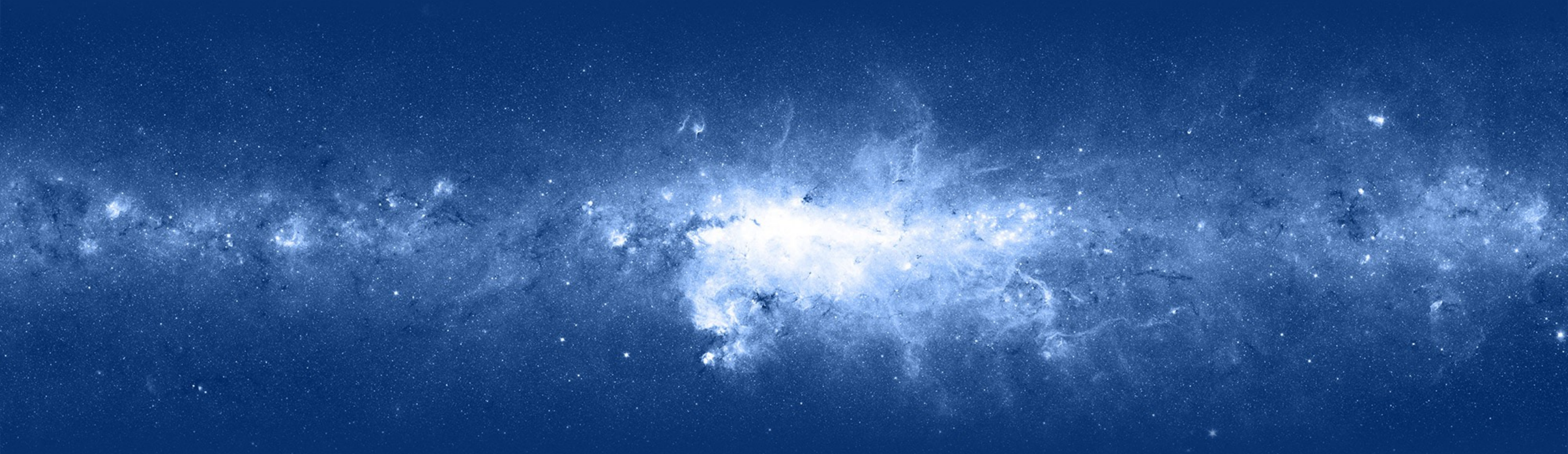


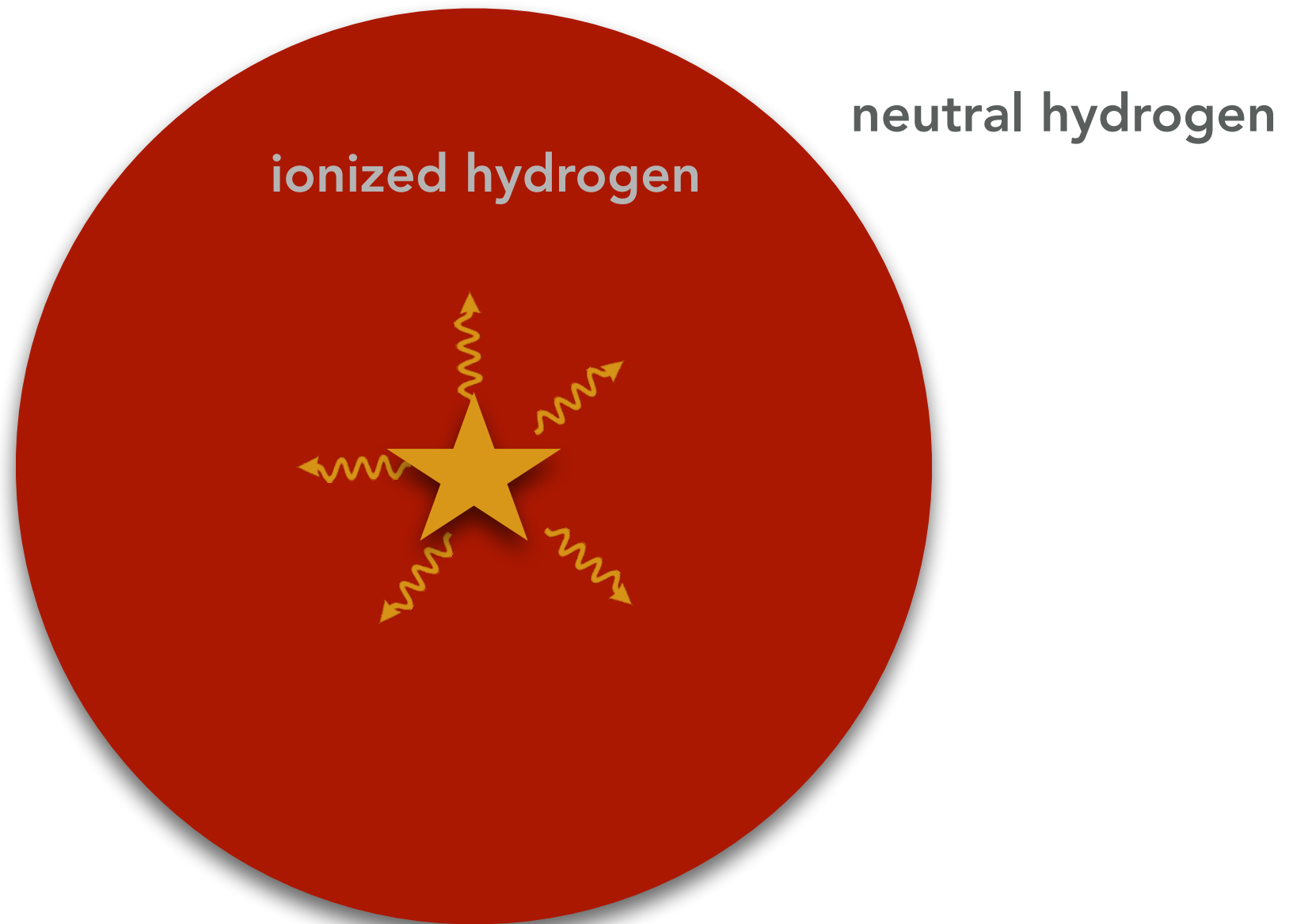
ASTR 670: Interstellar medium and gas dynamics

Prof. Benedikt Diemer



Chapter 5 • Atomic physics II: Photoionization and recombination

Ionization balance



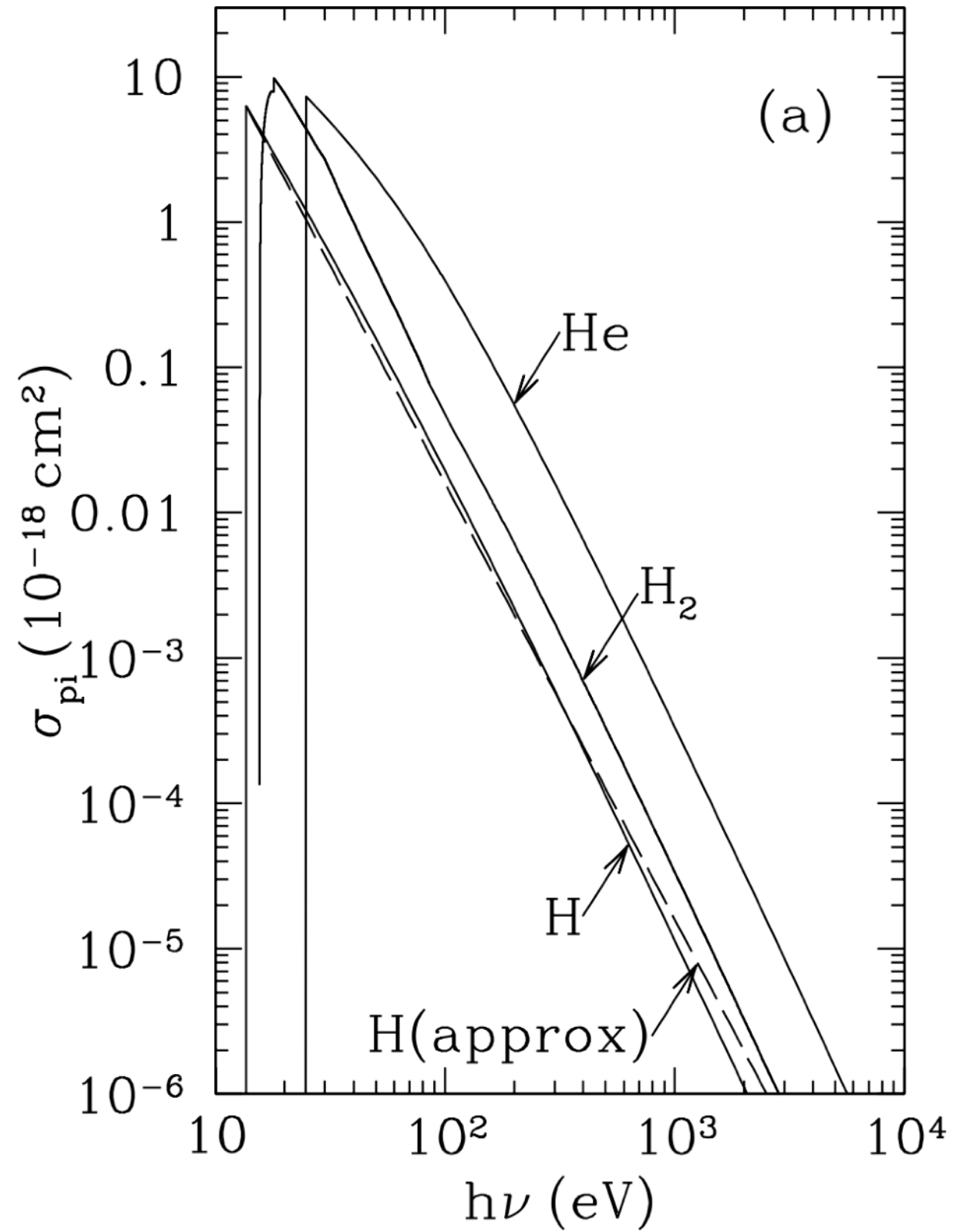
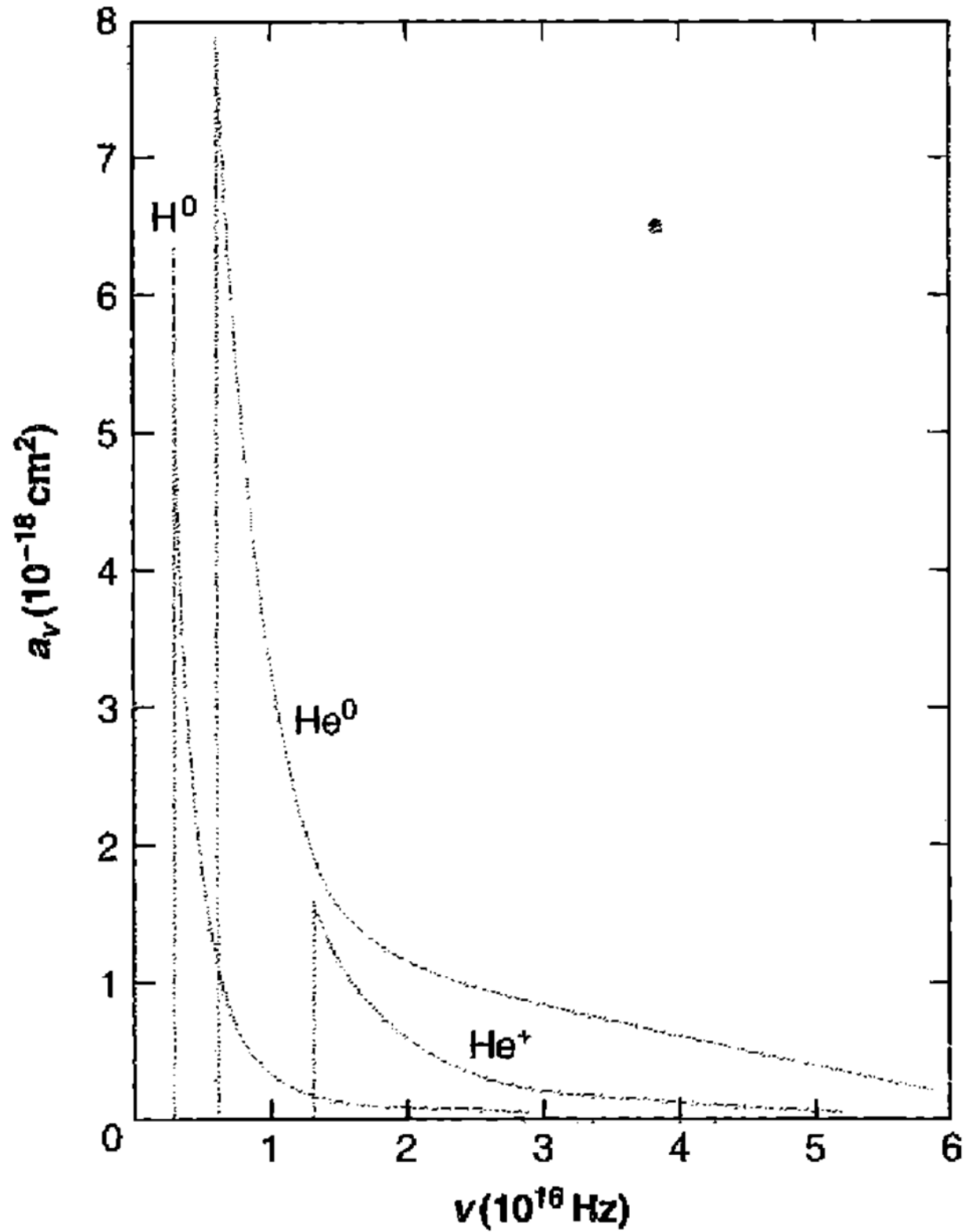
H II region



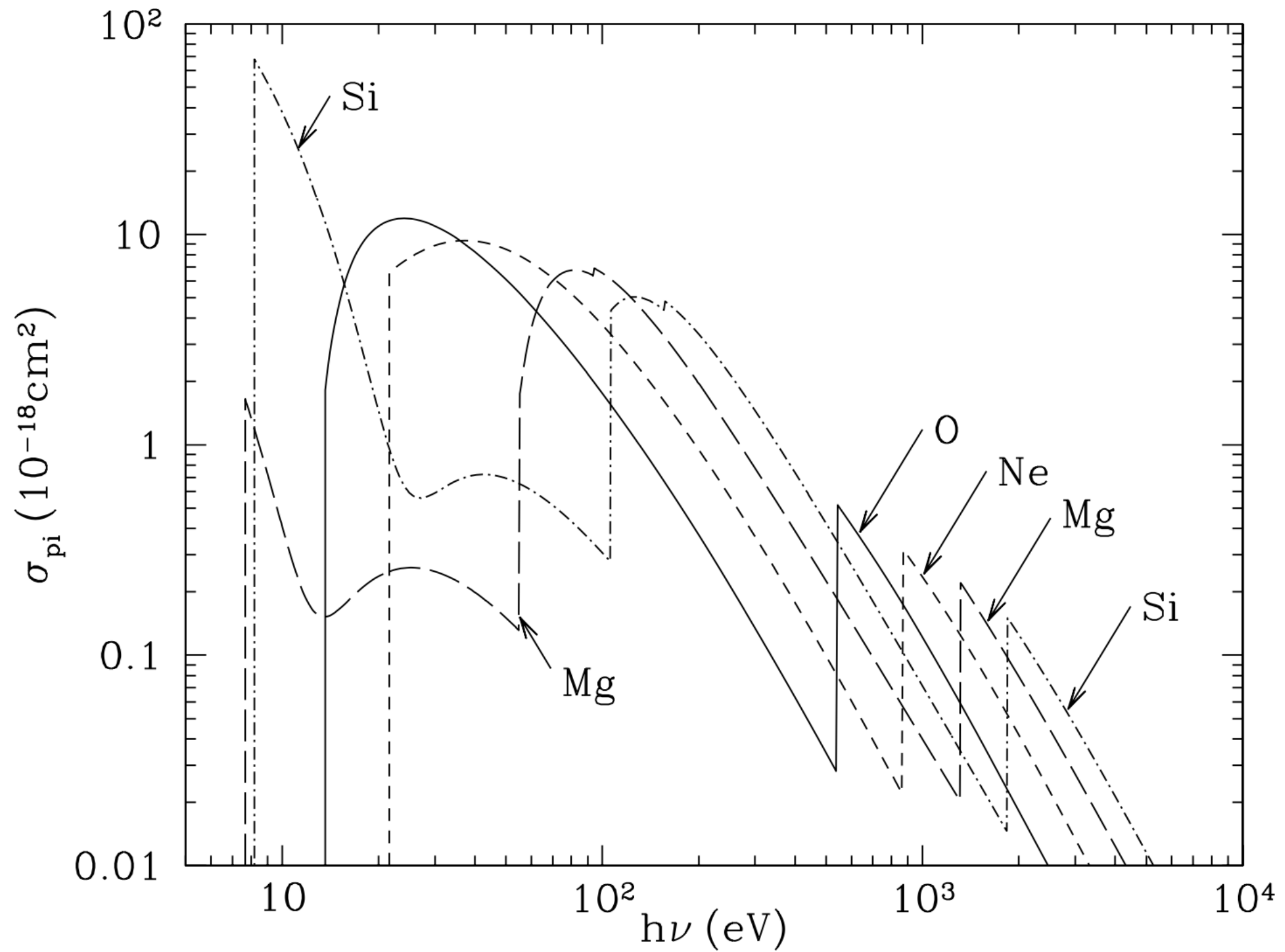
Rosette Nebula

§5.1 • Photoionization

Ionization cross-section



Ionization cross-section



§5.2 • Recombination

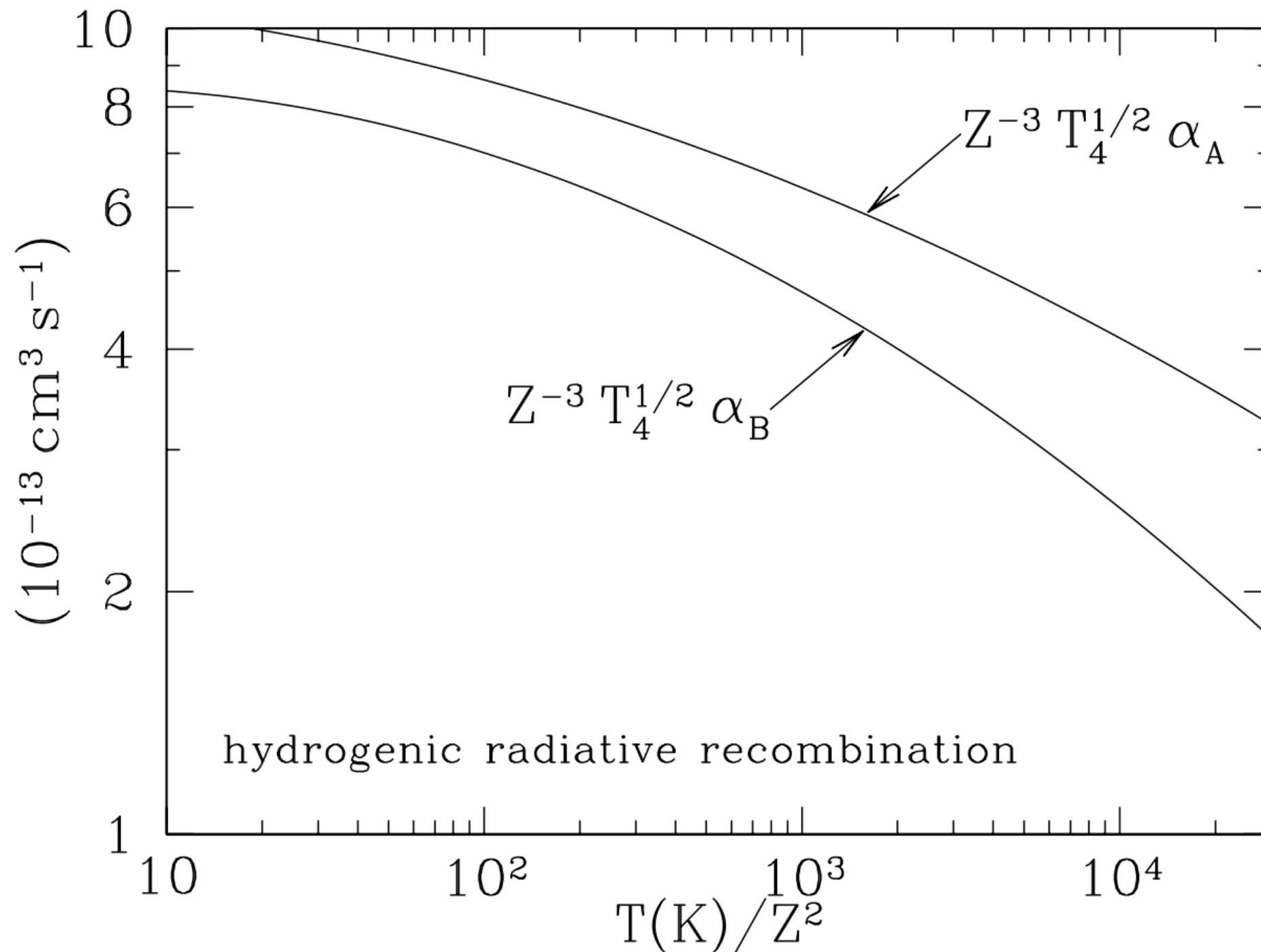
Radiative recombination coefficients

Table 14.1 Recombination Coefficients $\alpha_{n\ell}$ ($\text{cm}^3 \text{s}^{-1}$) for H.^a The approximation formulae are valid for $0.3 \lesssim T_4 \lesssim 3$. For a broader range of T , see Eq. (14.5,14.6).

$\alpha_n(^2L)$	Temperature T			
	$5 \times 10^3 \text{ K}$	$1 \times 10^4 \text{ K}$	$2 \times 10^4 \text{ K}$	approximation
α_{1s}	2.28×10^{-13}	1.58×10^{-13}	1.08×10^{-13}	$1.58 \times 10^{-13} T_4^{-0.539-0.014 \ln T_4}$
α_{2s}	3.37×10^{-14}	2.34×10^{-14}	1.60×10^{-14}	$2.34 \times 10^{-14} T_4^{-0.537-0.016 \ln T_4}$
α_{2p}	8.33×10^{-14}	5.35×10^{-14}	3.24×10^{-14}	$5.35 \times 10^{-14} T_4^{-0.681-0.061 \ln T_4}$
α_2	1.17×10^{-13}	7.69×10^{-14}	4.84×10^{-14}	$7.69 \times 10^{-14} T_4^{-0.637-0.045 \ln T_4}$
α_{3s}	1.13×10^{-14}	7.81×10^{-15}	5.29×10^{-15}	$7.81 \times 10^{-15} T_4^{-0.548-0.021 \ln T_4}$
α_{3p}	3.17×10^{-14}	2.04×10^{-14}	1.23×10^{-14}	$2.04 \times 10^{-15} T_4^{-0.683-0.068 \ln T_4}$
α_{3d}	3.03×10^{-14}	1.73×10^{-14}	9.09×10^{-15}	$1.73 \times 10^{-14} T_4^{-0.869-0.087 \ln T_4}$
α_3	7.33×10^{-14}	4.55×10^{-14}	2.67×10^{-14}	$4.55 \times 10^{-14} T_4^{-0.729-0.059 \ln T_4}$
α_{4s}	5.23×10^{-15}	3.59×10^{-15}	2.40×10^{-15}	$3.59 \times 10^{-15} T_4^{-0.562-0.028 \ln T_4}$
α_{4p}	1.51×10^{-14}	9.66×10^{-15}	5.81×10^{-15}	$9.66 \times 10^{-15} T_4^{-0.689-0.064 \ln T_4}$
α_{4d}	1.90×10^{-14}	1.08×10^{-14}	5.68×10^{-15}	$1.08 \times 10^{-14} T_4^{-0.871-0.081 \ln T_4}$
α_{4f}	1.09×10^{-14}	5.54×10^{-15}	2.56×10^{-15}	$5.54 \times 10^{-15} T_4^{-1.045-0.099 \ln T_4}$
α_4	5.02×10^{-14}	2.96×10^{-14}	1.65×10^{-14}	$2.96 \times 10^{-14} T_4^{-0.803-0.059 \ln T_4}$
α_A	6.82×10^{-13}	4.18×10^{-13}	2.51×10^{-13}	$4.18 \times 10^{-13} T_4^{-0.721-0.021 \ln T_4}$
α_B	4.54×10^{-13}	2.59×10^{-13}	1.43×10^{-13}	$2.59 \times 10^{-13} T_4^{-0.833-0.034 \ln T_4}$

^a From Hummer & Storey (1987)

Radiative recombination coefficients



Case A: optically thin,
recombination photons
escape

Case B: optically thick
above ionization energy,
recombination photons to
ground state ionize again

§5.3 • Ionization balance and the Milne equation

Deriving the Milne relation

Detailed balance:

$$n_0 \frac{4\pi J_\nu}{h\nu} \sigma_{\text{pi}}(h\nu) d\nu = n_e n_p u f_{\text{MB}}(u) \sigma_{\text{nl}}(u) [1 + \bar{n}_\gamma] du$$

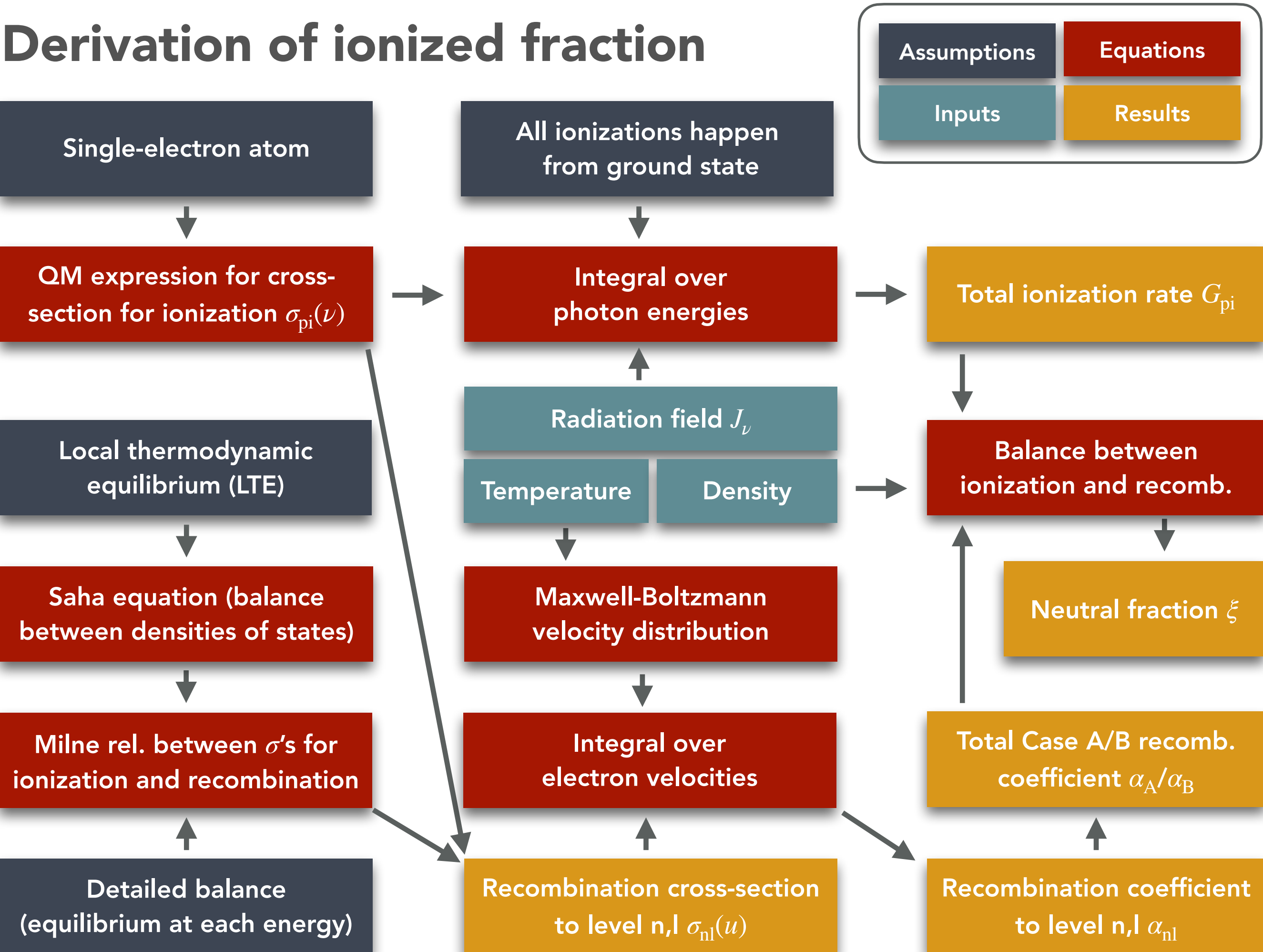
Electron energy:

$$\frac{m_e u^2}{2} = h(\nu - \nu_0) \quad \Rightarrow \quad u = \sqrt{\frac{2h(\nu - \nu_0)}{m_e}} \quad \Rightarrow \quad du = \frac{h}{m_e u} d\nu.$$

Ratio between
ionization and
recombination
cross-sections:

$$\begin{aligned} \frac{\sigma_{\text{nl}}(u)}{\sigma_{\text{pi}}(h\nu)} &= \frac{n_0}{n_e n_p} \frac{4\pi J_\nu}{h\nu} \frac{1}{(1 + \bar{n}_\gamma)} \frac{1}{u f_{\text{MB}}(u)} \frac{m_e u}{h} \\ &= \frac{g_n}{g_e g_p} \left(\frac{h^2}{2\pi m_e k_B T} \right)^{3/2} e^{h\nu_0/k_B T} \frac{4\pi}{h\nu} \frac{2h\nu^3}{c^2} \frac{1}{(e^{h\nu/k_B T} - 1)} \frac{1}{1 + \frac{1}{e^{h\nu/k_B T} - 1}} \\ &\quad \times \frac{\sqrt{\pi}}{4} \left(\frac{2k_B T}{m_e} \right)^{3/2} \frac{1}{u^2} e^{m_e u^2/2k_B T} \frac{m_e}{h} \\ &= \frac{g_n}{g_e g_p} \frac{2\pi^{3/2} \nu^2 m_e}{hc^2} \left(\frac{h^2}{\pi m_e^2} \right)^{3/2} \frac{1}{u^2} \exp \left[\frac{h\nu_0}{k_B T} - \frac{h\nu}{k_B T} + \frac{h(\nu - \nu_0)}{k_B T} \right]. \end{aligned}$$

Derivation of ionized fraction



Reading

Draine

- §13.1
- §14.1-2

O&F

- §2.1-2