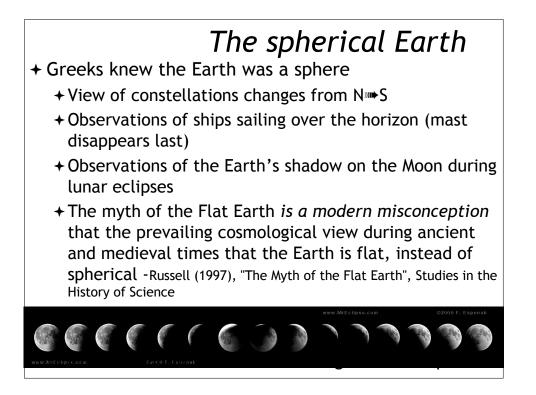
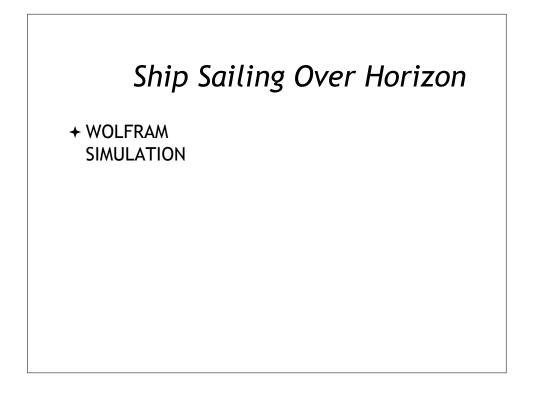
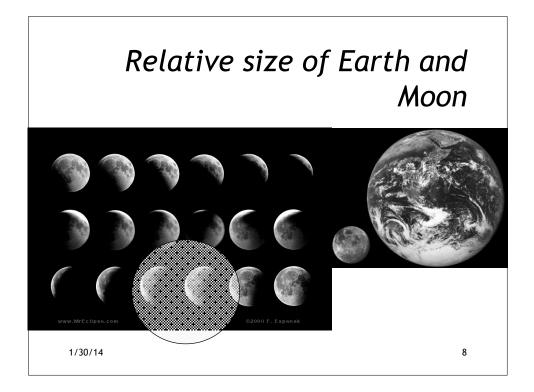


# Line constructionState in the second problem in the se







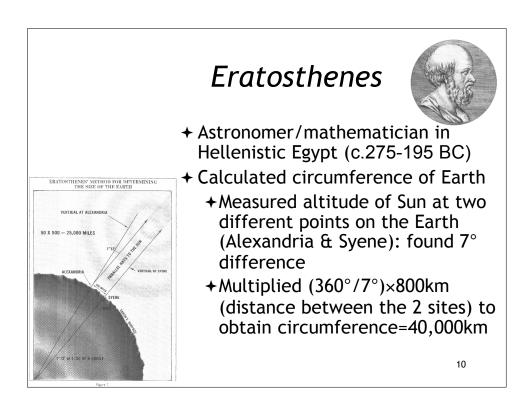
### Aristarchus of Samos (310-230 B.C.)

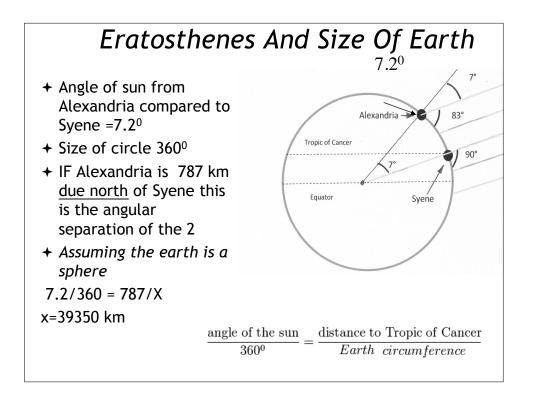
- + Using eclipse data and geometry:
  - + Measured relative sizes of Earth, Moon
  - + Measured distance to Moon (how\*)
- + Attempted to measure distance to Sun
  - + Need to measure (using time interval ratios) the angle of Sun when Moon is exactly at 1st or 3rd quarter
  - + Then use trigonometry and <u>known</u> Earth-Moon distance to get Sun's distance
  - + Very difficult measurement... He deduced that Sun is 20 times further from Earth than Moon... actual answer is that Sun is 400 times further.

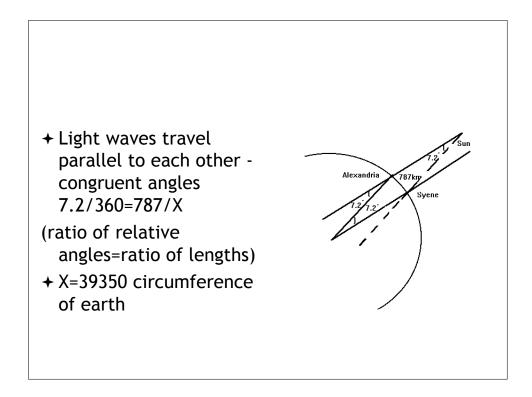
9

- + First to propose a heliocentric model!
- \*Aristarchus around 270 BC derived the Moon's distance from the duration of a lunar eclipse (next slide)

1/30/14







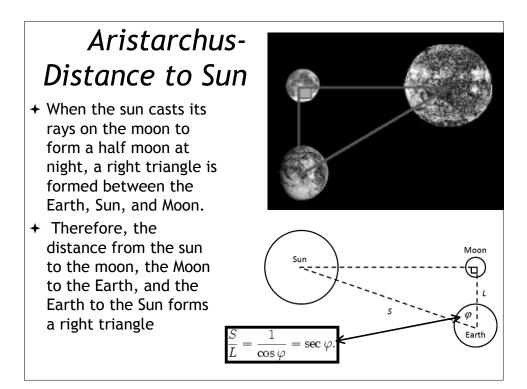
# Distance of Moon From Earth

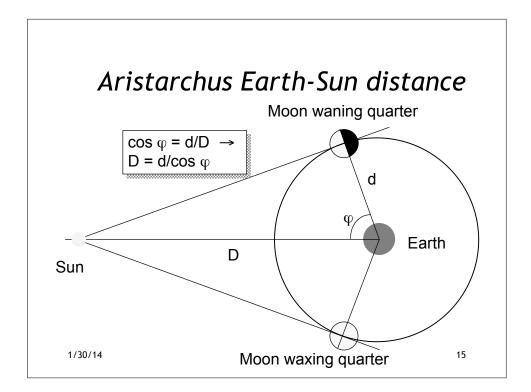
- + Let the Earth be a sphere and the Moon moves in a large circle around Earth.
- + Let R be the radius of that circle and T the time it takes the Moon to go around once, about one month (720hours).
- + In that time the Moon covers a distance of  $2\pi R$
- + An eclipse of the Moon occurs when the Moon passes through the shadow of the Earth, on the opposite side from the Sun (therefore, it must be a full Moon).
- + r is the radius of the Earth, the shadow's width is ~ 2r.
- + Let t be the time it takes the Moon to cross the shadow, (~3 hours )

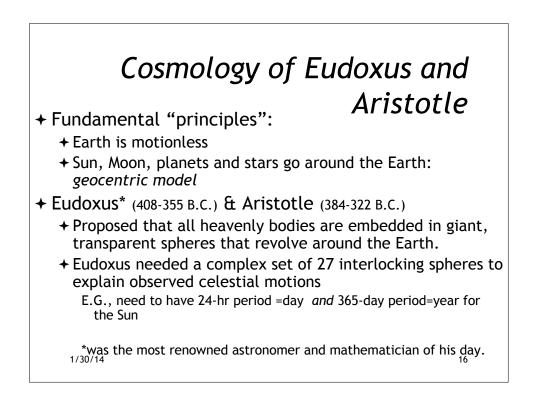
It takes time T to cover  $2\,\pi\,R$  and time t to cover 2r--then

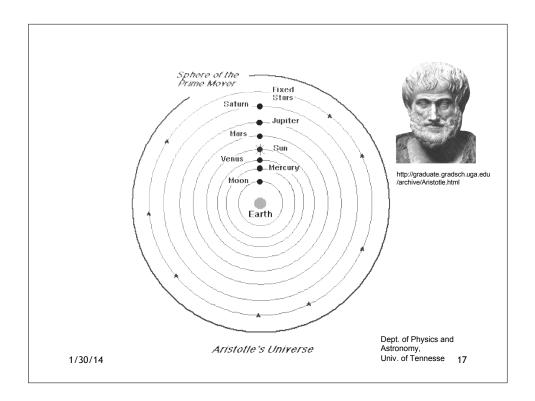
- +  $2\pi R/2$  r = T/t and Aristarchus obtained R/r ~ 60
- + which fits the average distance of the Moon, 60 Earth radii.

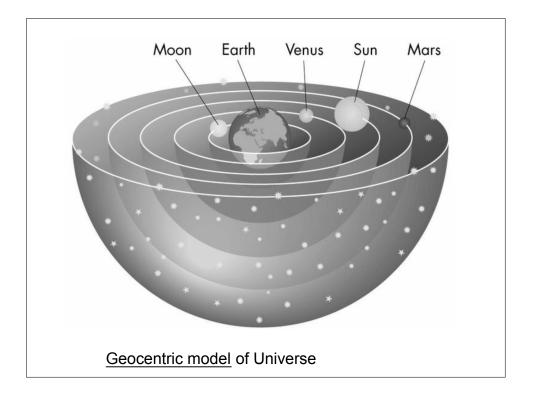
http://en.wikipedia.org/wiki/On\_the\_Sizes\_and\_Distances\_(Aristarchus)

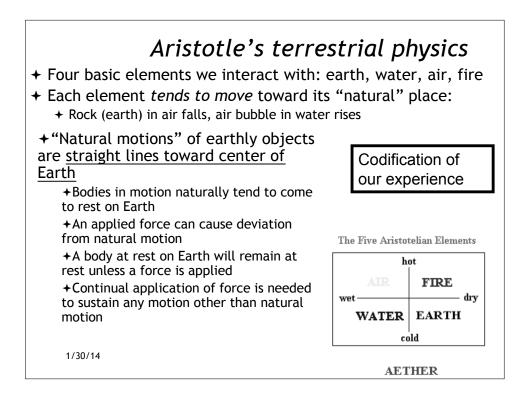


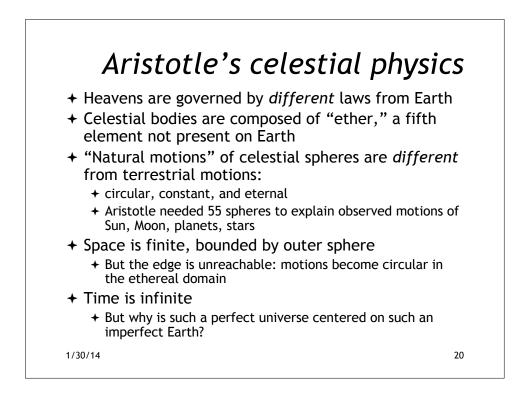


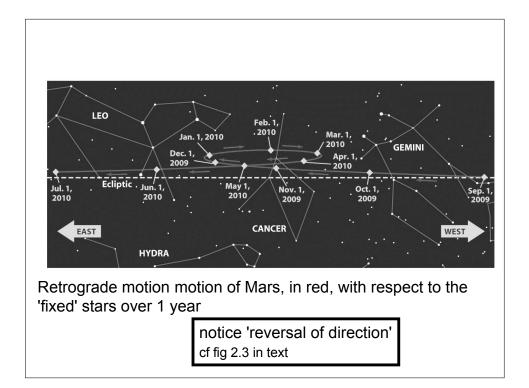


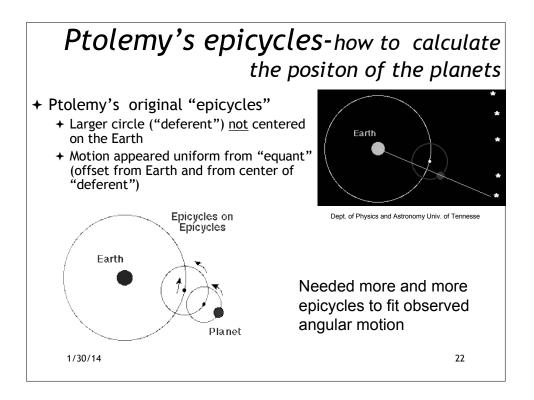


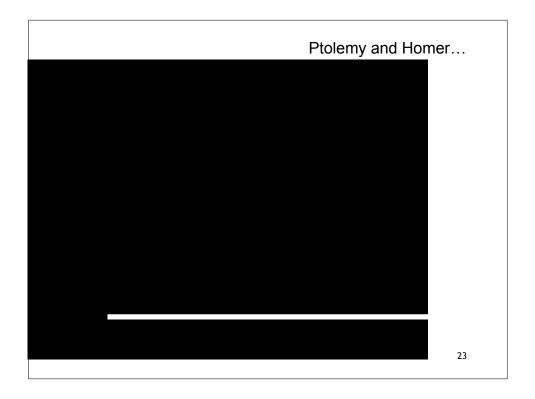


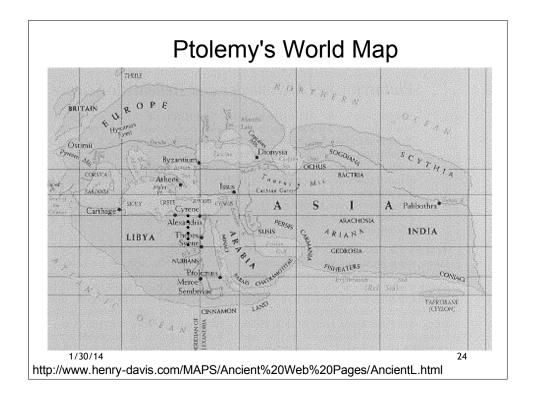








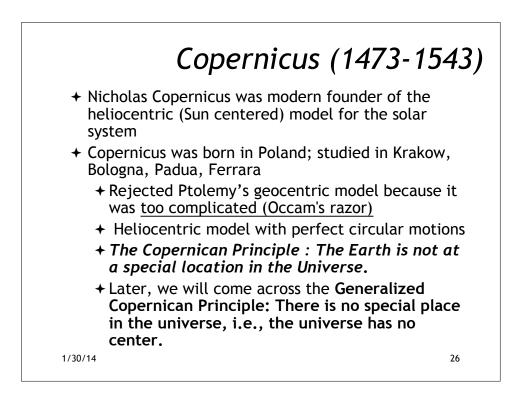




## II : The Renaissance

25

- During European "dark ages," Arab astronomers preserved and extended Ptolemy's work (Muslim astronomy was the most advanced in the world at this time).
- + Aristotelian/Ptolemaic view prevailed in Europe, through 1400's
  - +Geocentric model
  - +Creation at finite time in past, for consistency with Christian theology
  - + Earth known to be round (Columbus battling against flat Earthers is myth!)



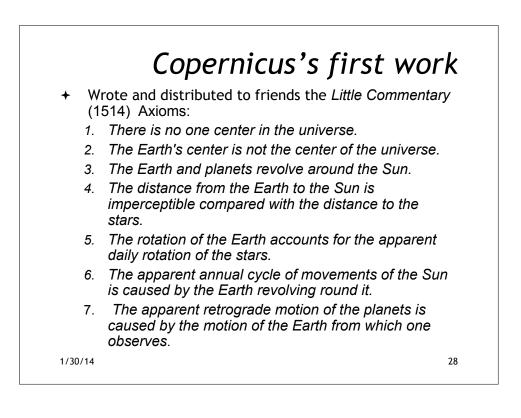
- Even though the mathematics was not any simpler than Ptolemy's, <u>it</u> required fewer basic assumptions.
- By postulating only
  - + the rotation of the Earth,
  - + revolution about the sun,
  - + tilt of Earth's rotational axis, Copernicus could explain the observed motion of the heavens.
- + However, because Copernicus retained circular orbits, his system required the inclusion of epicycles. Out of fear that his ideas might get him into trouble with the church, Copernicus delayed publication of them.

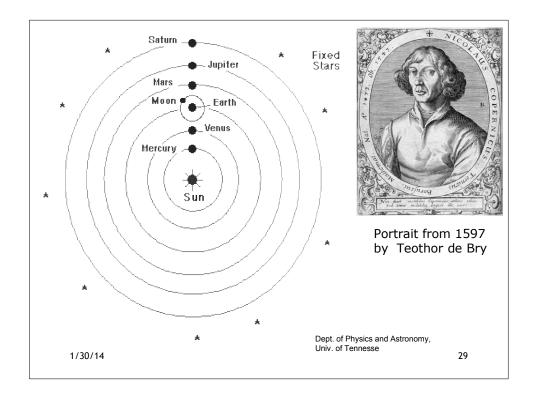
# Copernicus

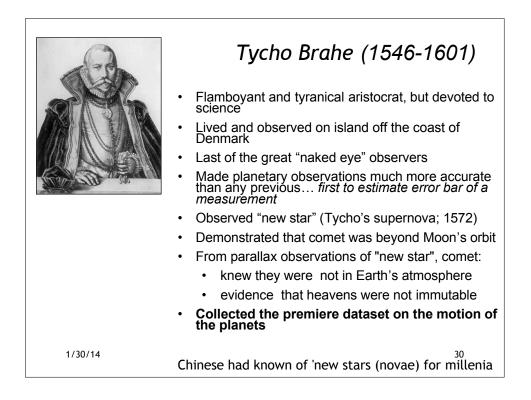
Copernicus adapted physics to the demands of astronomy, believing that <u>the principles</u> <u>of Ptolemy's system were</u> <u>incorrect</u>, not the math or observations.

He was the first person in history to create a complete and general system, combining mathematics, physics, and cosmology (**the theme of this class !**)

http://scienceworld.wolfram.com/biography/Copernicus.html







### Tycho Brahe and the Origin of Scientific Funding

- + He received an annual pension of five hundred dalers.
  - This was far more than the income of any other man of learning in Europe, and even for an aristocrat, it was a substantial income. Tycho set a new European standard for the financial support of scientific research.
  - + It is estimated that Brahe's observatory cost about 1% of the Danish government budget during construction.



Uraniborg

