

## Class 2 : Hubble's Law Renewed & the Friedmann Equation

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- This class...
  - Recap of Cosmological Principles
  - A new view of Hubble's law
  - The Friedmann Equation

## O : Recap The Cosmological Principle

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- **The Cosmological Principle** states that the Universe looks the same where-ever you are and which-ever direction you look
- Two immediate consequences of the Cosmological Principle...
  - Universe is **homogeneous** : every place in the universe has the same conditions as every other place, on average.
  - Universe is **isotropic** : there is no preferred direction in the universe, on average.

## I : Fresh look at Hubble's Law

- Hubble's Law (original form) : A galaxy a distance  $R$  from us will be moving away from us with velocity  $V$  given by

$$V = H_0 R$$

- Notes:

- The constant  $H_0$  is the **Hubble constant**. Subscript "0" refers to the current cosmic time.
- Current estimates are  $H_0 = 67.8 \pm 0.9$  km/s/Mpc (Planck Collaboration)
- As written, this is only true in the non-relativistic limit ( $V \ll c$ ). Once you start to look far enough out, the value of Hubble's constant changes! For this reason, it is best called the **Hubble parameter,  $H$** .

- We can think of the Hubble flow as an overall homogeneous & isotropic expansion of space
- Imagine a coordinates that expand with the space...
  - These are called **co-moving coordinates**
  - Can then write

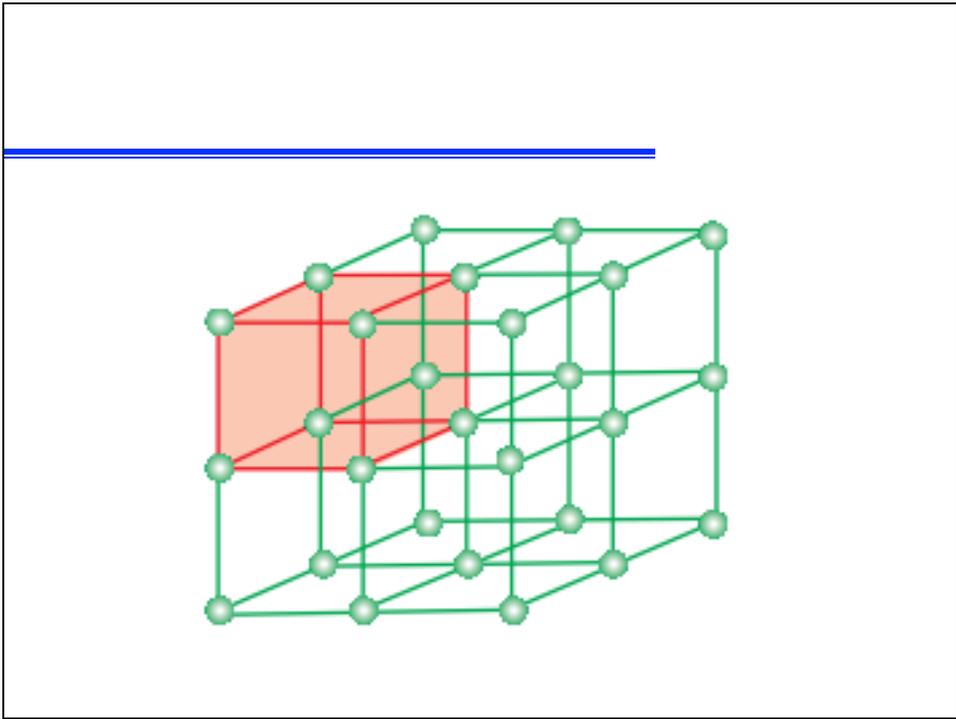
$$R = a x$$

proper distance
Scale factor
comoving distance

- Cosmological Principle  $\rightarrow$  scale factor is purely function of  $t$
- So, a galaxy's velocity can be decomposed into two pieces:

$$V \equiv \frac{dR}{dt} = \frac{da}{dt} x + a \frac{dx}{dt} = \left( \frac{1}{a} \frac{da}{dt} \right) R + a \frac{d}{dt} \left( \frac{R}{a} \right)$$

Hubble Flow
Peculiar velocity



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- We define the **Hubble parameter** by

$$H \equiv \frac{1}{a} \frac{da}{dt}$$

- So, if a galaxy has zero peculiar velocity, then it is carried along by the Hubble Flow and we have

$$V = \left( \frac{1}{a} \frac{da}{dt} \right) R = HR$$

- Notes

- We have recovered Hubble's Law
- Hubble parameter can be a function of time (depending upon the mathematical form of the scale factor)

## II : The Friedmann Equation

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- Friedmann equation is the basic (differential) equation that describes the time evolution of the scale factor

$$\left(\frac{1}{a} \frac{da}{dt}\right)^2 = \frac{8\pi G}{3} \rho - \frac{kc^2}{a^2}$$

- Here,  $\rho$  is total equivalent mass density of the Universe
- $k$  is a constant ( $-$ ,  $0$ , or  $+$ )
- $c$  is speed of light
- From now on, we will use "dot" for the time derivative
- Full derivation comes from applying assumptions of homogeneity and isotropy to the Field Equations of General Relativity
- With a little cheating, we can also derive the Friedmann equation from Newtonian arguments... [see discussion on board]