

## Lecture 6 : *The extragalactic universe I*

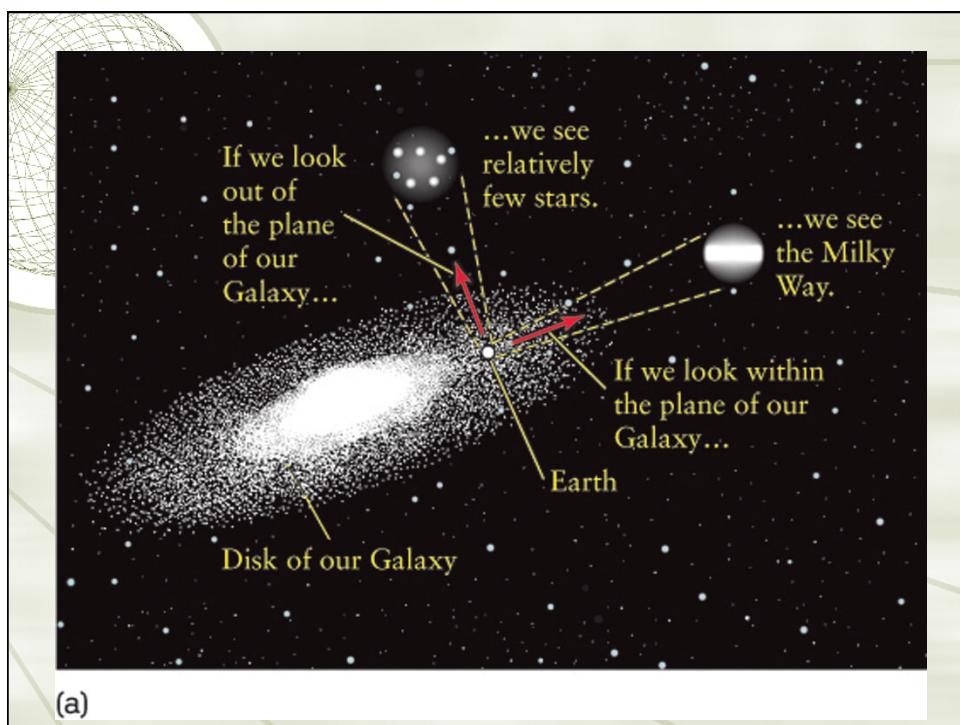
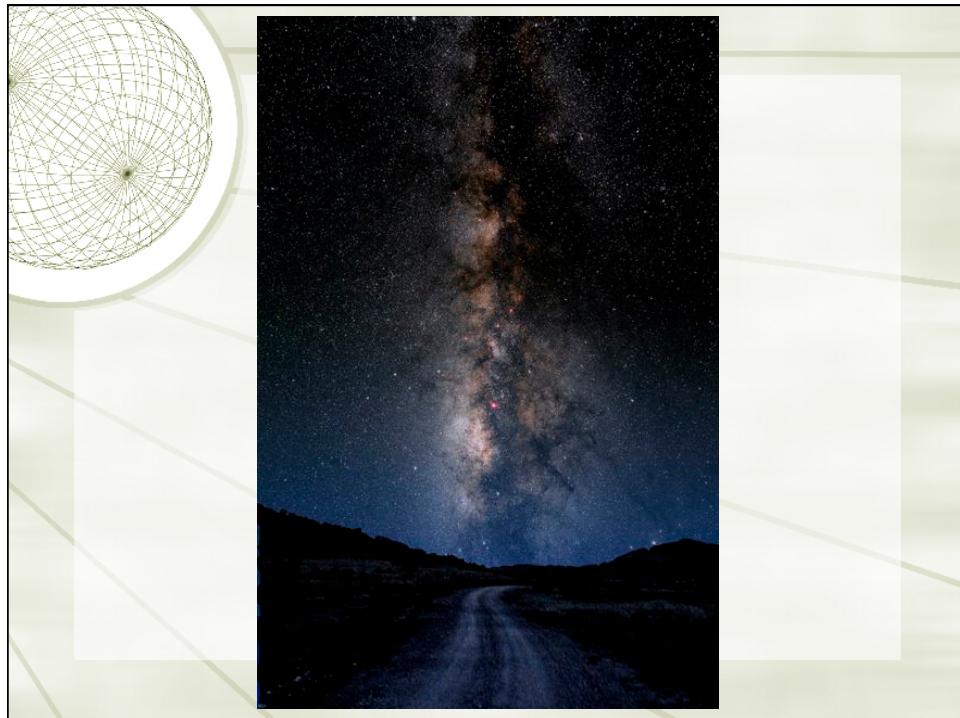
- ❖ Our place in the Galaxy
- ❖ The Great Debate
- ❖ Measuring distance in astronomy
  - ❖ Parallax and Cepheids
- ❖ Hubble's (first) great discovery

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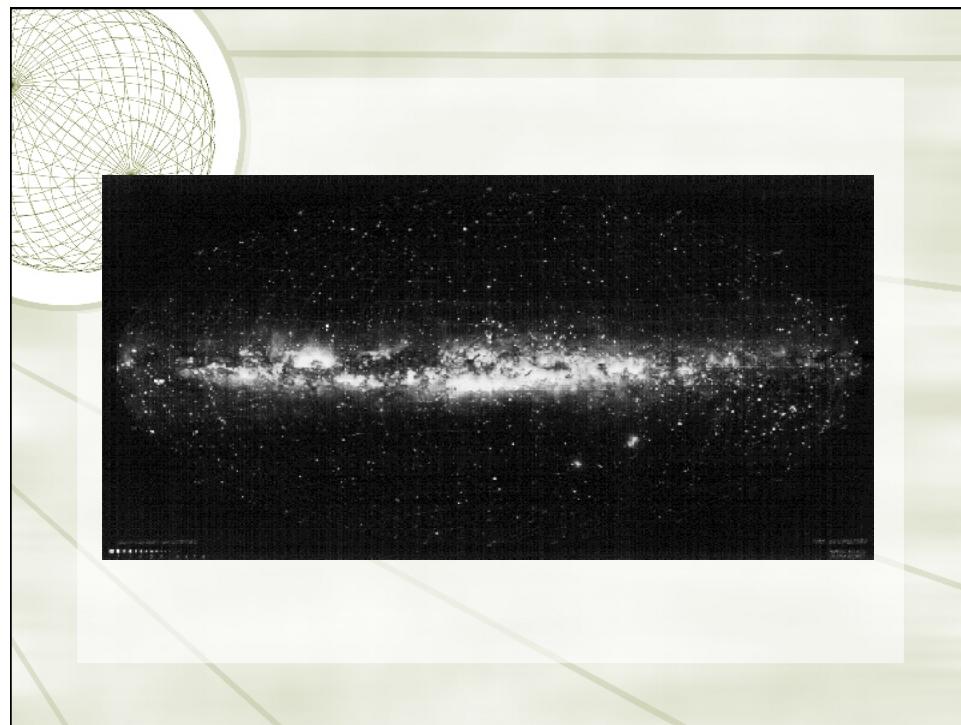
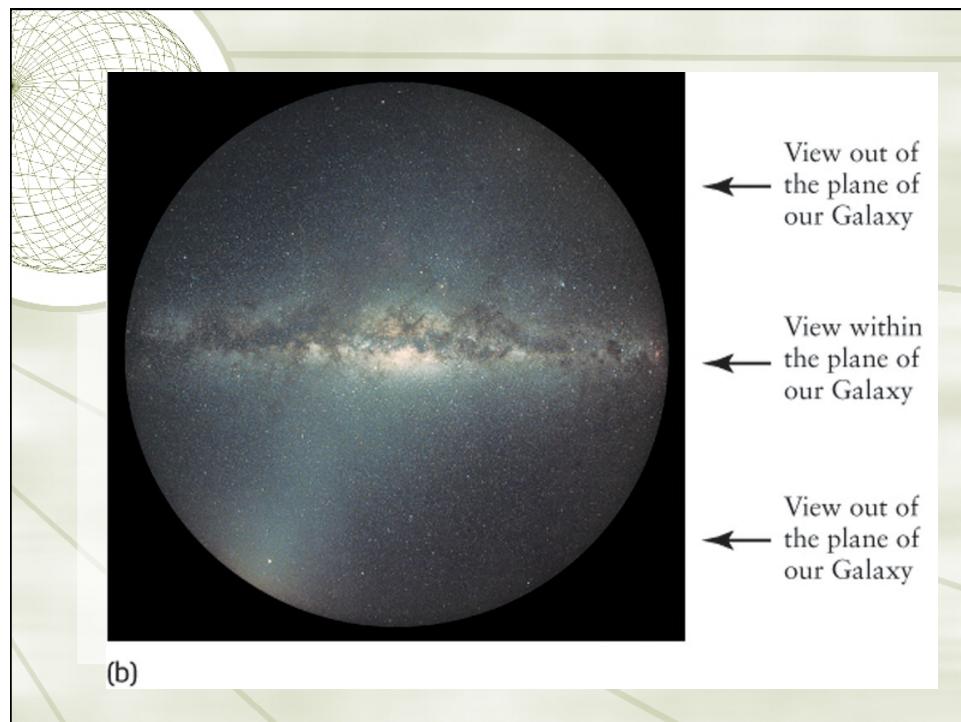


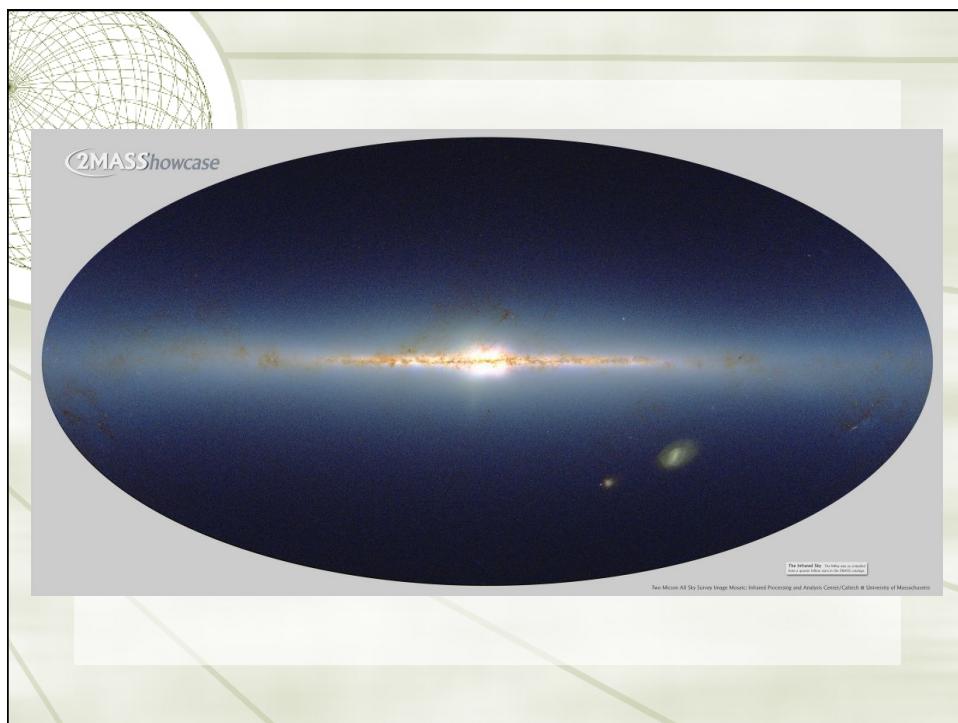
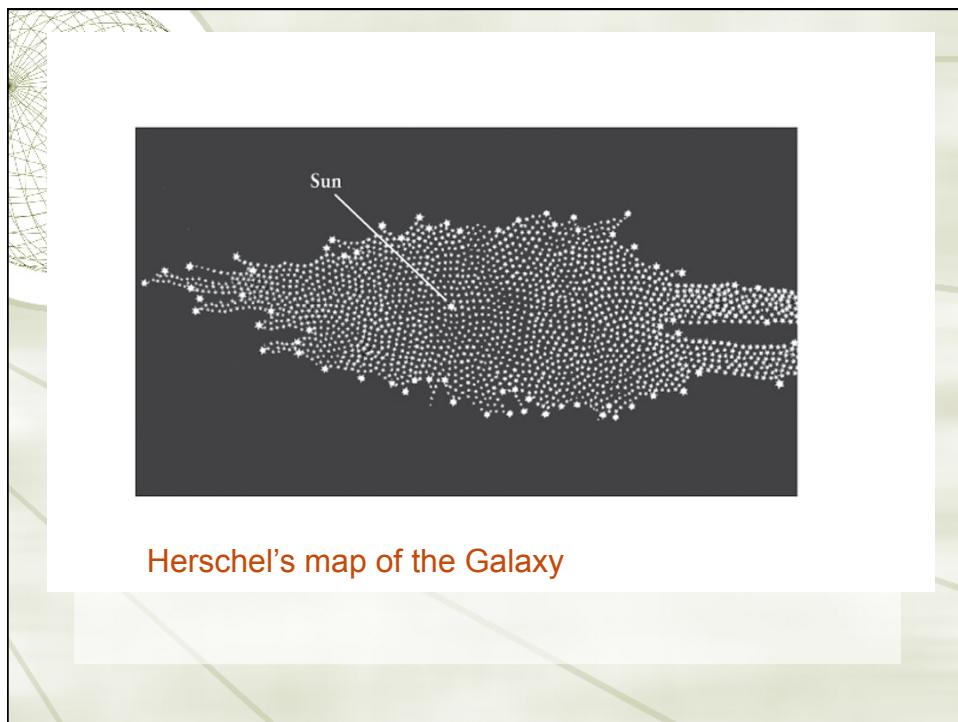
### I : Our place in the Galaxy

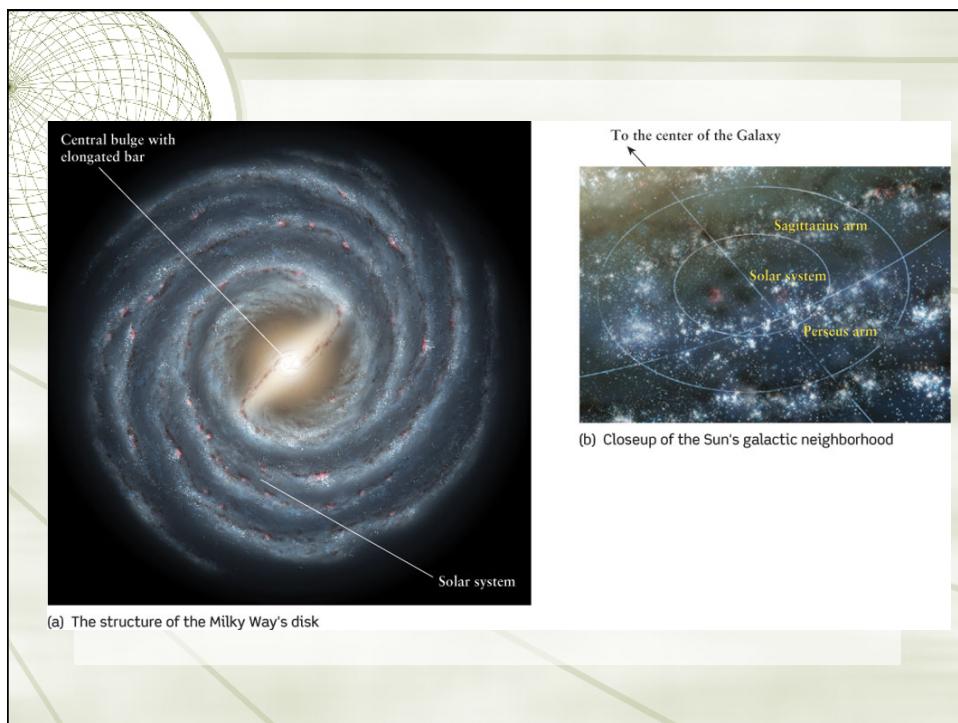
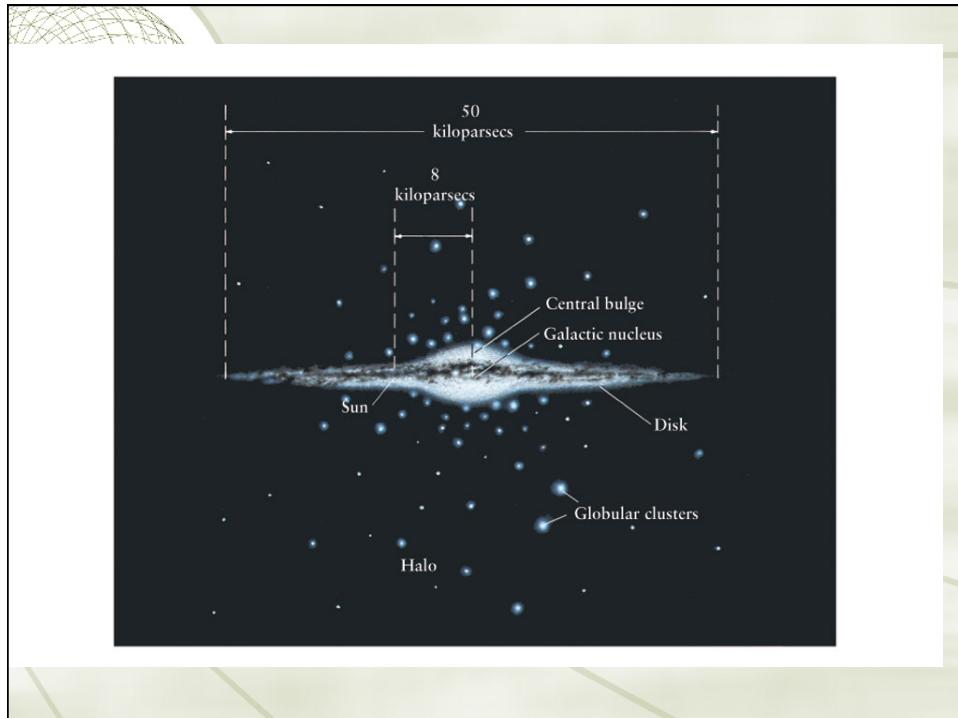
- ❖ We live in a large disk galaxy
  - ❖ We live in the disk, towards the edge (8kpc~25,000lyr out)
  - ❖ Projected onto the sky, this disk of stars looks like a band of light that rings the sky... the Milky Way
- ❖ This realization came somewhat slowly...
  - ❖ Disk-like nature of galaxy realized by Thomas Wright (1780); refined by Kant
  - ❖ First attempt to map out galaxy made by William Herschel (1785); refined by Kapteyn in 1920
  - ❖ Herschel came to the conclusion that we sit at the center of the Galactic disk. In fact, he was wrong... had not accounted absorption by dust!



(a)







## II : The Great Debate... “what are spiral nebulae?”

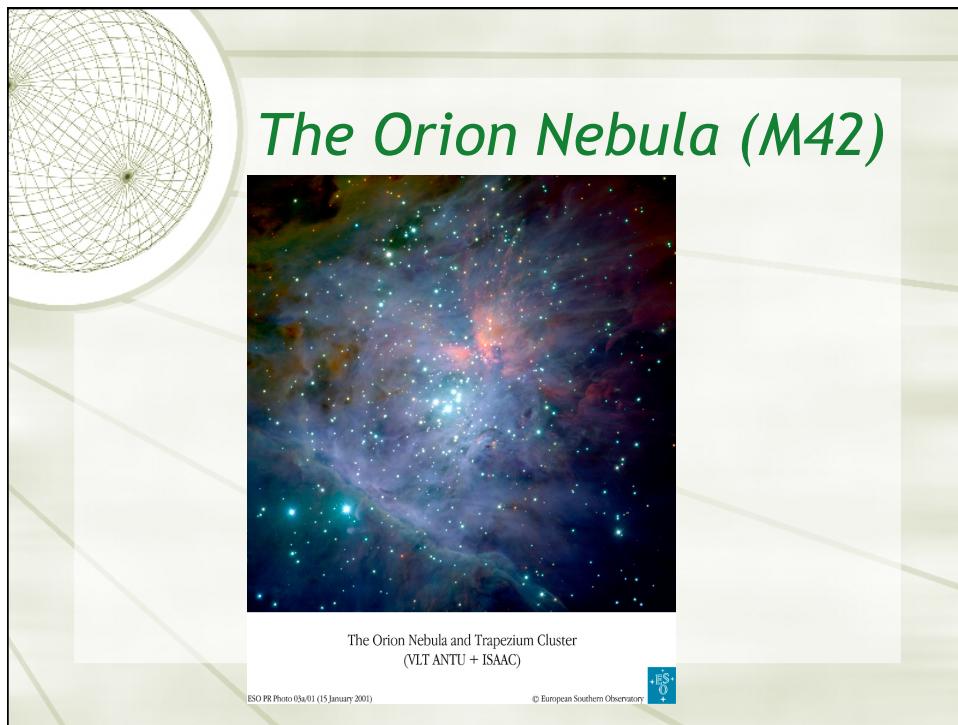
❖ Early 20th century...

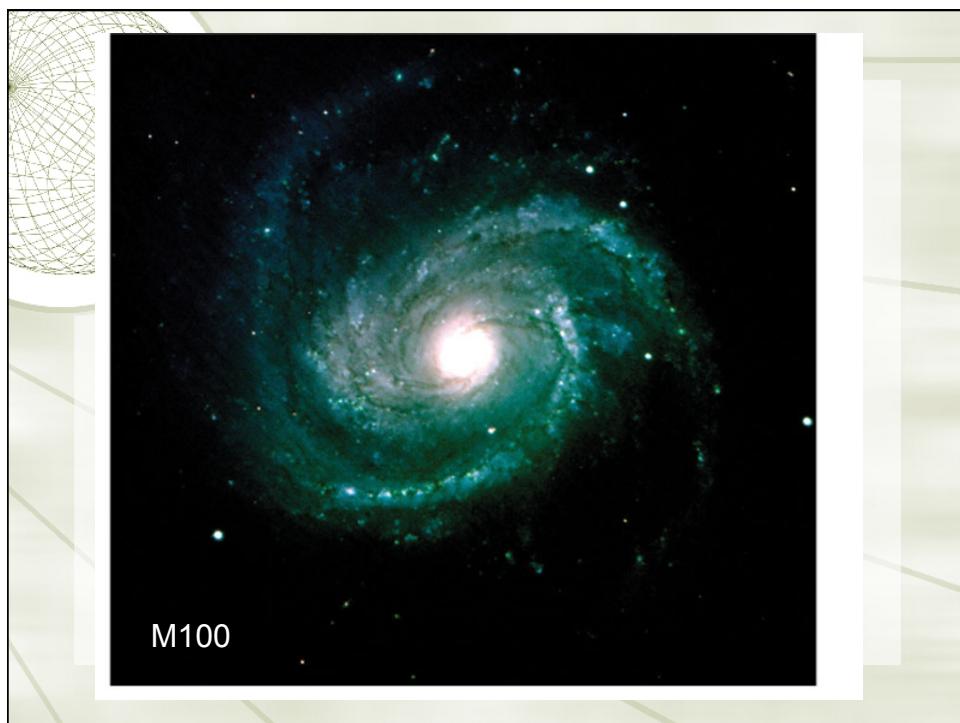
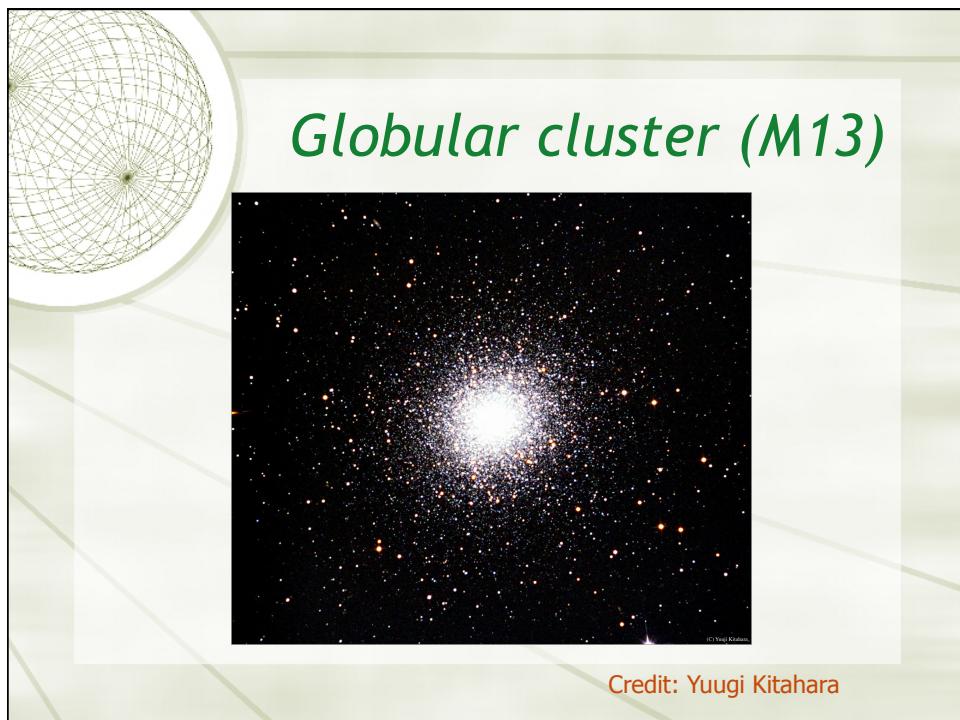
- ❖ Knew that we lived in a large disk galaxy
- ❖ But what was the nature of the larger Universe?
- ❖ Two opposing ideas:
  - ❖ Our galaxy is alone, sitting in the middle of otherwise empty space
  - ❖ Our galaxy is one of many galaxies that fill space (so-called “Island Universes”)
- ❖ The debate rapidly focused on the nature of nebulae

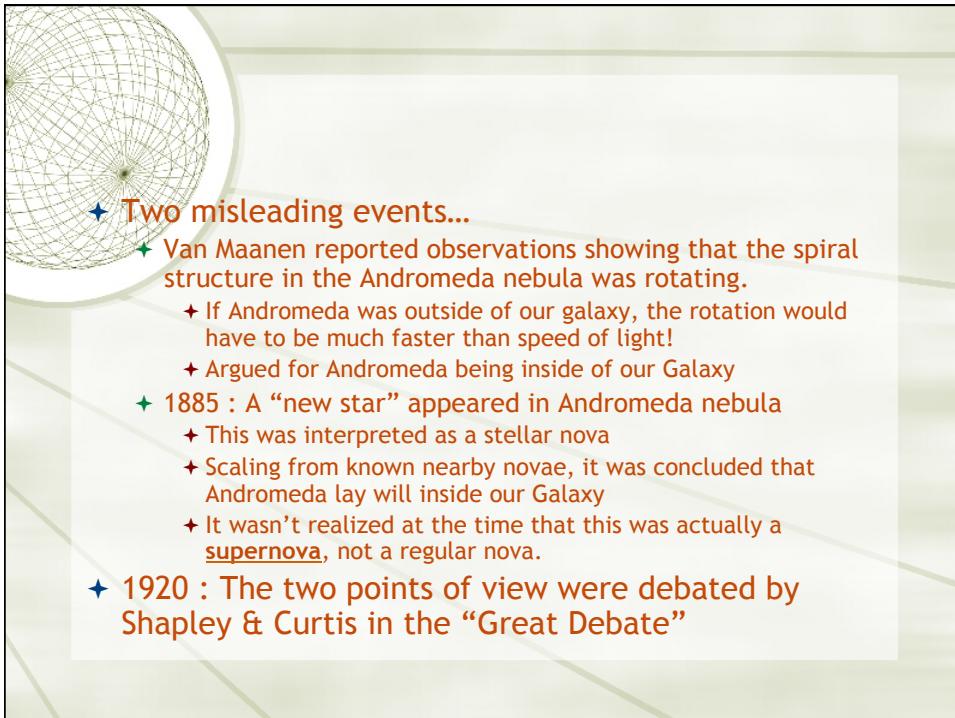
❖ Nebulae have been studied for ages

❖ Messier (1780)

- ❖ Systematically catalogued over 100 bright nebulae
- ❖ Main reason for doing this was so that comet hunters knew which “fuzzy patches” to ignore!
- ❖ But what were these nebulae? Possibilities:
  - ❖ Glowing patches of gas (e.g., Orion)
  - ❖ Clusters of many stars within our Galaxy (e.g., the Globular cluster M13)
  - ❖ Whole other galaxies!!
- ❖ Of special interest were the “spiral nebulae” that showed Milky Way like spiral structure... the brightest spiral nebulae was the Andromeda nebula

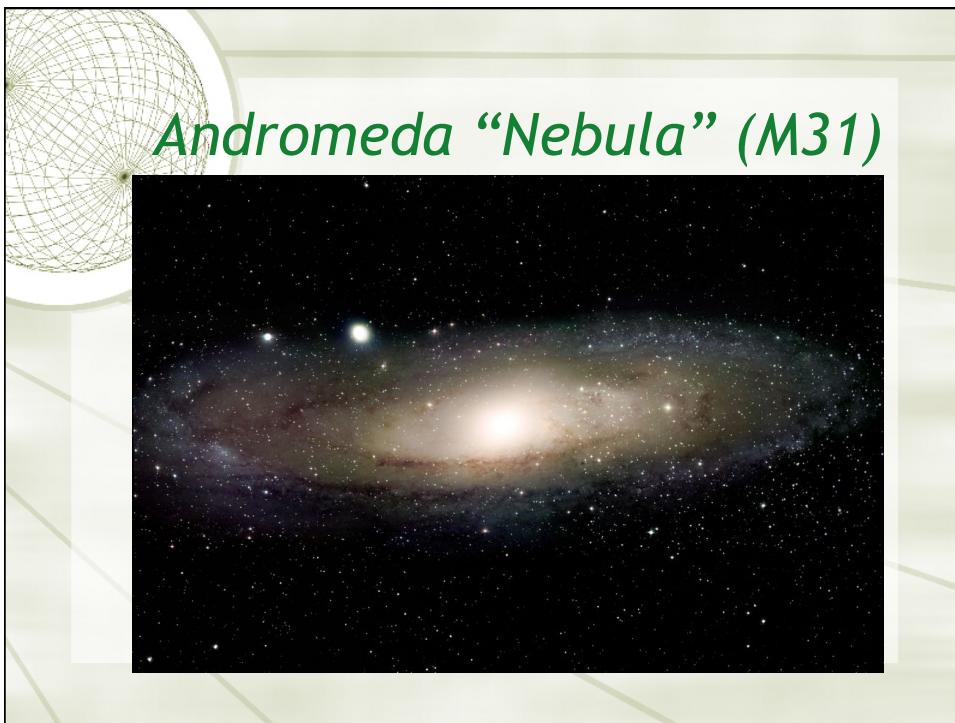






❖ Two misleading events...

- ❖ Van Maanen reported observations showing that the spiral structure in the Andromeda nebula was rotating.
  - ❖ If Andromeda was outside of our galaxy, the rotation would have to be much faster than speed of light!
  - ❖ Argued for Andromeda being inside of our Galaxy
- ❖ 1885 : A “new star” appeared in Andromeda nebula
  - ❖ This was interpreted as a stellar nova
  - ❖ Scaling from known nearby novae, it was concluded that Andromeda lay well inside our Galaxy
  - ❖ It wasn’t realized at the time that this was actually a supernova, not a regular nova.
- ❖ 1920 : The two points of view were debated by Shapley & Curtis in the “Great Debate”

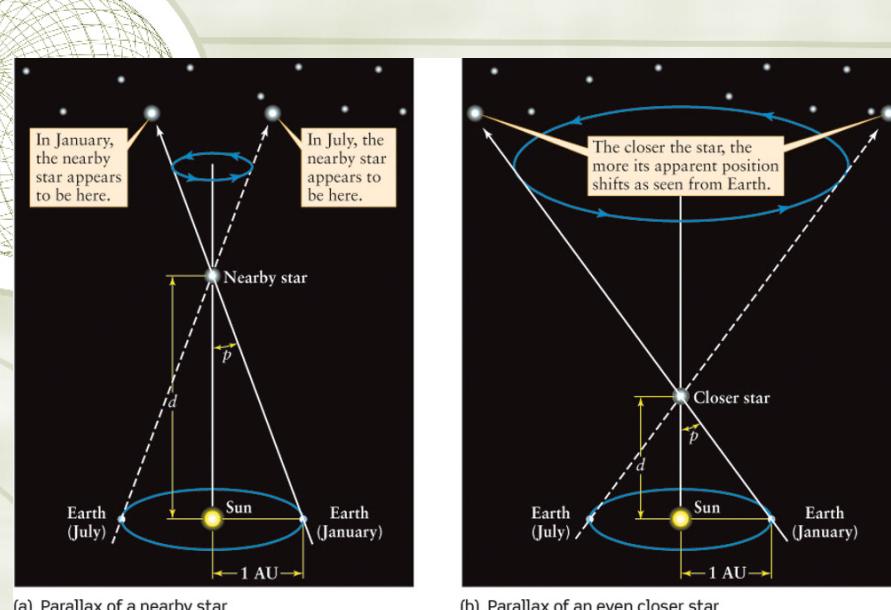


## Andromeda “Nebula” (M31)



### III : Measuring distances in astronomy

- ◆ The distance to any astronomical object is the most basic parameter we want to know
  - ◆ Require knowledge of distance in order to calculate just about any other property of the object
  - ◆ Distance is often difficult to determine!
- ◆ Most direct method for measure distances to “nearby” stars uses an effect called parallax
  - ◆ As Earth orbits Sun, we view a star along a slightly different line of sight
  - ◆ This causes the star to appear to move slightly with respect to much more distant stars
  - ◆ We can currently use this technique to measure stellar distances out to ~3000 light years from Earth



If star wobbles with amplitude of 1 arc-second, then it is at distance of **1 parsec** (definition of parsec).

- 1pc = 3.26 lt-yr
- In general,

$$D(\text{pc}) = \frac{1}{\theta_{\text{wobble}}(\text{arcsec})}$$

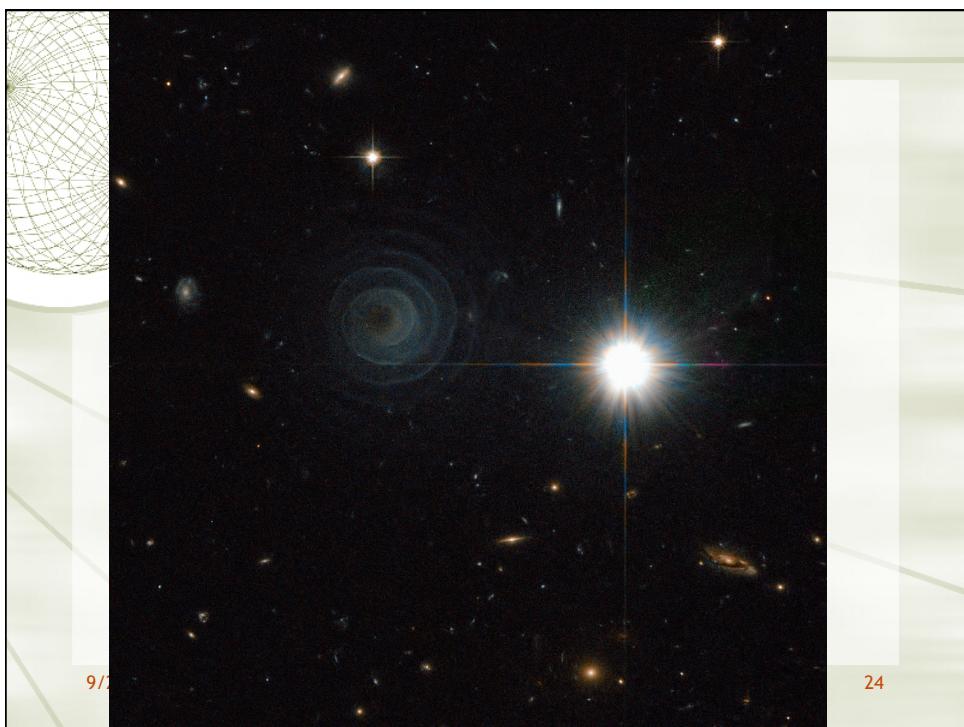
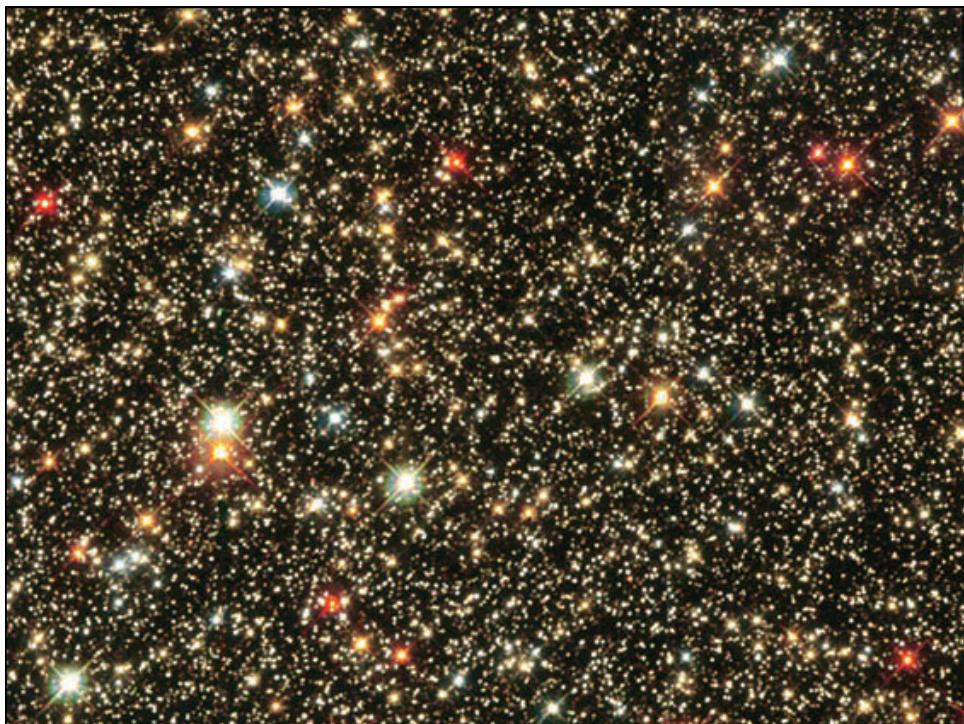
9/23/10 21

Until 1990s, could only detect parallax out to 50pc.

- Hipparcos satellite
  - Designed to measure parallax of stars
  - Can detect wobble out to distance of about 1kpc (1000pc)
  - Used to map out locations of nearby stars.
- GAIA satellite
  - Due to launch 2012
  - Can map out positions and motions of stars across the whole galaxy!!

Hipparcos (ESA)

9/23/10



## Beyond parallax?

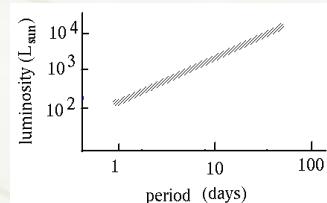
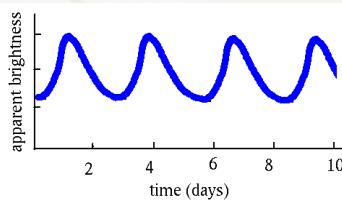
- ◆ Definition : The observed flux of a star is the energy received from the star per unit time per unit area.
  - ◆ Definition : The luminosity of a star is the energy per unit time (i.e. power) emitted by the star
  - ◆ If the star is at distance D and emits equally in all directions (i.e. it emits isotropically), then the observed flux F and luminosity L are related by
- $$L = 4\pi D^2 F \quad \text{or} \quad F = \frac{L}{4\pi D^2}$$
- ◆ Suppose we know the luminosity of some object... then we can use its measured flux to determine the distance! Objects with known luminosities are called standard candles.



Henrietta Leavitt  
ph.credit: AAVSO

## Cepheid variables

- ◆ Henrietta Leavitt discovered (1912) that a certain class of variable stars called Cepheids had properties that meant they could be used as standard candles
  - ◆ She studied Cepheids that are close enough for parallax to be measured... found that the luminosity is related to the period of fluctuations in brightness
  - ◆ So, if you measure the period of a Cepheid, you can determine its luminosity. Measuring flux then gives you distance, even if its too far for parallax!



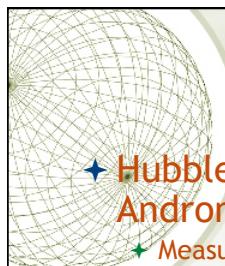


*Edwin Hubble*

Hale Observatories, courtesy AIP

9/23/10 27

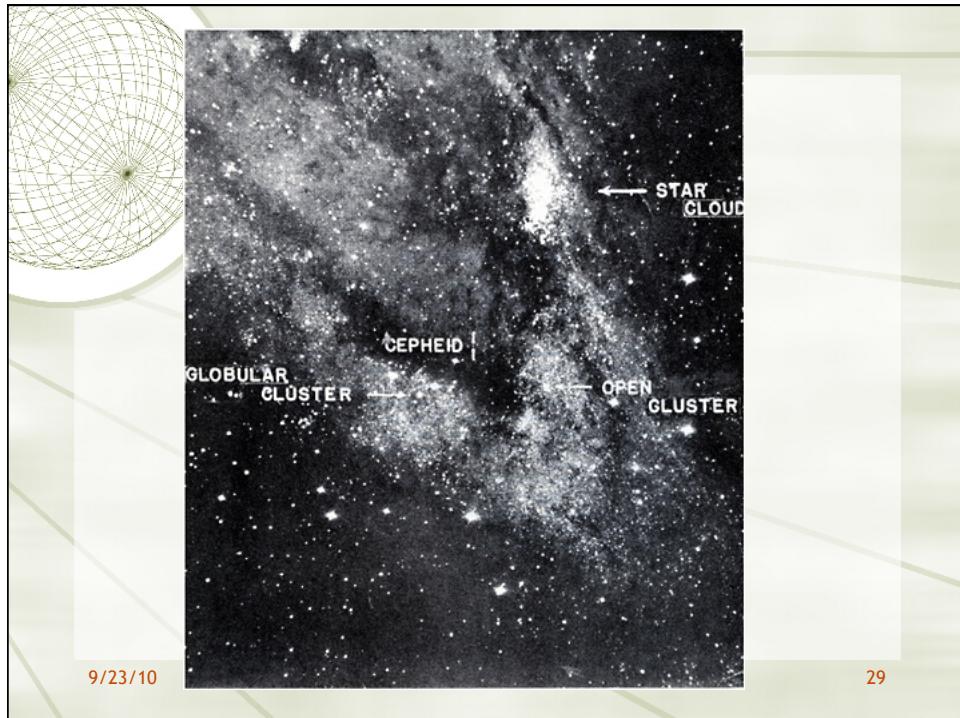
A black and white portrait of Edwin Hubble, an astronomer, sitting at a desk and looking towards the camera. He is holding a large astronomical photograph of a galaxy, likely Andromeda, in his hands. The background shows some scientific equipment and papers.



## *Hubble's observations*

- ♦ Hubble found Cepheid Variables in Andromeda
  - ♦ Measured period and flux, and hence distance
  - ♦ Concluded that Andromeda must be well outside of the Milky Way Galaxy
  - ♦ Thus, the Great Debate was settled... the MW is just one of many many many galaxies
- ♦ Modern measurements
  - ♦ Distance to Andromeda 2 million lyr
  - ♦ About 20x MW diameter

9/23/10 28



9/23/10

29

