Hubble's (first) great discovery

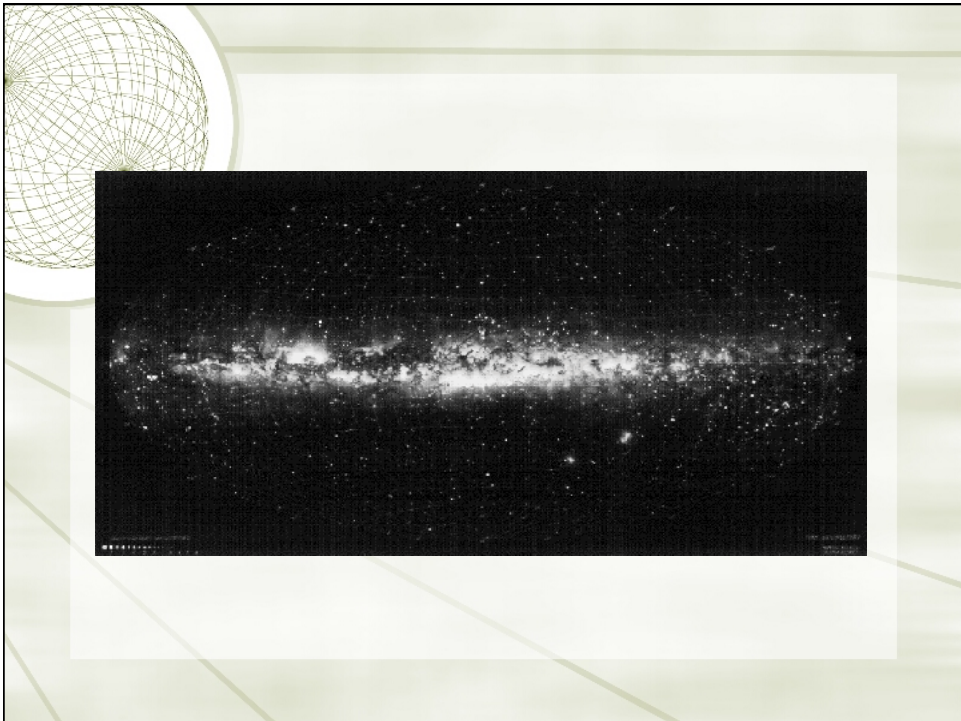
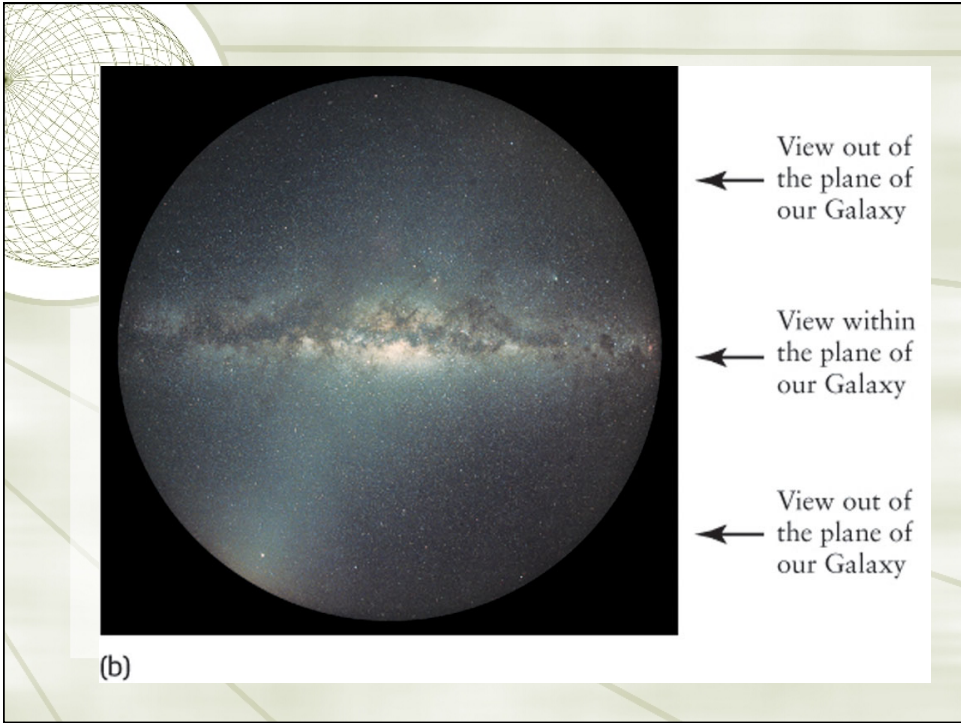
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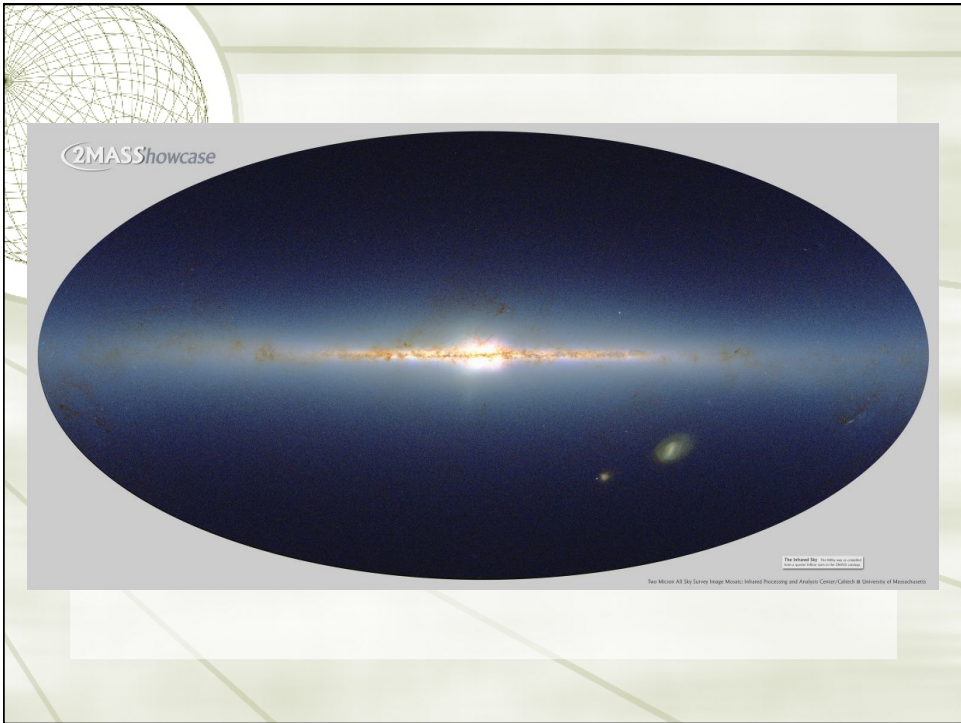
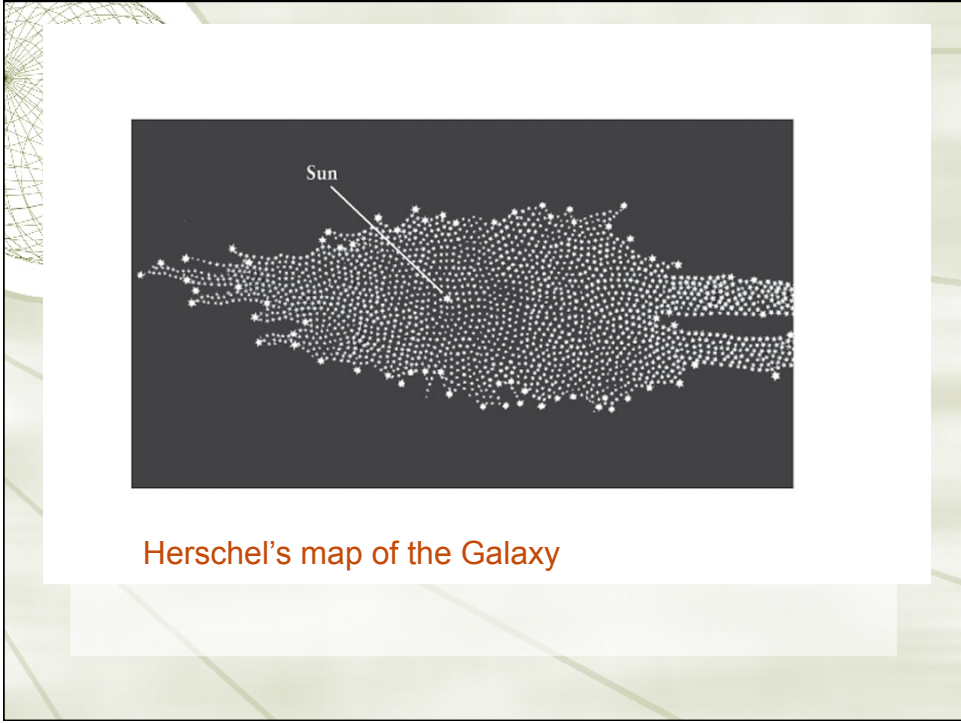
## I : Our place in the Galaxy

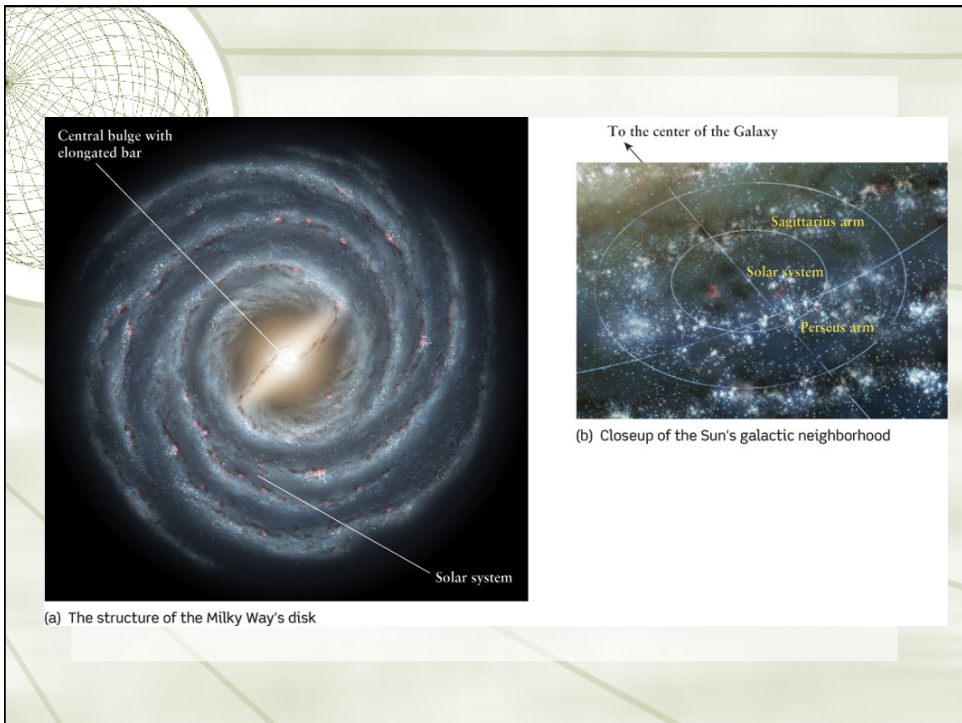
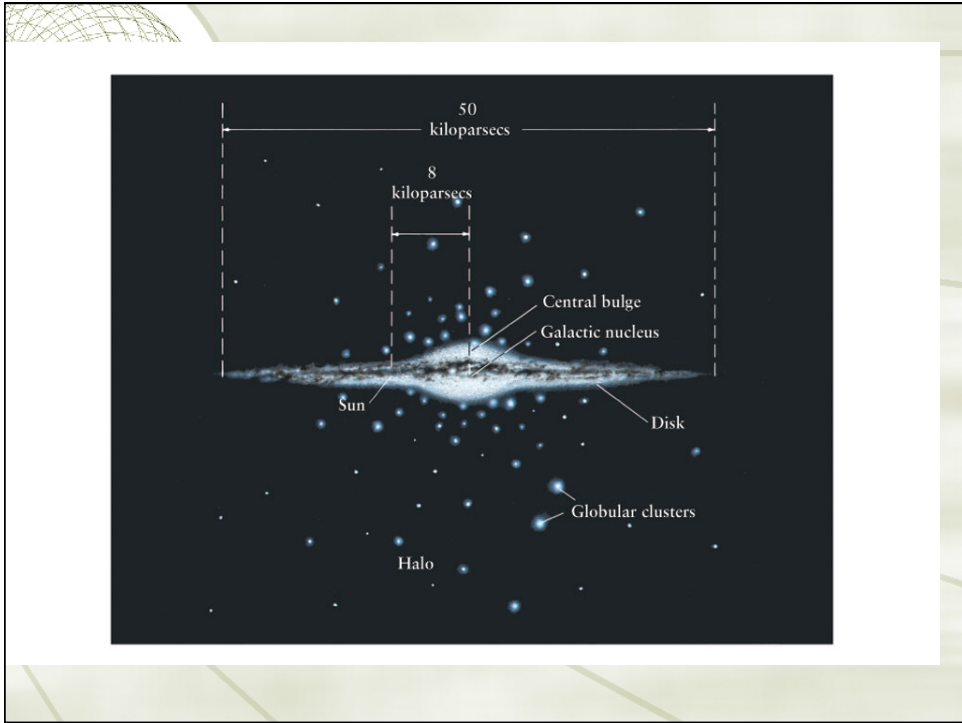
- ★ We live in a large disk galaxy
  - ✦ We live in the disk, towards the edge (8kpc-25,000lyr out)
  - ✦ Projected onto the sky, this disk of stars looks like a band of light that rings the sky... the Milky Way
- ★ This realization came somewhat slowly...
  - ✦ Disk-like nature of galaxy realized by Thomas Wright (1780); refined by Kant
  - ✦ First attempt to map out galaxy made by William Herschel (1785); refined by Kapteyn in 1920
  - ✦ Herschel came to the conclusion that we sit at the center of the Galactic disk. In fact, he was wrong... had not accounted absorption by dust!








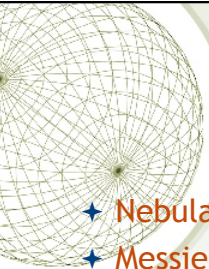






## II : The Great Debate... “what are spiral nebulae?”

- ★ Early 20th century...
  - ✦ Knew that we lived in a large disk galaxy
  - ✦ But what was the nature of the larger Universe?
  - ✦ Two opposing ideas:
    - ✦ Our galaxy is alone, sitting in the middle of otherwise empty space
    - ✦ Our galaxy is one of many galaxies that fill space (so-called “Island Universes”)
  - ✦ The debate rapidly focused on the nature of nebulae



- ★ Nebulae have been studied for ages
- ★ Messier (1780)
  - ✦ Systematically catalogued over 100 bright nebulae
  - ✦ Main reason for doing this was so that comet hunters knew which “fuzzy patches” to ignore!
  - ✦ But what were these nebulae? Possibilities:
    - ✦ Glowing patches of gas (e.g., Orion)
    - ✦ Clusters of many stars within our Galaxy (e.g., the Globular cluster M13)
    - ✦ Whole other galaxies!!
  - ✦ Of special interest were the “spiral nebulae” that showed Milky Way like spiral structure... the brightest spiral nebulae was the Andromeda nebula




*Messier actually cared about comets*



Credit : A. Dimai

This slide features a decorative wireframe sphere in the top-left corner. The main image shows a comet with a bright nucleus and a long, faint tail streaking across a dark night sky filled with stars. The foreground consists of dark, silhouetted mountains under a clear sky.

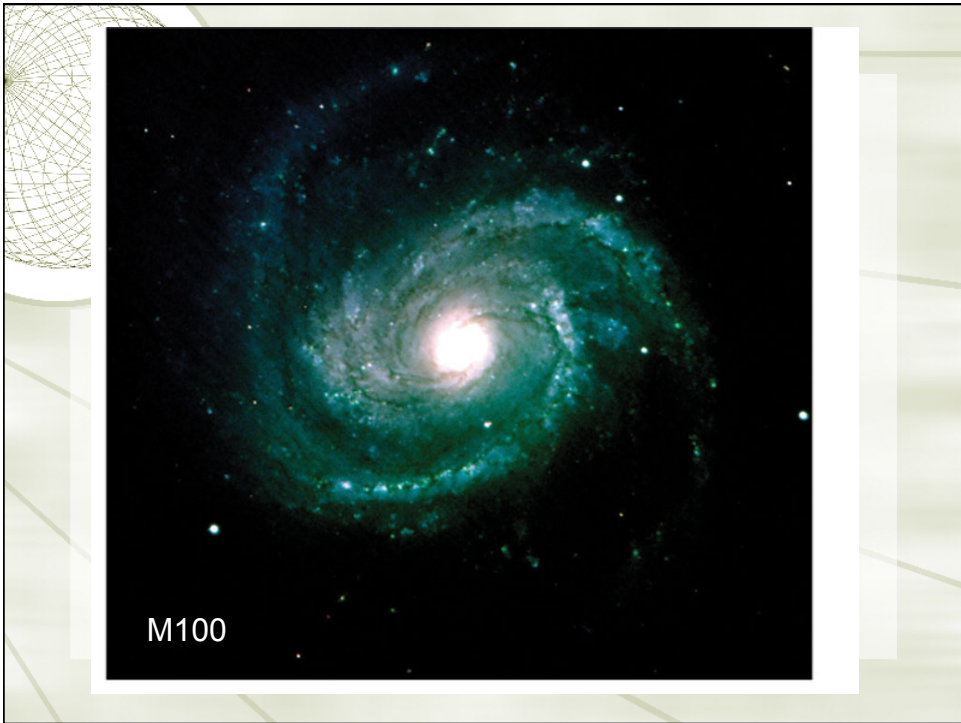
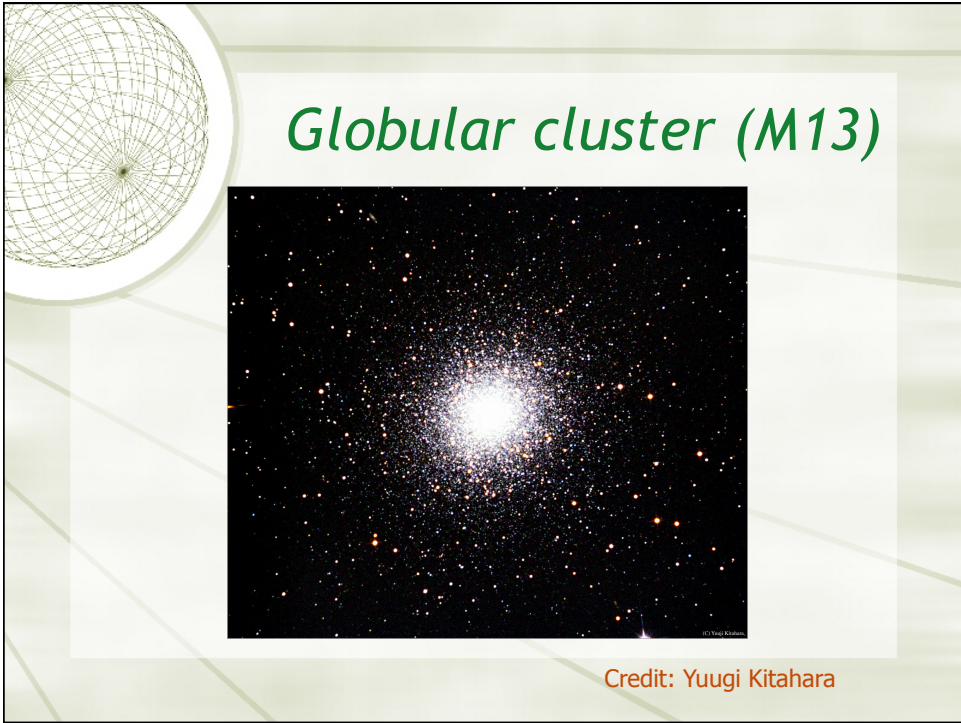
*The Orion Nebula (M42)*



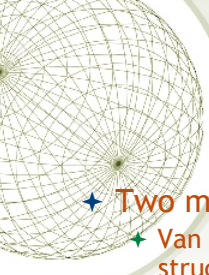
The Orion Nebula and Trapezium Cluster  
(VLT ANTU + ISAAC)

ESO PR Photo 03a/01 (15 January 2001) © European Southern Observatory

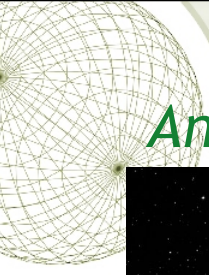
This slide also features a decorative wireframe sphere in the top-left corner. The main image is a detailed, multi-colored view of the Orion Nebula, showing intricate patterns of gas and dust in shades of blue, green, and red, with numerous bright stars scattered throughout. Below the image is a caption and credit information.








- ★ **Two misleading events...**
  - ★ Van Maanen reported observations showing that the spiral structure in the Andromeda nebula was rotating.
    - ★ If Andromeda was outside of our galaxy, the rotation would have to be much faster than speed of light!
    - ★ Argued for Andromeda being inside of our Galaxy
  - ★ 1885 : A “new star” appeared in Andromeda nebula
    - ★ This was interpreted as a stellar nova
    - ★ Scaling from known nearby novae, it was concluded that Andromeda lay will inside our Galaxy
    - ★ It wasn’t realized at the time that this was actually a supernova, not a regular nova.
  - ★ 1920 : The two points of view were debated by Shapley & Curtis in the “Great Debate”

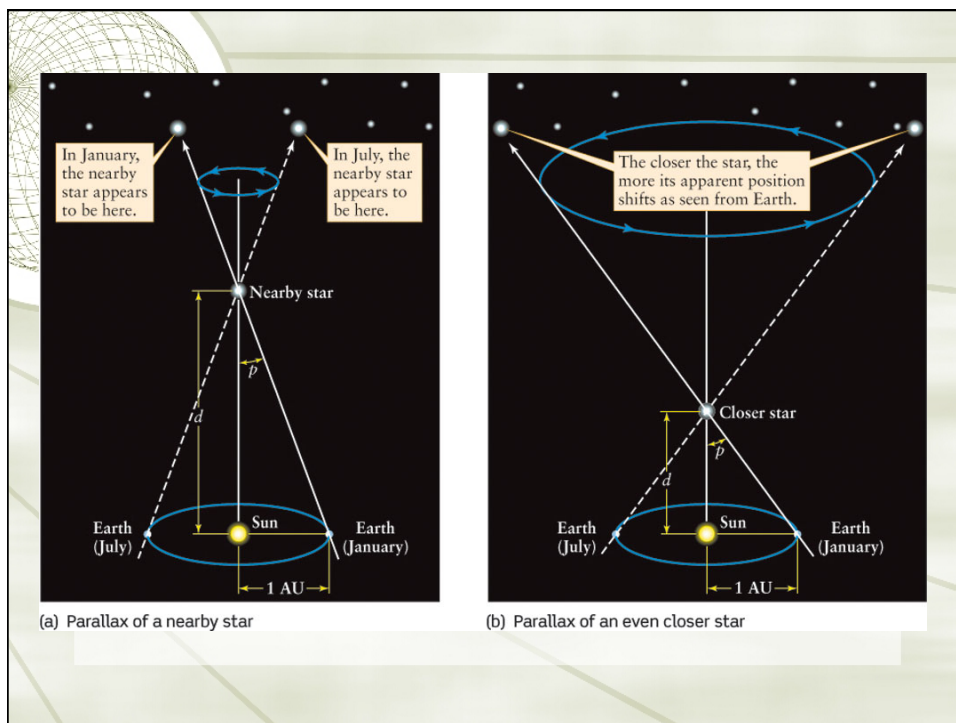


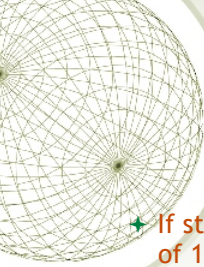
## Andromeda “Nebula” (M31)



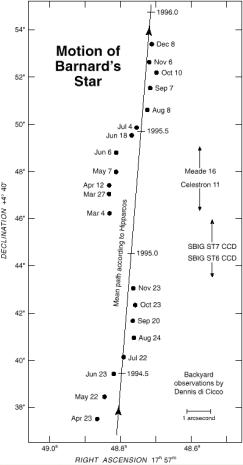
### III : Measuring distances in astronomy

- ★ The distance to any astronomical object is the most basic parameter we want to know
  - ★ Require knowledge of distance in order to calculate just about any other property of the object
  - ★ Distance is often difficult to determine!
- ★ Most direct method for measure distances to “nearby” stars uses an effect called parallax
  - ★ As Earth orbits Sun, we view a star along a slightly different line of sight
  - ★ This causes the star to **appear** to move slightly with respect to much more distant stars
  - ★ We can currently use this technique to measure stellar distances out to ~3000 light years from Earth





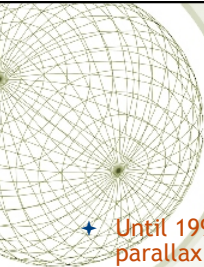
- ✦ If star wobbles with amplitude of 1 arc-second, then it is at distance of 1 parsec (definition of parsec).
- ✦ 1pc = 3.26 lt-yr
- ✦ In general,

$$D(\text{pc}) = \frac{1}{\theta_{\text{wobble}} (\text{arcsec})}$$


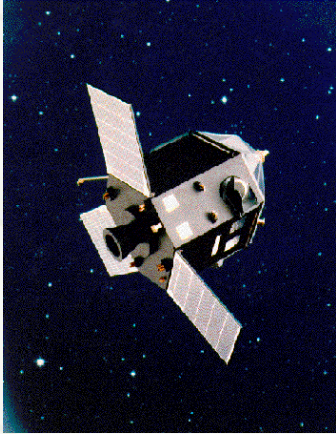
**Motion of Barnard's Star**

The plot shows the star's path in the sky from 1945 to 1996. The y-axis is Declination +4' 27" (38 to 54) and the x-axis is Right Ascension 17' 52" (48.0 to 48.6). A dashed line represents the mean path. Data points are labeled with dates and years: Apr 23 (1945), May 22 (1945), Jun 23 (1945), Jul 22 (1945), Aug 24 (1945), Sep 20 (1945), Oct 23 (1945), Nov 23 (1945), Dec 8 (1945), Jan 4 (1946), Feb 4 (1946), Mar 4 (1946), Apr 12 (1946), May 7 (1946), Jun 6 (1946), Jun 18 (1946), Jul 4 (1946), Aug 8 (1946), Sep 7 (1946), Oct 10 (1946), Nov 6 (1946), Dec 8 (1946). Other labels include 'Meade 16', 'Celestron 11', 'SBIG ST7 CCD', 'SBIG ST6 CCD', and 'Backyard observations by Dennis & Cicco'. A 1 arcsecond scale bar is shown at the bottom right.

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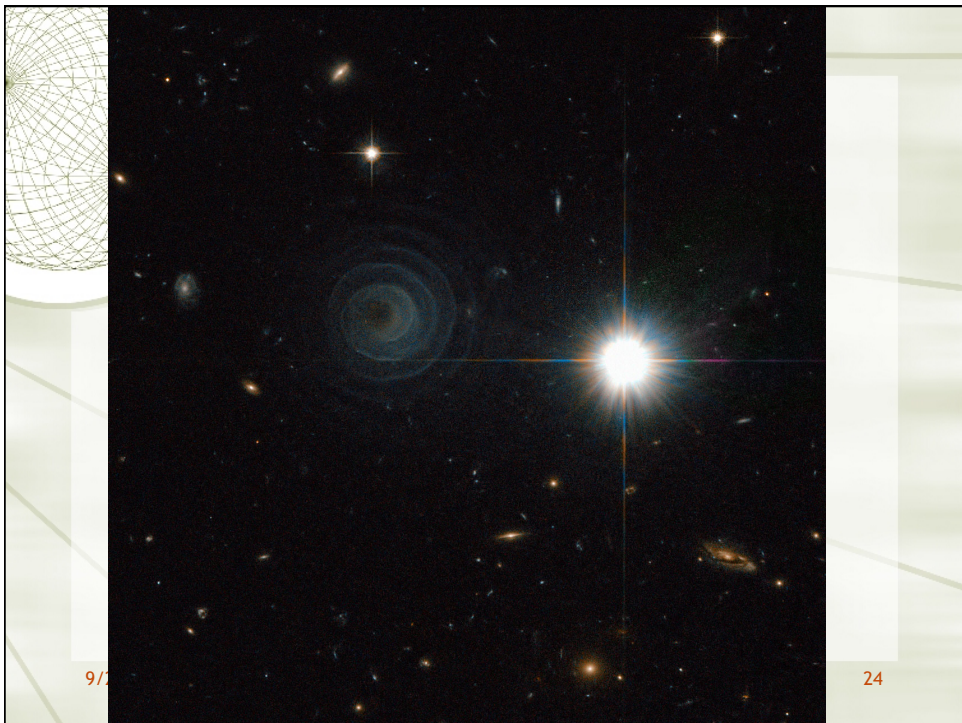
- ✦ Until 1990s, could only detect parallax out to 50pc.
- ✦ Hipparcos satellite
  - ✦ Designed to measure parallax of stars
  - ✦ Can detect wobble out to distance of about 1kpc (1000pc)
  - ✦ Used to map out locations of nearby stars.
- ✦ GAIA satellite
  - ✦ Due to launch 2012
  - ✦ Can map out positions and motions of stars across the whole galaxy!!

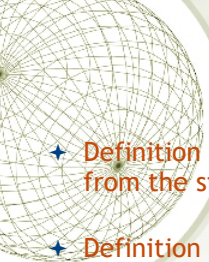


**Hipparcos (ESA)**

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## Beyond parallax?

- ✦ Definition : The **observed flux** of a star is the energy received from the star per unit time per unit area.
- ✦ Definition : The **luminosity** of a star is the energy per unit time (i.e. power) emitted by the star
- ✦ If the star is at distance D and emits equally in all directions (i.e. it emits isotropically), then the observed flux F and luminosity L are related by

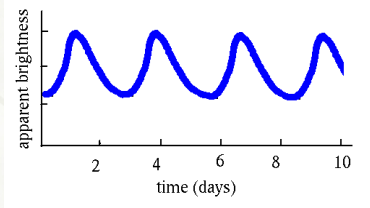
$$L = 4\pi D^2 F \quad \text{or} \quad F = \frac{L}{4\pi D^2}$$

- ✦ Suppose we know the luminosity of some object... then we can use its measured flux to determine the distance! Objects with known luminosities are called **standard candles**.



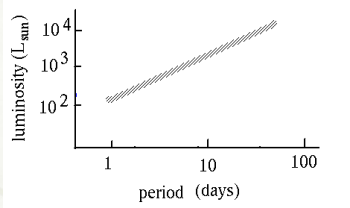
## Cepheid variables

- ✦ Henrietta Leavitt discovered (1912) that a certain class of variable stars called Cepheids had properties that meant they could be used as standard candles
  - ✦ She studied Cepheids that are close enough for parallax to be measured... found that the luminosity is related to the period of fluctuations in brightness
  - ✦ So, if you measure the period of a Cepheid, you can determine its luminosity. Measuring flux then gives you distance, even if its too far for parallax!



apparent brightness

time (days)

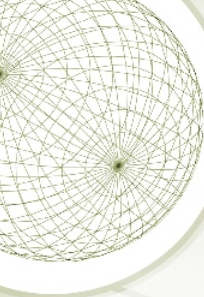


luminosity ( $L_{\text{sun}}$ )


period (days)

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graphics: University of Oregon Astronomy Dept
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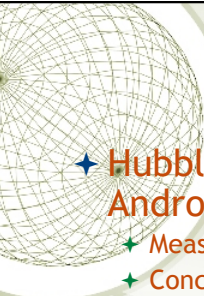


## Edwin Hubble



Hale Observatories, courtesy AIP

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## Hubble's observations

- ◆ Hubble found Cepheid Variables in Andromeda
  - ◆ Measured period and flux, and hence distance
  - ◆ Concluded that Andromeda must be well outside of the Milky Way Galaxy
  - ◆ Thus, the Great Debate was settled... the MW is just one of many many many galaxies
- ◆ Modern measurements
  - ◆ Distance to Andromeda 2 million lyr
  - ◆ About 20x MW diameter

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