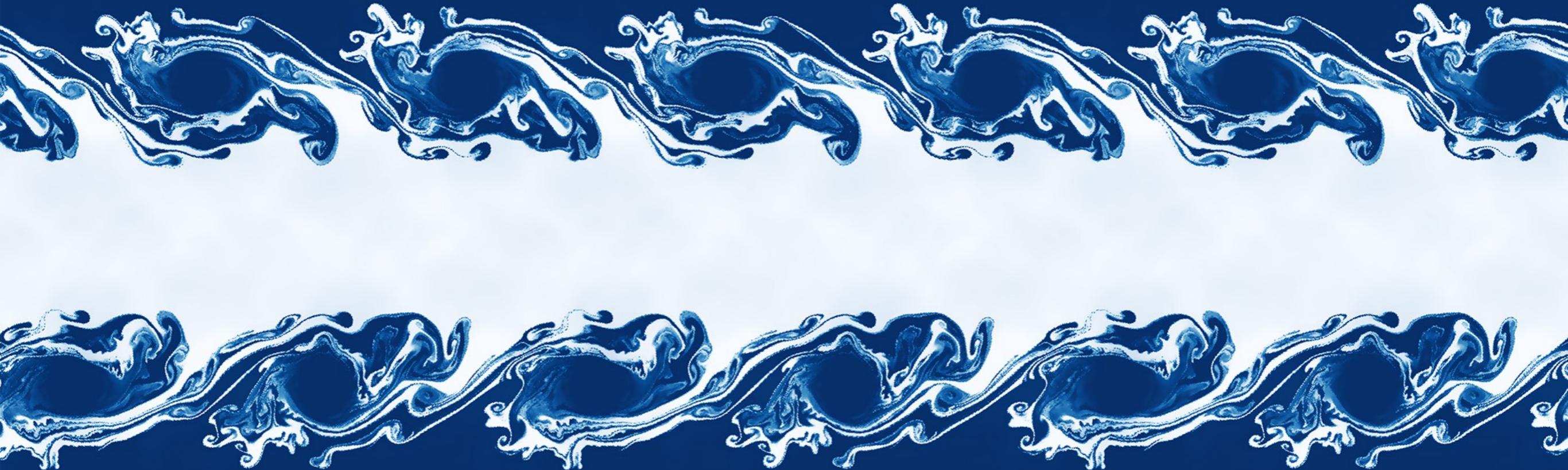


ASTR 670: Interstellar medium and gas dynamics

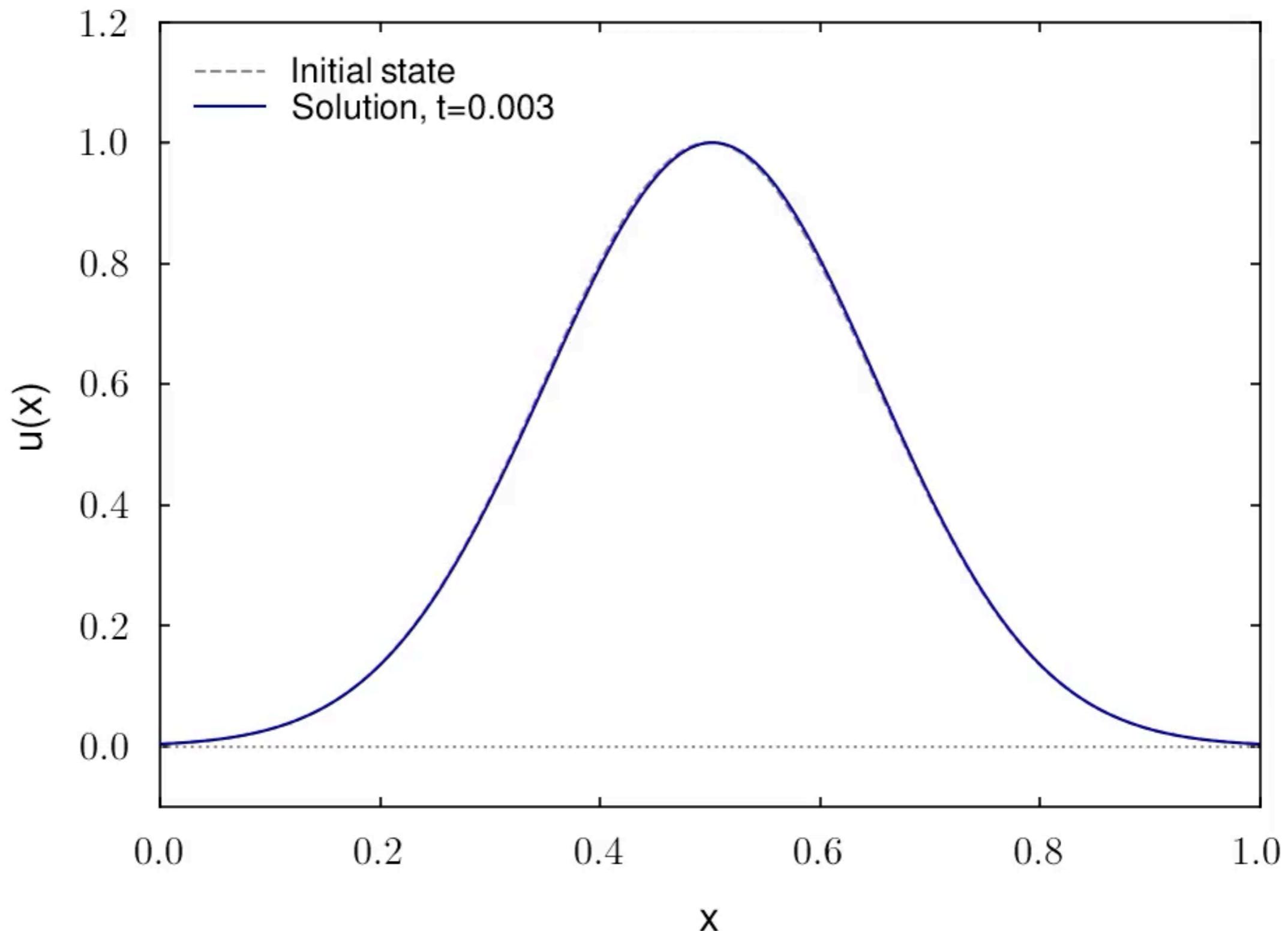
Prof. Benedikt Diemer



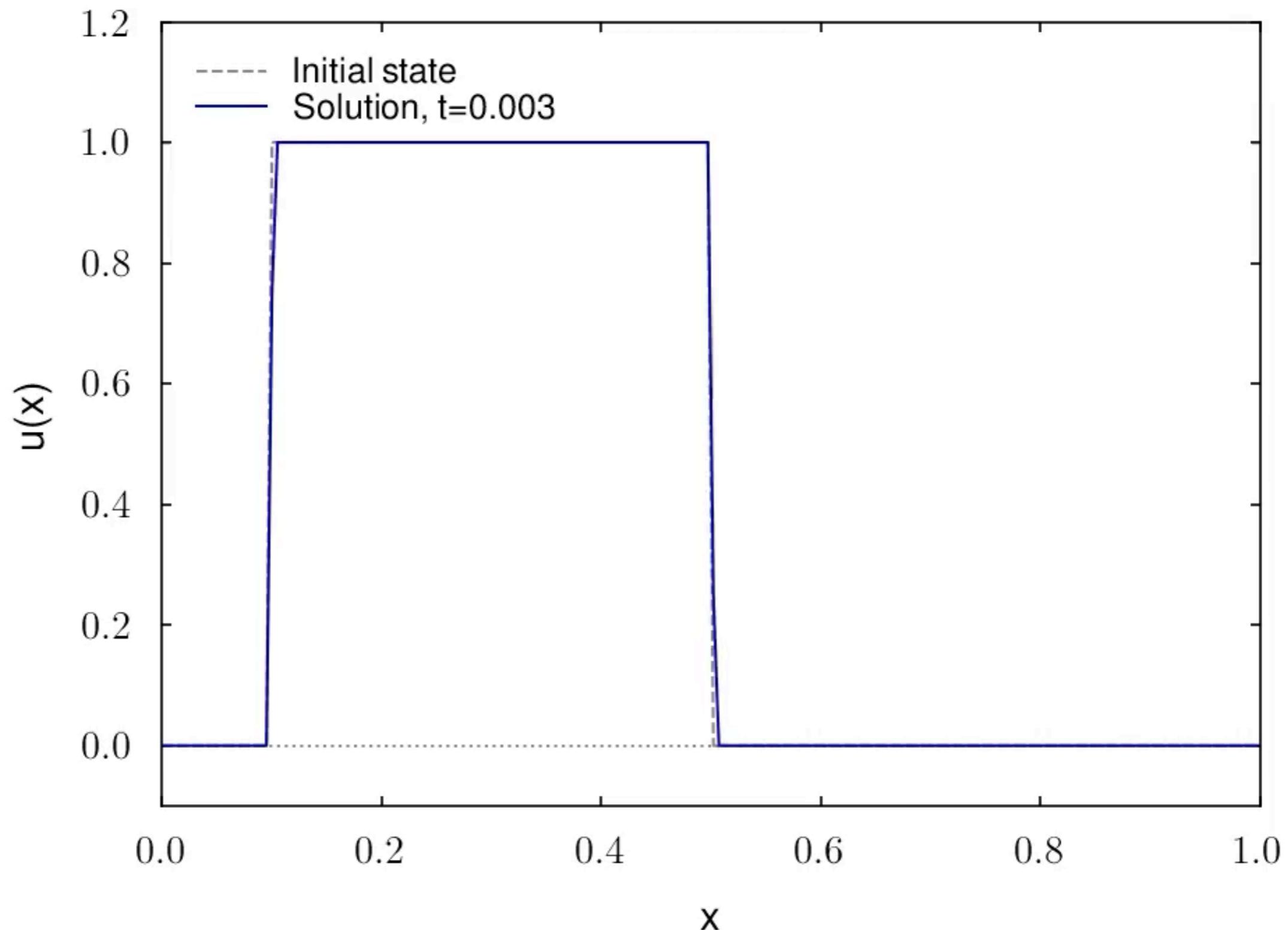
Chapter 7 • Shocks

Why shocks?

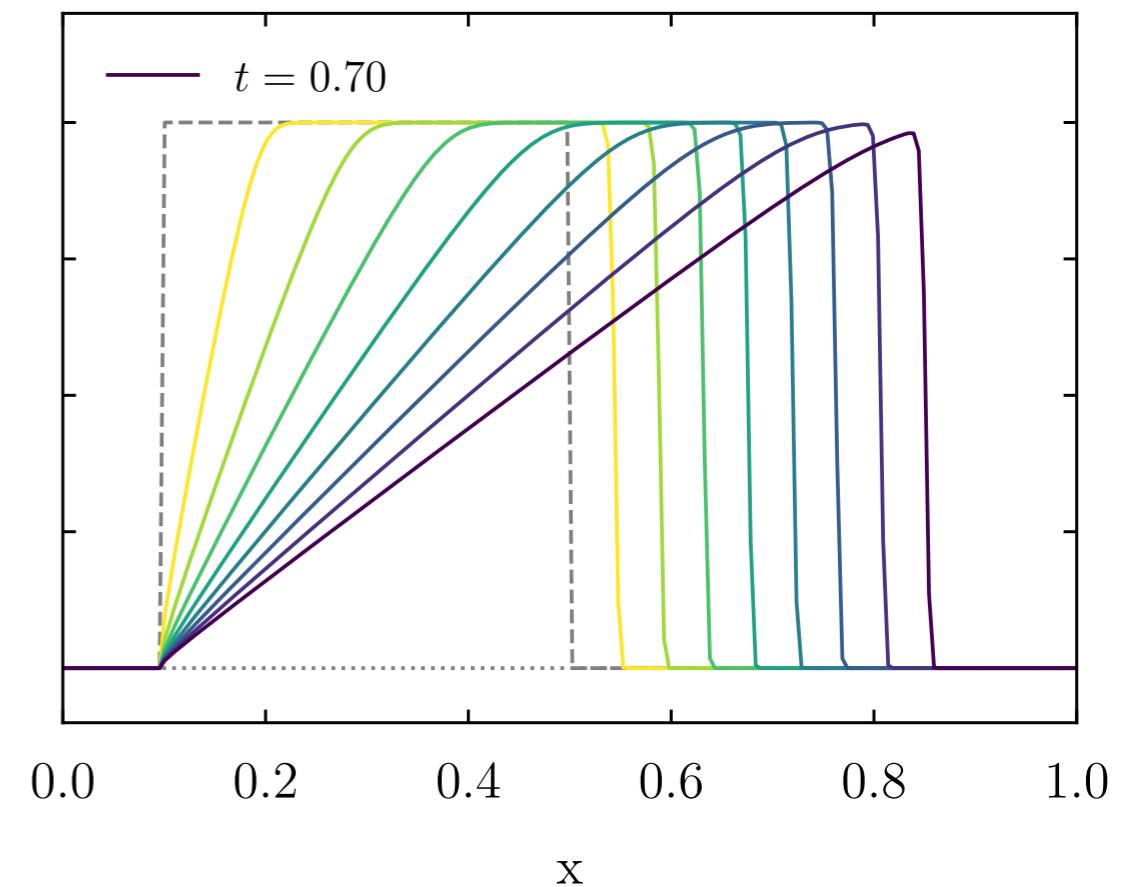
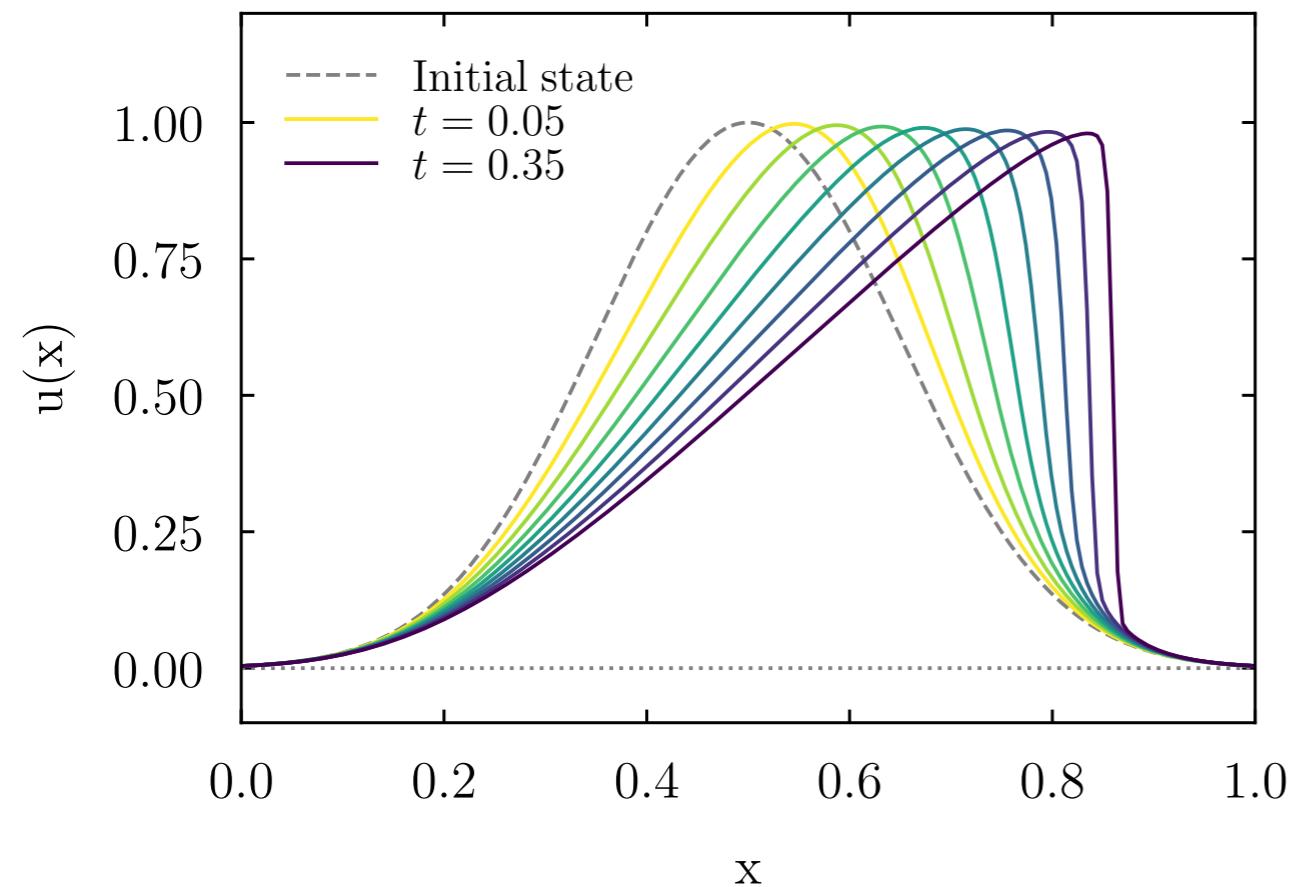
Burgers' equation



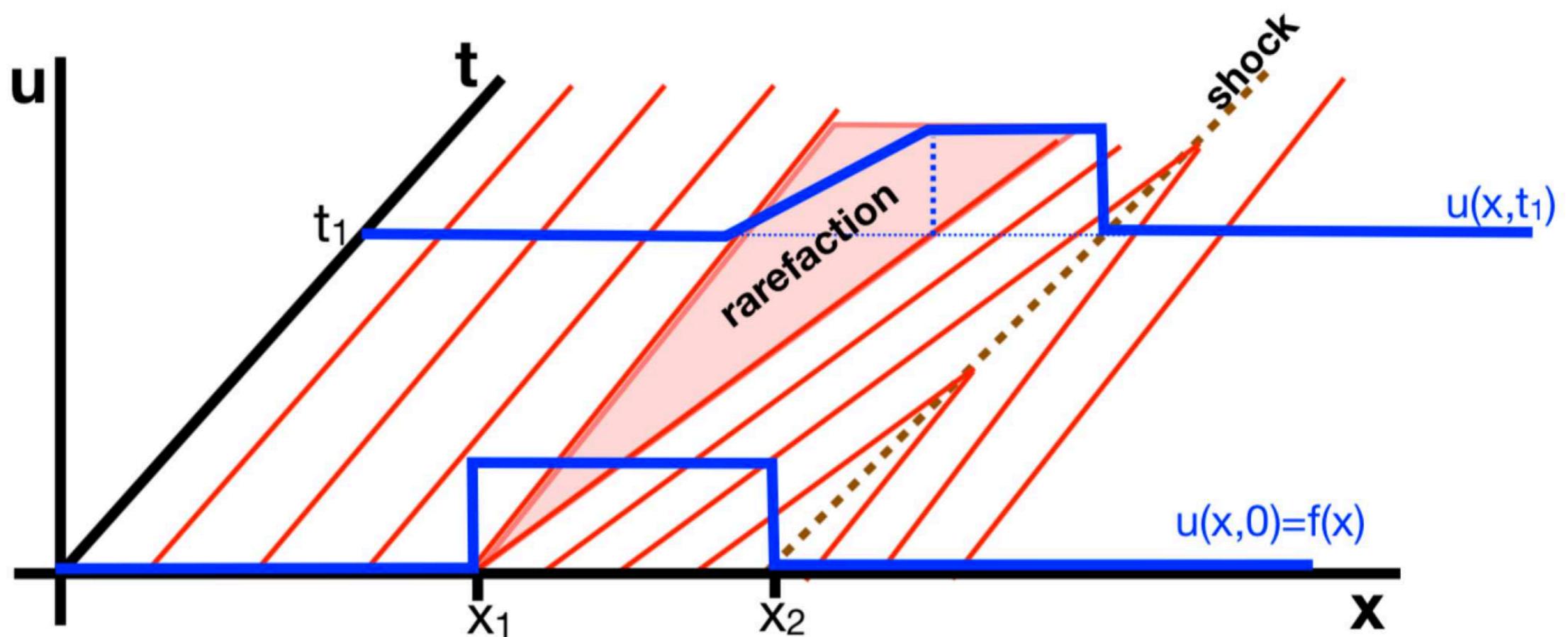
Burgers' equation



Burgers' equation

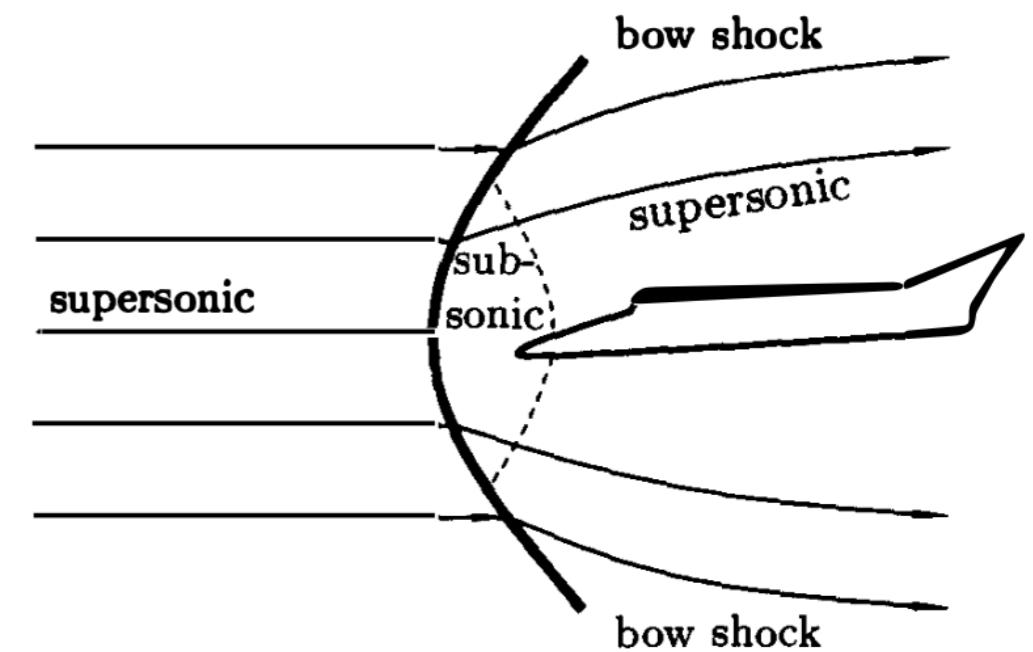
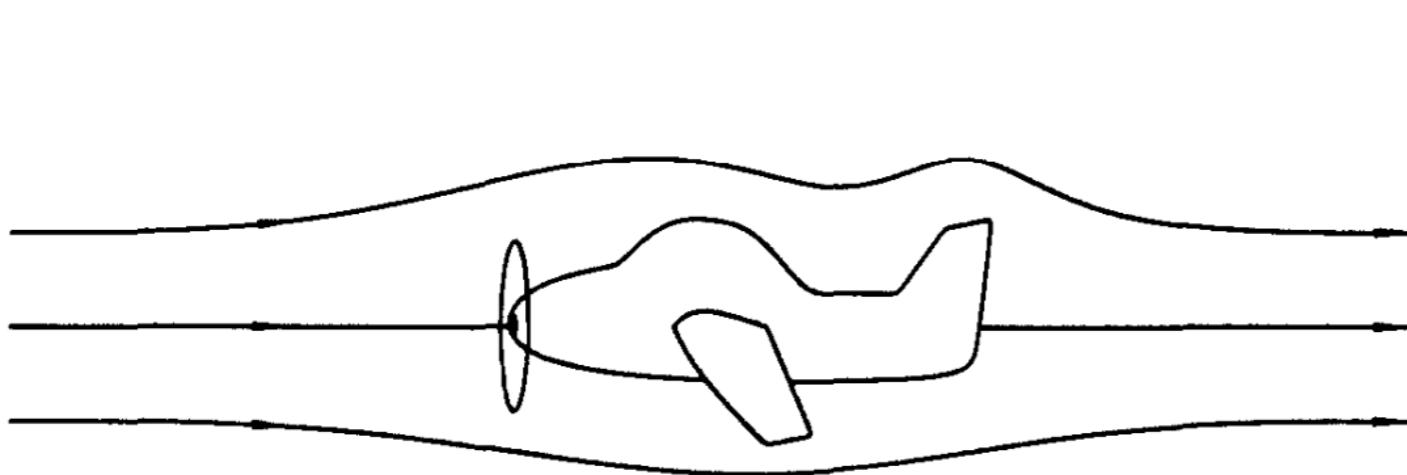


Burgers' equation

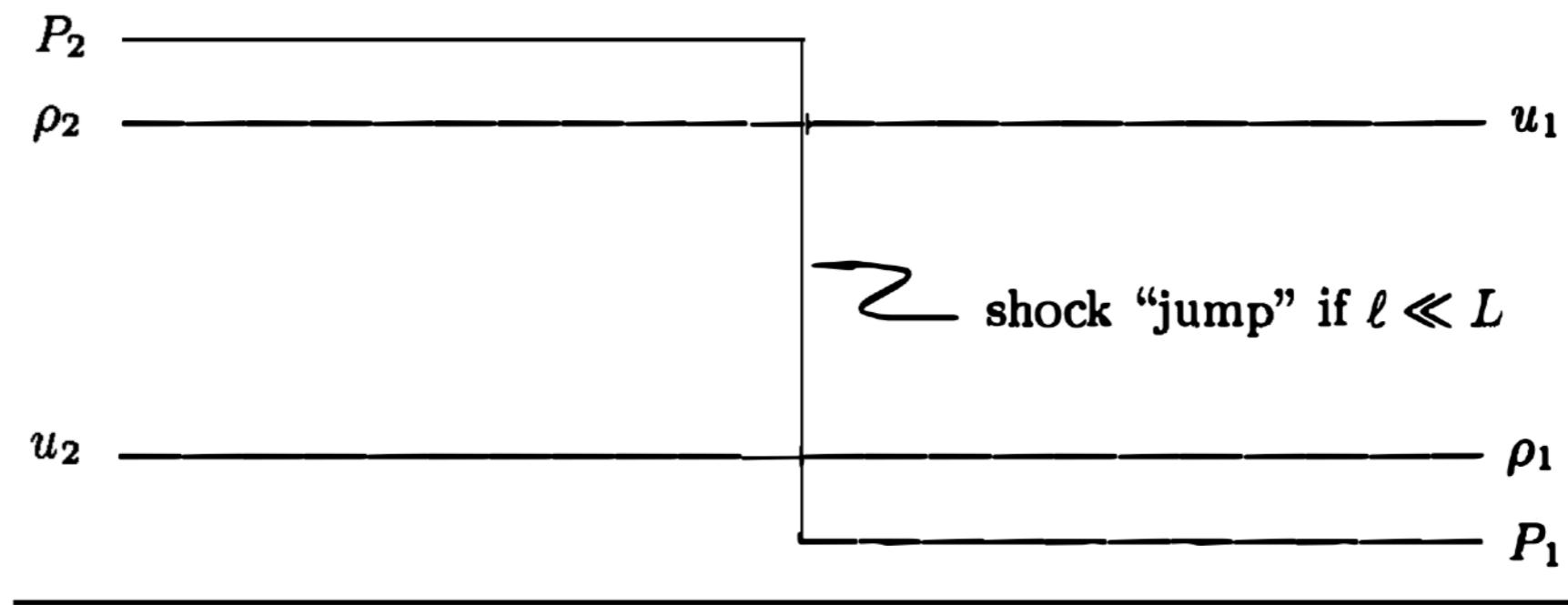
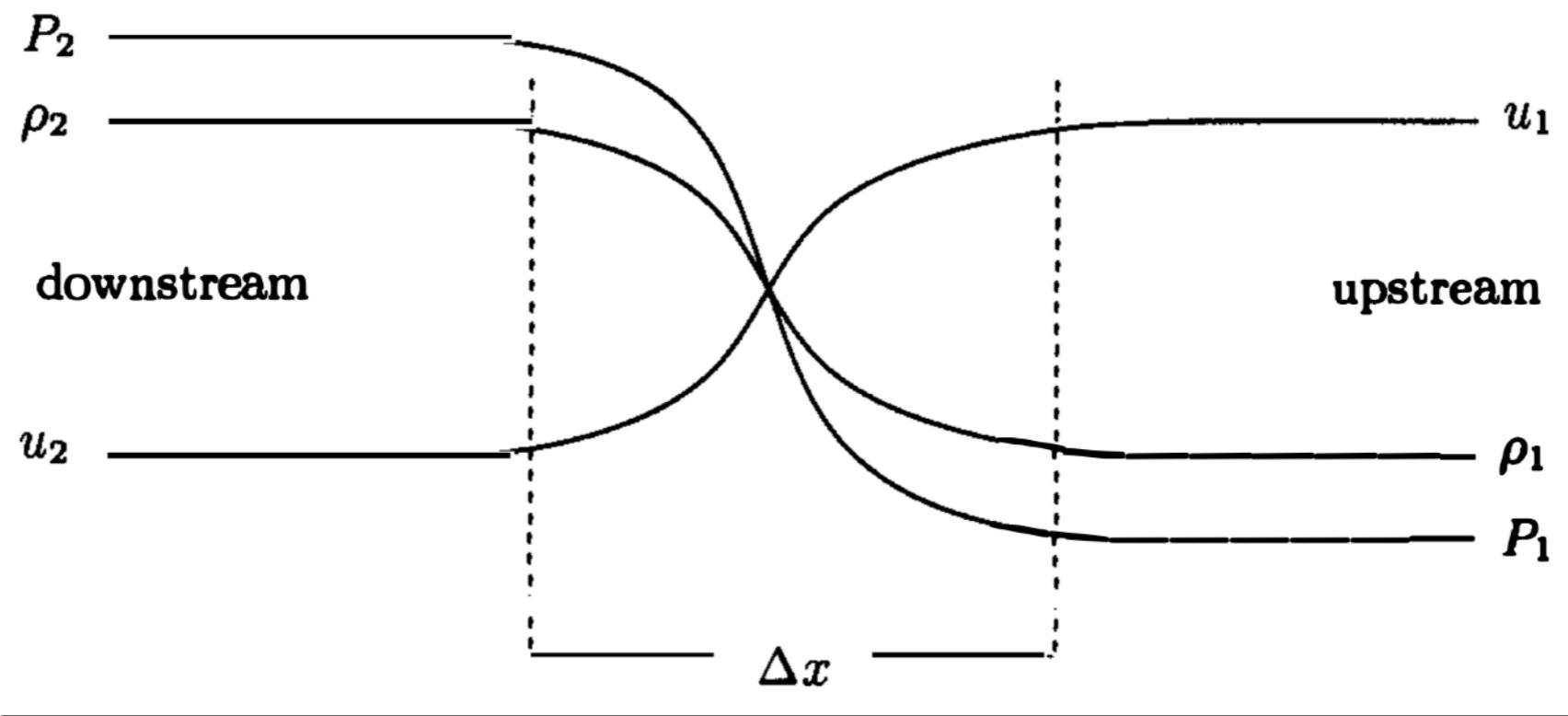


Shock physics

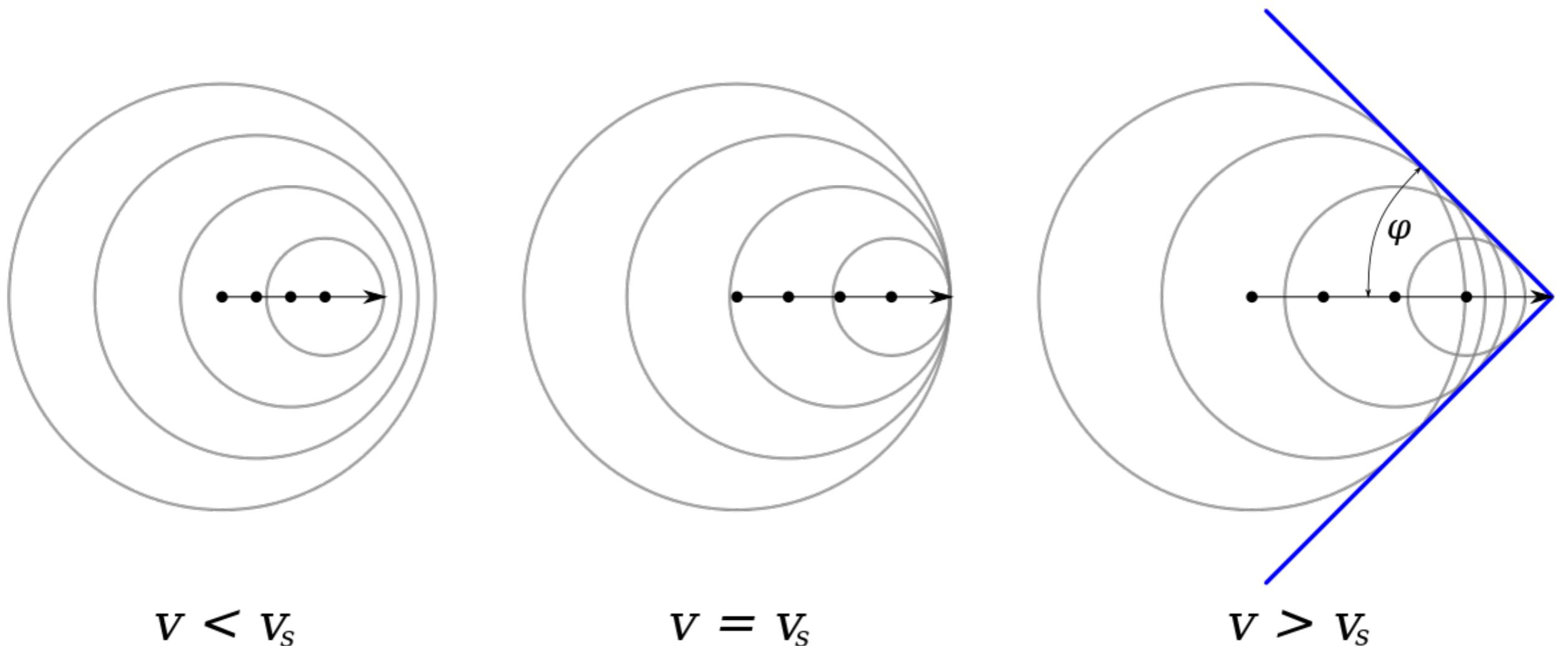
Subsonic vs. supersonic



Shocks

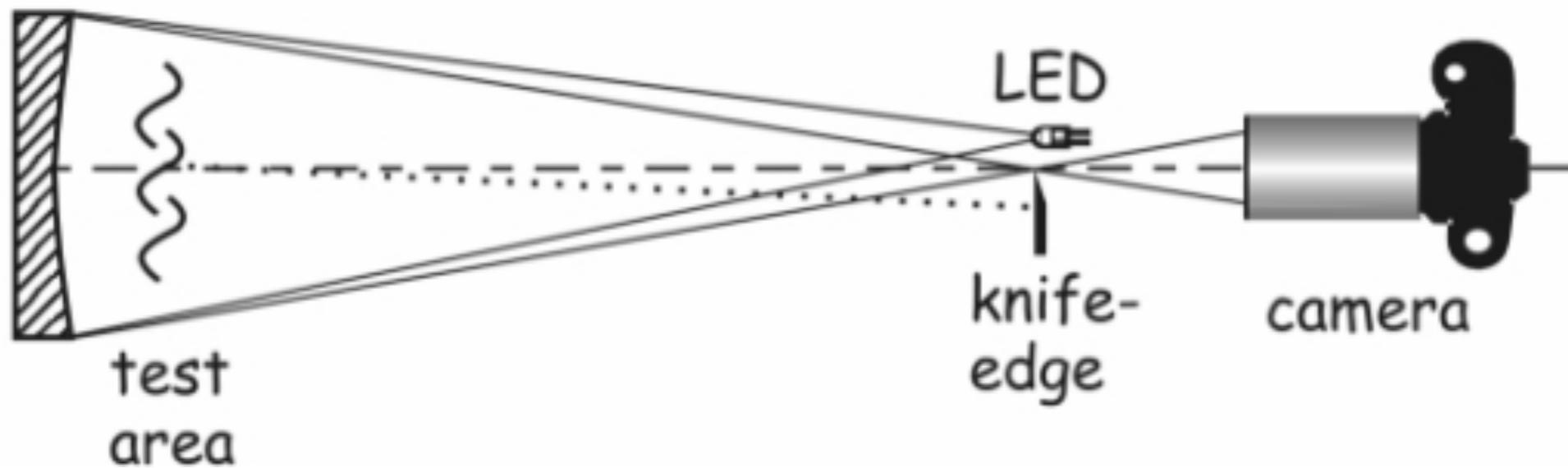


Mach cone

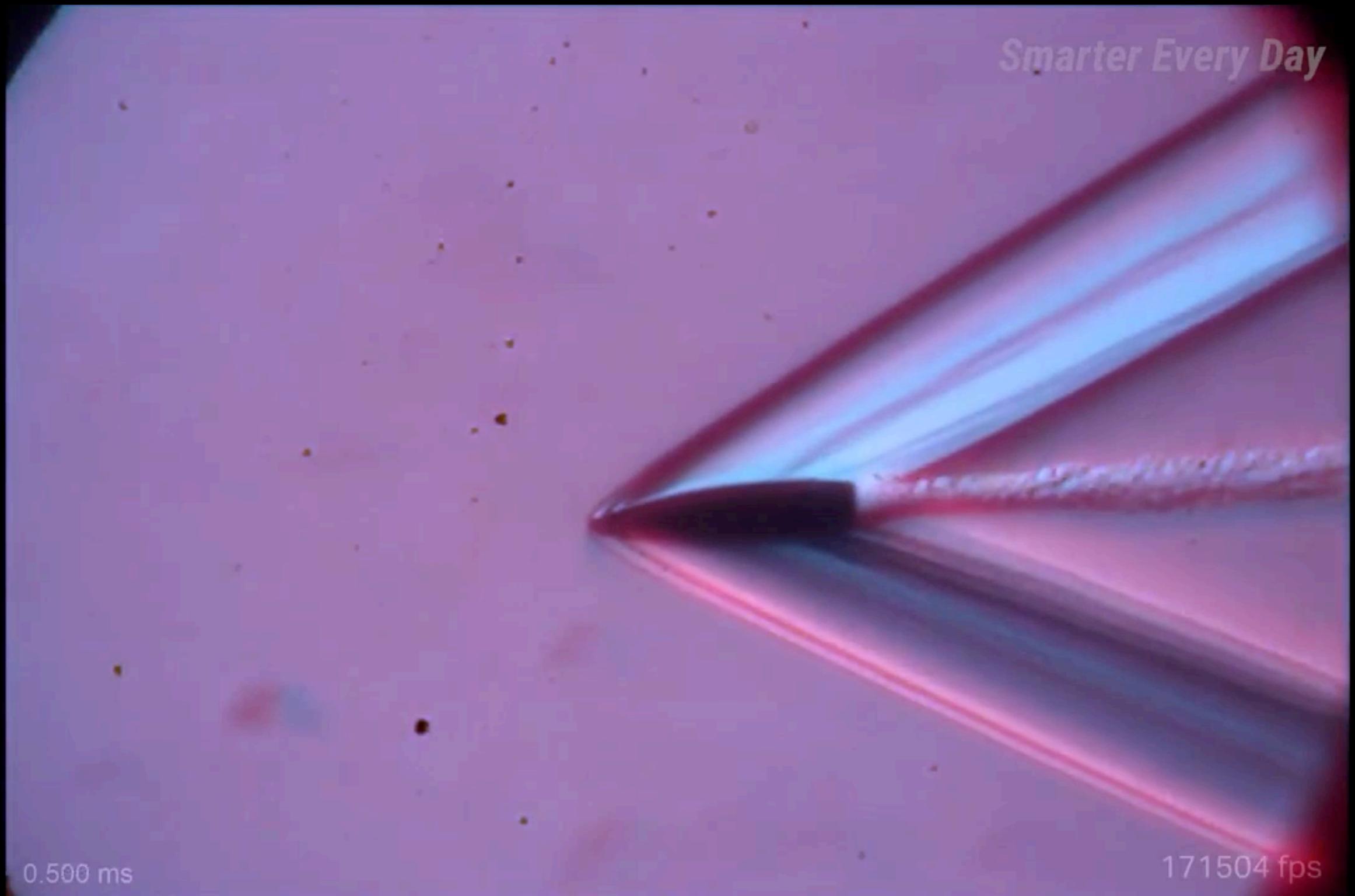


Schlieren imaging

spherical
mirror



Schlieren imaging of shocks

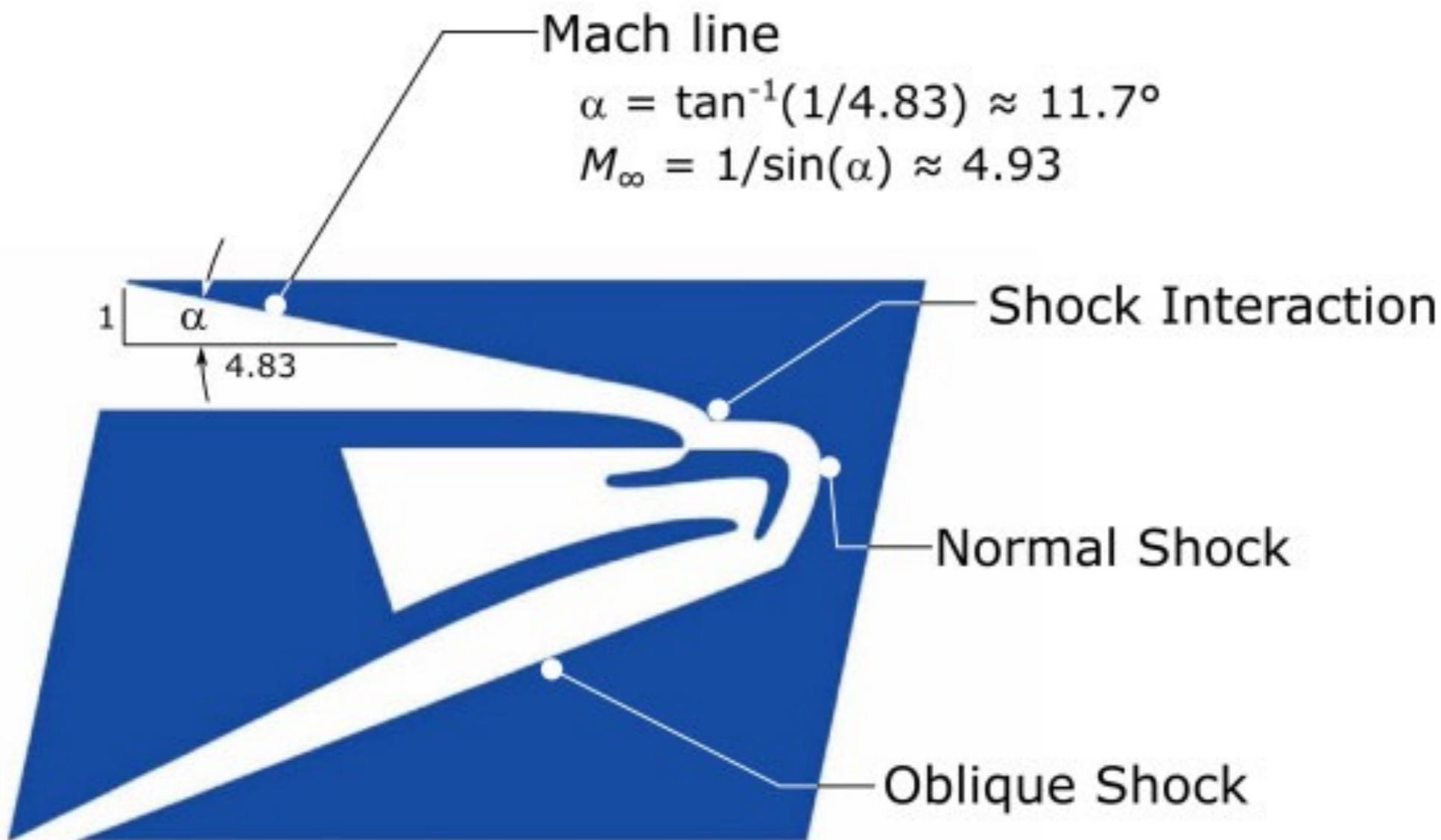


Mach cone



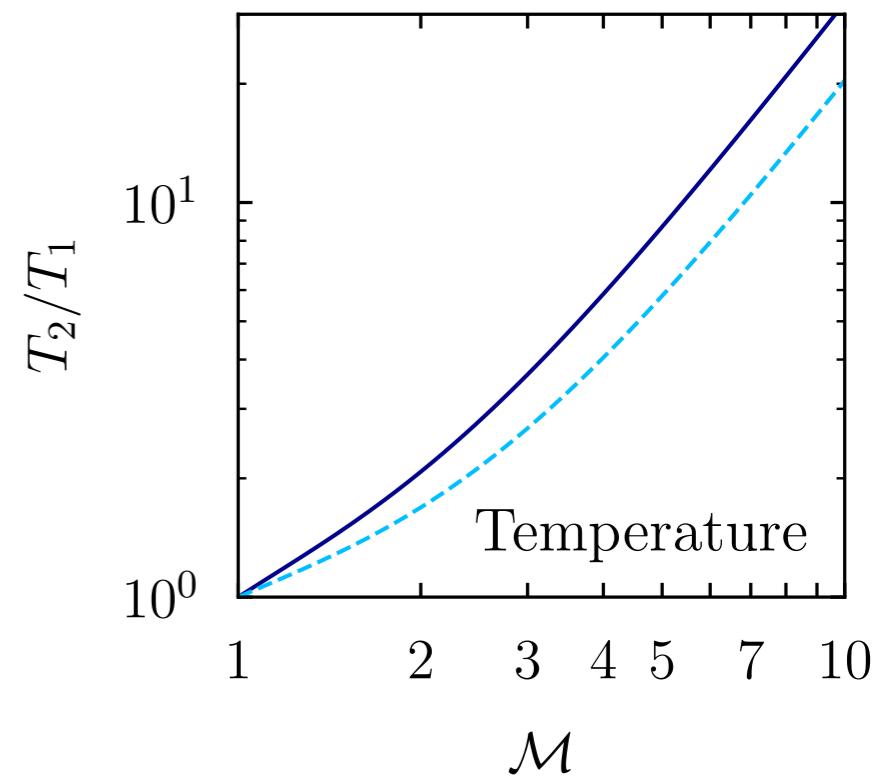
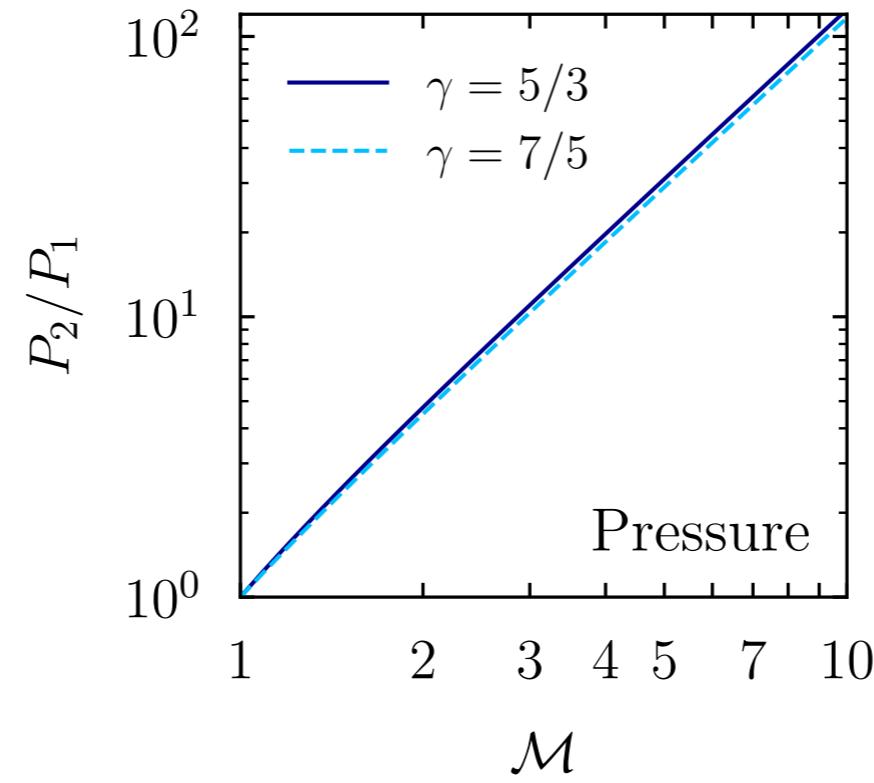
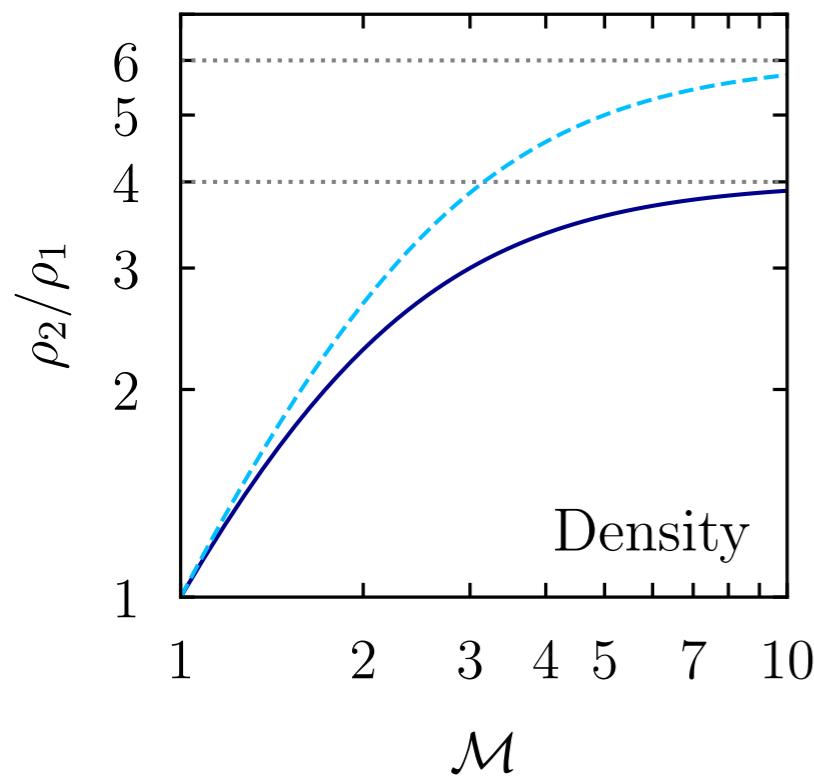
Image: NASA

Mach cone

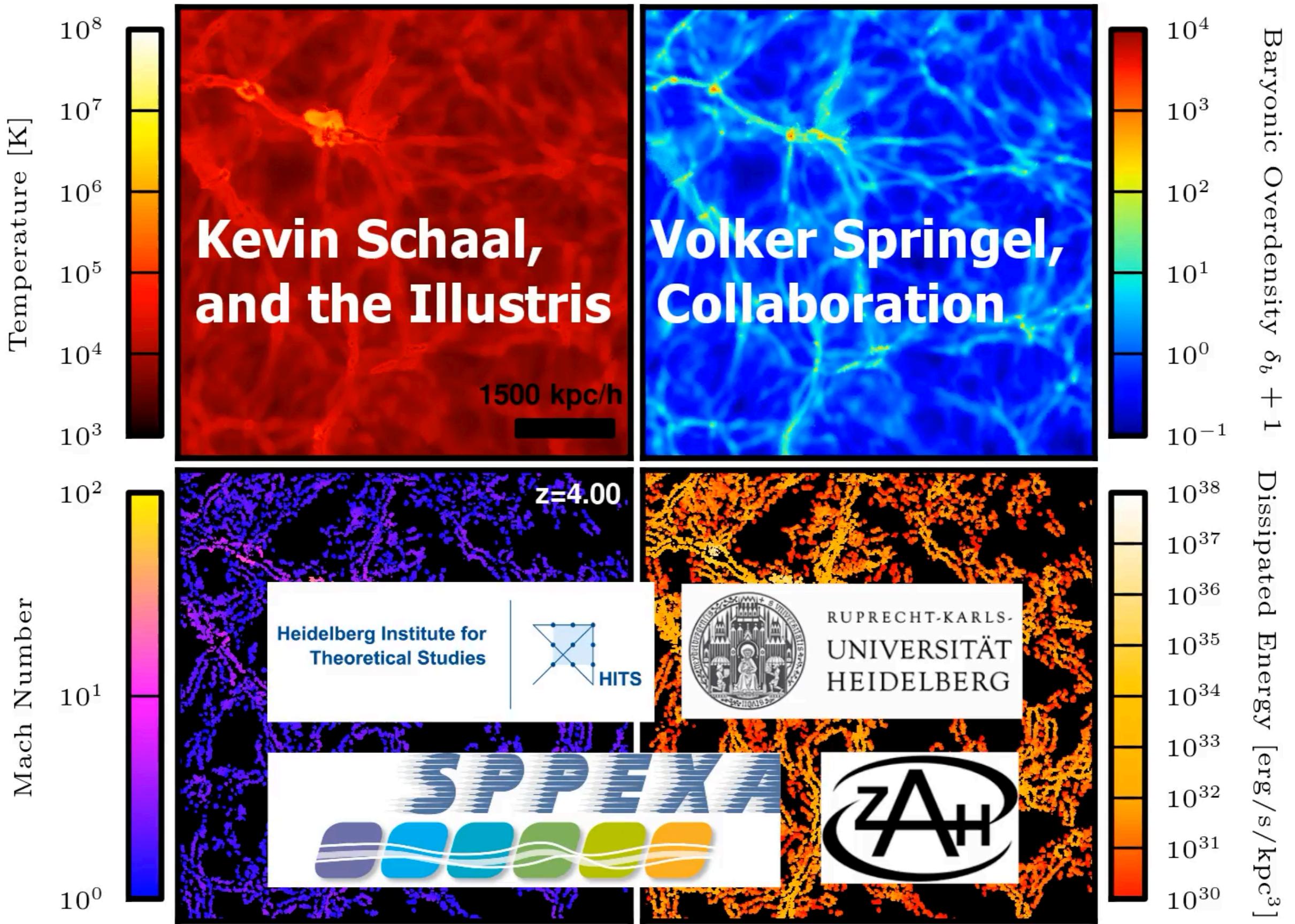


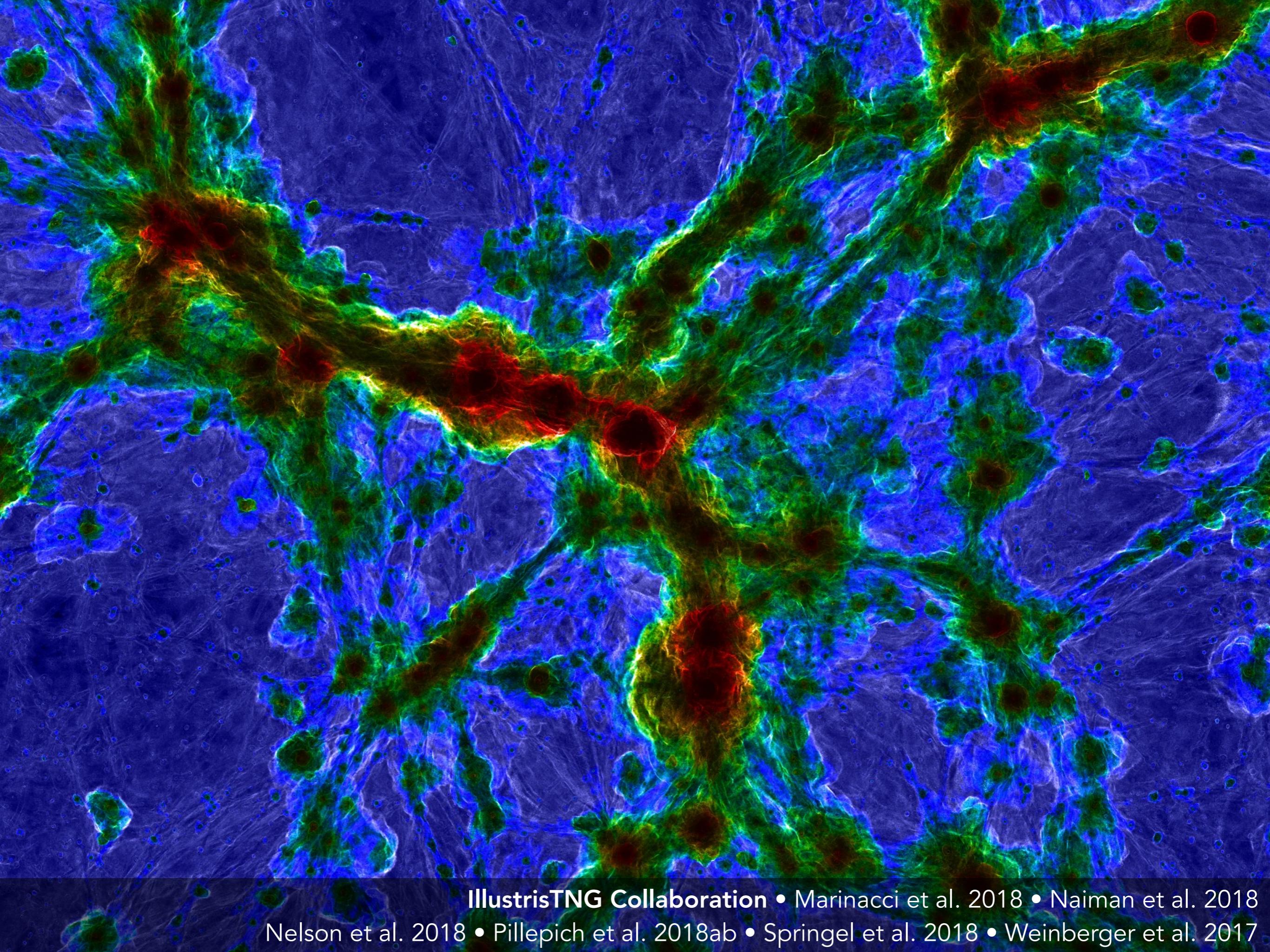
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Ideal gas shocks

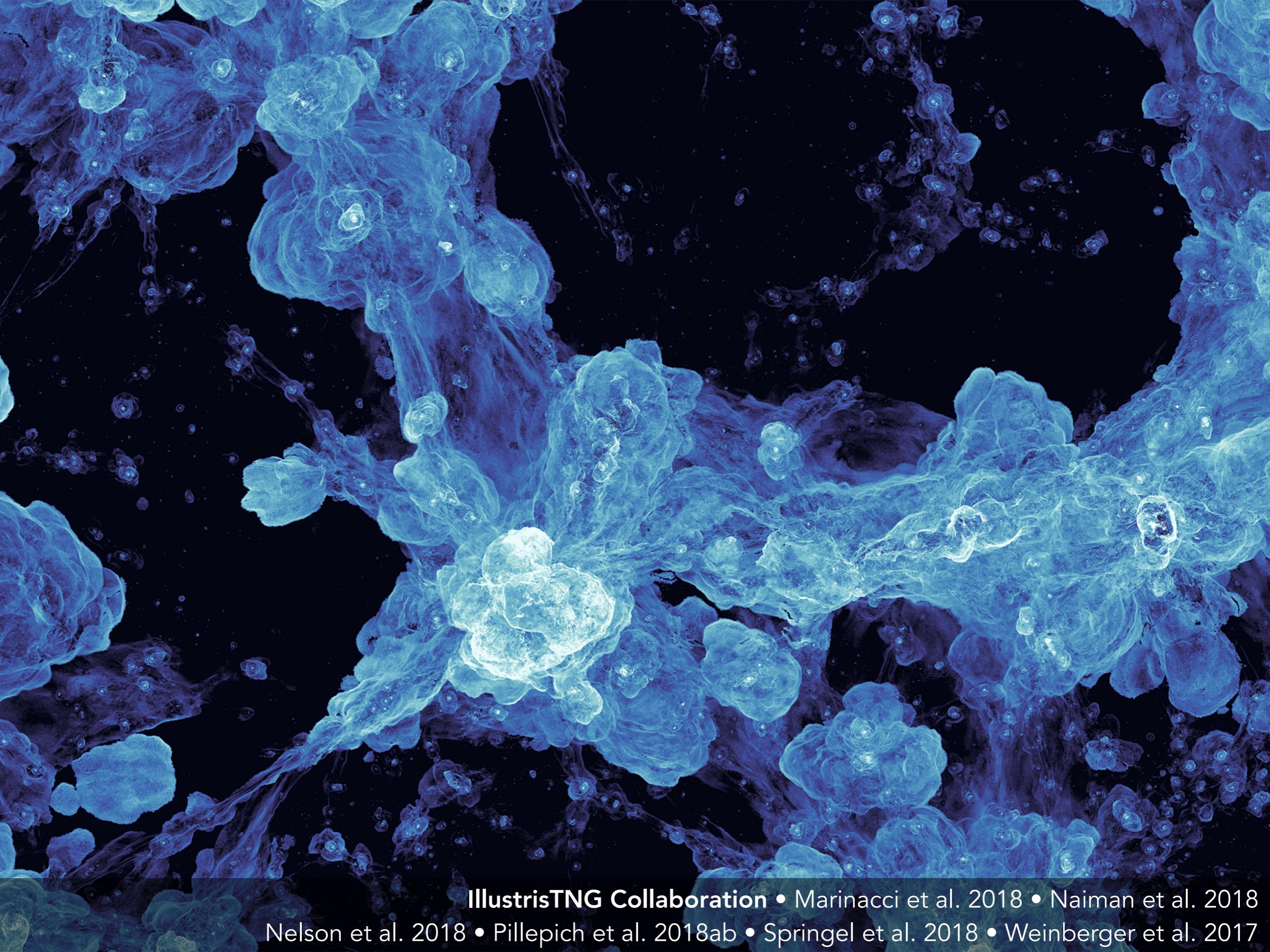


Accretion shocks





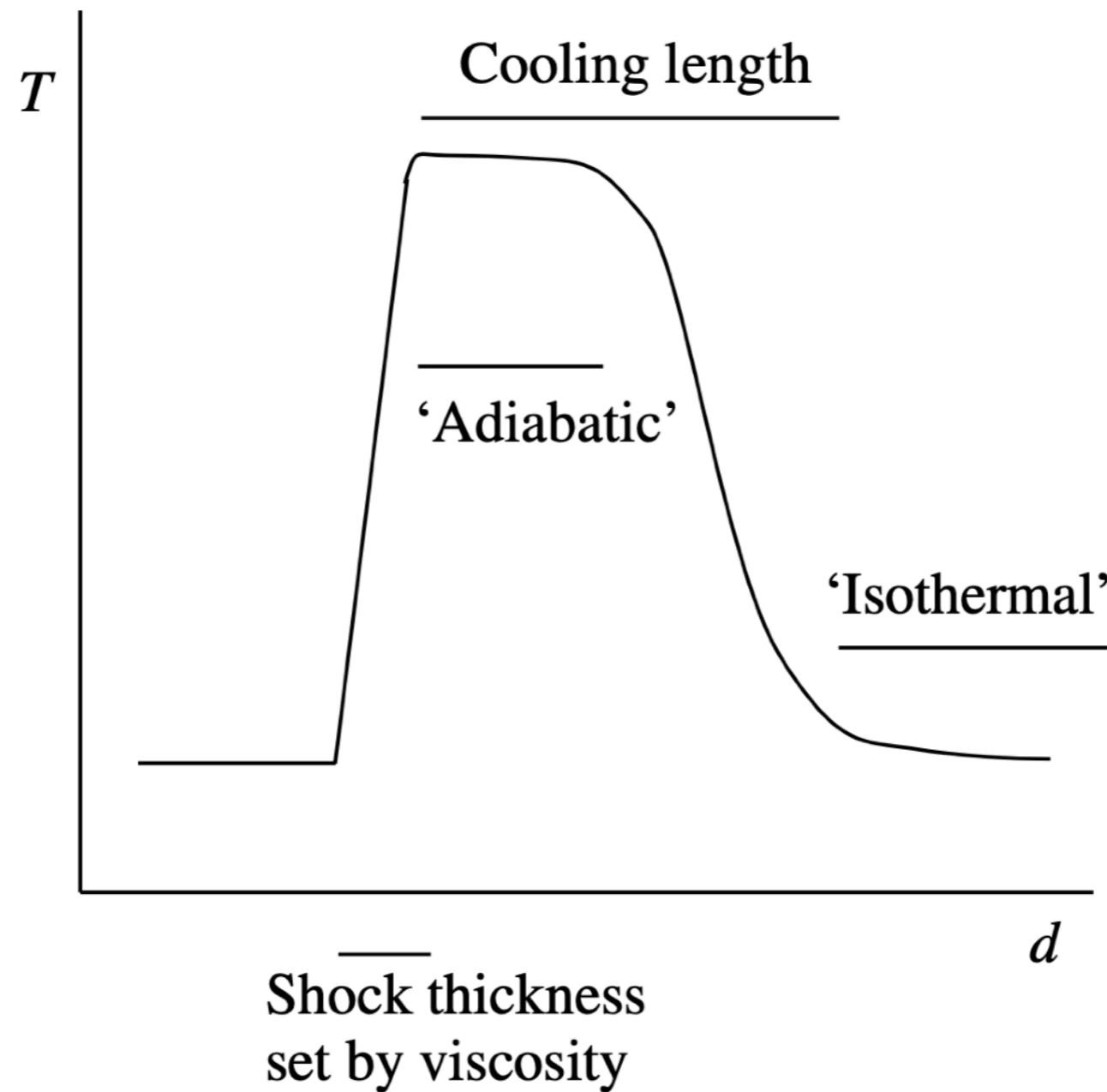
IllustrisTNG Collaboration • Marinacci et al. 2018 • Naiman et al. 2018
Nelson et al. 2018 • Pillepich et al. 2018ab • Springel et al. 2018 • Weinberger et al. 2017



IllustrisTNG Collaboration • Marinacci et al. 2018 • Naiman et al. 2018
Nelson et al. 2018 • Pillepich et al. 2018ab • Springel et al. 2018 • Weinberger et al. 2017

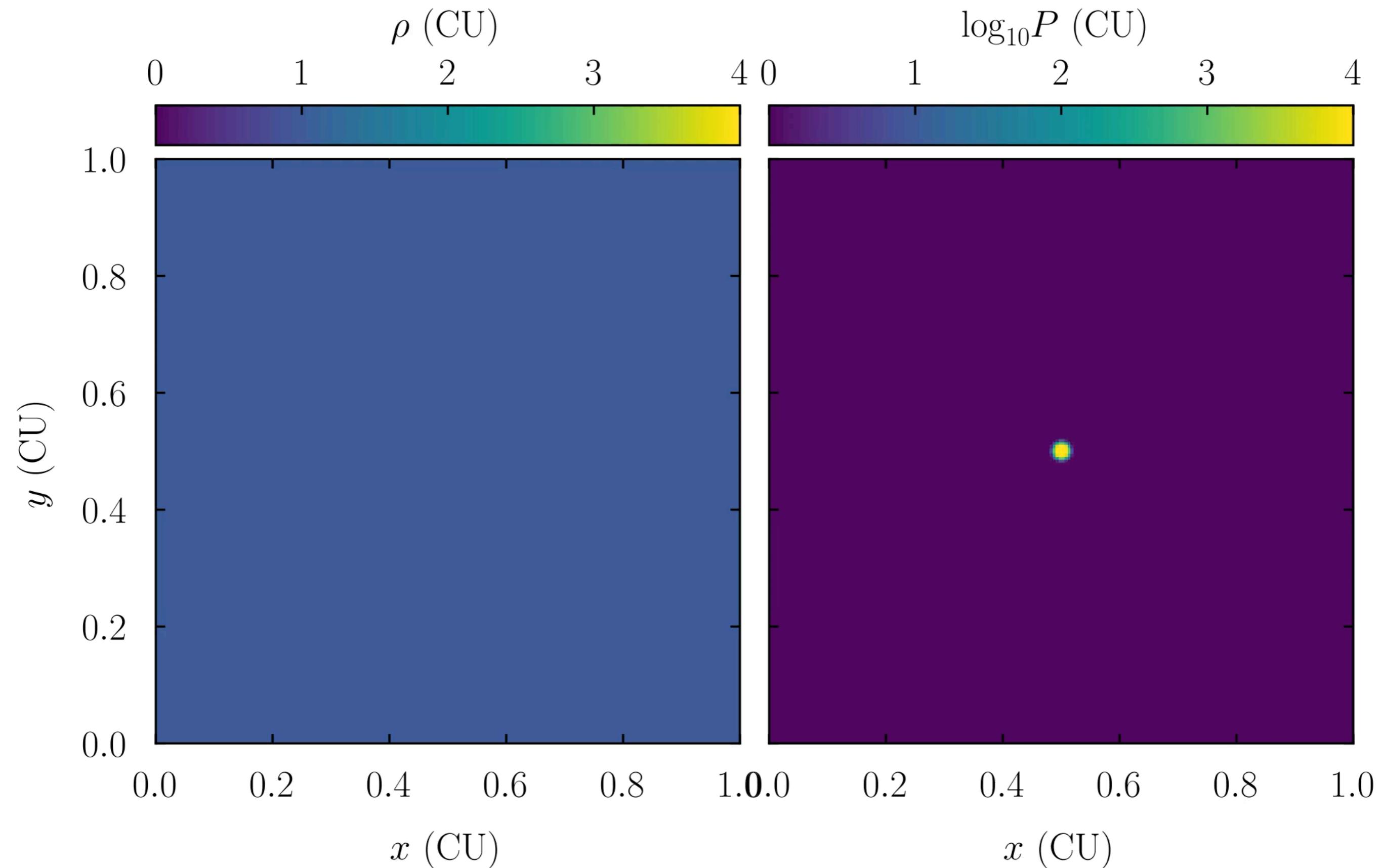
Isothermal shocks

"Isothermal" shocks

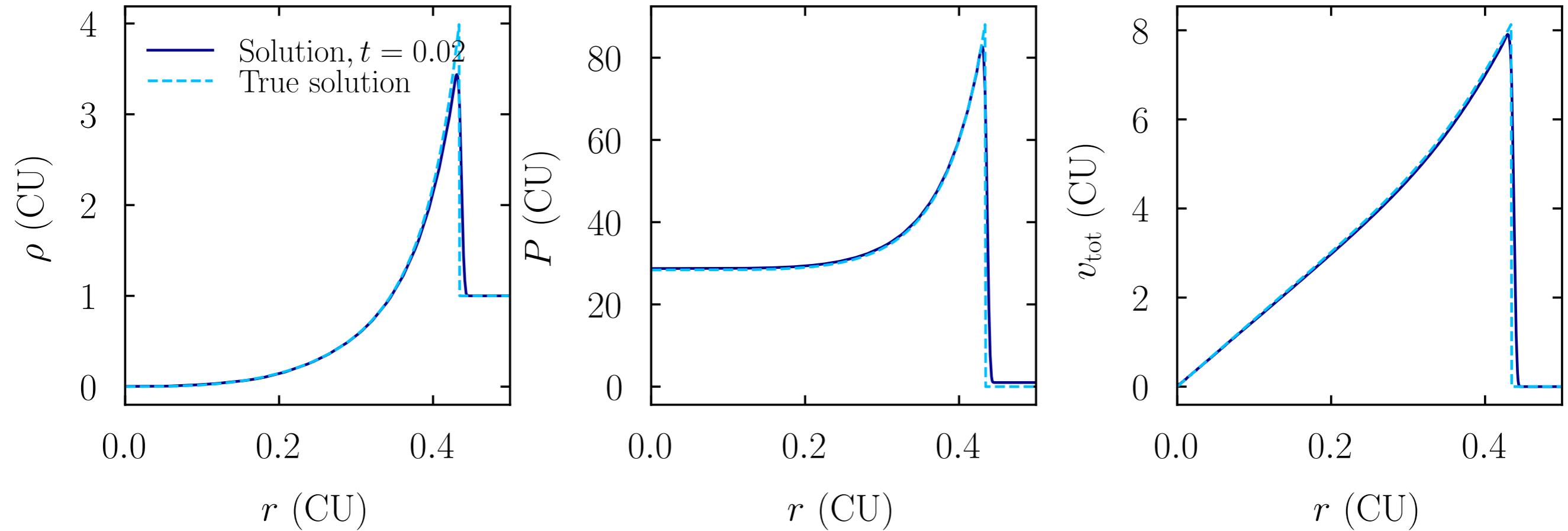


Blast waves

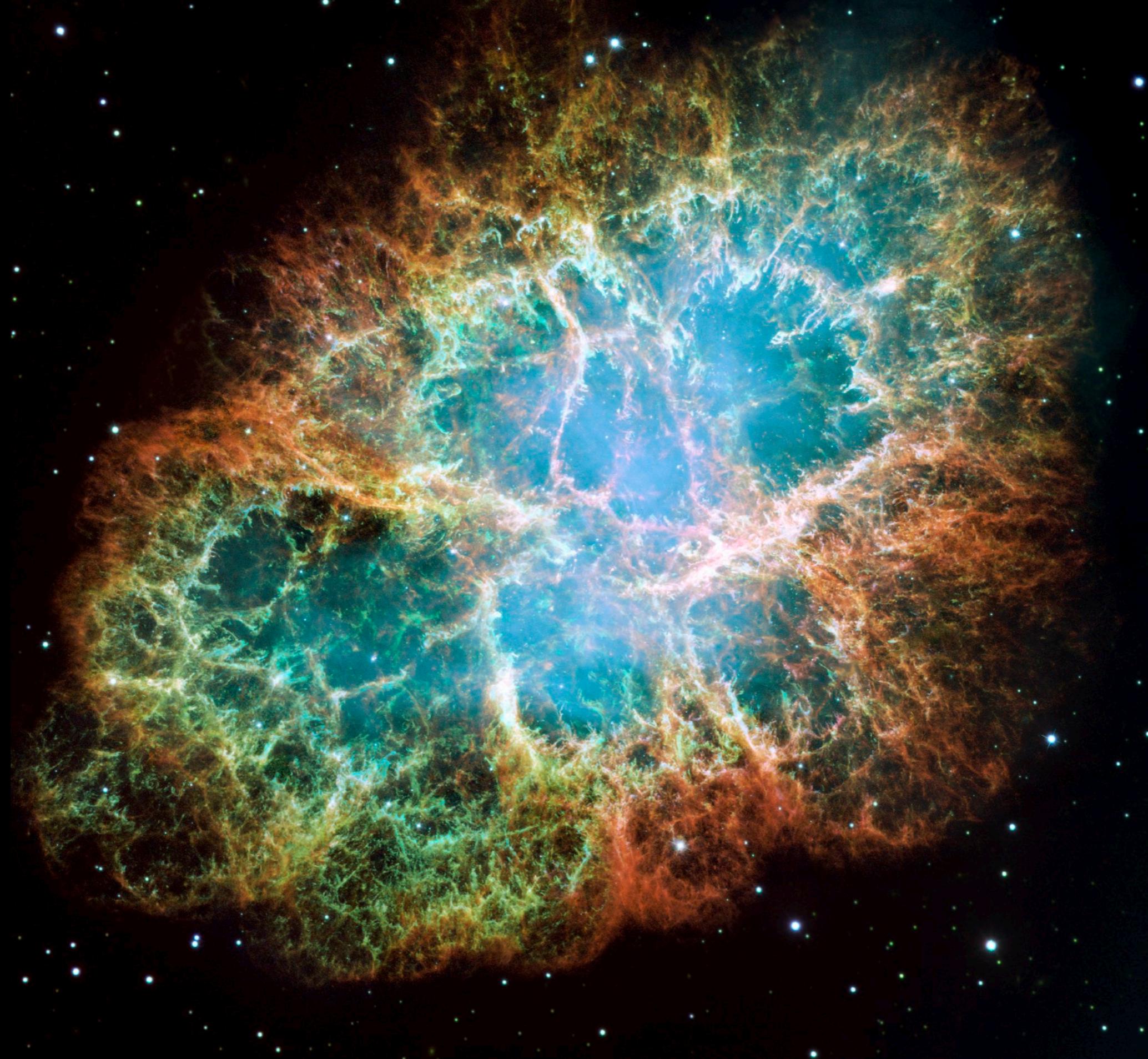
Sedov-Taylor blastwave



Sedov-Taylor blastwave

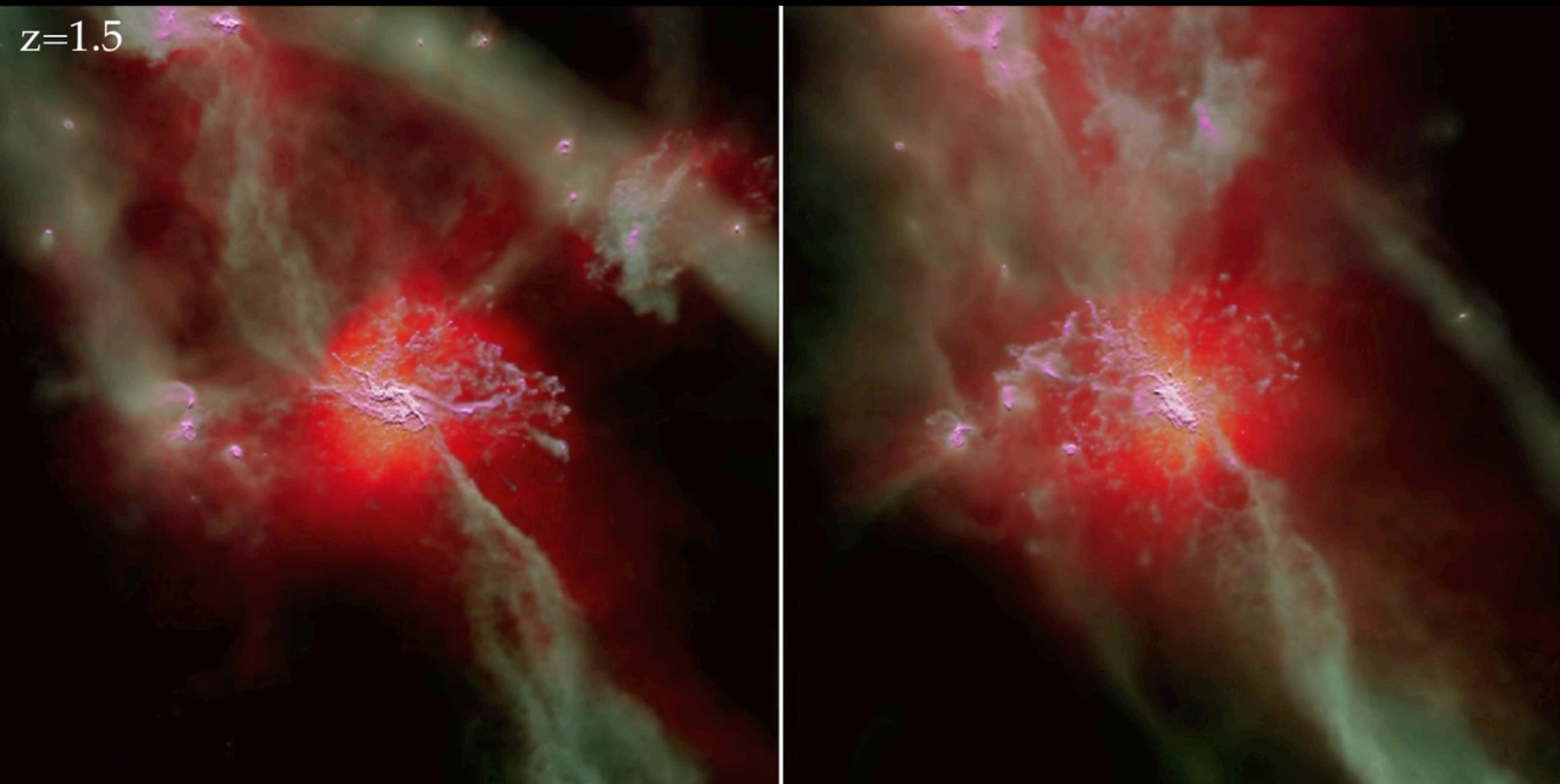








$z=1.5$



Reading

- Recommended: CC §7, §8.1-3, Zingale §6.1
- Additional: Shu §15, §16, §17