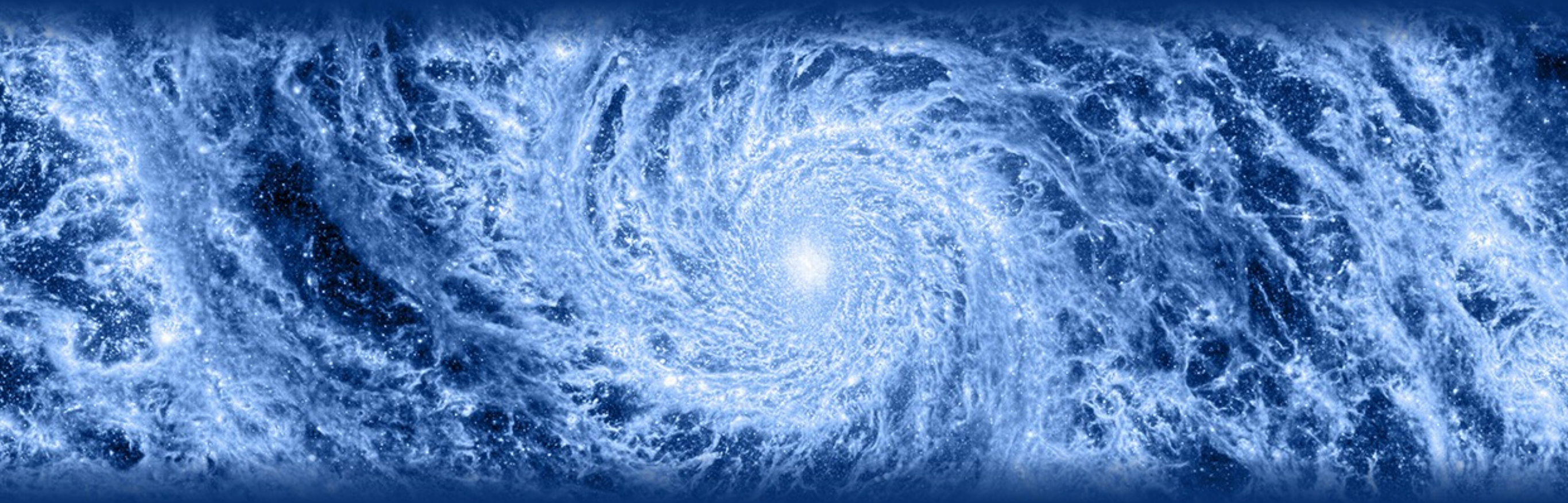


Galaxies

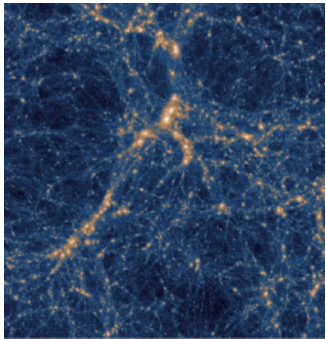
Prof. Benedikt Diemer



Chapter 8 • The co-evolution of halos, gas, stars, and metals

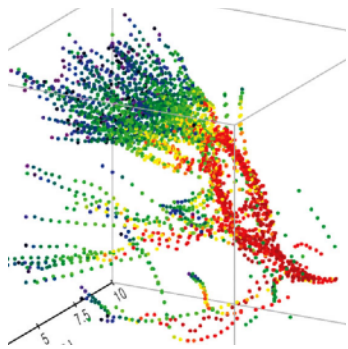
§8.1 • Semi-analytical models of galaxy evolution

Models of galaxy formation

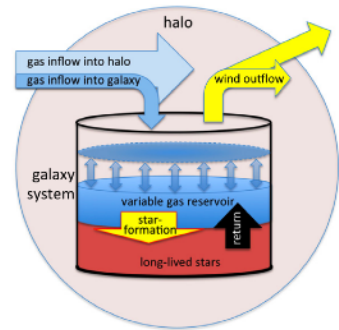


Hydrodynamical simulations

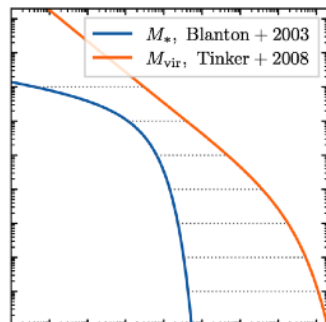
Most complex



Semi-analytical models



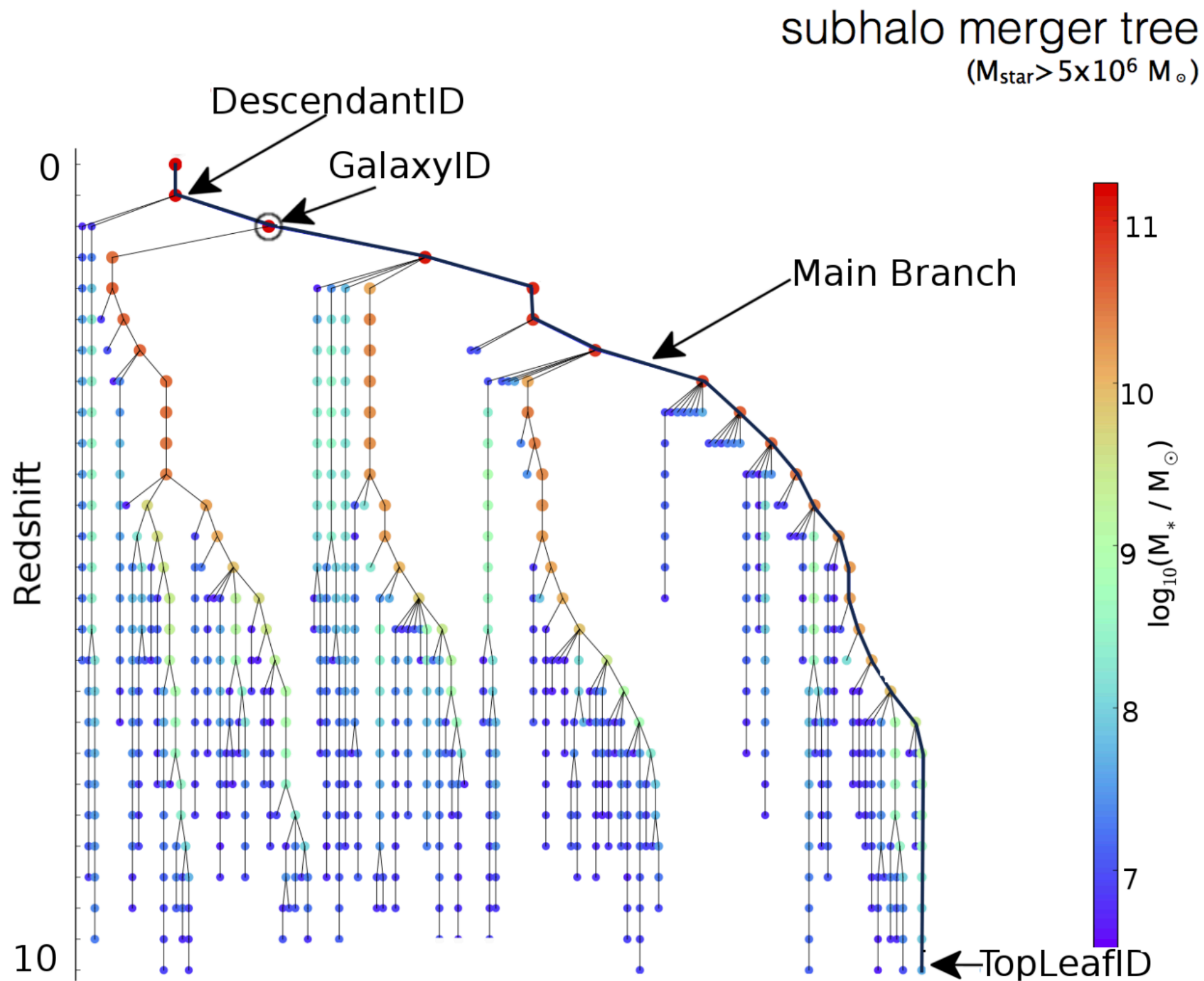
Regulator/bathtub models



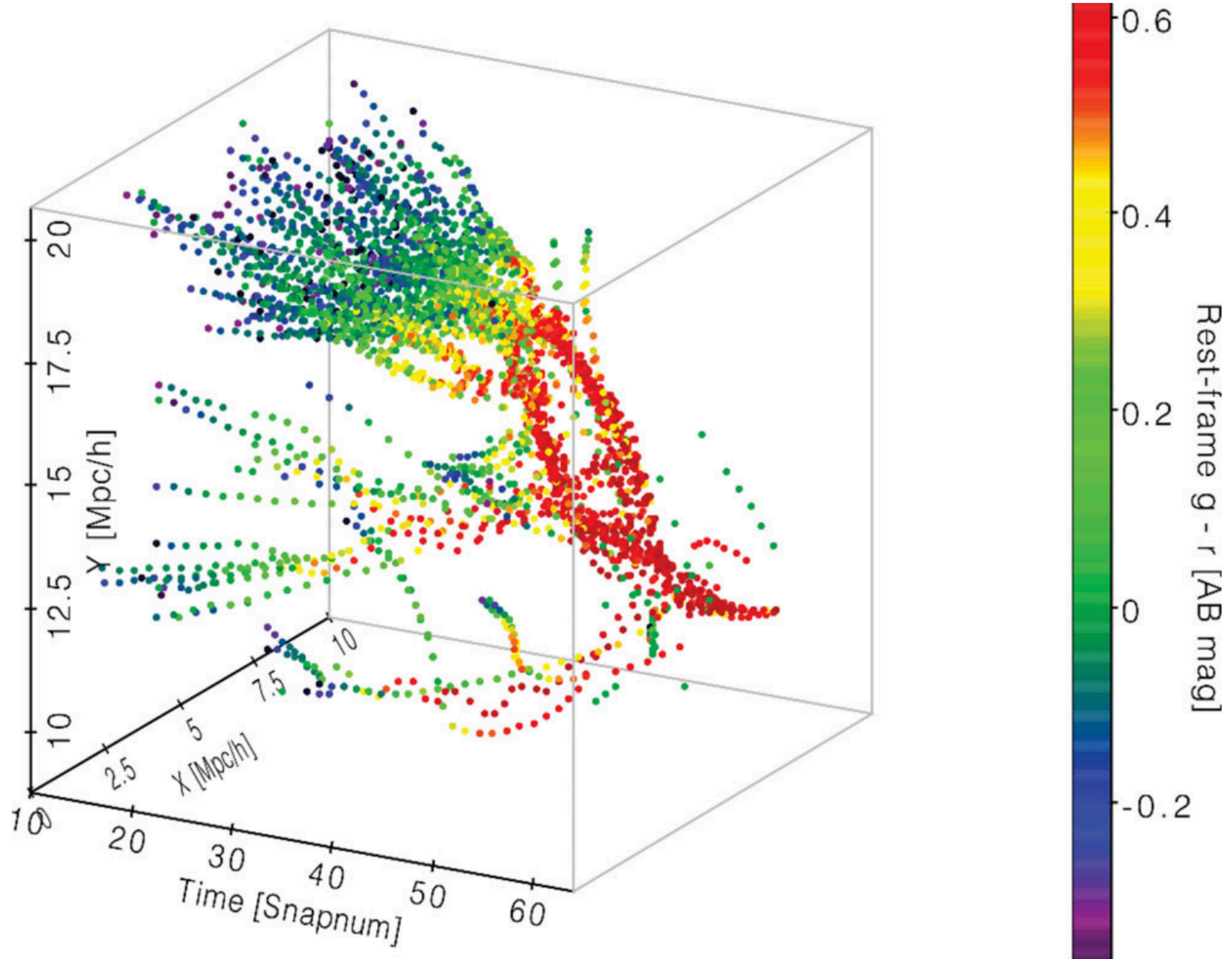
Empirical models
(abundance matching etc.)

Least complex

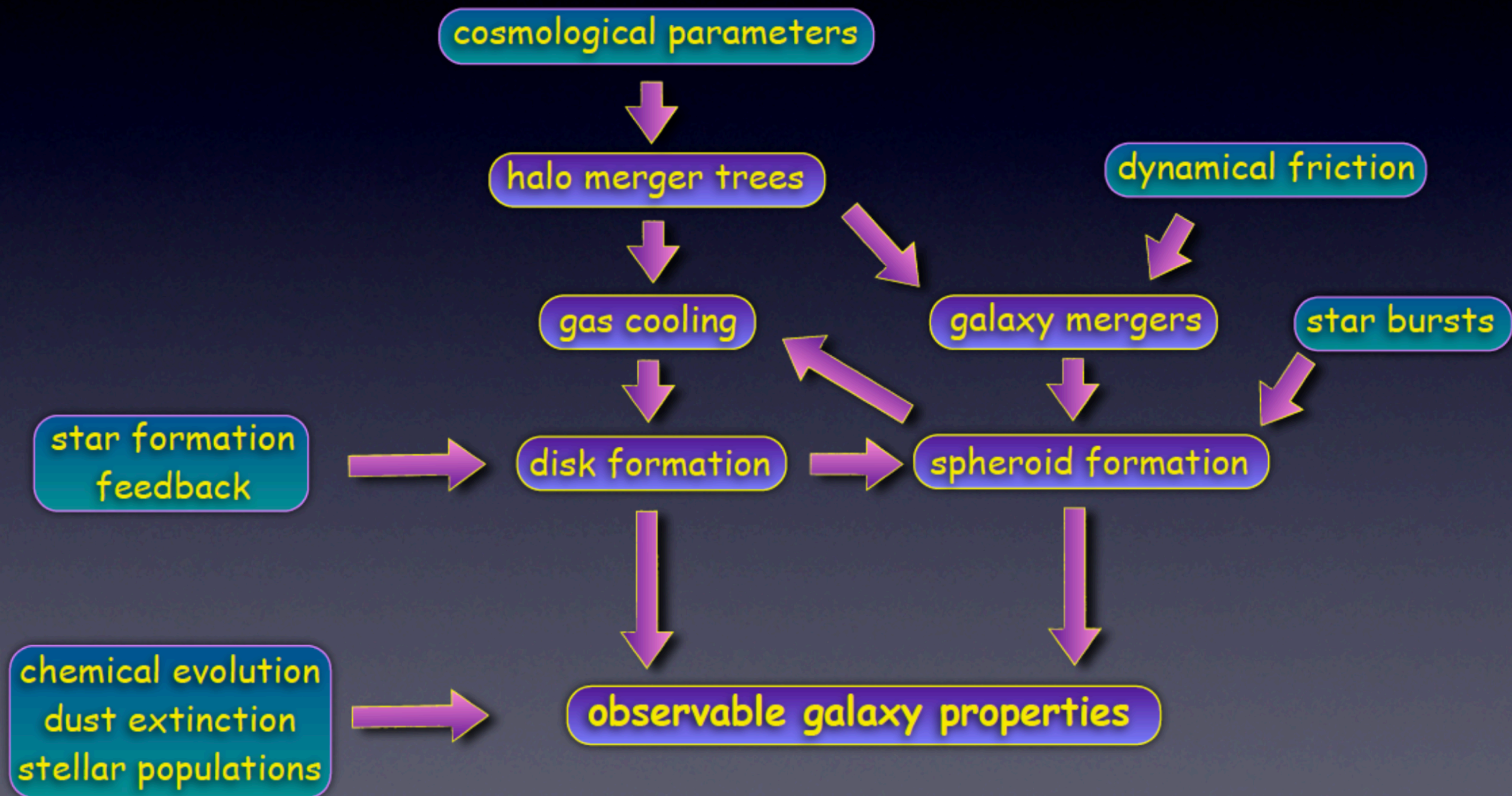
Merger trees



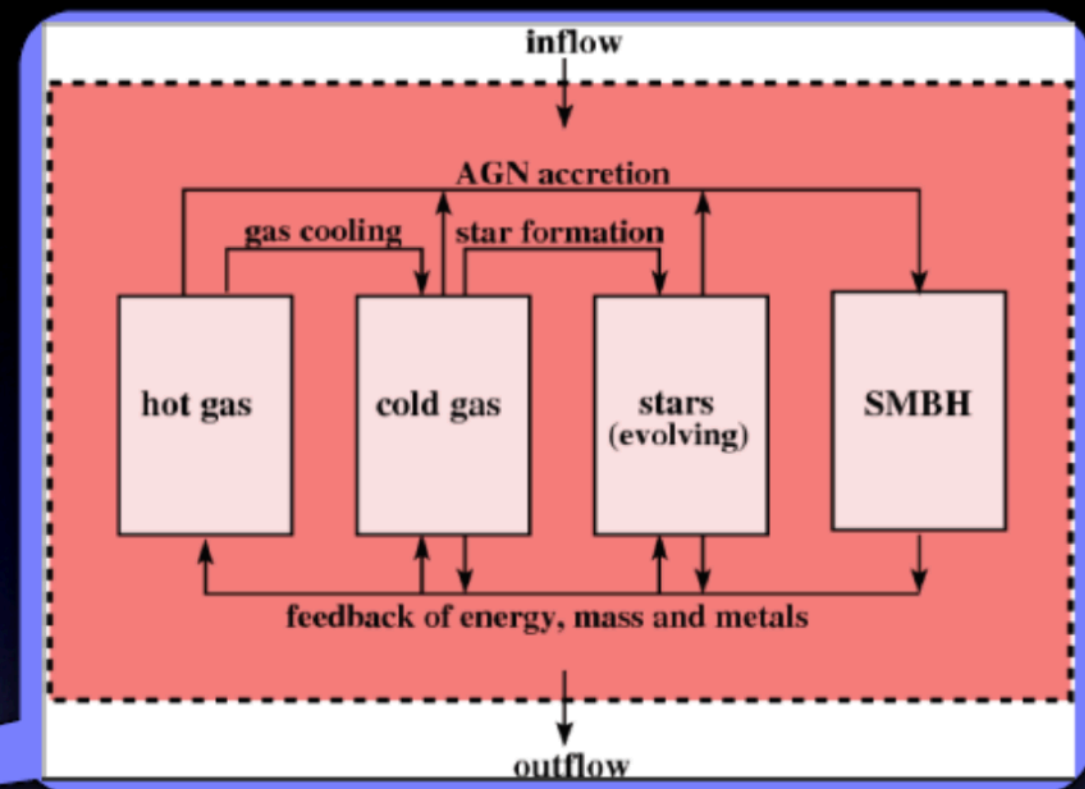
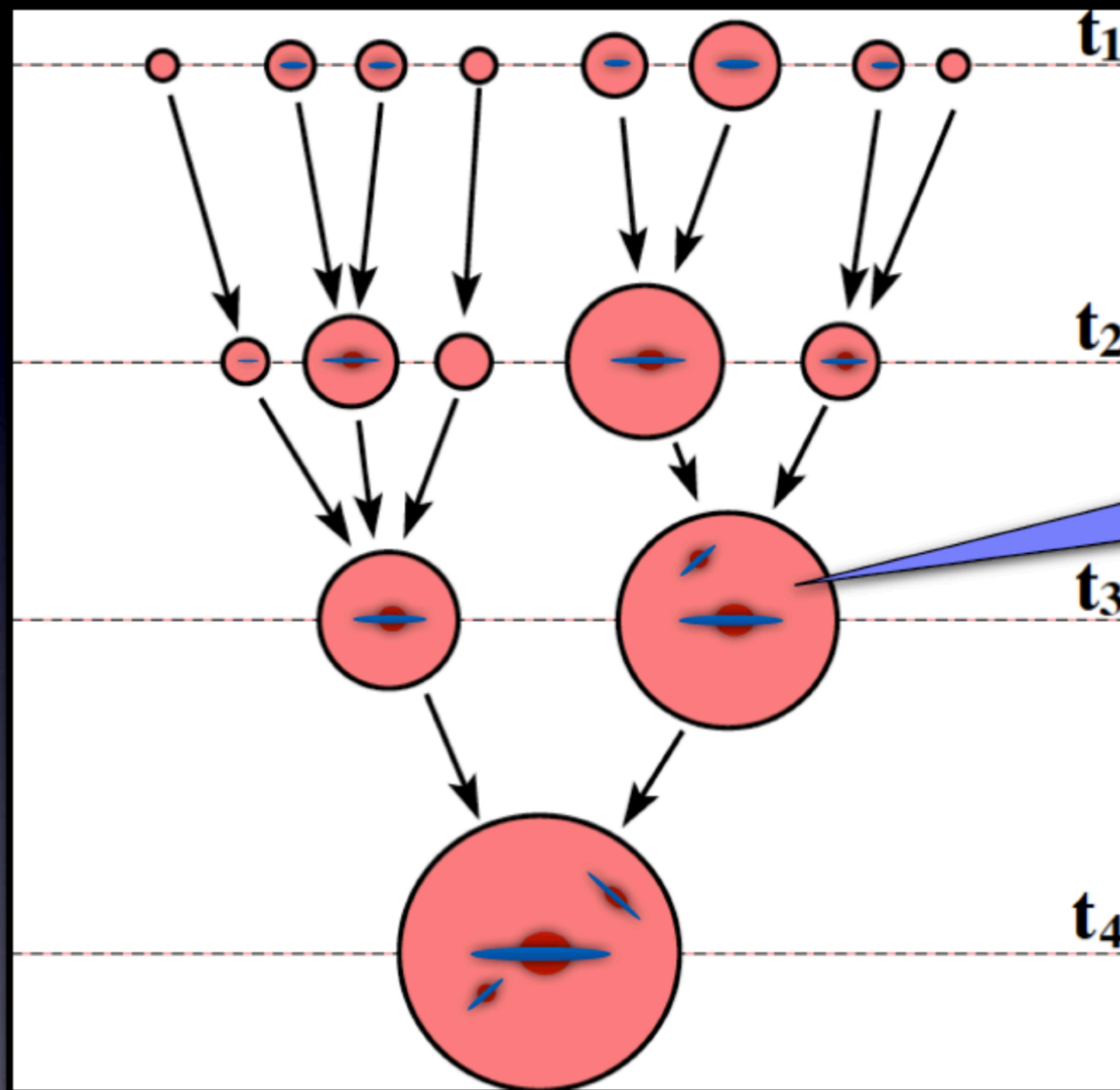
Merger trees



Semi-analytical models



Semi-analytical models



$$M_{\text{vir}} = M_{\text{DM}} + M_{*} + M_{\text{cold}} + M_{\text{hot}}$$

$$\dot{M}_{\text{hot}} = -\dot{M}_{\text{cool}} + \dot{M}_{\text{reheat}} + f_{\text{bar}} \dot{M}_{\text{vir}}$$

$$\dot{M}_{\text{cold}} = \dot{M}_{\text{cool}} - \dot{M}_{\text{reheat}} - \dot{M}_{*}$$

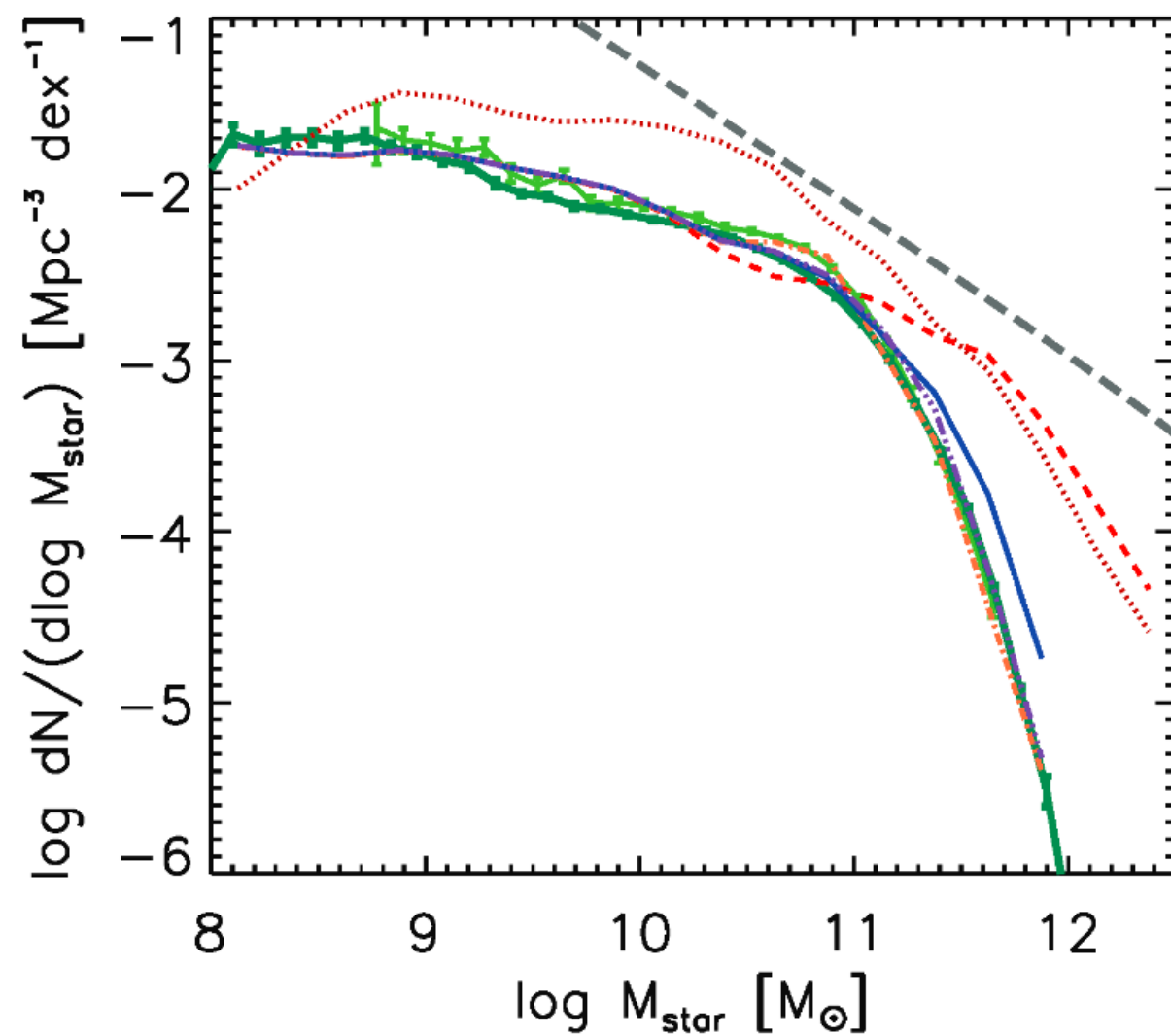
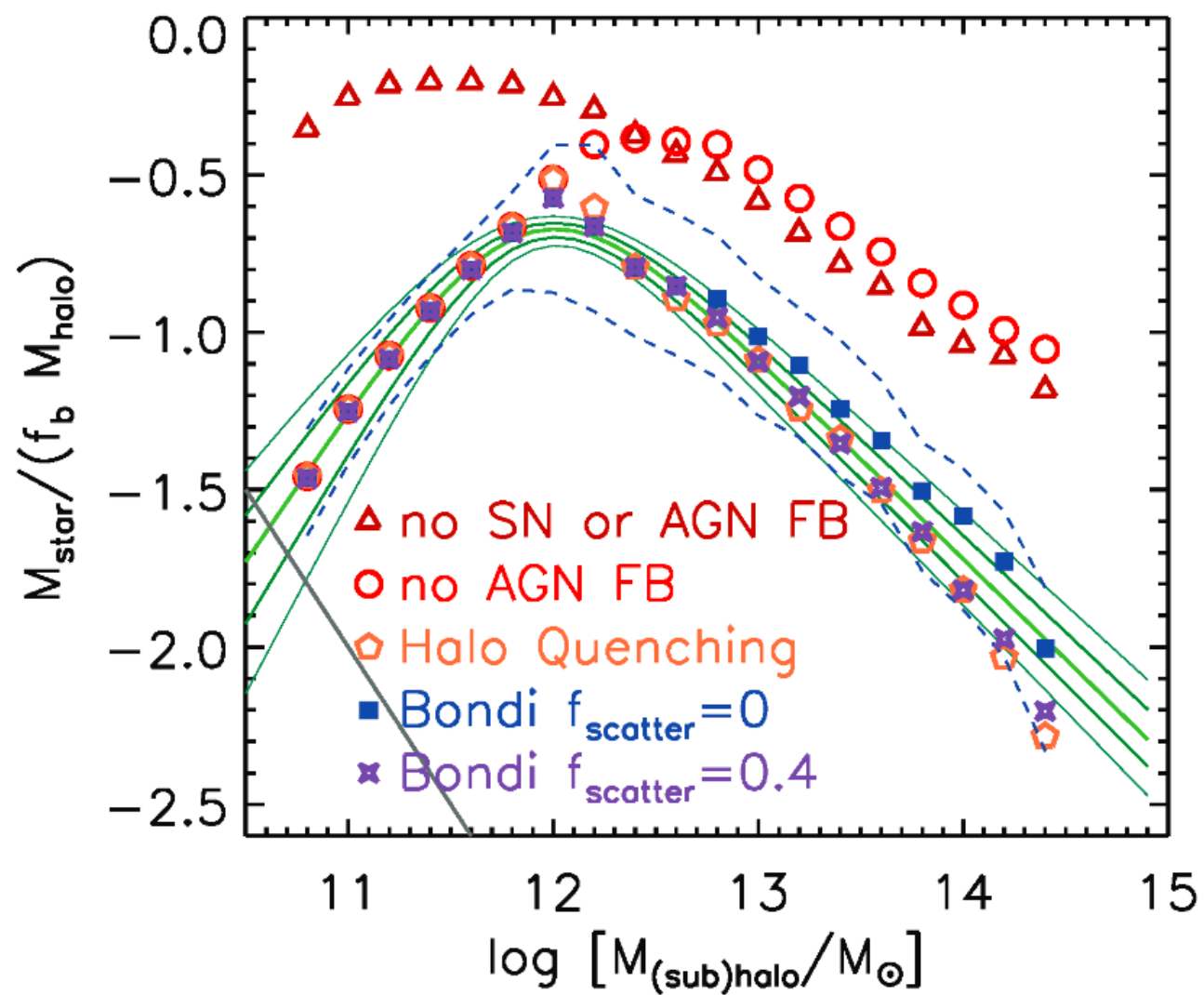
$\dot{M}_{\text{vir}}$ $\longleftrightarrow$ mass accretion

$\dot{M}_{\text{cool}}$ $\longleftrightarrow$ cooling

$\dot{M}_{*}$ $\longleftrightarrow$ star formation

$\dot{M}_{\text{reheat}}$ $\longleftrightarrow$ feedback

Semi-analytical models



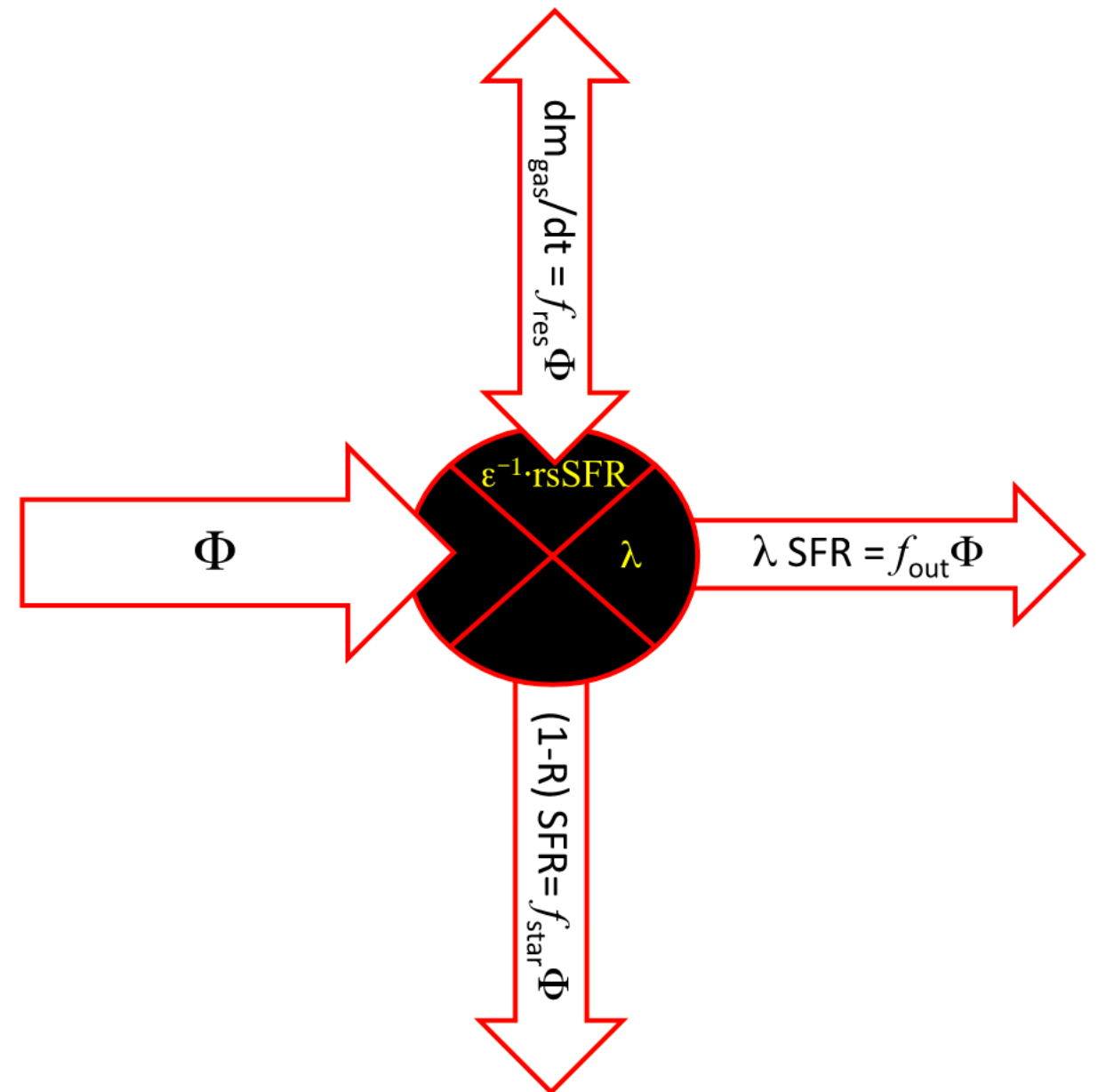
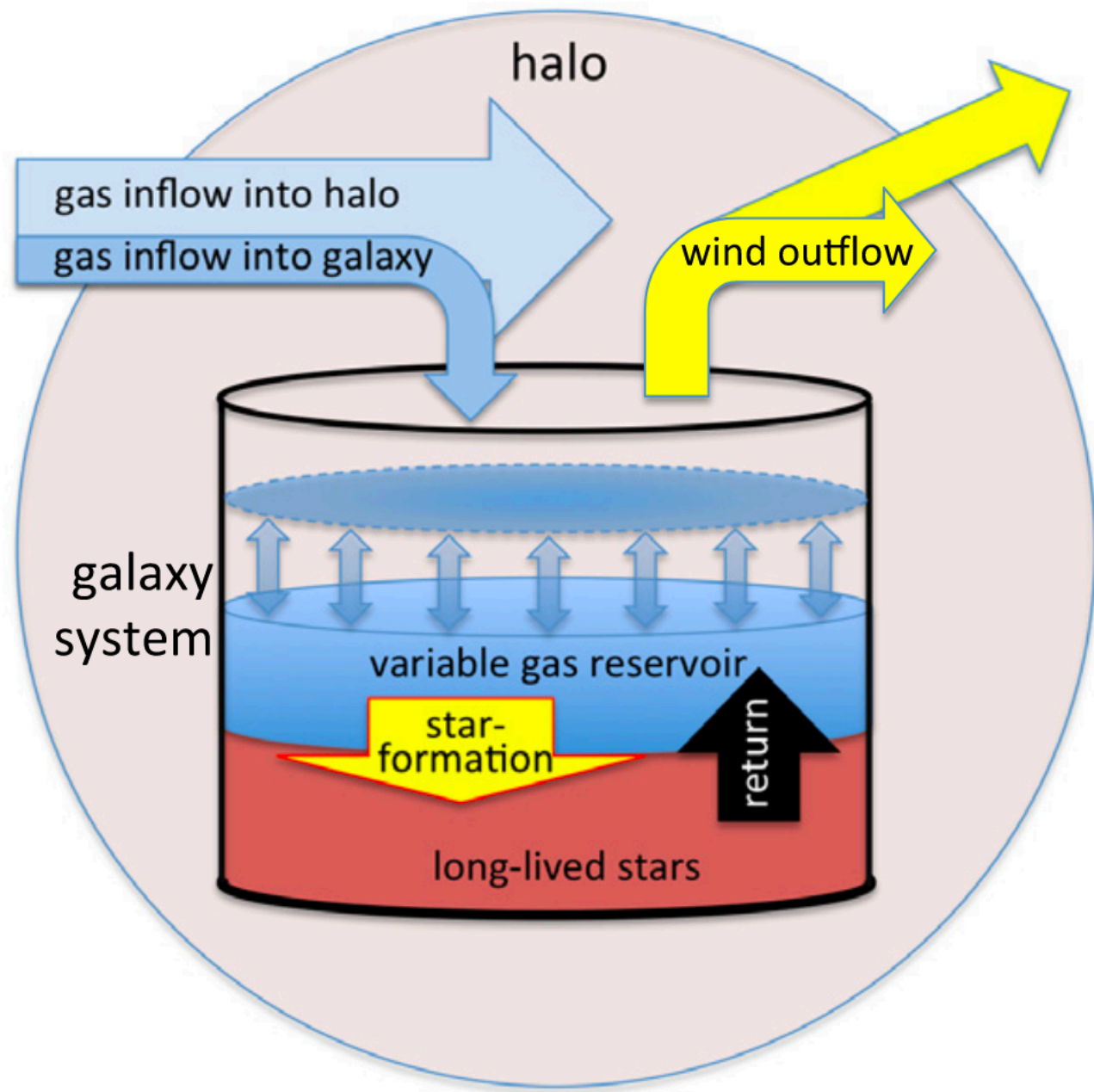
Semi-analytical models

Parameter	Description	Fiducial value	Fixed/adjusted
Photoionization squelching (Section 2.3)			
$z_{\text{overlap}}, z_{\text{re-ionize}}$	Redshift of overlap/re-ionization	11, 10	F
Quiescent star formation (Section 2.5.1)			
A_{Kenn}	Normalization of Kennicutt law ($M_{\odot} \text{ yr}^{-1} \text{ kpc}^{-2}$)	8.33×10^{-5}	A
N_{K}	Power-law index in Kennicutt law	1.4	F
χ_{gas}	Scale radius of gas disc, relative to stellar disc	1.5	A
Σ_{crit}	Critical surface density for star formation ($M_{\odot} \text{ pc}^{-2}$)	6.0	A
Burst star formation (Section 2.5.2)			
μ_{crit}	Critical mass ratio for burst activity	0.1	F
$e_{\text{burst},0}$	Burst efficiency for 1:1 merger	Equation (9)	F
γ_{burst}	Dependence of burst efficiency on mass ratio	Equation (8)	F
τ_{burst}	Burst time-scale	Equation (10)	F
Merger remnants and morphology (Section 2.6)			
f_{sph}	Fraction of stars in spheroidal remnant	Equation (11)	A
f_{scatter}	Fraction of scattered satellite stars	0.4	A
SN feedback (Section 2.7)			
ϵ_{SN}^0	Normalization of reheating function	1.3	A
α_{rh}	Power-law slope of reheating function	2.0	A
V_{eject}	Velocity scale for ejection of reheated gas (km s^{-1})	120	A
$\chi_{\text{re-infall}}$	Time-scale for re-infall of ejected gas	0.1	A
Chemical evolution (Section 2.8)			
y	Chemical yield (solar units)	1.5	A
R	Recycled fraction	0.43	F
Black hole growth (Section 2.9)			
η_{rad}	Efficiency of conversion of rest mass to radiation	0.1	F
M_{seed}	Mass of seed BH ($M_{\odot}$)	100	F
$f_{\text{BH,final}}$	Scaling factor for mass of BH at end of merger	2.0	A
$f_{\text{BH,crit}}$	Scaling factor for ‘critical mass’ of BH	0.4	F
AGN-driven winds (Section 2.1.0)			
ϵ_{wind}	Effective coupling factor for AGN-driven winds	0.5	F
Radio mode feedback (Section 2.1.1)			
κ_{radio}	Normalization of ‘radio mode’ BH accretion rate	3.5×10^{-3}	A
κ_{heat}	Coupling efficiency of radio jets with hot gas	1.0	F

Semi-analytical models

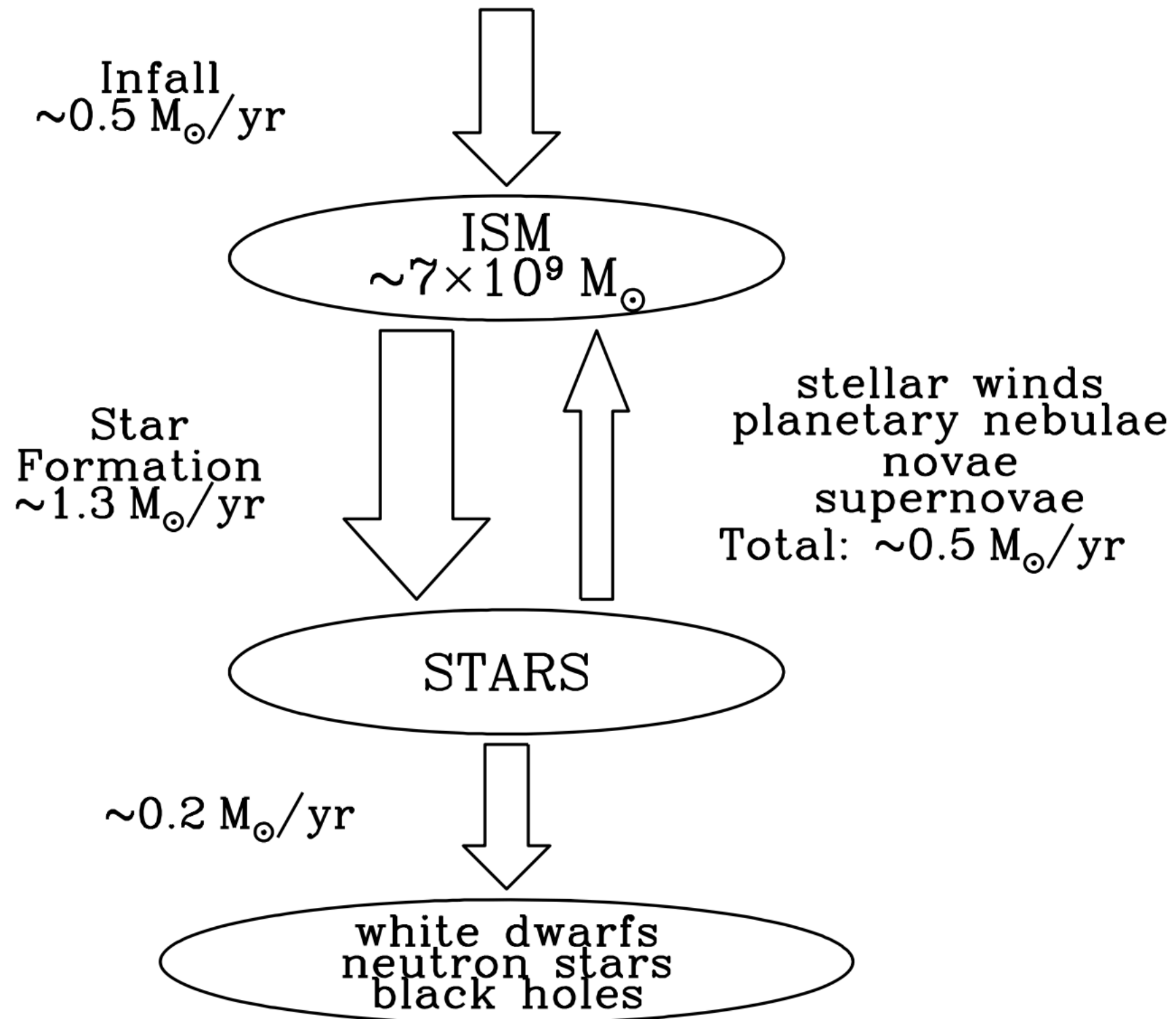
Parameter	Value	Reference
[H_O]	70.2 km/s	Section 4.2; (Komatsu et al., 2010)
[Omega_O]	0.2725	Section 4.2; (Komatsu et al., 2010)
[Omega_DE]	0.7275	Section 4.2; (Komatsu et al., 2010)
[Omega_b]	0.0455	Section 4.2; (Komatsu et al., 2010)
[T_CMB]	2.72548 K	Section 4.2; (Komatsu et al., 2010)
[accretionDisksMethod]	ADAF	Section 4.3
[adafAdiabaticIndex]	1.444	Section 4.3
[adafEnergyOption]	Pure ADAF	Section 4.3
[adafRadiativeEfficiency]	0.01	Section 4.3
[adafViscosityOption]	Fit	Section 4.3
[adiabaticContractionGnedinA]	0.8	Section 4.8
[adiabaticContractionGnedinOmega]	0.77	Section 4.8
[barInstabilityMethod]	ELN	Section 4.7
[blackHoleSeedMass]	100	Section 3.1.2
[blackHoleWindEfficiency]	0.001	Section 3.1.2
[bondiHoyleAccretionEnhancementHotHalo]	1	Section 3.1.2
[bondiHoyleAccretionEnhancementSpheroid]	1	Section 3.1.2
[bondiHoyleAccretionTemperatureSpheroid]	100	Section 3.1.2
[coolingFunctionMethod]	Atomic CIE Cloudy	Section 4.5.1
[coolingTimeAvailableAgeFactor]	0	Section 4.5.5
[coolingTimeSimpleDegreesOfFreedom]	3	Section 4.5.4
[darkMatterProfileMethod]	NFW	Section 4.6.1
[darkMatterProfileMinimumConcentration]	4	Section 3.8.2
[diskOutflowExponent]	2	Section 4.2.3
[diskOutflowVelocity]	200 km/s	Section 4.2.3
[effectiveNumberNeutrinos]	4.34	Section 4.4.2
[galacticStructureRadiusSolverMethod]	Adiabatic	Section 4.8
[haloMassFunctionMethod]	Tinker2008	Section 4.4.6
[haloSpinDistributionMethod]	Bett2007	Section 4.6.3
[hotHaloOutflowReturnRate]	1.26	Section 3.2.2
[imfSalpeterRecycledInstantaneous]	0.39	Section 4.12.2
[imfSalpeterYieldInstantaneous]	0.02	Section 4.12.2
[imfSelectionFixed]	Salpeter	Section 4.12.1
[isothermalCoreRadiusOverVirialRadius]	0.1	Section 4.10
[majorMergerMassRatio]	0.1	Section 4.9.1
[mergerRemnantSizeOrbitalEnergy]	1	Section 4.9.2
[mergerTreeBuildCole2000AccretionLimit]	0.1	Section 4.16
[mergerTreeBuildCole2000MassResolution]	$5 \times 10^9 M_\odot$	Section 4.16
[mergerTreeBuildCole2000MergeProbability]	0.1	Section 4.16
[mergerTreeConstructMethod]	Build	Section 4.14
[minorMergerGasMovesTo]	Spheroid	Section 4.9.1
[modifiedPressSchechterFirstOrderAccuracy]	0.1	Section 4.15
[modifiedPressSchechterG0]	0.57	Section 4.15
[modifiedPressSchechterGamma1]	0.38	Section 4.15
[modifiedPressSchechterGamma2]	-0.01	Section 4.15
[powerSpectrumIndex]	0.961	Section 4.4.1; (Komatsu et al., 2010)
[powerSpectrumReferenceWavenumber]	1 Mpc^{-1}	Section 4.4.1; (Komatsu et al., 2010)
[powerSpectrumRunning]	0	Section 4.4.1; (Komatsu et al., 2010)
[randomSpinResetMassFactor]	2	Section 3.7.2
[reionizationSuppressionRedshift]	9	Section 4.1
[reionizationSuppressionVelocity]	30 km/s	Section 4.1
[satelliteMergingMethod]	Jiang2008	Section 4.22.1
[sigma_8]	0.807	Section 4.4.1 & Section 4.4.2
[spheroidEnergeticOutflowMassRate]	1	Section 3.4.2
[spheroidOutflowExponent]	2	Section 4.2.3
[spheroidOutflowVelocity]	50 km/s	Section 4.2.3
[spinDistributionBett2007Alpha]	2.509	Section 4.6.3
[spinDistributionBett2007Lambda0]	0.04326	Section 4.6.3
[stabilityThresholdGaseous]	0.9	Section 4.7
[stabilityThresholdStellar]	1.1	Section 4.7
[starFormationDiskEfficiency]	0.01	Section 4.17
[starFormationDiskMinimumTimescale]	0.001 Gyr	Section 4.17
[starFormationDiskVelocityExponent]	-1.5	Section 4.17
[starFormationSpheroidEfficiency]	0.1	Section 4.17
[starFormationSpheroidMinimumTimescale]	0.001 Gyr	Section 4.17
[starveSatellites]	True	Section 3.2.2
[stellarPopulationPropertiesMethod]	Instantaneous	Section 4.18
[summedNeutrinoMasses]	0	Section 4.4.2
[transferFunctionMethod]	Eisenstein + Hu	Section 4.4.2
[virialDensityContrastMethod]	Spherical top hat	Section 4.4.5

Regulator models

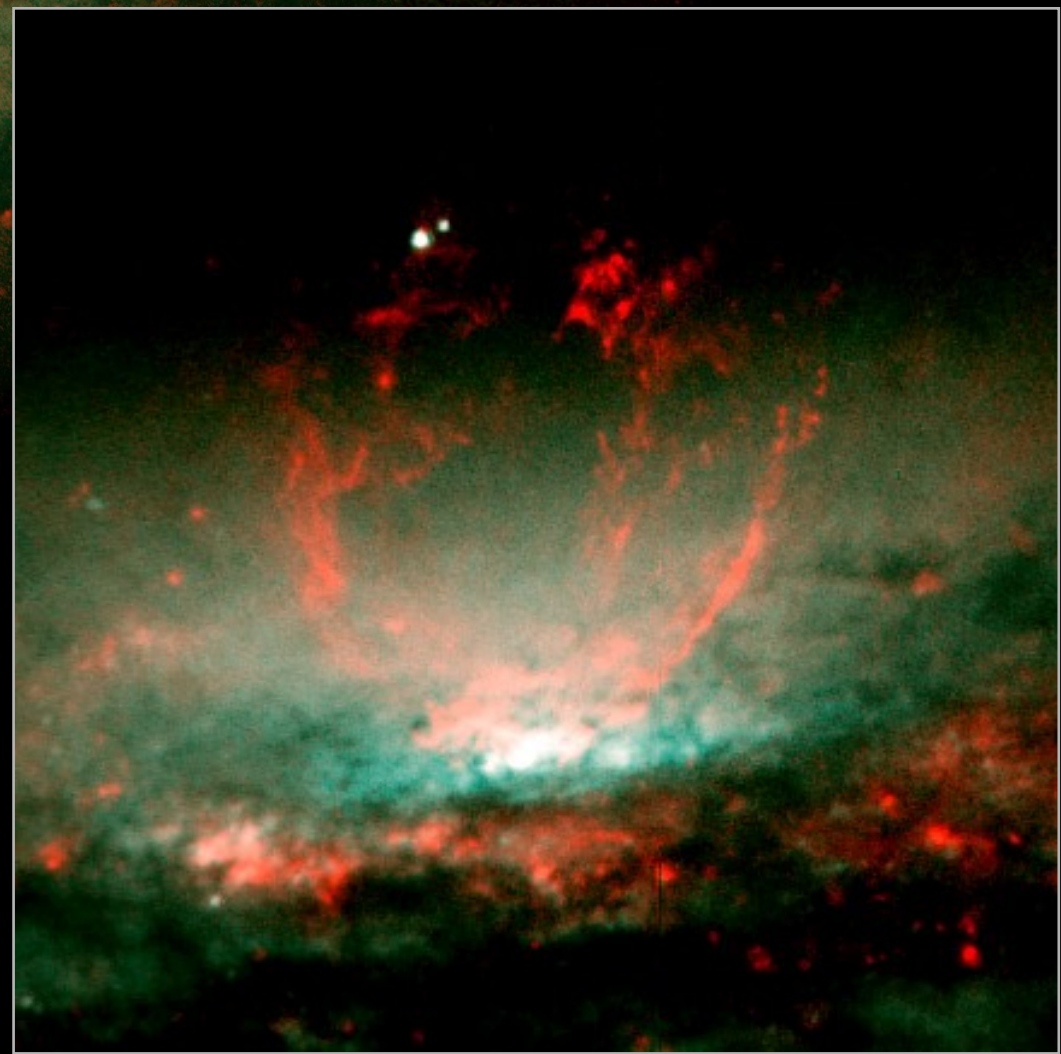
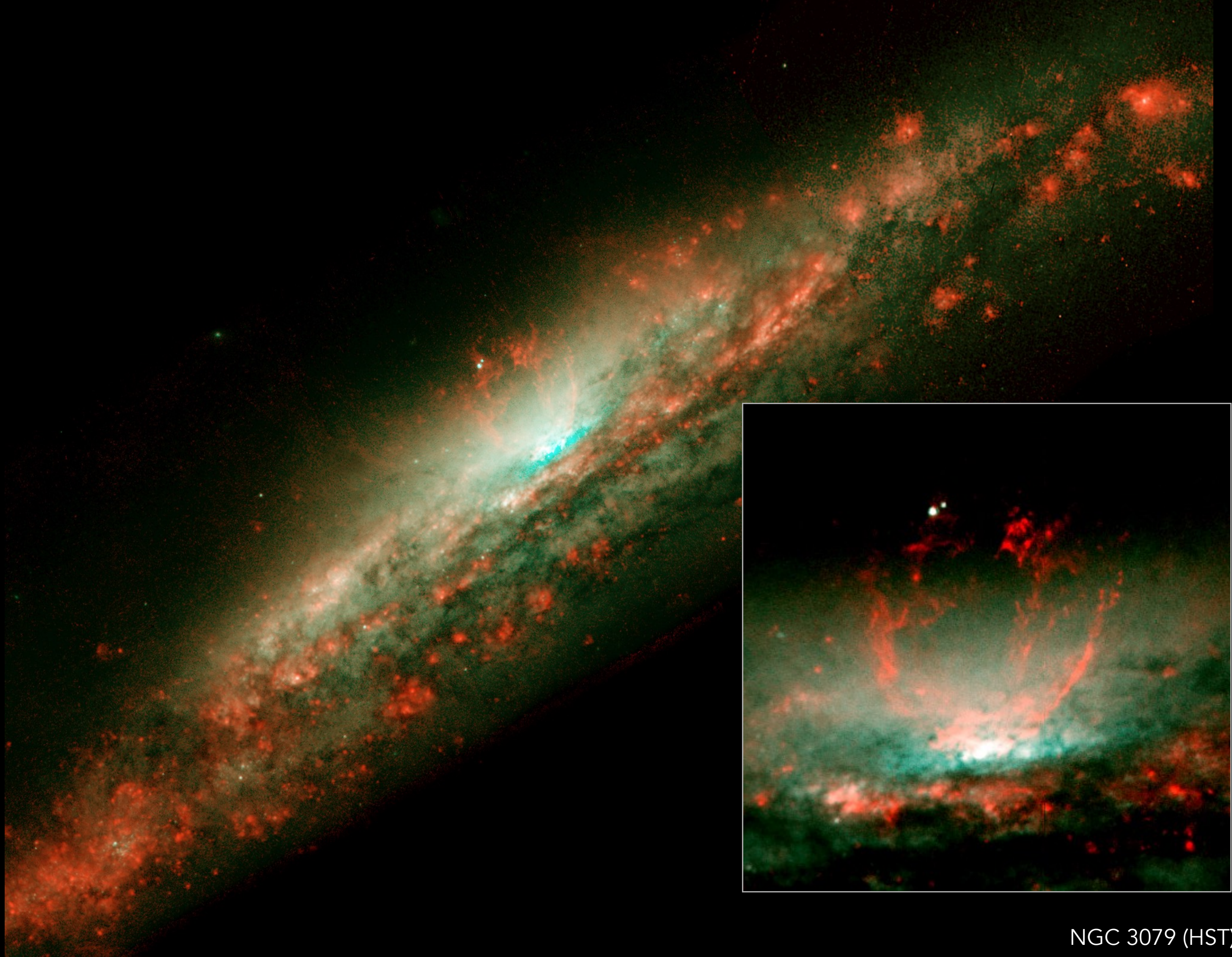


§8.2 • Star formation and gas recycling

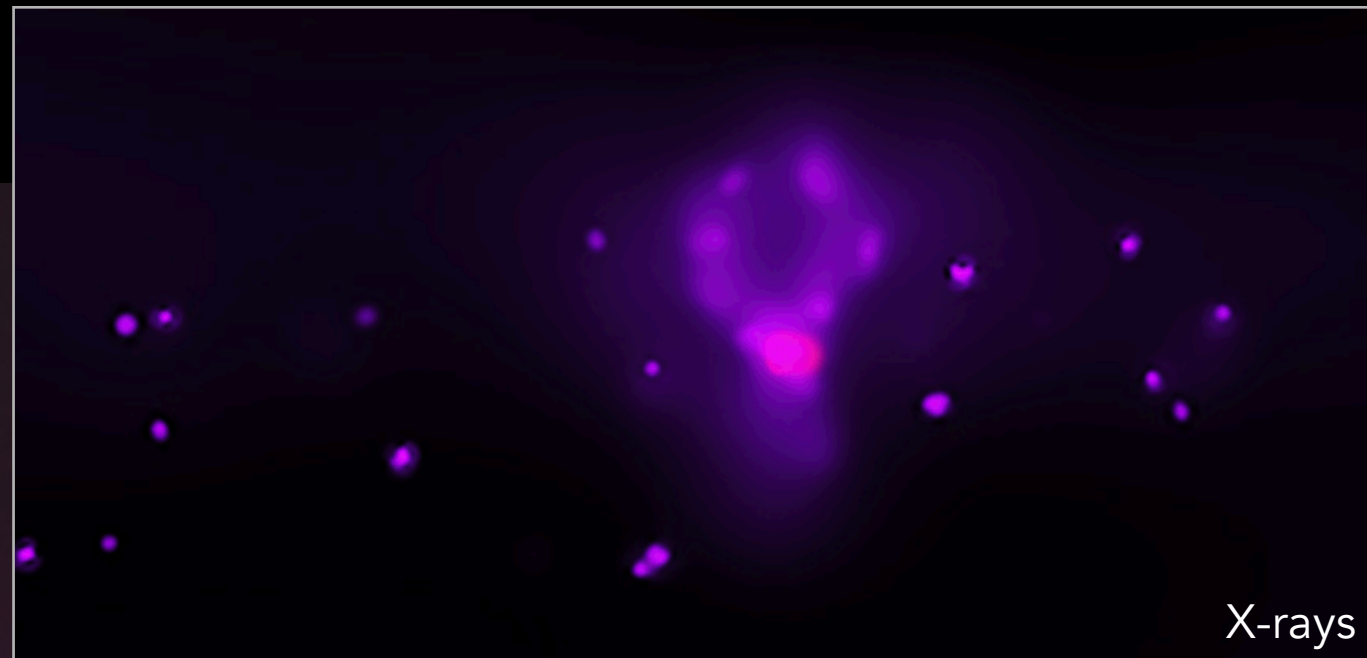
Flow of baryons in the MW



§8.4 • Feedback and galactic winds



NGC 3079 (HST)



NGC 3079 (HST + Chandra)



M82 (HST)

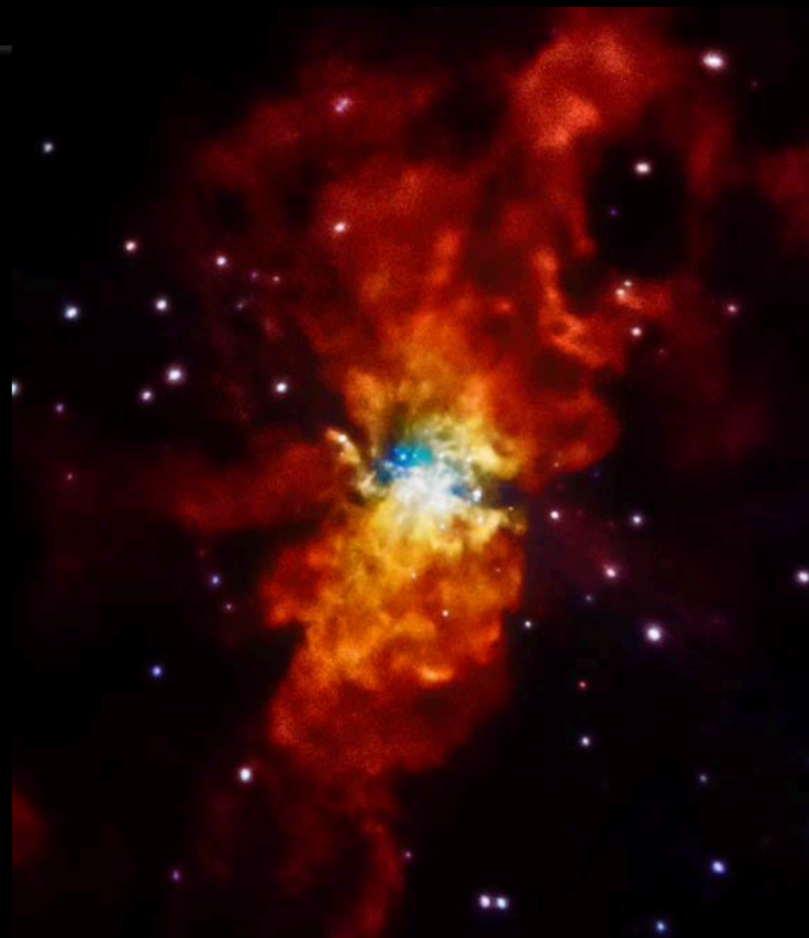
Multiphase outflows (M82)



H α

Warm ionized gas

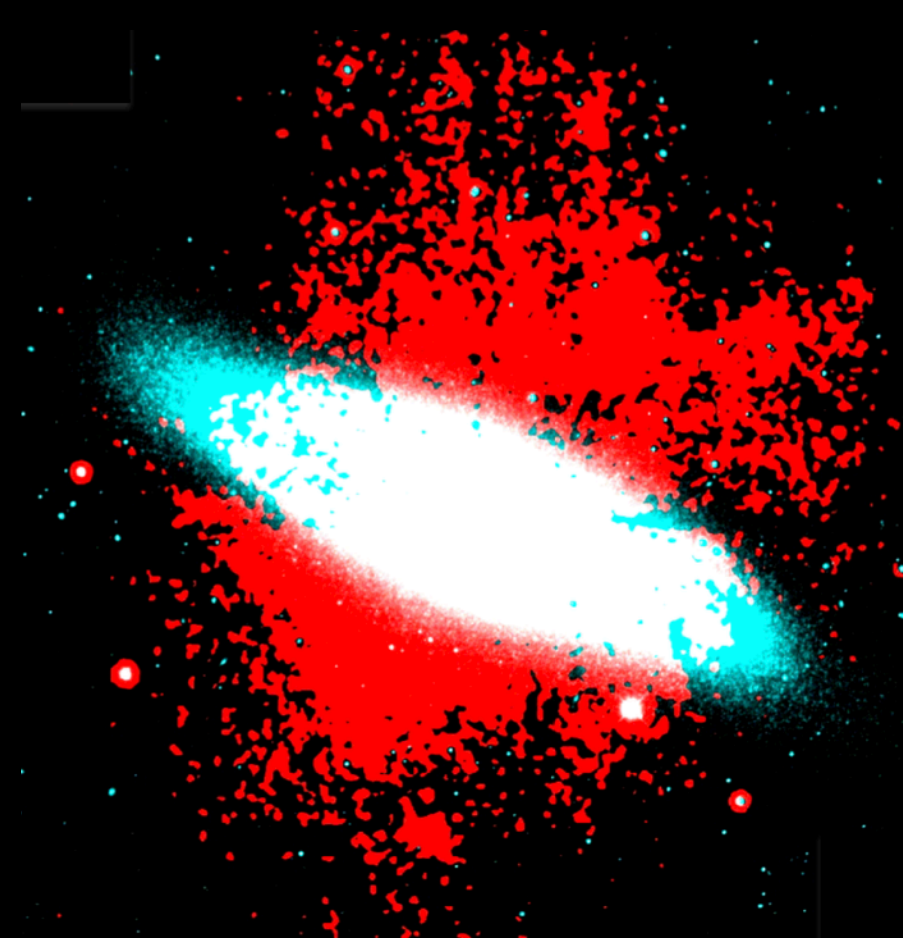
$$T \sim 10^4 \text{ K}$$



X-rays

Hot ionized gas

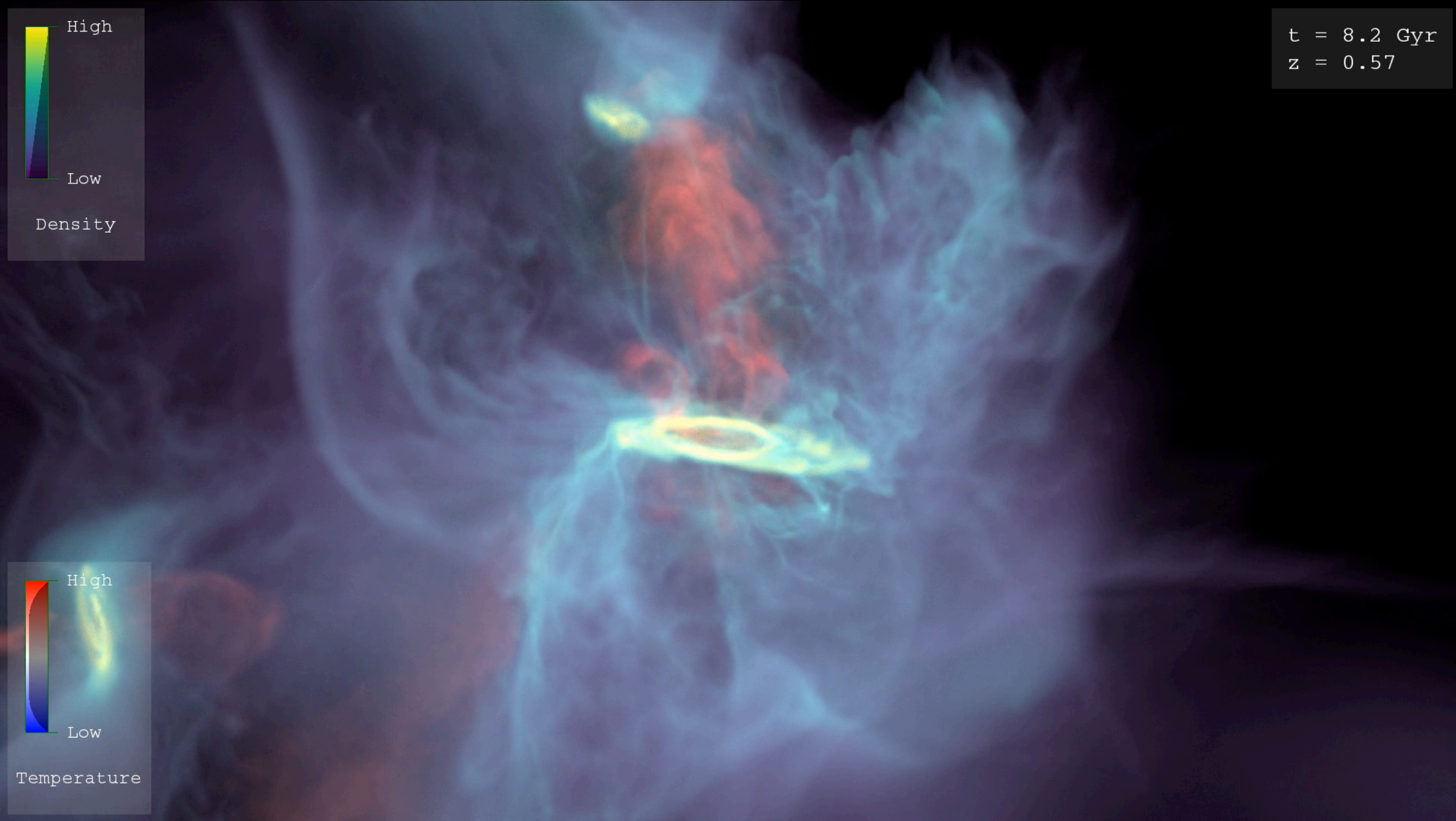
$$T \sim 10^{6-7} \text{ K}$$



2.12 μm

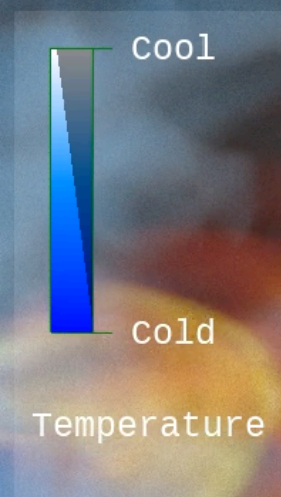
Warm Molecular H₂

$$T \sim 10^3 \text{ K}$$





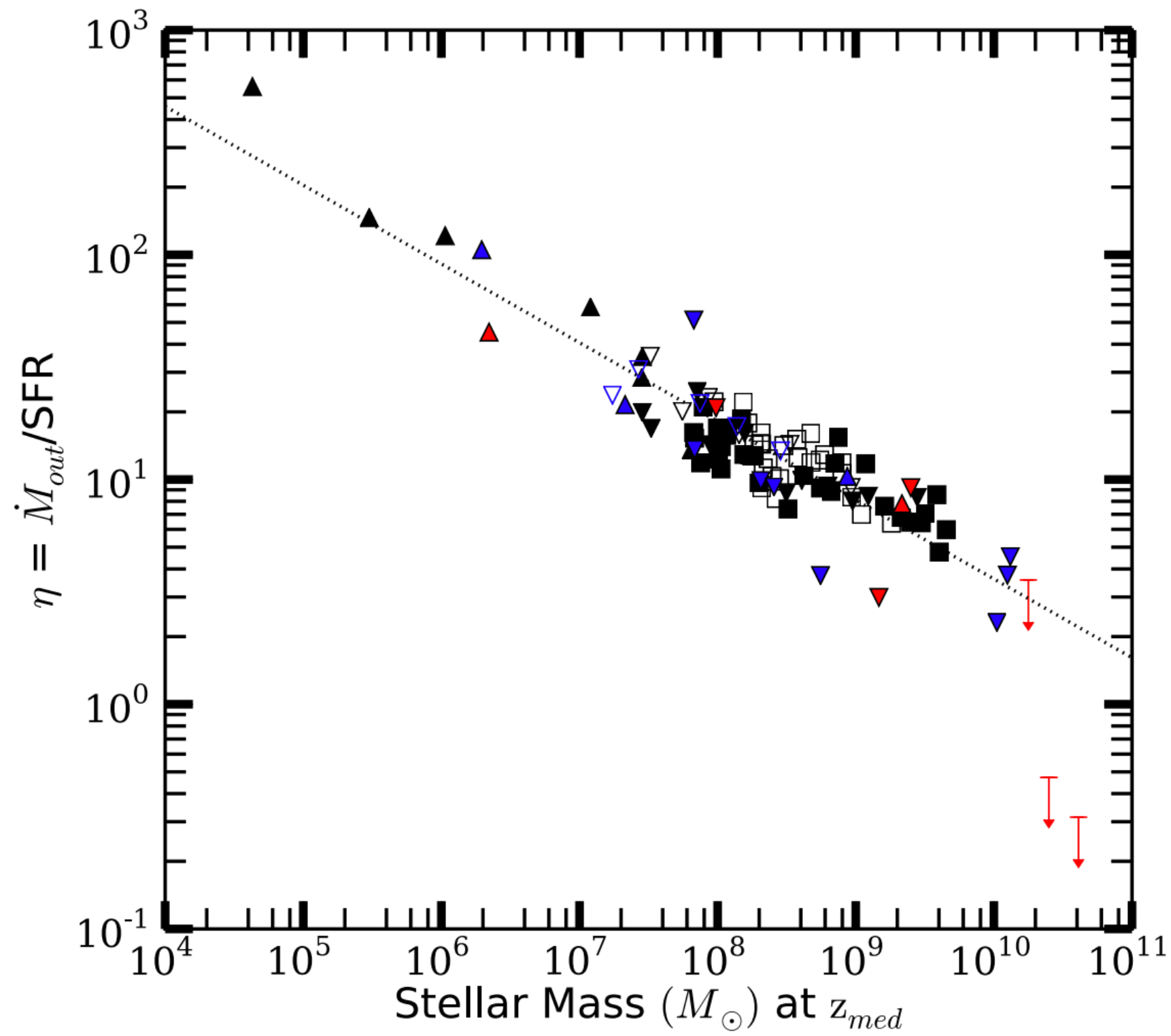
$t = 7.9 \text{ Gyr}$
 $z = 0.62$



$z=0.82$



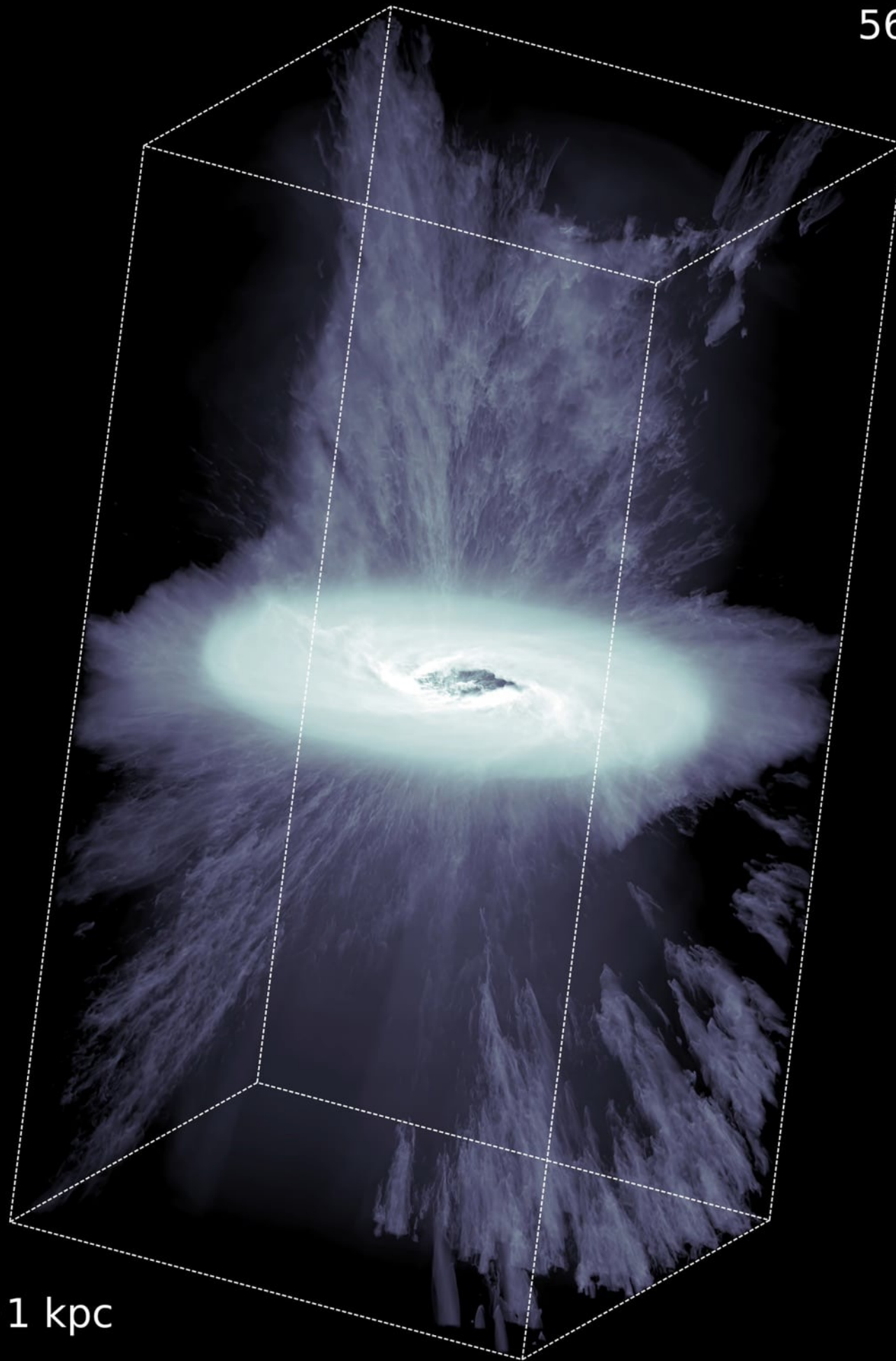
Mass loading factor



56 Myr

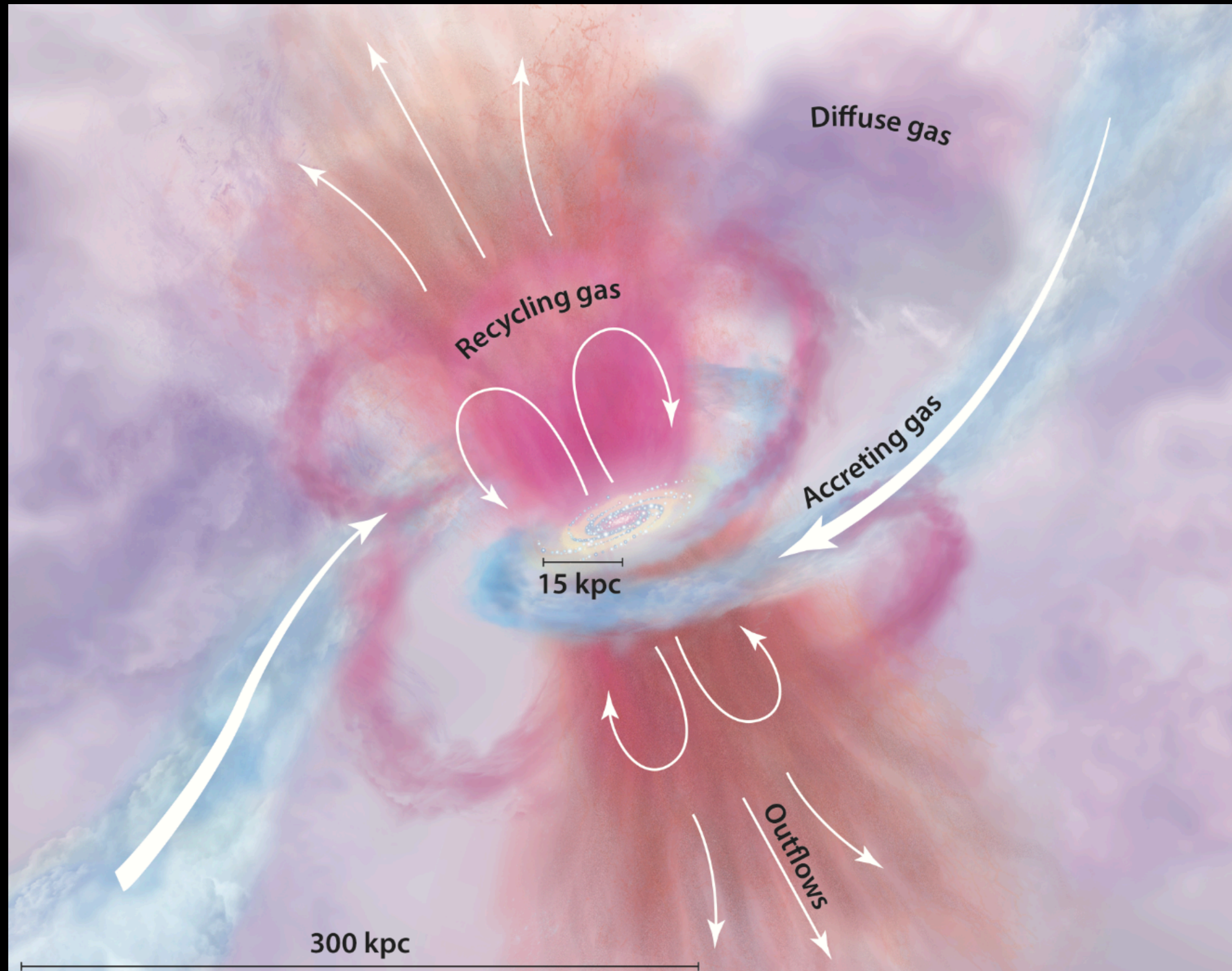
— 1 kpc

Video by Evan Schneider ([vimeo](#))

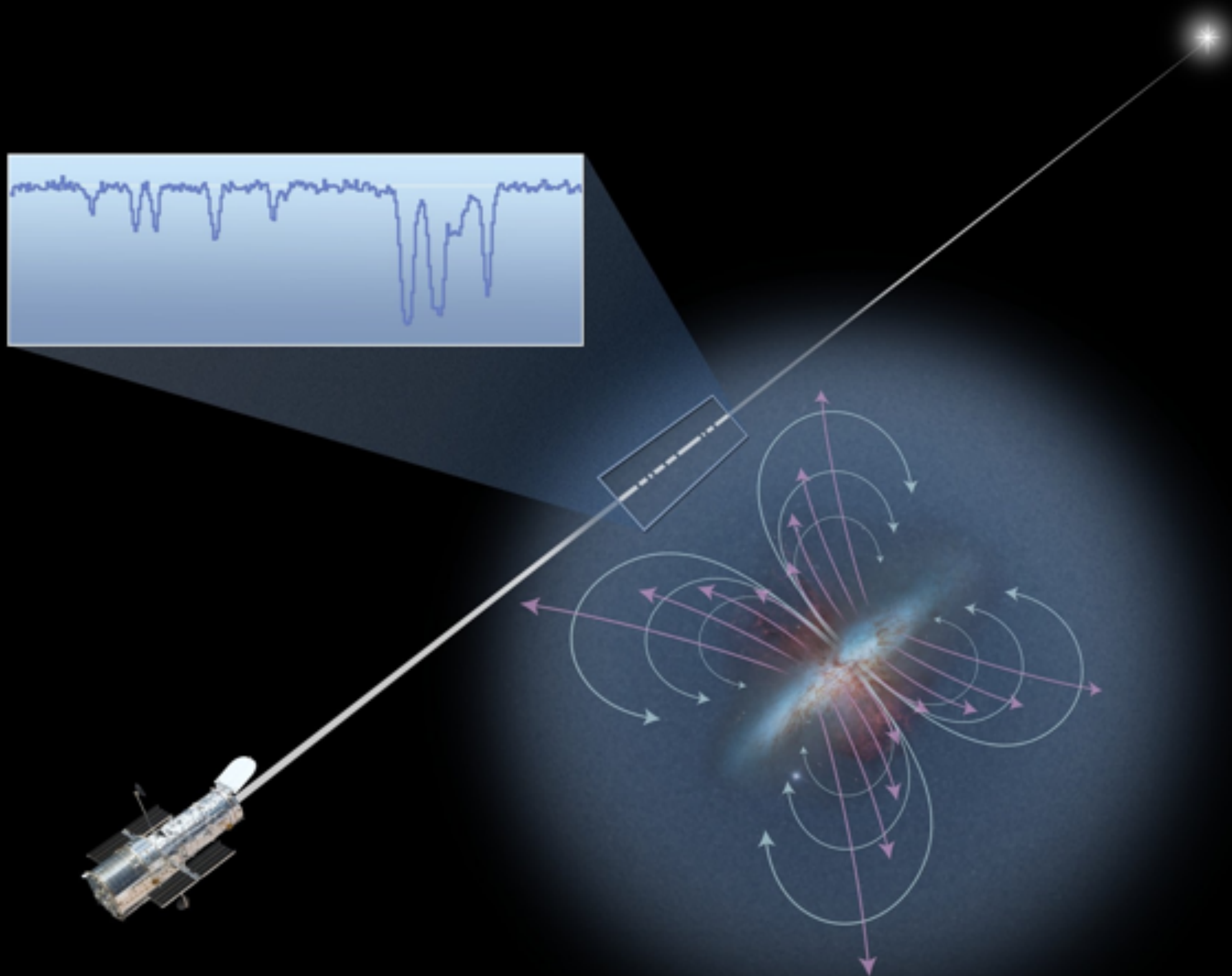




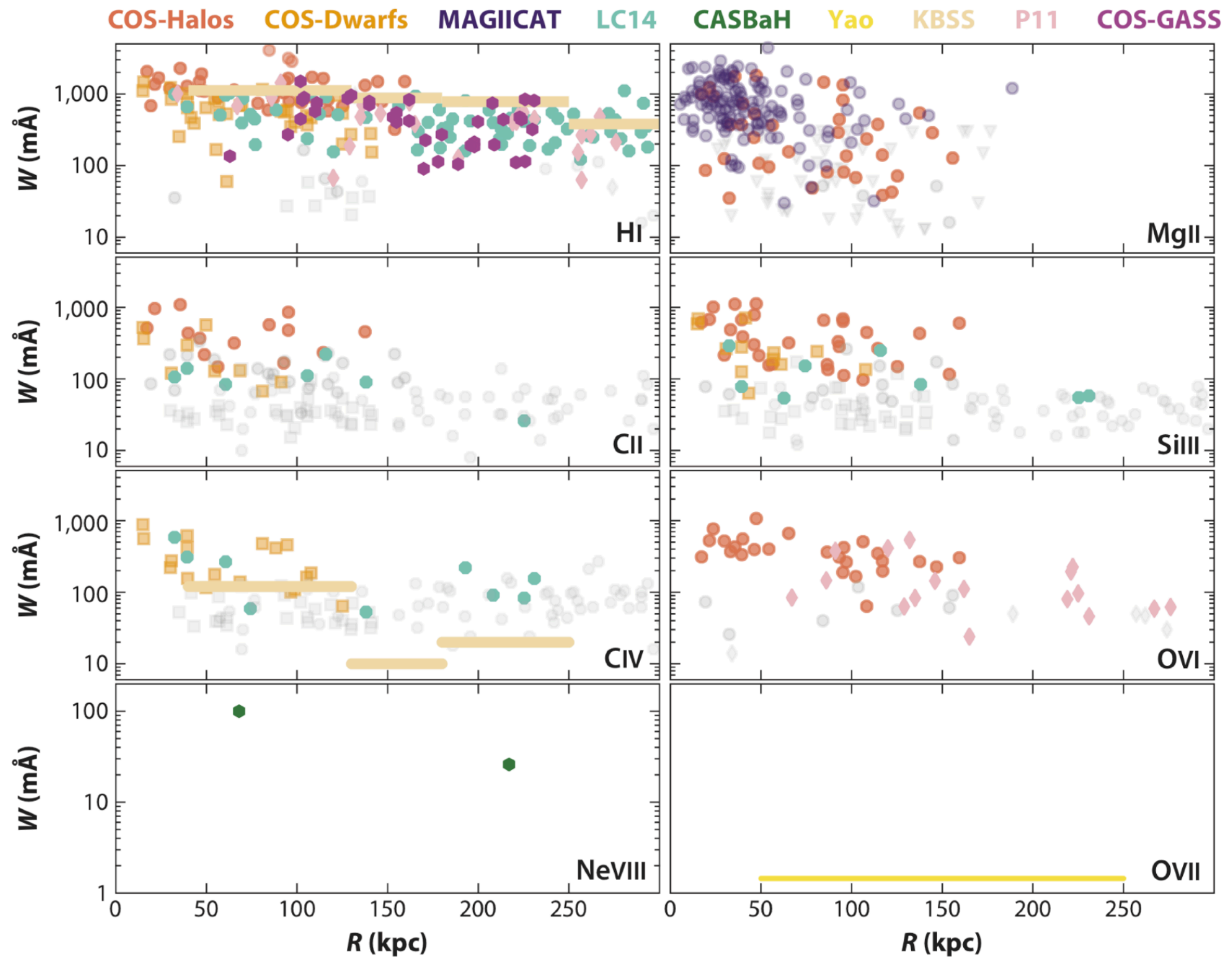
The circum-galactic medium (CGM)



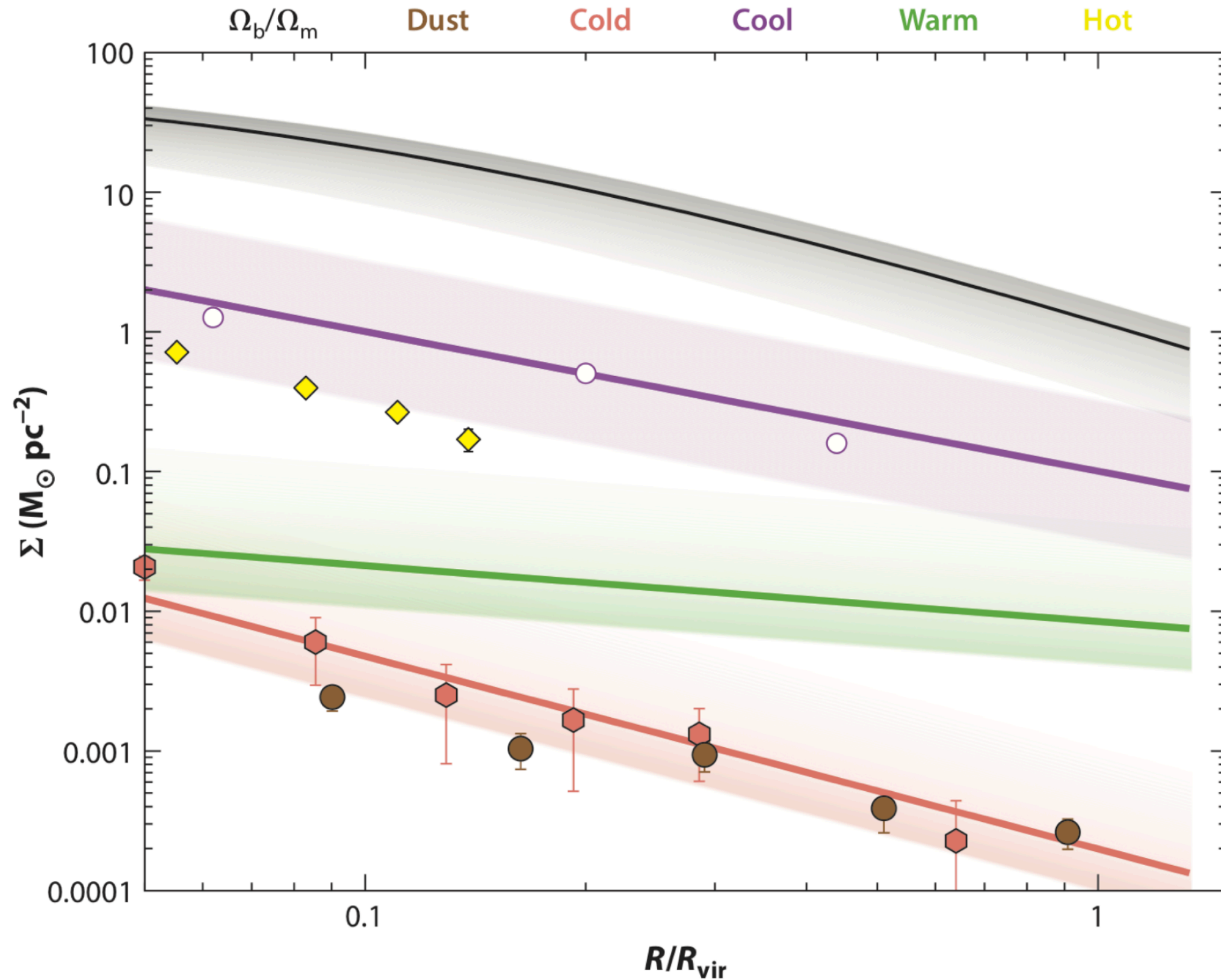
The circum-galactic medium (CGM)



The circum-galactic medium (CGM)



The circum-galactic medium (CGM)



§8.5 • Chemical evolution

Stellar yields

Table 8.1 Stellar yields			
Initial mass ($\mathcal{M}_{\text{ini}}$) ($\mathcal{M}_{\odot}$)	Ejection phase	Species	$y_i \mathcal{M}_{\text{ini}}$ ($10^{-3} \mathcal{M}_{\odot}$)
11–40 ^a	SN II	^{16}O	53–5720
		^{20}Ne	31–1240
		^{28}Si	17–345
		^{12}C	24–259
		^{24}Mg	≤ 235
		^{32}S	≤ 159
		^{56}Fe	11–26
		^{40}Ca	≤ 10
7–8	AGB	^{14}N	68–88
		^{13}C	~ 1
4–6 ^b	AGB	^{14}N	3–52
		^{12}C	< 19
2.5–4	AGB	^{12}C	4–20
		^{14}N	0.5–7
$\lesssim 1.4$ ^c	SN Ia	^{56}Fe	610
		^{28}Si	160
		^{54}Fe	140
		^{24}Mg	90
		^{32}S	80
		^{58}Ni	60

Origin of the elements

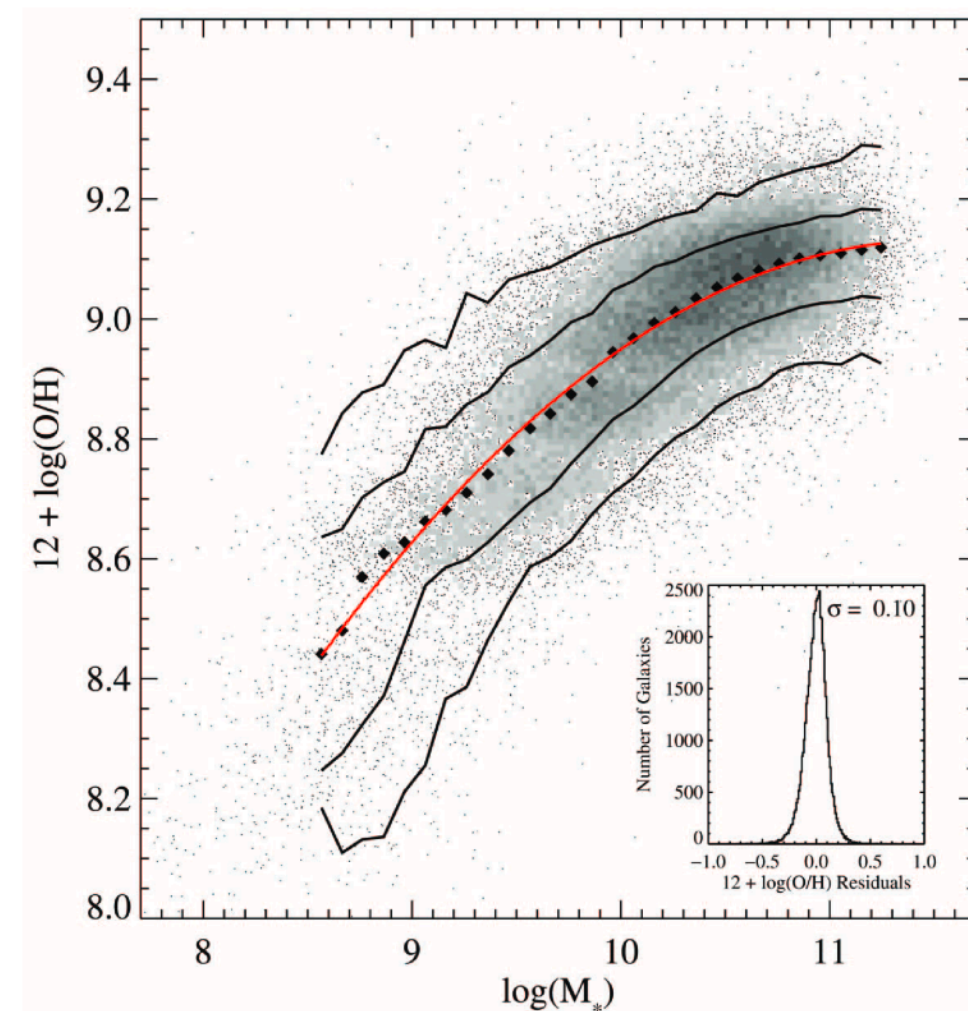
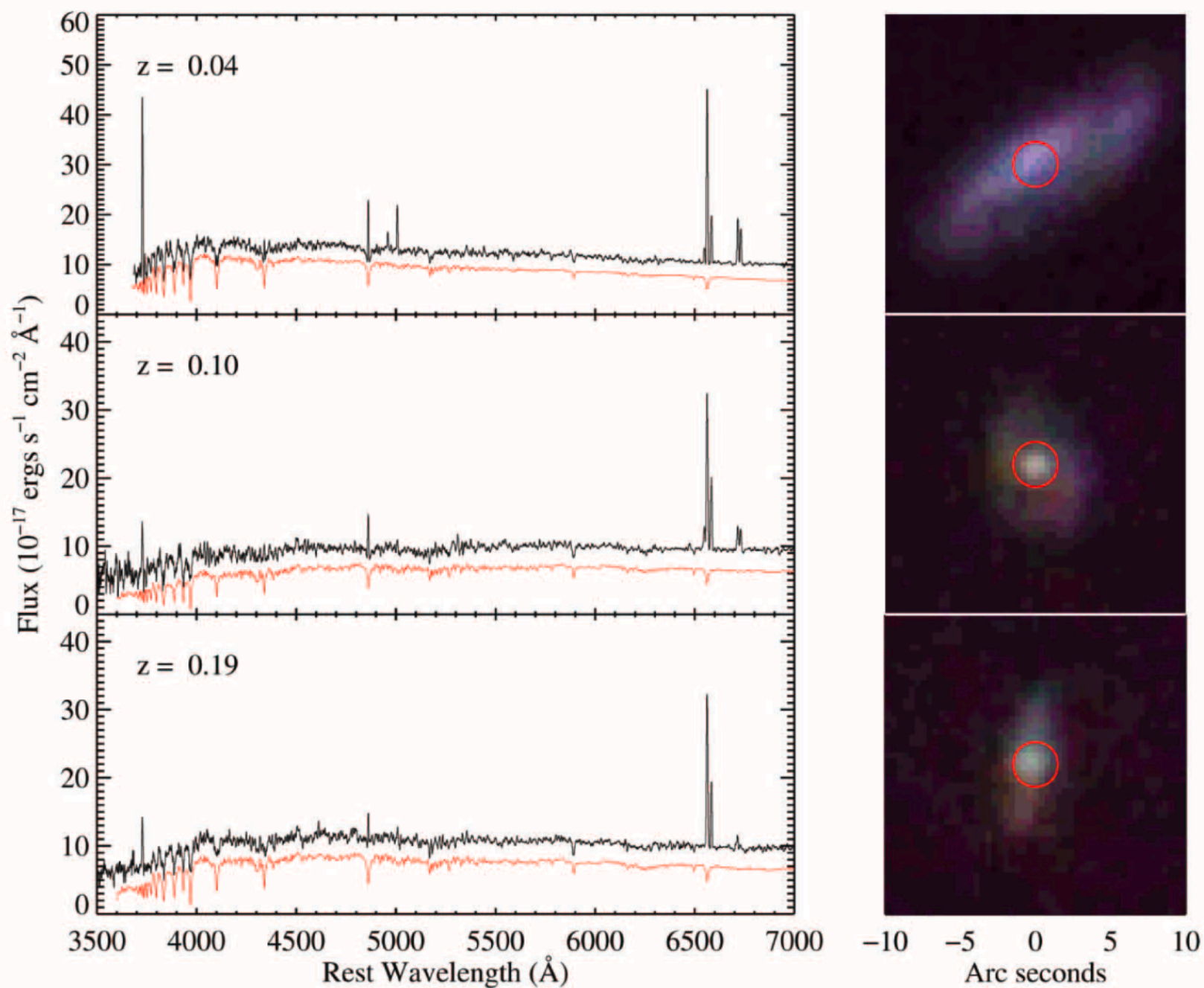
1 H																	2 He	
3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne	
11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar	
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr	
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe	
55 Cs	56 Ba			72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
87 Fr	88 Ra																	
			57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu	
			89 Ac	90 Th	91 Pa	92 U												

Merging Neutron Stars
Dying Low Mass Stars

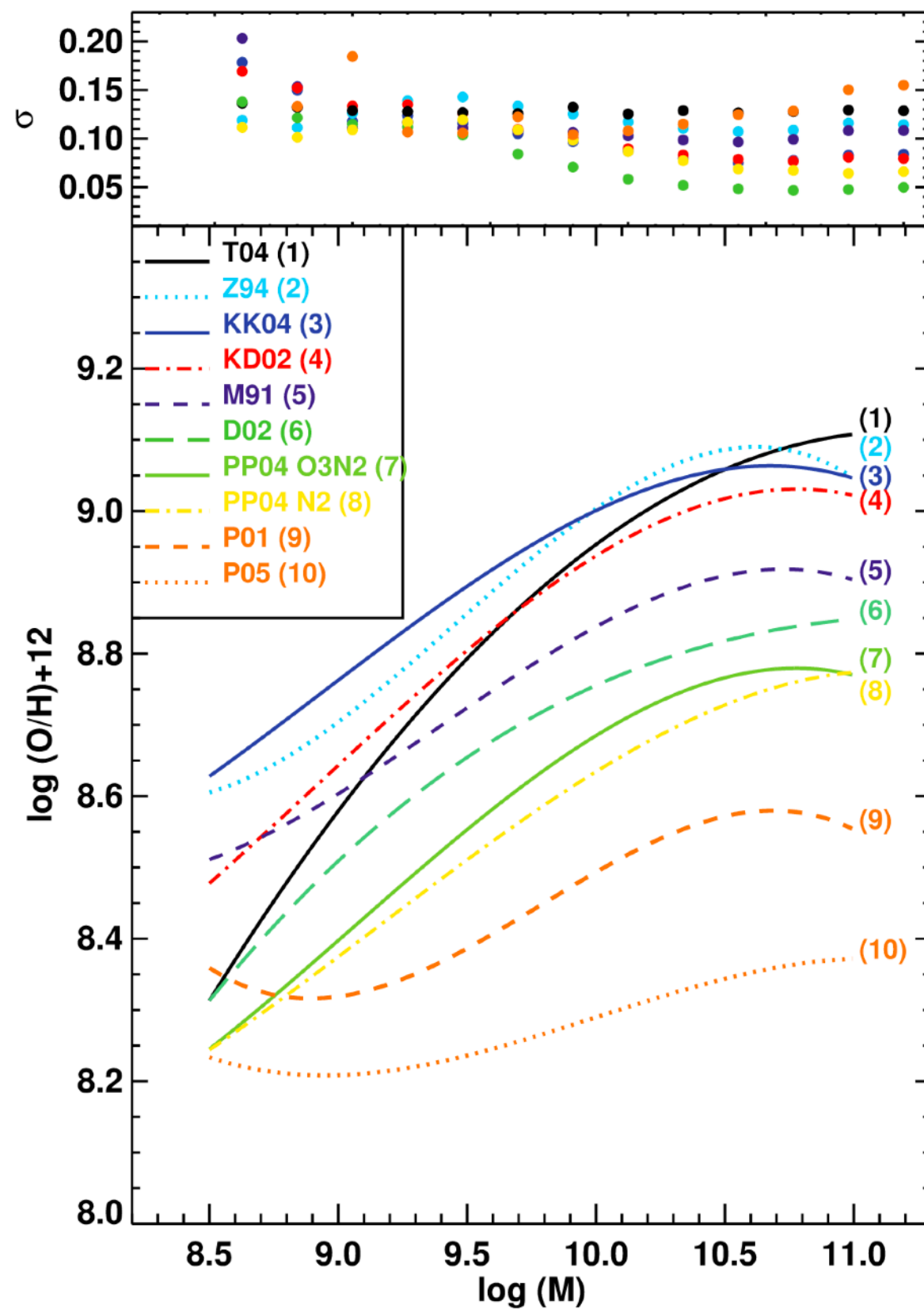
Exploding Massive Stars
Exploding White Dwarfs

Big Bang
Cosmic Ray Fission

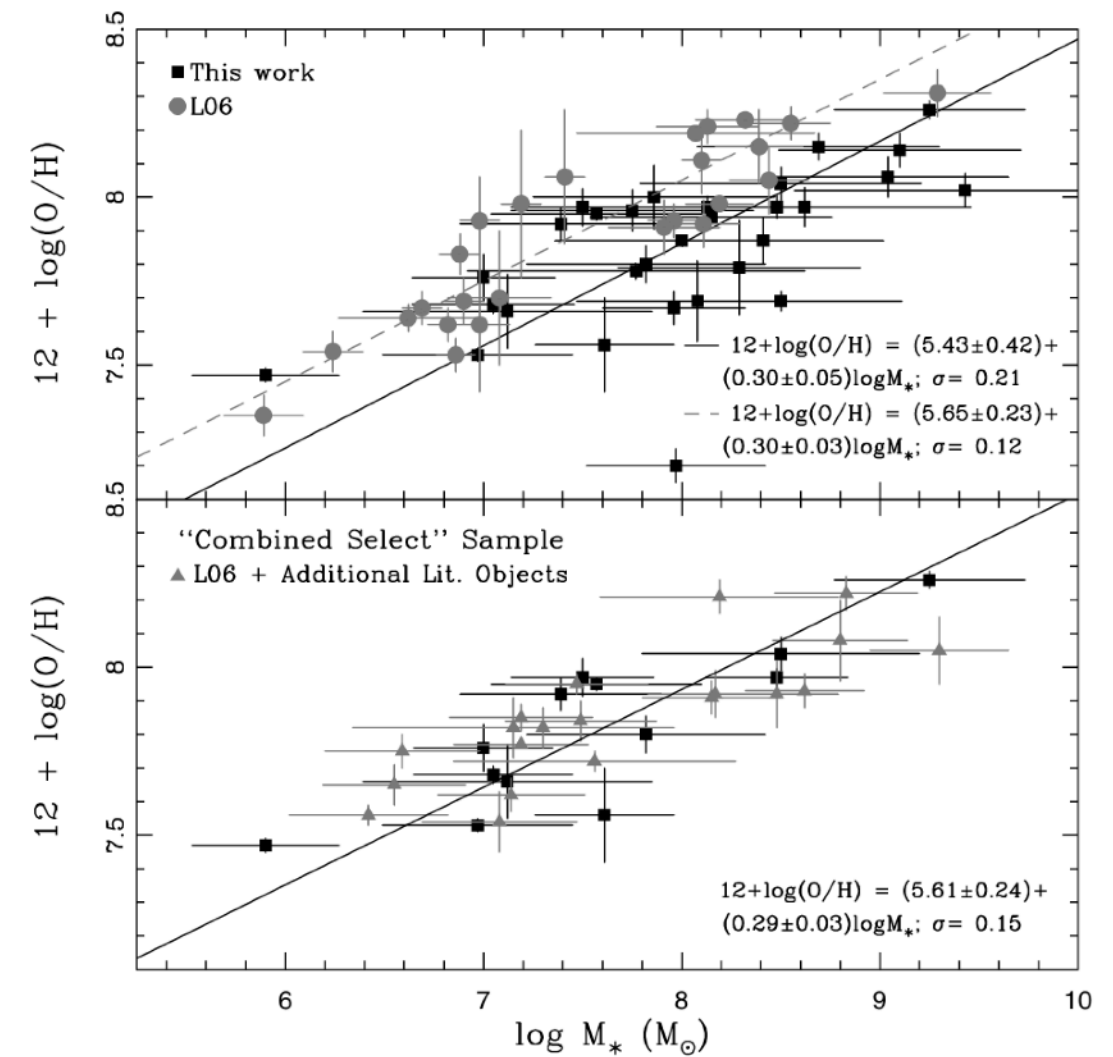
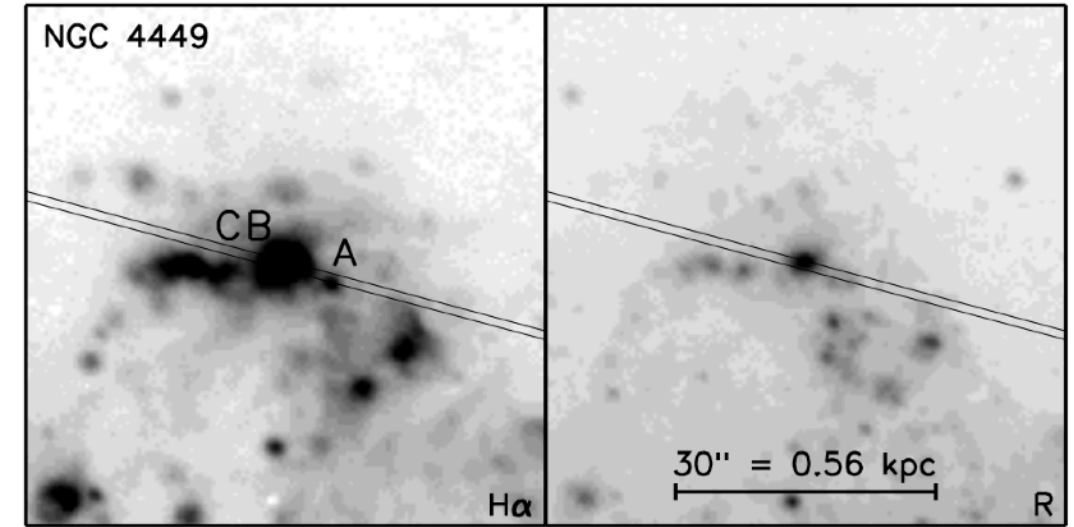
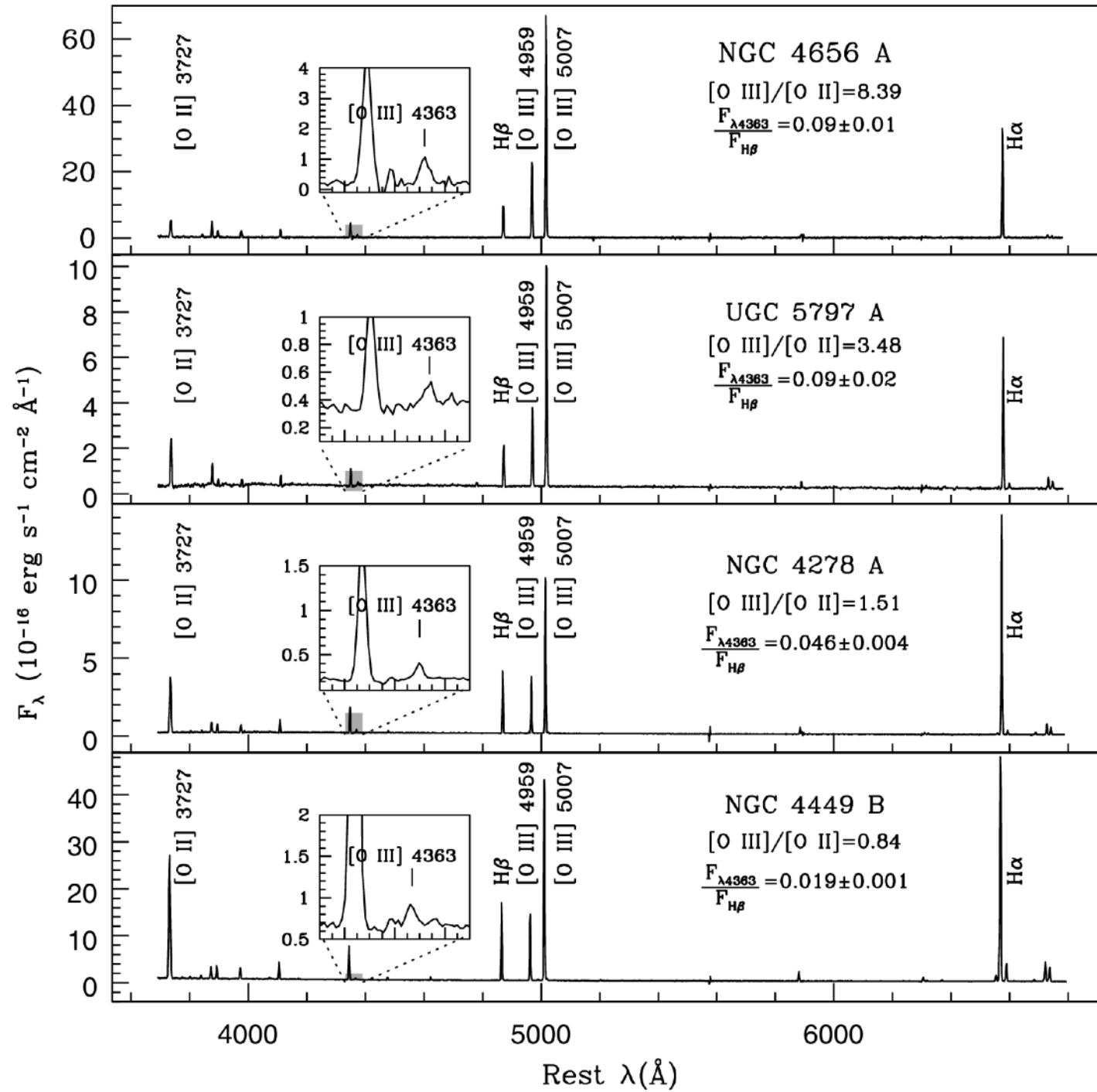
Mass-metallicity relation



Mass-metallicity relation

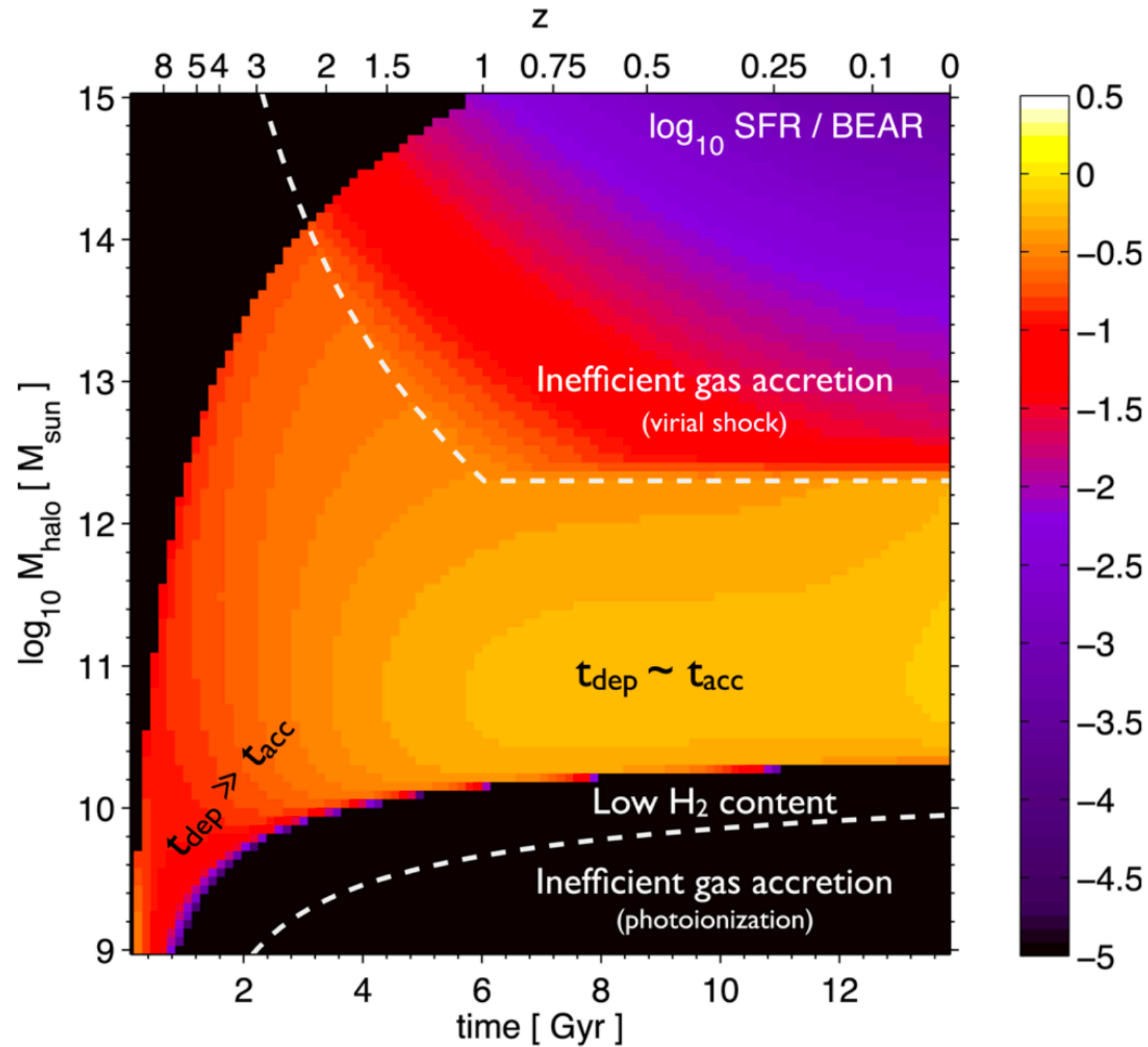


Mass-metallicity relation

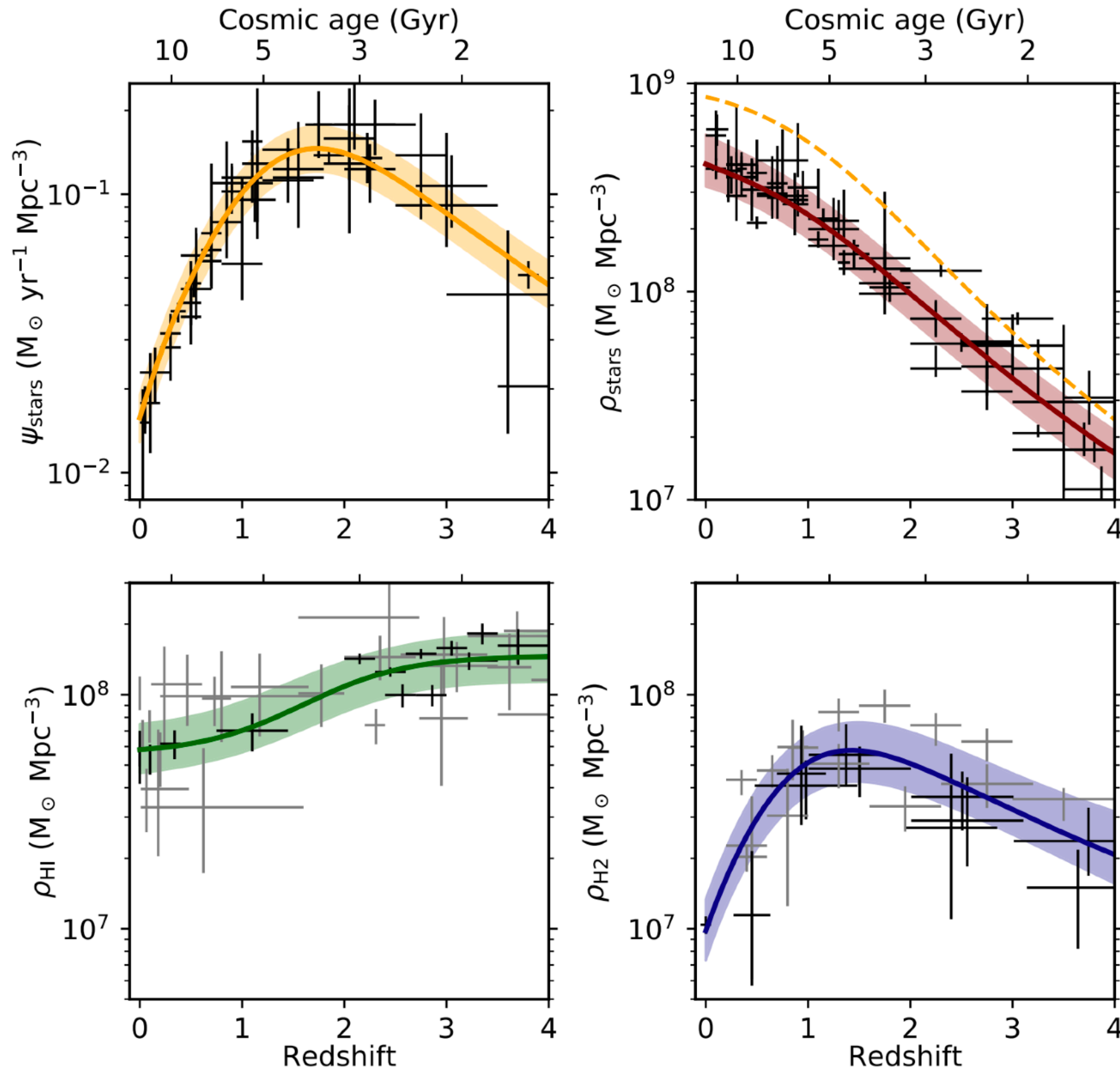


Model summary

Model summary



Evolution of gas abundance over cosmic time



Reading

- CFN §8.4-8.5, §8.7
- MvdBW §10.4, §11.8.1, §15.7.1