The background of the slide is a high-resolution aerial photograph of a dry riverbed on Mars. The riverbed is a complex network of light-colored, sandy channels and ridges, winding across a darker, more textured terrain. The lighting creates strong shadows, highlighting the three-dimensional structure of the riverbed. The overall color palette is a mix of light tan, beige, and dark brown.

An Examination of the Possibility of
Flowing Liquid Water on Mars

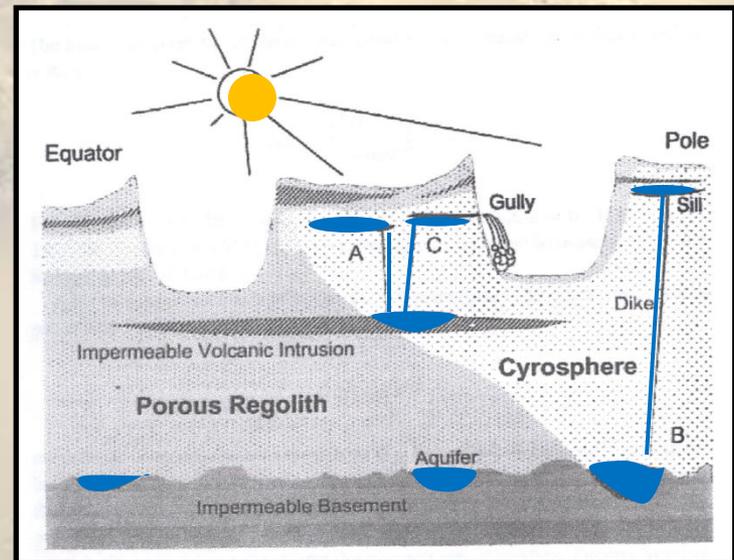
By
Michael Brazelton

TERPS Conference 2012
University of Maryland

Source Paper

Cryovolcanism and the Recent Flow of Liquid Water on Mars by Eric J. Gaidos

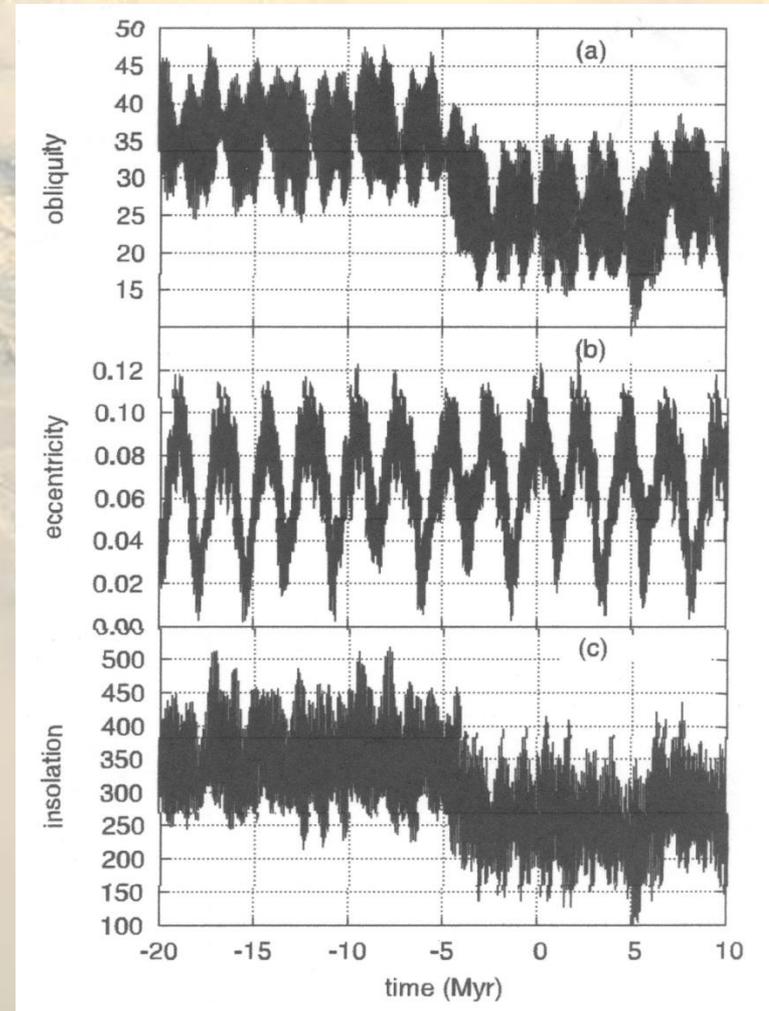
- Presupposes set of geological conditions that would lead to surface eruption of underground water.
- Complicated circumstances and process
- One of many papers on the subject of flowing liquid water





Mars Obliquity

- Obliquity as high at 47 degrees
 - Vaporizing the water at N. Pole
 - Coldest near equator
- Eccentricity as high as .12
 - Increasing the distance from sun at aphelion
- After high obliquity, Mars received less energy from the sun
 - Preserving non-polar ice fields





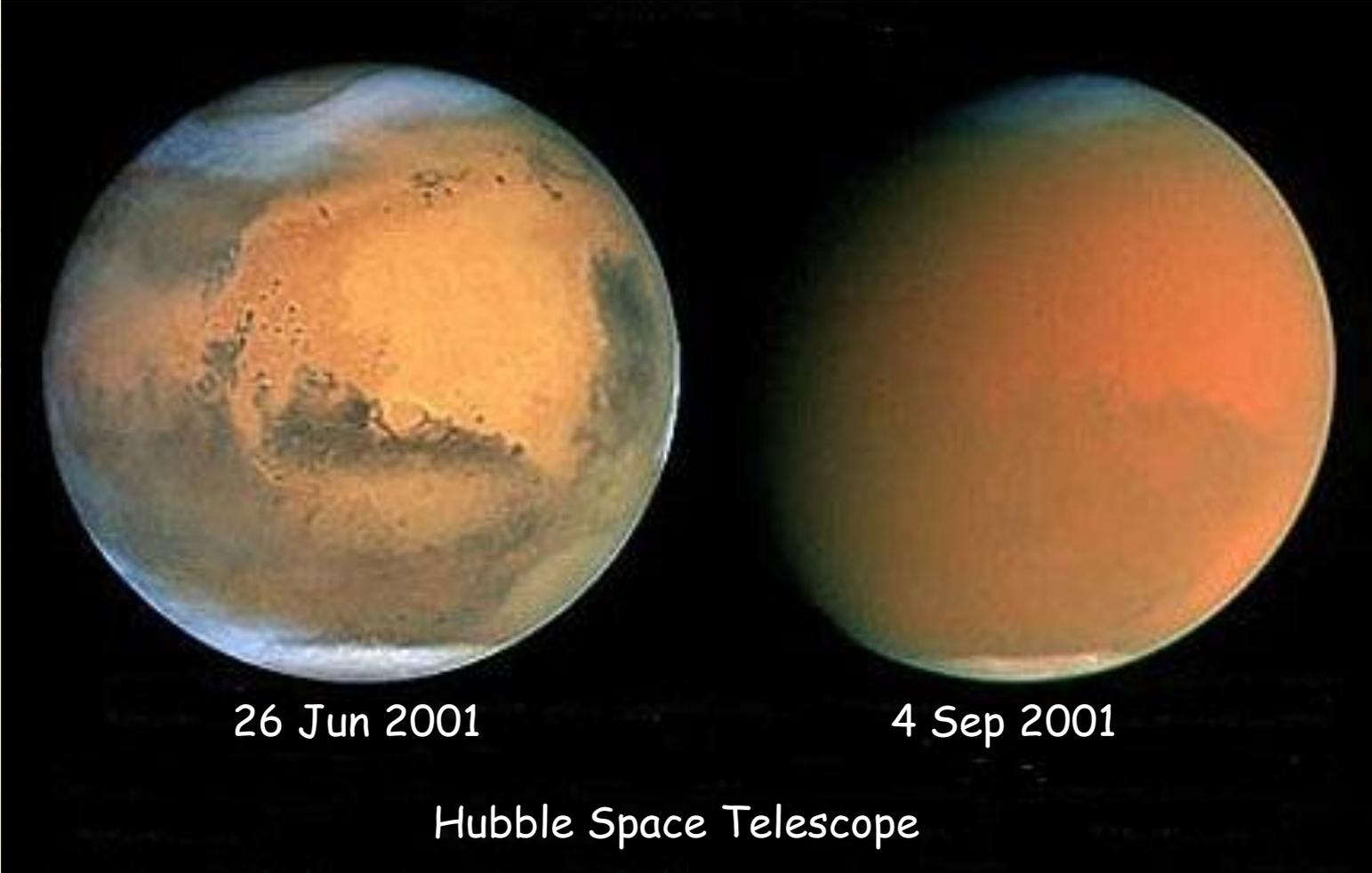
Sources of Water on Mars

- North Pole
- South Pole
- (Black) Glaciers
- (Black) Ice Sheets
- (Black) Lakes
- Permafrost
- Hot Springs
- Underground Water/Ice

- Alternatives to flowing water
 - Dry flows
 - CO_2 crystals or Liquid CO_2
 - Ice crystals



Martian Dust Storms



26 Jun 2001

4 Sep 2001

Hubble Space Telescope



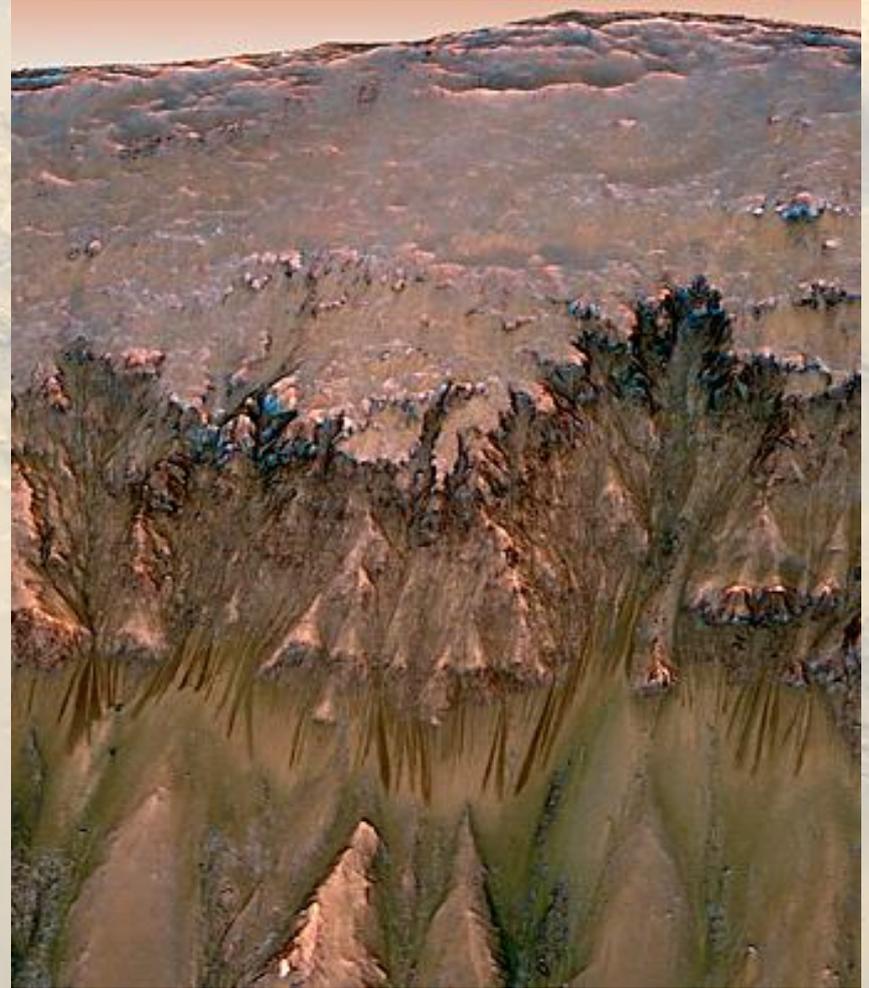
Thermal Properties of Soil

- Dirty coverings on ice generally cause the snow to melt faster due to greater thermal absorption
- Until the covering gets thick enough to act as an insulating blanket
- On Earth, 20cm of soil can hold the subsurface at the average of the day/night temperatures $\pm 2^{\circ}\text{C}$
- The summertime temperature on Mars is $+15^{\circ}\text{C}$ to -120°C . The average temperature would be approximately -55°C (218K)



Recent Flowing Water on Mars

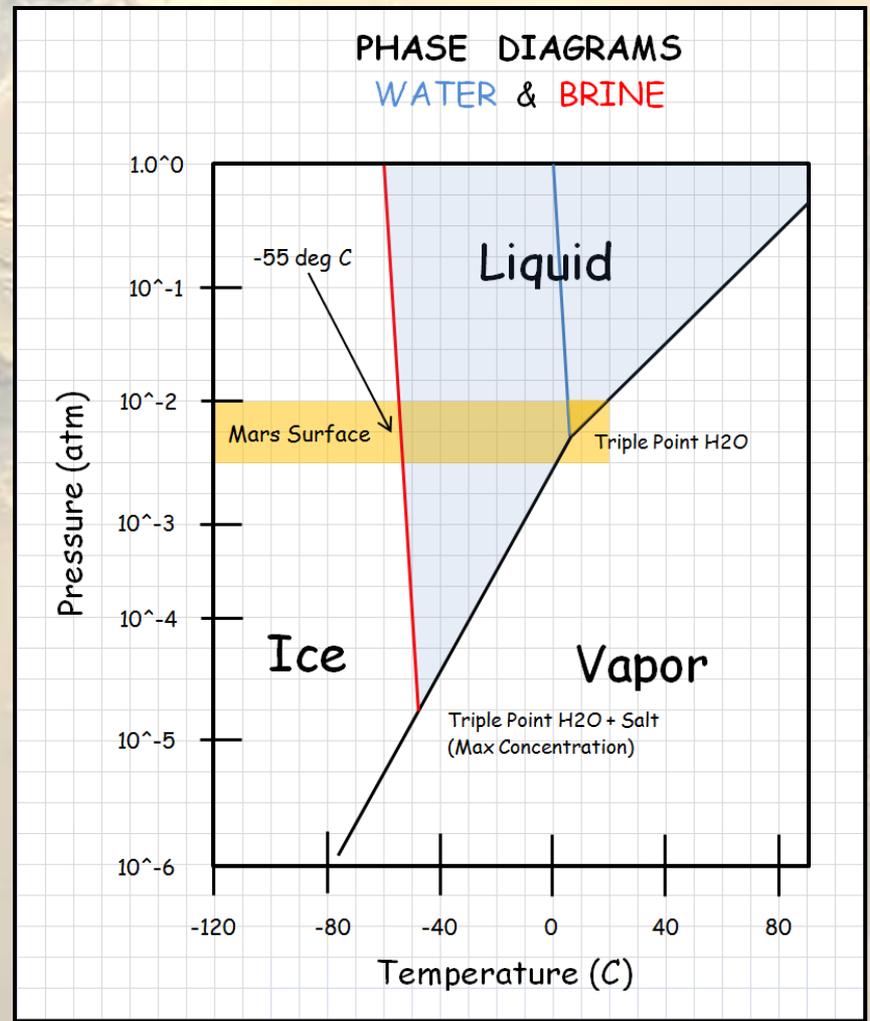
- Repeatable
- Observed reoccurring for 3 years
- Seasonal
 - During the warm season in southern hemisphere
- Sloping terrain
- Pole-facing slopes
- Multiple outlets
- Multiple locations





Triple Point of Water & Brine

- The absorption of salts drastically lowers the boiling point of water making it liquid at a much lower temperature and pressure
- Salts are known to be present in Martian soil (NaCl, MgCl, CaCl)





Carbon Dioxide/Salt *Clathrate*

- Carbon dioxide reacts with salt water to form a solid phase carbon dioxide clathrate
- Decomposition of this phase produces carbon dioxide gas in equilibrium with water - like carbonated salt water
- Melting will begin at the carbon dioxide-brine eutectic point
 - Eutectic Point: **Boiling point** of a mixture of several components, which melts sharply below the melting point of any individual component
- **The decomposition of a calthrate brine produces brine + a large volume of gas "capable of considerable mechanical energy"**



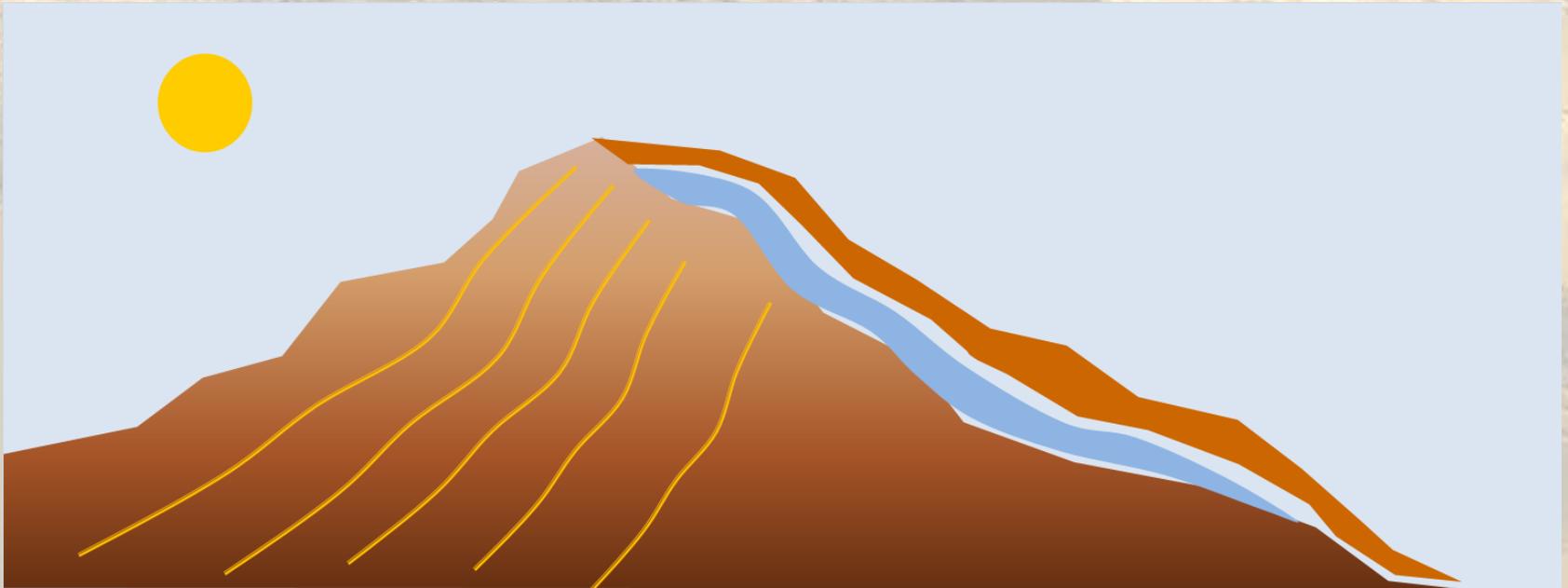
Hypothesis

- Water source: Black glaciers or huge insulated ice deposits
- Liquid: Water, Brine or brine & Carbon-dioxide clathrate
- Process:
 - When the southern hemisphere heats up, the underside of the ice flow reaches its melting point. Briny water, above its triple point fills the lower part of the cavity which is blocked with an ice plug.
 - Carbon dioxide comes out of solution and pressurizes the under-ice cavities
 - Gravity and CO₂ pressure force water out between sloped ridge surface and black ice cover
 - When the water is drained, the pressure is relieved, the drainage holes are plugged with frozen water
 - The process then begins all over.



Protected Ice Sheet

- Sun heats sunward side of ridge
- Ridge transmits heat to underside of ice field



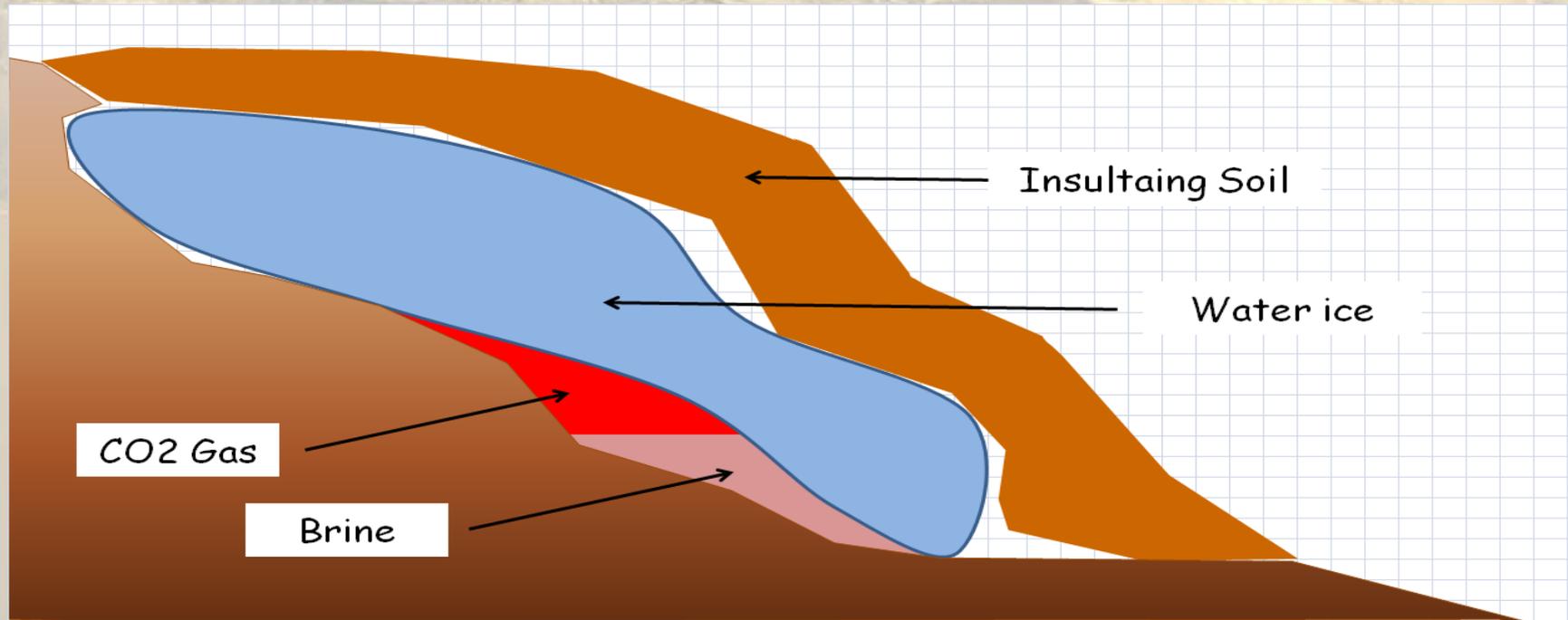


Water Flowing Under Ice Sheet



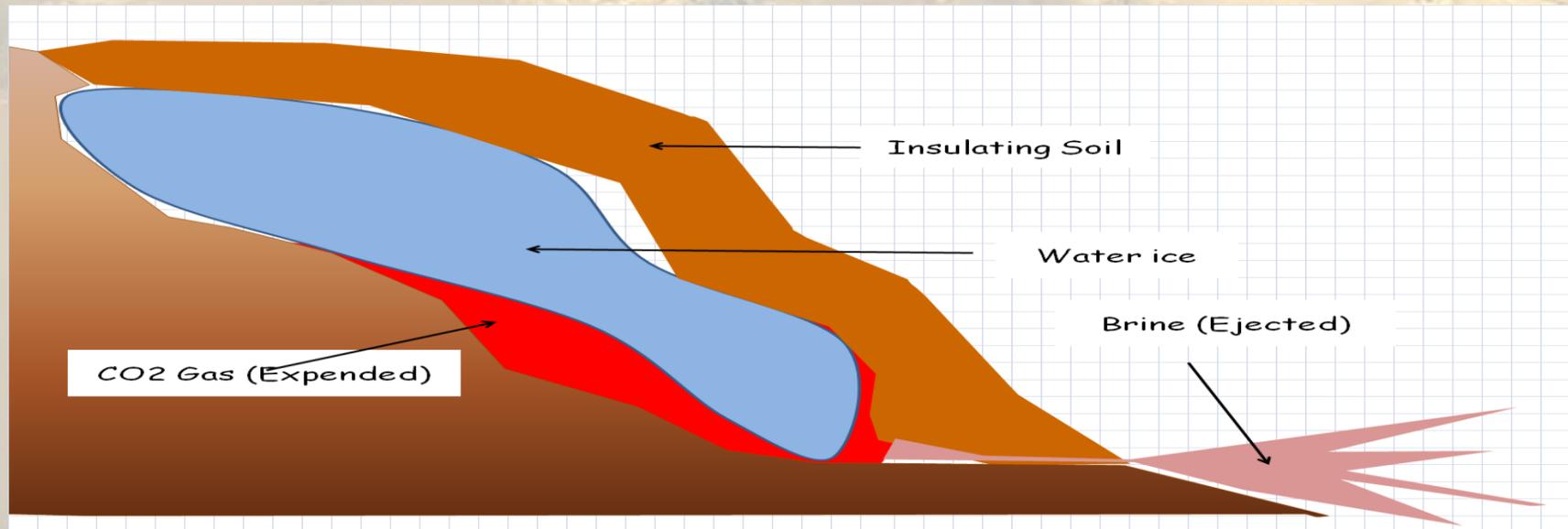
Build-up of Pressure

- Heat melts water ice/brine and releases CO_2 gas which remain trapped while weight of brine and pressure of gas build up



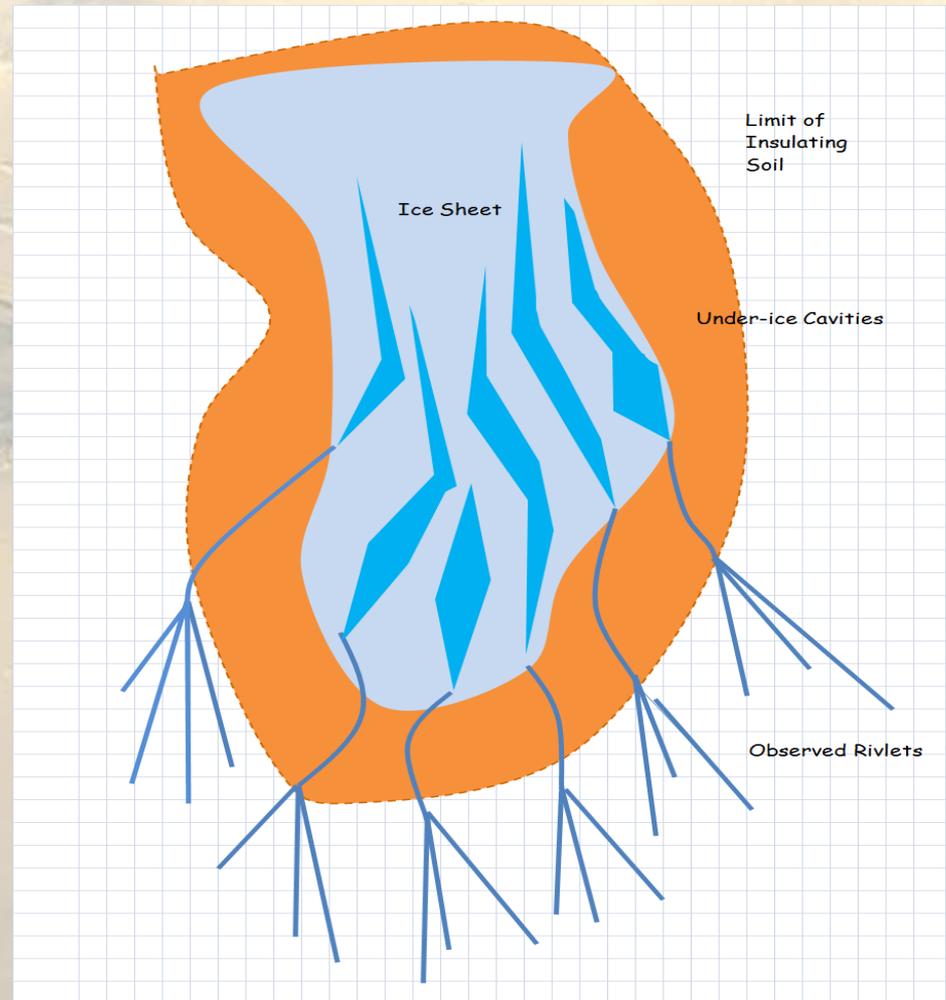
Ejection of Brine

- Pressure of gas and brine push through the crust and into the open to flow until evaporated
- Upon release of pressure, exit to atmosphere is re-blocked by ice and process begins to repeat



Flowing Water on Mars

- Brine gushes out of regolith covering with enough mass and volume to form gullies and alluvial fan formations before evaporating in low pressure atmosphere





Questions?

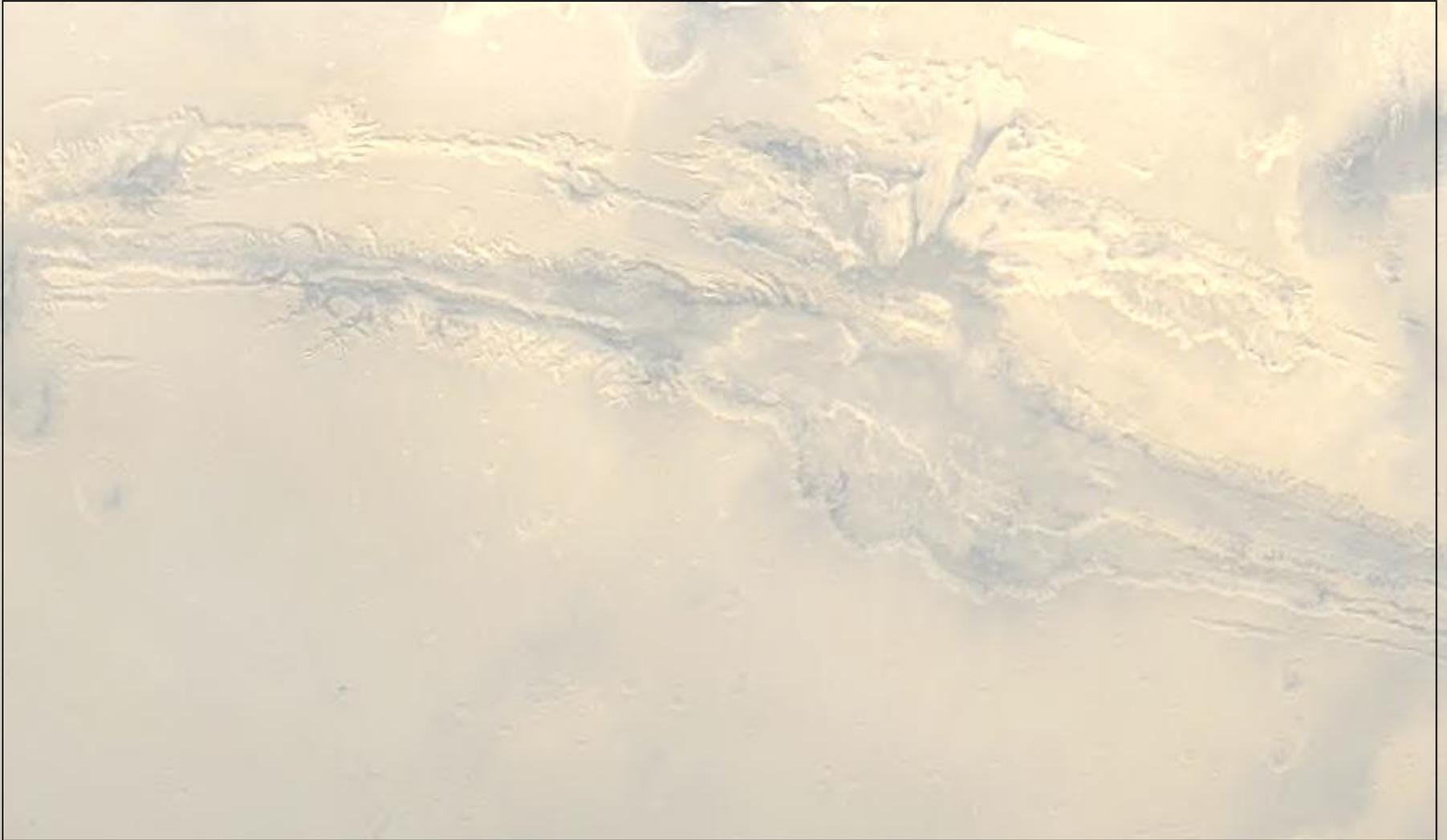




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Additional Slides

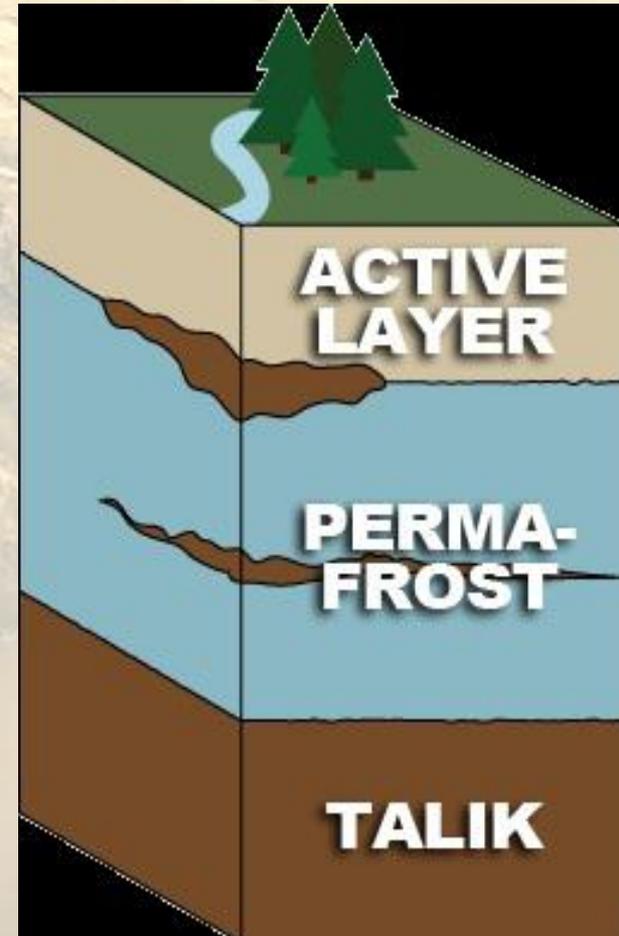
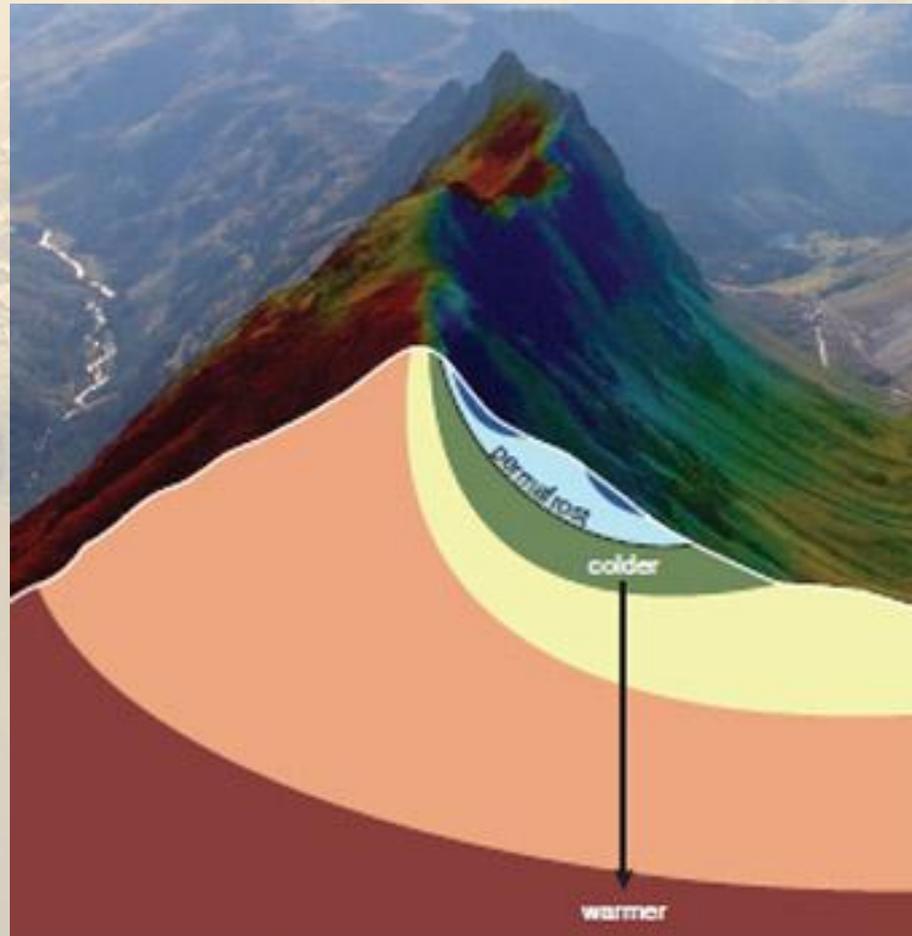




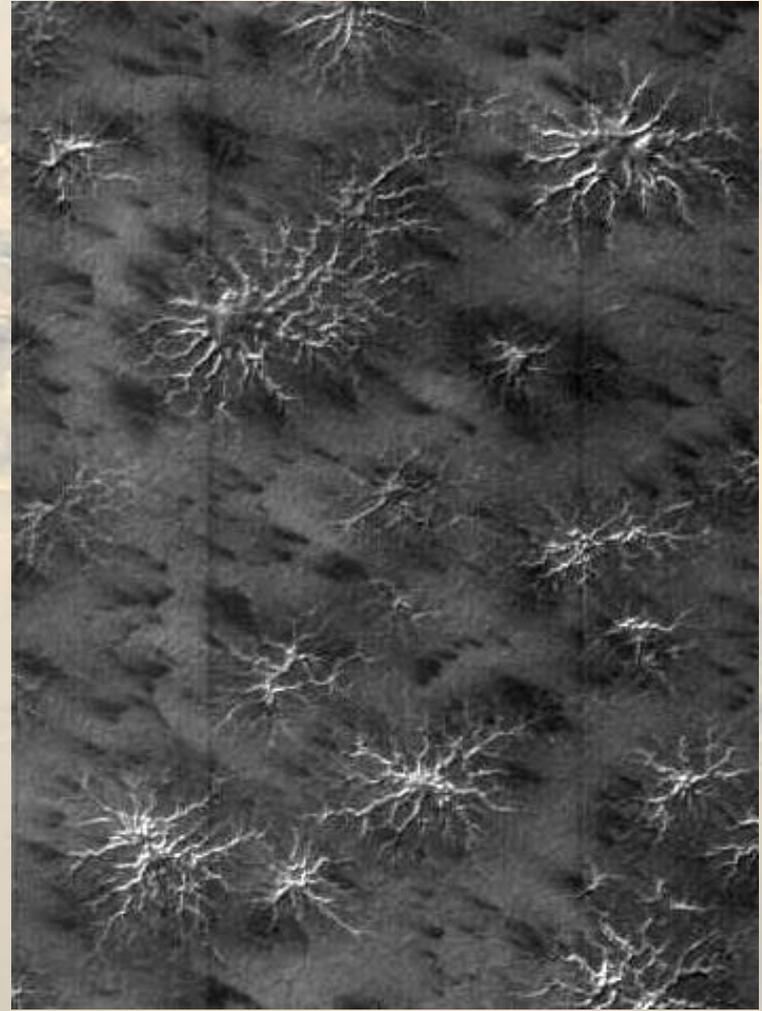
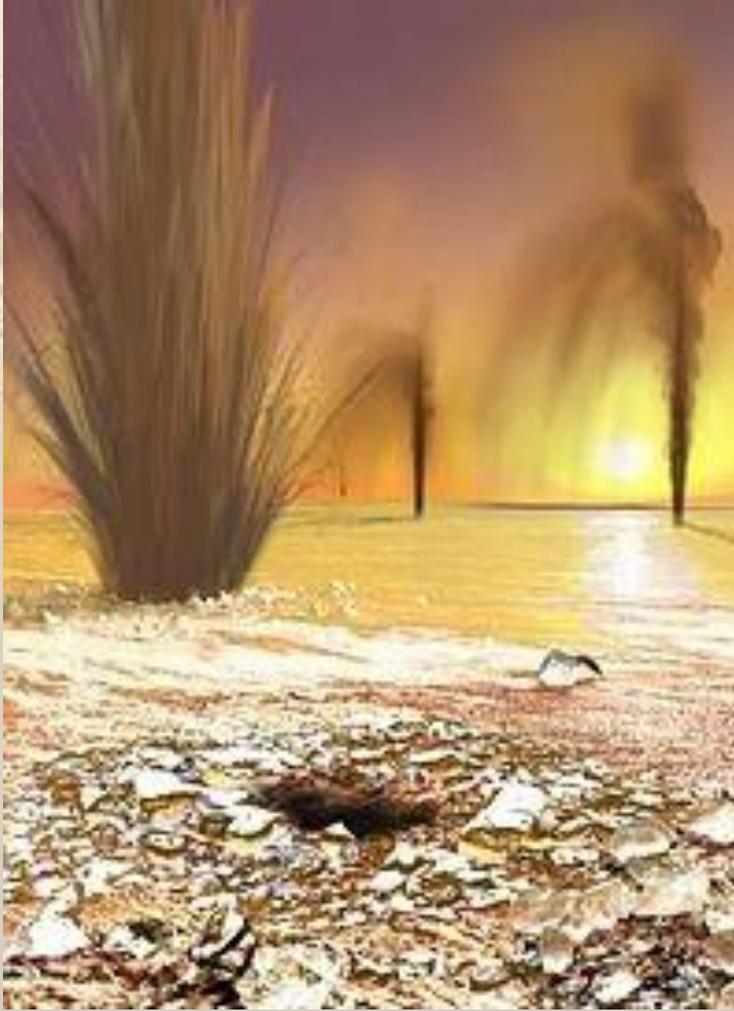
Why Finding Water is Important

- Locate potential life forms on Mars
- Support a long-term manned Mars exploration program
 - Drinking/bathing/laundry/food preparation
 - Source of oxygen
 - For breathing
 - For fuel cells to power ground equipment
 - For rocket propellant oxidizer
 - Source of hydrogen
 - For fuel cells
 - To combine with atmospheric CO_2 to produce methane for rocket fuel (CH_4)

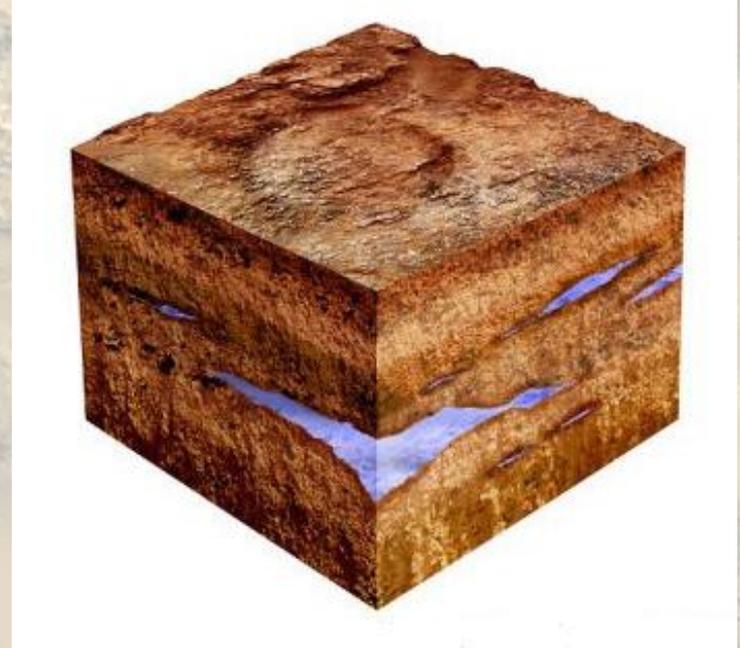
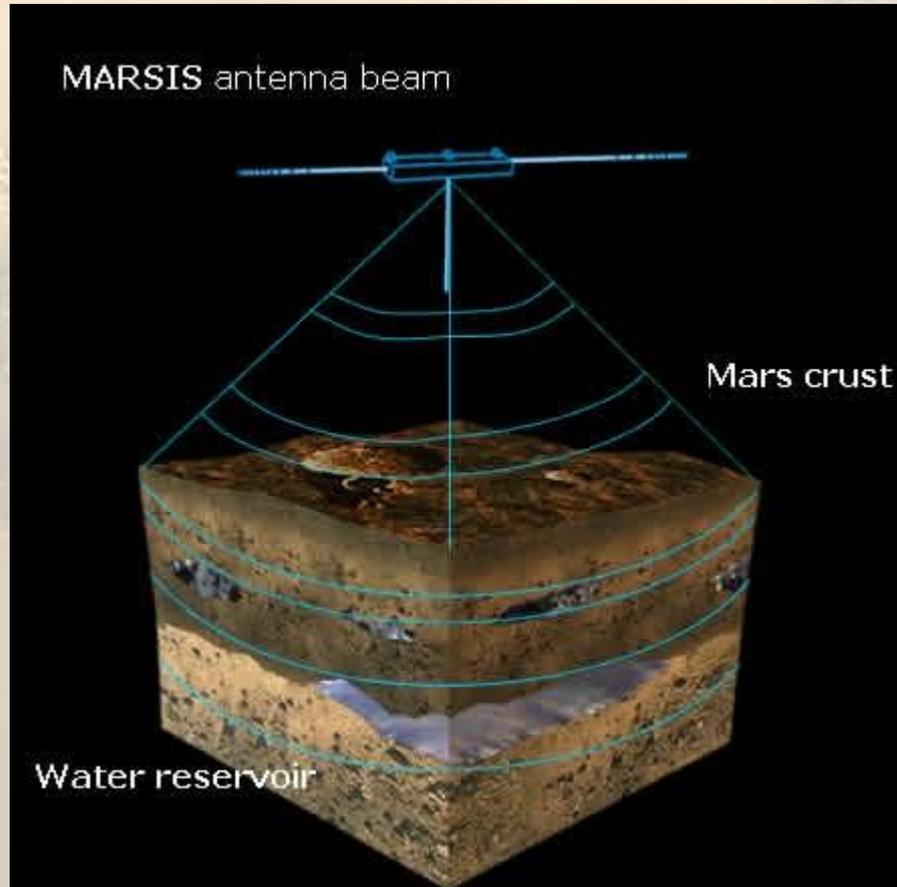
Permafrost



Geysers on Mars



Underground Water

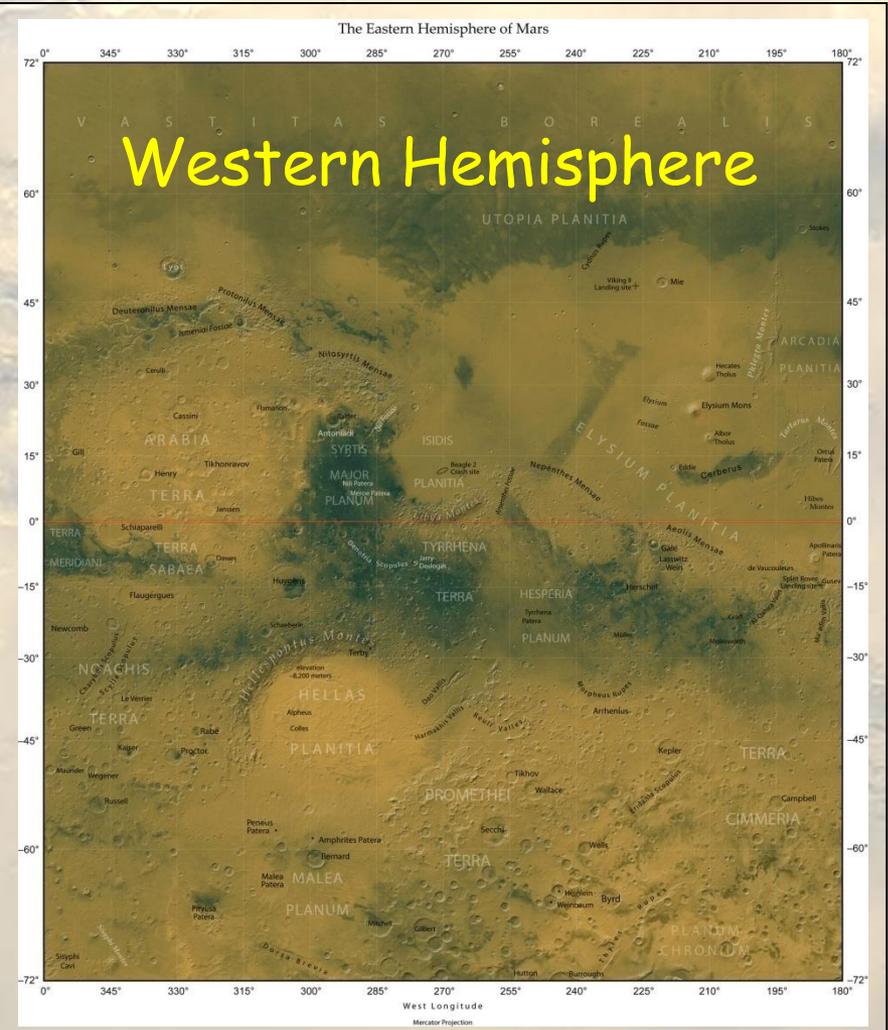
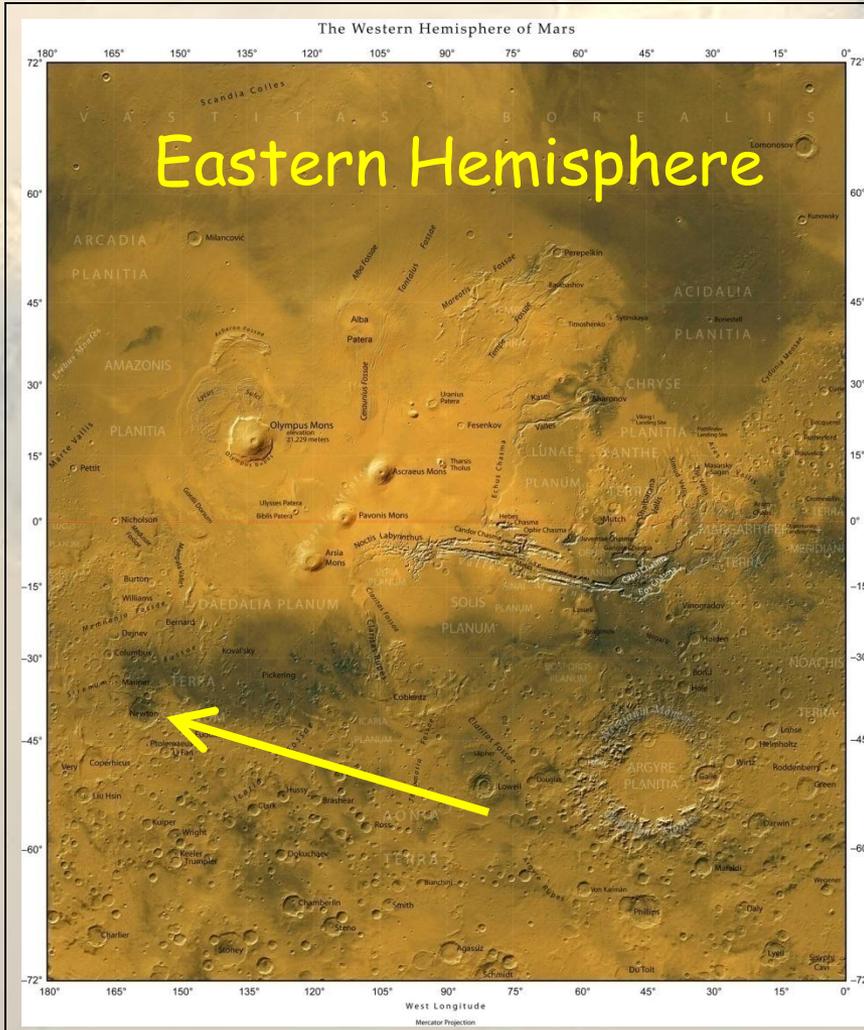


Black Glaciers / Ice Sheets

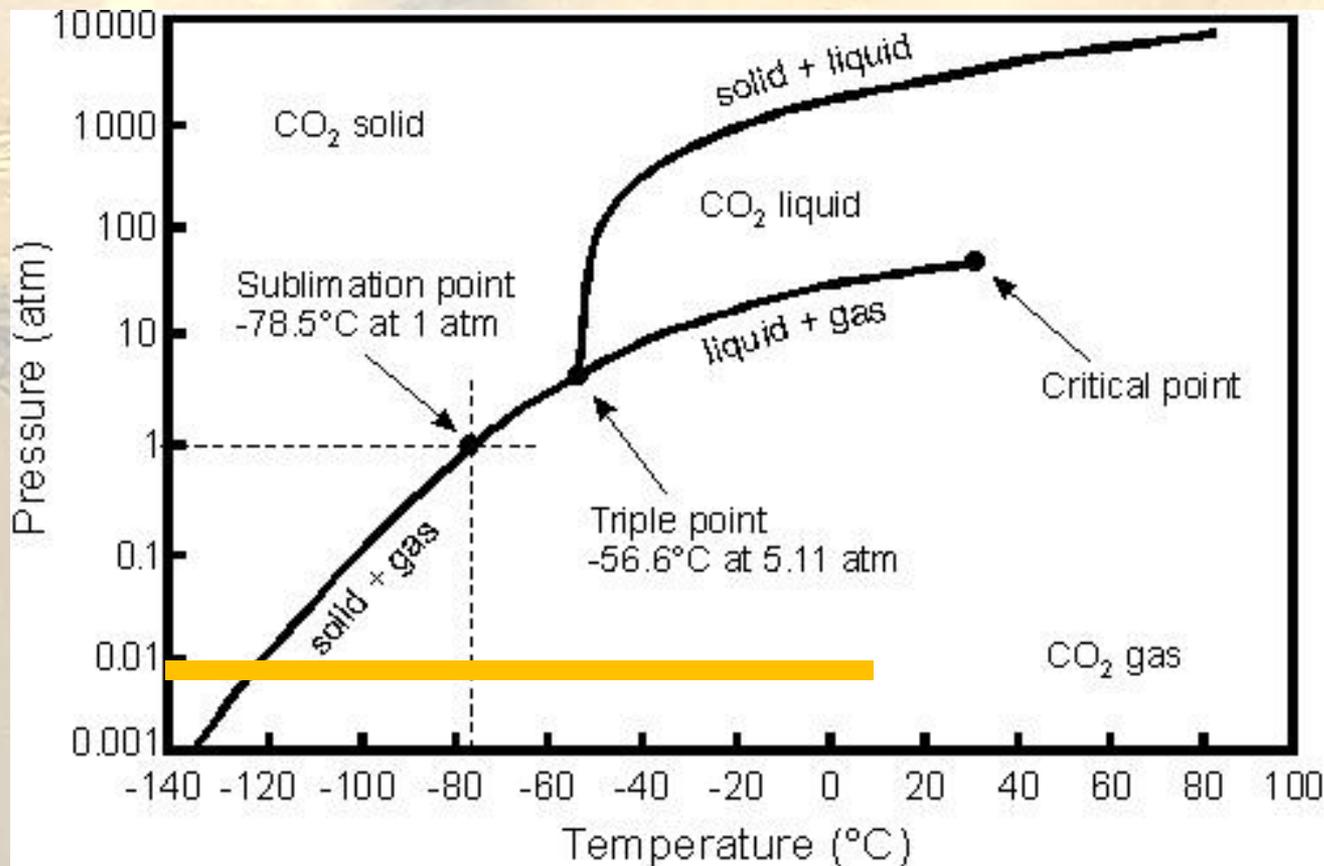
- On Earth:
Partially covered
by avalanches,
forests, etc.
- On Mars:
Completely
covered with
Martian soil
from wind &
storms for
millions of years



Newton Crater

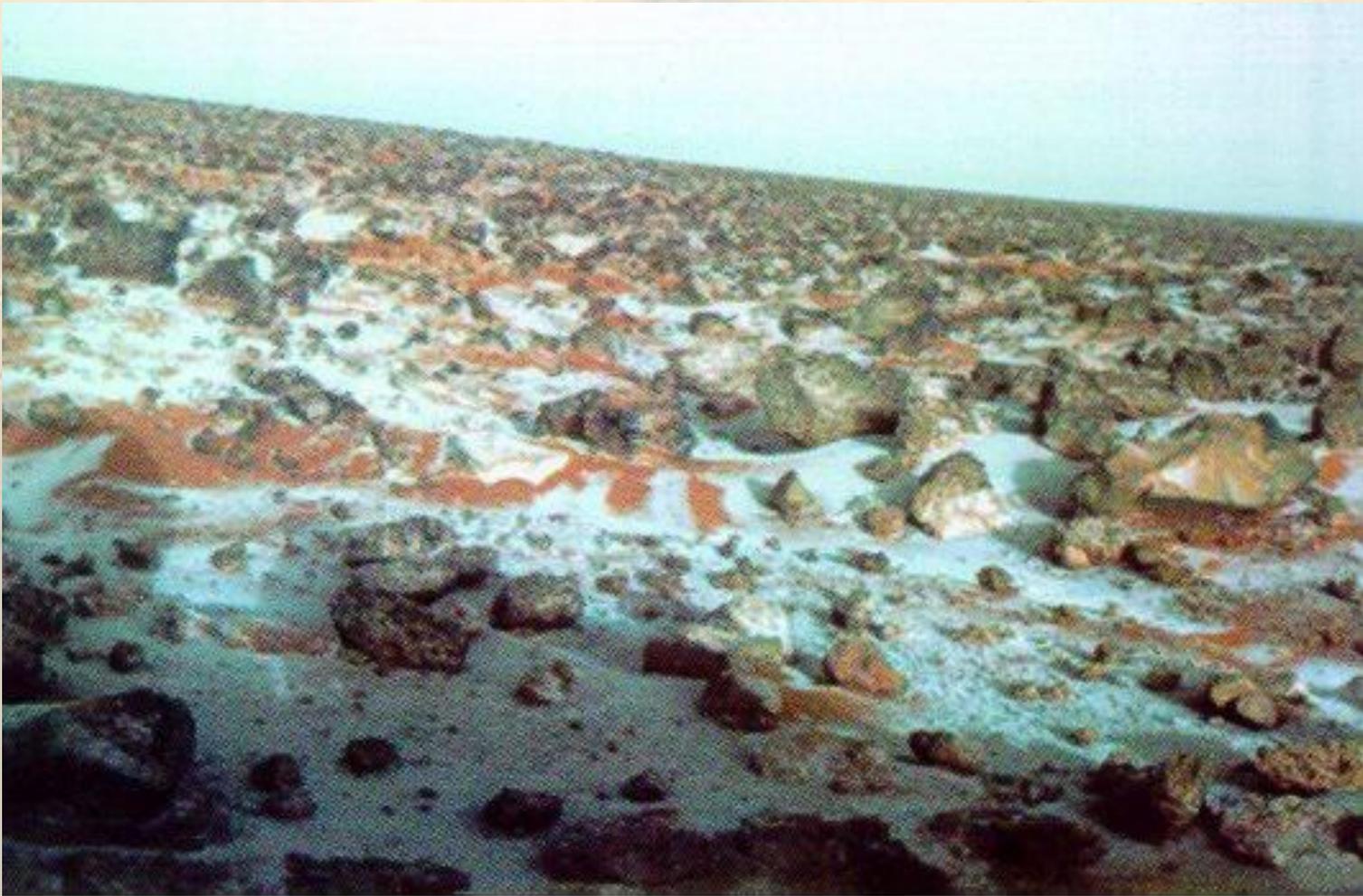


Phase Diagram - CO₂



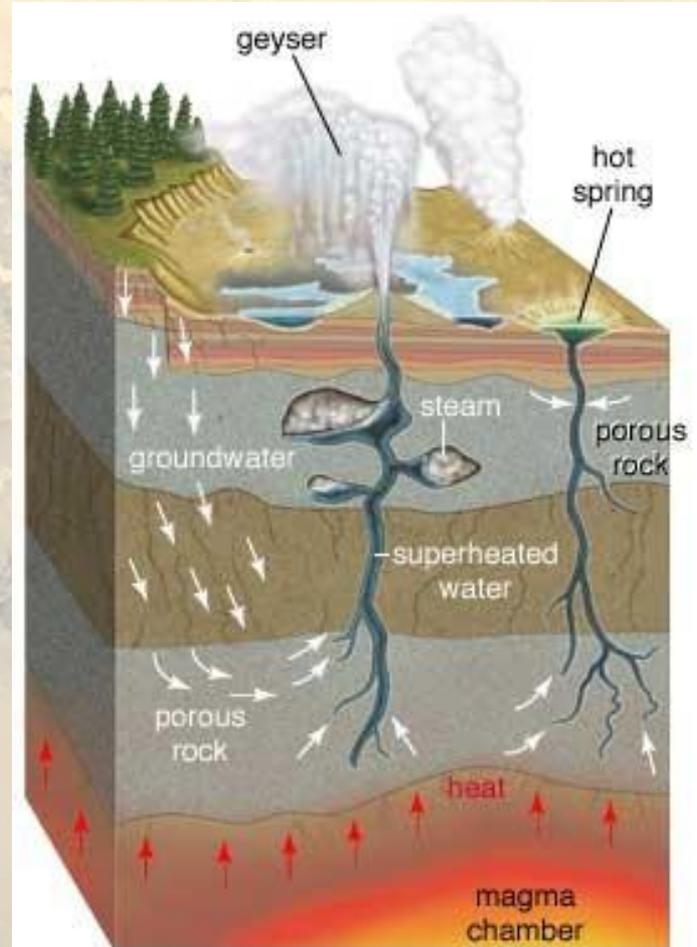
Pressure-Temperature phase diagram for CO₂.

Frost on Surface of Mars



Old Faithful

- Completely unique set of circumstances that results in a singularly unique event
- Closest similar events are steam vents and hot springs



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Running Water / Freezing Conditions



Ice Sheet

