Research with Jian-Yang Li

WHAT DO I (WE) DO IN PLANETARY ASTRONOMY

Planetary Astronomy

- Study planetary system(s)
 - Solar system formation and evolution
 - Other planetary systems



Solar system small bodies

- Building blocks of planetary system
- Vehicles of volatiles and organics
- Potential impact hazardous



Impact hazardous





Comets







Comets



Photo credit: Hannes Pieterse, Astronomical Society - Bloemfontein © Hannes Pieterse, Bloemfontein, South Africa e-mail: pieterse@polka.co.za Comet C/2006 P1 McNaught 20 km West from Bloemfontein. Friday, 19/01/2007; 20:39. 30 Second exposure.

Comets



NASA Discovery Missions

Eleven Discovery Missions have been chosen to date. > NEAR > Mars Pathfinder > ASPERA - 3 > MOON MINERAL On February 17, 1996, Near Farth Asteroid The Mars Pathfinder mission demonstrated a low-cost Moon Mineralogy Moper, or M3, will fly aboard ndia's to deep Oce mission, Chandrayaan-1, and pper, or M3, will fly aboard ASPERA-3 (Analyzer of Space Plasma and Rennezvou NEAR as the inst Discovery Program sparecraft to be launched and in became the first ever method of delivering a set of science instruments and Energetic Atoms), an instrument flying on the Mars the first free-ranging rover to the surface of Mars. will characterize and map the mineral composition of the Express spacecraft, is studying the interaction between it and land on an asteroid 5000 2 More on Pathfinder.. the solar wind and the Martian atmosphere. oon at high resolution. More on NEAR... More on ASPERA-3... More on M3... > Lunar Prospector > Stardust > Stardust-NExT The science data returned from Lunar Prospector has The Stardust spacecraft collected interstellar dust and EPOXI is an extension of the Deep Impact ardust New Exist ration of Tempel 1, or enabled scientists to create detailed maps of the comet dust during a close encounter with Com exT, prison will se the Stardust spaceraft to revisit will search for planets around other stars and (gravity, magnetic properties and chemical composition and returned the particles to Earth for analysis by new comet. S 16 of the Moon's entire surface. scientists worldwide. A. nission. More on Prospector... More on Star ore on EPOXI.. ore on Stardust-NExT... > Genesis CONTO > Strofio The ometocleus our, or CONTOUR, mission was going to encounter and study two diverse comets The Genesis spacecraft spent two years collecting Strofio will use a unique mass spectrometer to study atoms of solar wind before returning them to Earth in ac the atoms and molecules that compose Mercury's September 2004 for detailed analysis during their periodic visits to the solar system. atmosphere to reveal the composition of the planet's More on Genesis More on CONTOUR. surface. More on Strofio... MESSENGER Пеер н MESSENGER (MErcury Surface, Space ENvironm The Deep Impact mission propelled a large copp GEochemistry, and Ranging) mission is a scientific projectile into the surface of a comet, creating investigation of the planet Mercury crater and yielding new information about composition and structure of More on MESSENGER. Led by > DAWN > Kepler The Drwn mission is on its Ny to the oldest and most Kepler will use a unique telescope to search for Earthmassive astocids in cosolar sistem, to give us a glinipse of conditions and processes at the dawn of our size habitable planets around stars beyond our solar system. - Elet systen More on Dawn... More on Kepler... > GRAIL We are The Gravity Recovery and Interior Laboratory, or GRAIL, mission will put twin satellites into orbit around the Moon to map the Moon's interior and variations in its gravitational pull & to reconstruct its thermal history. More on GRAIL

Five Discovery Missions of Opportunity have been selected.

EPOXI

Epoch

- Drake Deming (last week?)
- - Jessica Sunshine (next week?)





Dawn Mission

Photographs copyright Randy Pollock (2007) Stiching copyright Johan Kiviniemi (2007)

Dawn Mission



•Double-rendezvous: Vesta and Ceres



- Science Payload:
 - Framing Camera (visible)
 - Visible Infrared Mapping
 Spectrometer (0.25 5 µm)
 - Gamma Ray and Neutron
 Spectrometer

Dawn will characterize the surfaces of two complementary protoplanets, Vesta and Ceres, and probe their internal structure

- ♦ Map the geologic units
- Create detailed shape models
- Determine how and when the bodies formed
- Understand the internal and external forces that shaped them

Mission Itinerary



Vesta and Ceres



Vesta and Ceres

- Oceres is wet
 - Show evidence of compositional water (25%)
 - Phyllosilicates, carbonates, magnesium hydroxide
 - Water has played an important role in its evolution
- Vesta is dry
 - Undergone intense heating, driven out all volatiles
 - Differentiation, basaltic surface
 - Source of HED meteorites
- Two objects have distinctively different formation environments and paths
 - Ceres more like outer solar system small bodies and icy bodies
 - Vesta more like inner solar system terrestrial planet – the smallest terrestrial planet









Potential projects

Phase function of cometary nuclei

- Separation of cometary nuclei and coma
- Construct and model phase function
- Vesta surface mapping
 - Develop general mapping tool (not from scratch)
 - Map Vesta with low-resolution Dawn data
- Vesta rotation
 - Precisely measure the rotation of Vesta using Dawn data
- High-contrast imaging on asteroids
 - Search for companions
 - New for extended sources

Photometry of cometary nuclei

Separation of nucleus from coma



Photometry of cometary nuclei

Construct and model the phase function



Vesta mapping

Output Photometric properties of Vesta

- How reflectance changes with illumination and viewing geometry
- Fundamental properties of the surface, albedo, particle impurity, roughness
- Photometric mapping (albedo/color, roughness)
- Mission support
 - Photometric correction for images and spectra
 - Data archive

Naked-eye Asteroid



Will be 5.6 mag in Aug 2011 during Dawn's arrival in the constellation of Capricornus

HST/WFPC2 2007





http://hubblesite.org/newscenter/archive/releases/2007/27/

HST/WFC3 2010



http://hubblesite.org/newscenter/archive/releases/2010/33/image/

Current maps



Vesta rotation

Revised recently from all available data (Li et al., Icarus in press)

- Limited by spatial resolution
- Uncertainty 1.5° in Dec and 3° in RA
- Statistically significant
- Precise measurement from Dawn
 - Pixel scale down to 25 m (0.006° on Vesta)
 - One year time baseline
 - Precisely measures the pole and potentially wobble (nodding)
 - Rotation evolution
- Data available from June 2011

High-contrast imaging

Adaptive optics

- Decrease the size of point spread function (PSF)
- Put diffraction pattern back into the central lobe
- Stable PSF
- Coronagraph
 - Block the light from the center
 - Suppress speckles
- Post-processing
 - Remove speckles
 - Enhance contrast

Technique for direct imaging of exoplanets



HR 8799 (Marois et al., 2008)

Adaptive optics + coronagraph





- Use a disk to block the light from center source
- Use funny shapes and transparency patterns to suppress diffraction patterns
 - Band-limited mask
 - Phase mask
 - Optical vortex mask

Post-processing

Direct speckle subtractionAngular differential imaging



High-contrast imaging for asteroids

Only the brightest ones
Large angular size (!)
Irregular shape
Fast moving companions



Eugenia – Two known satellites

Petit-Prince

- Discovered in 1998 by CFHT AO, first time by groundbased telescope
- Diameter 13 km
- Semi-major axis 1184 km, or 11 radii of Eugenia
- S/2004 (45) 1
 - Discovered in 2004 by VLT
 - Diameter 6 km
 - Semi-major axis ~700 km(?)





Merline et al. (1998)

Eugenia

- V=12.5 mag, Ang. Diam=0.130"
- No calibration star was taken
- Remove speckle using Adeona (V=11.8, Ang. Diam=0.126"



Metis – occulted



Metis

Calibration star

Convolve PSF with target disk and subtract



Contrast ratio and limiting size



Contrast ratio map

Limiting size map

Limiting size



Ks-band