

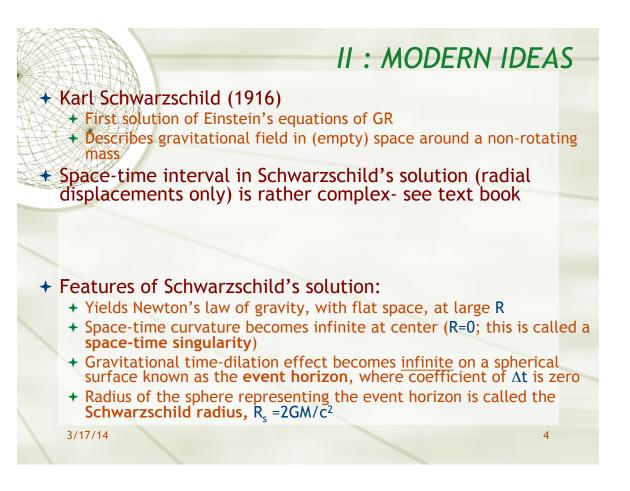
## 18<sup>th</sup> Century ideas

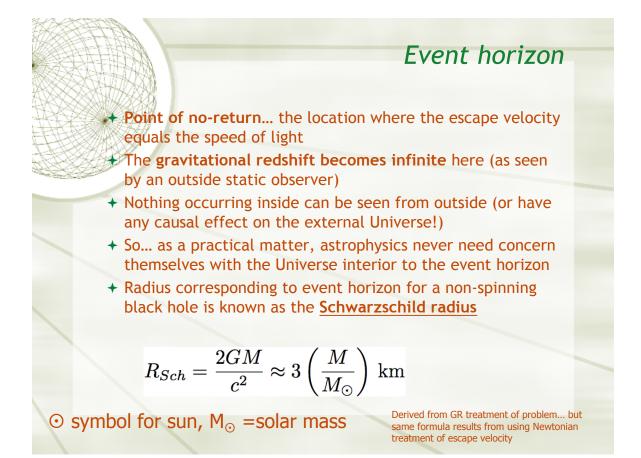
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 By making M larger and R smaller, V<sub>esc</sub> increases
Idea of an object with gravity so strong that light cannot escape first suggested by Rev. John Mitchell in 1783

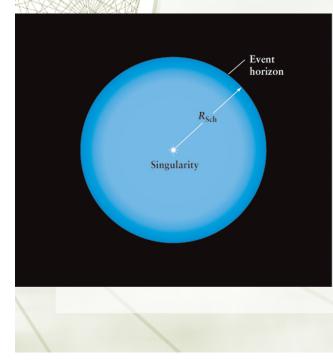
Laplace (1798) - "A luminous star, of the same density as the Earth, and whose diameter should be two hundred and fifty times larger than that of the Sun, would not, in consequence of its attraction, allow any of its rays to arrive at us; it is therefore possible that the largest luminous bodies in the universe may, through this cause, be invisible."

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Gravitational redshift outside of a spherical object with mass M is

$$\nu_{obs} = \left(1 - \frac{2GM}{rc^2}\right)^{1/2} \nu_{emit}$$

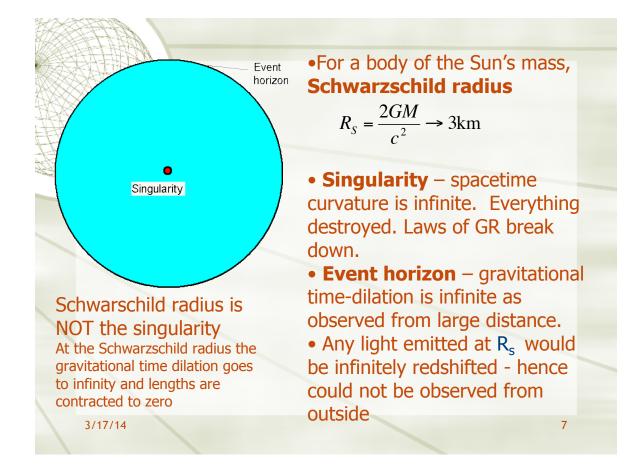
Gravitational length contraction

 $L'=L (1-R_s/R)^{1/2}$ 

Gravitational time dilation  $t'=t/(1-R_s/R_s)^{1/2}$ 

 $R_{s=} 2GM/c^2$ 

As  $R \rightarrow R_s$  time goes to  $\infty$ length goes to zero, wavelength of emitted radiation goes to zero



## More features of Schwarzschild black hole

- Events inside the event horizon are causally-disconnected from events outside of the event horizon (i.e. no information can be sent from inside to outside the horizon)
- Observer who enters event horizon would only "feel" "strange" gravitational effects if the black hole mass is small, so that R<sub>s</sub> is comparable to observer's size
- + Once inside the event horizon, future light cone always points toward singularity (any motion must be inward)
- Stable, circular orbits are not possible inside 3R<sub>s</sub>: inside this radius, orbit must either be inward or outward but not steady
- + Light ray passing BH tangentially at distance 1.5R<sub>s</sub> would be bent around into a circle
- + Thus black hole would produce "shadow" on sky

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