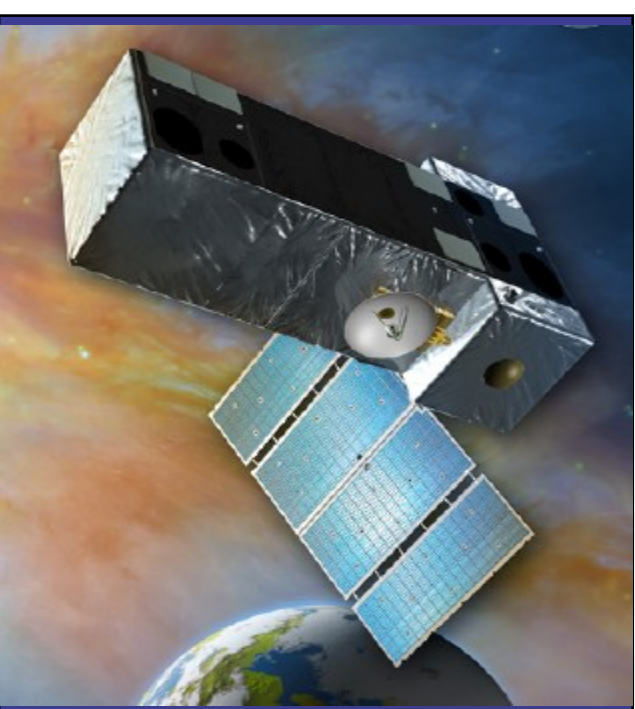


Searching for Solar System Analogs with SIM Lite

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SIM Lite ASTROMETRIC OBSERVATORY

• **Solar System Analogs, SOASAs,** have a “Jupiter” &/or “Saturn”

• To determine orbit:
 $P_{\text{ORBIT}} < \sim 2 \times T_{\text{MISSION}}$

• All outer planets have
 $P_{\text{ORBIT}} > 2 \times T_{\text{MISSION}}$

Planet	a [AU]	Period [yr]	Mass [M_{JUP}]
Jupiter	5.2	11.9	318
Saturn	9.5	29.4	95
Uranus	19.2	84.0	15
Neptune	30.1	164.0	17

• **SOSAs & Heavy SOSAs**

SOSAs: $P \in [11.9, 165]$ yr
 $M \in [0.05, 1] M_{\text{JUP}}$

HOSAs: $P \in [11.9, 165]$ yr
 $M \in [1, 13] M_{\text{JUP}}$

• Fraction of Plan. Systems:
 [Tabachnik & Tremaine (2002) or Cumming et al (2008)]

– **SOSAs: 13 %**

– **HOSAs: (17 ± 3)%**

• **8% of Sun-like stars**

Migration Theories predict

• Inward migration (known “RV” planets)

• Outward migration (Uranus & Neptune)

– Outer edge: 50-100 AU (350 – 1,000 yr) [Ida & Lin, 2004]

• Predict massive long-period planets

• requires massive disks

– Without migration: 30-40 AU (165-250 yr)

– **MUCH, MUCH, MUCH longer than $2T_{\text{MISSION}}$**

Expected values for size, velocity & acceleration

$$a_0 = 95/d_{10\text{pc}} (P^{+2} M_{\text{TOT}}^{-2})^{1/3} M_{\text{C,J}} [\mu\text{as}]$$

$$|\mu| = 600/d_{10\text{pc}} (P^{-1} M_{\text{TOT}}^{-2})^{1/3} M_{\text{C,J}} [\mu\text{as/yr}]$$

$$|d\mu/dt| = 3800/d_{10\text{pc}} (P^{-4} M_{\text{TOT}}^{-2})^{1/3} M_{\text{C,J}} [\mu\text{as/yr}^2]$$

10	$M_{\text{JUPITER @}}$	20	pc		
Period [yr]	a_0 [μas]	$ \mu $ [$\mu\text{as/yr}$]	$ d\mu/dt $ [$\mu\text{as/yr}^2$]		Comment
10	2,191	1,376	864.9		5 yr; Std. Orbit Fitting
20	3,478	1,093	343.2		
40	5,521	867	136.2		
80	8,763	688	54.1		
160	13,911	546	21.5		GAIA 5yr ($3\text{-}\sigma_{\text{acc}} = 14.5 \mu\text{as/yr}^2$)
320	22,082	434	8.5		SIM 5yr ($3\text{-}\sigma_{\text{acc}} = 7.5 \mu\text{as/yr}^2$)
640	35,053	344	3.4		
1,280	55,644	273	1.3		S+G 5yr ($3\text{-}\sigma_{\text{acc}} = 1.5 \mu\text{as/yr}^2$)

Standard method:

Proper-motion difference between epochs

SIM vs. Hipparcos?? $3 \times \epsilon_{\mu}(\text{HIP}) \sim 3 \text{ mas/yr} \implies \text{⊗}$

Next-Gen Astrometry missions:

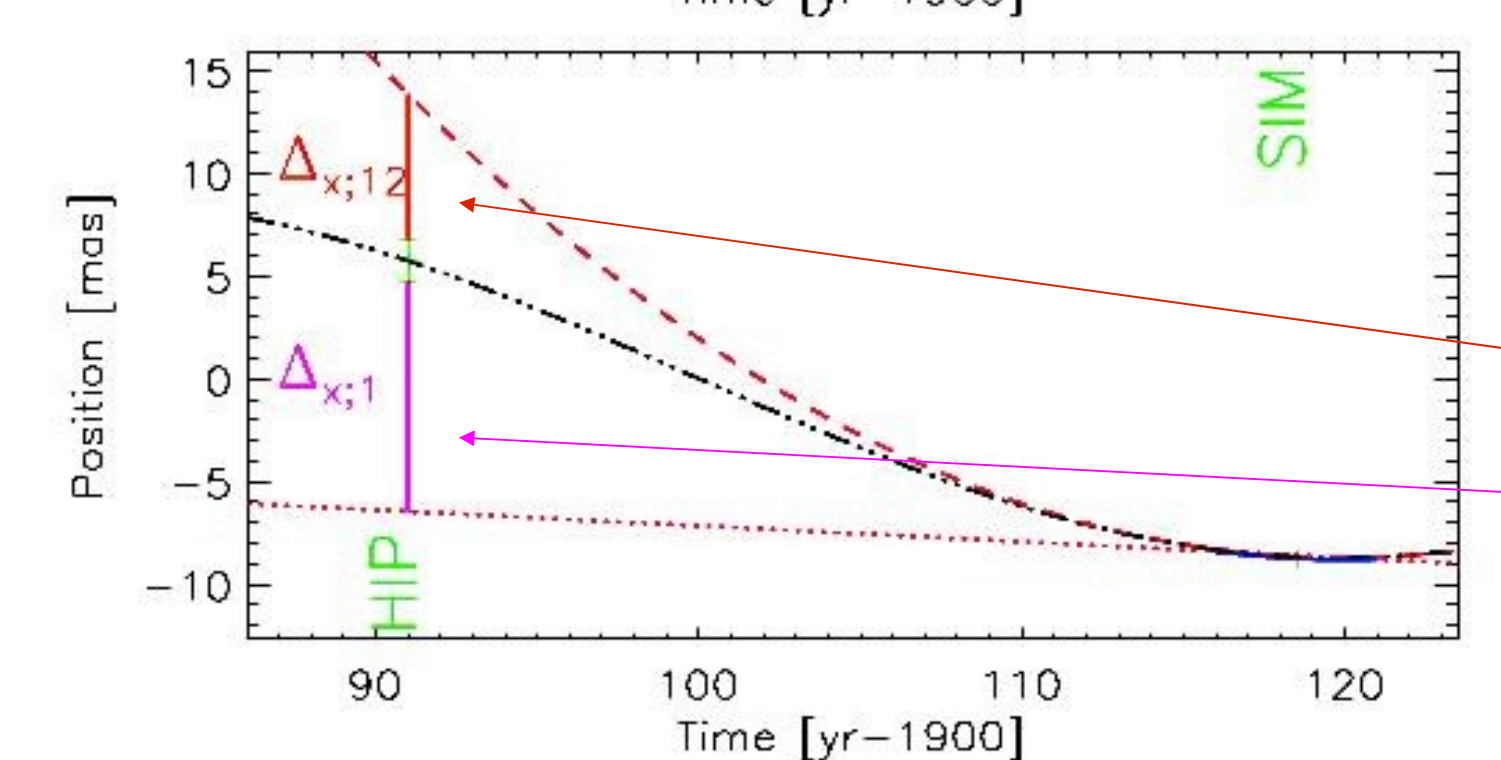
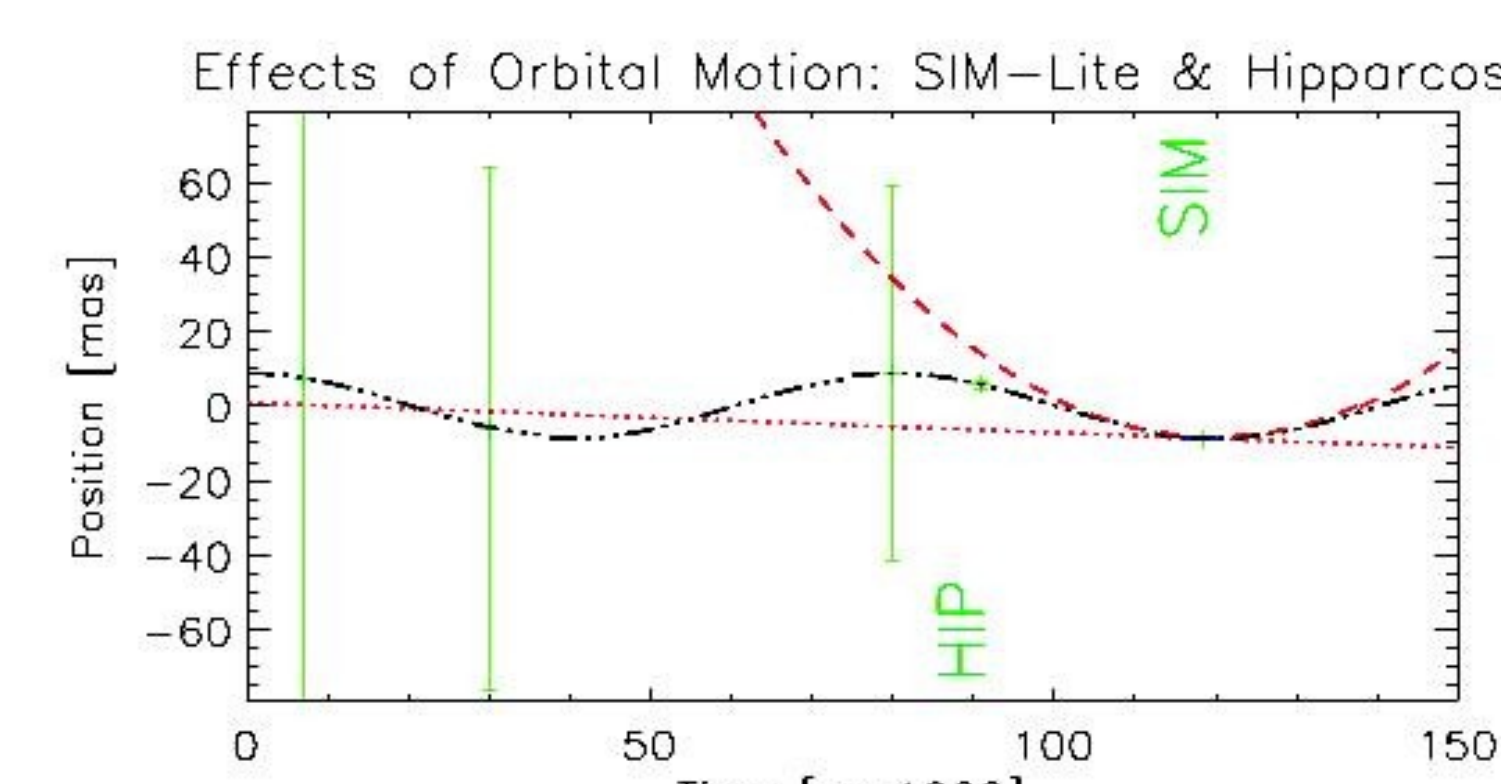
Detection of acceleration down to:

~160 yr: GAIA only

~320 yr: SIM only

~1,280 yr: GAIA + SIM data

$M=10 M_{\text{JUP}}, P=80 \text{ yr}, D=20 \text{ pc}, a_0=8.8 \text{ mas}, \mu_{\text{ORBIT}}=0.69 \text{ mas/yr}$



Our Approach:

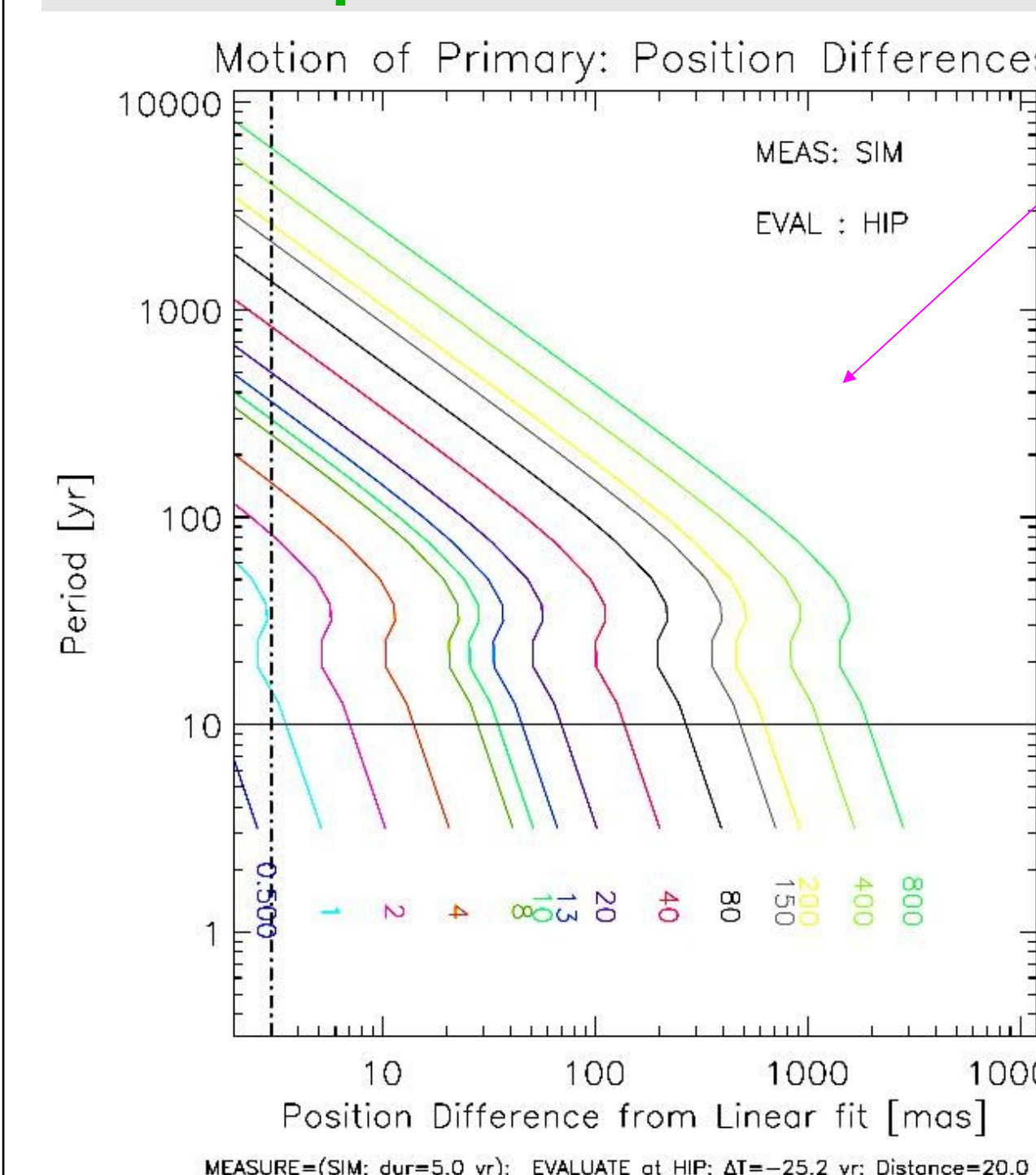
Extrapolate SIM astrometry to earlier (Hipparcos) epoch and compare positions [Olling, 2007]

Difference between: backtrapolations:

Quadratic: $\Delta_{x,12}$
 Linear: $\Delta_{x,1}$

tells Period & Mass

Extrapolate SIM astrometric model to HIPPARCOS position



LINER FITS:
 Position Differences are degenerate:
 multiple Masses & Periods at given pos.dif

QUADRATIC FITS:
 Lift Degeneracy
 Analytically proven (face-on, circular)
 PD depends only on orbital motion:
 center-of-mass motion drops out

