# Homework #10 Solutions

#### Problem # 1 (11.3)

t\_/tH=15/17=.88>2/3

From table 11.1 the hyperbolic k=1 model will be the acceptable standard model.

The addition of a negative(attractive)  $\Lambda$  results always in recollapse (see fig. 11.6). A positive  $\Lambda$  will not affect the hyperbolic model, except in making them expand exponentially.

### Problem # 2(11.4)

 $z=(R_{now}/R_{then})-1=4$ . The Universe was 1/5 compared to now.

The lookback time is 7.37 billion years

The factor  $2/3H_o$  will be replaced a factor smaller than that in spherical model (see table 11.1) resulting in smaller lookback time

## Problem # 3 (11.5)

 $\Omega = 1 + kc^2/H^2R^2$ . It varies inversely proportionally to  $H^2$ .

For the case where  $\Omega=1$ , the density parameter is equal to 1 at all times, k=0 and the Universe is flat.  $\Omega$  is independent of the value of the Hubble constant.

## Problem 4 (11.3)

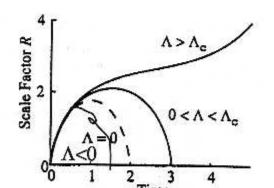


Fig. 11.10 The Lemaître model has spherical geometry and a  $\Lambda$  value slightly greater than the Einstein critical value  $\Lambda_c$ . This model features a hovering period, during which the scale factor remains nearly constant over a lengthy time interval. Following the hovering period, expansion continues at an accelerating rate with q < 0.