HYPERVELOCITY STARS (HVSs)

SOME HVSs HAVE PLANETS, TOO

A MASSIVE BLACK HOLE CAN DISRUPT A BINARY STAR SYSTEM AND CREATE A HVS FROM ONE OF THE BINARY STARS; CAPTURE THE OTHER COMPANION STAR

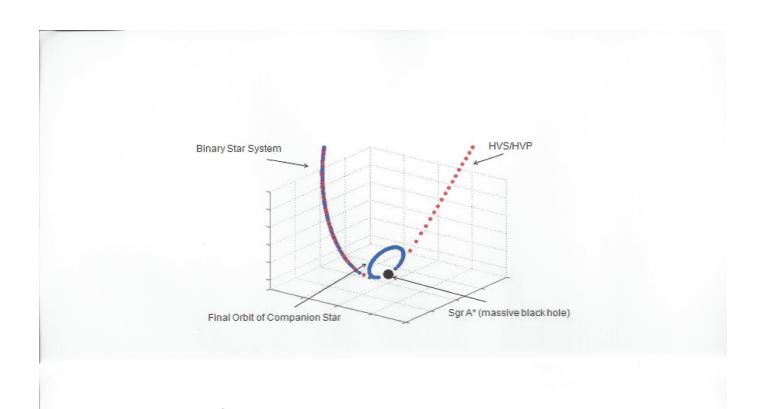
OUTLINE OF TALK

- BACKGROUND
- BASIC THEORY OF THE ORIGIN OF HVSs
- SIMULATION RESULTS
- SOME SIMPLE CALCULATIONS
- WHAT IS THE FATE OF HVSs WITH PLANETS?
- WHY I CAN'T PROVE CONSERVATION OF ENERGY

BACKGROUND

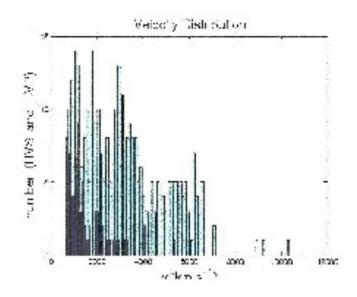
- WHEN A BINARY STAR SYSTEM APPROACHES A MBH, THE SYSTEM CAN BE TIDALLY DISRUPTED
- THINK: <u>THE ROCHE LIMIT</u> and <u>TIDAL FORCES</u>
- ONE OF THE STARS PICKS UP KINETIC ENERGY+ TO BECOME UNBOUND; SECOND STAR ORBITS MBH
- EJECTION SPEEDS OF HVSs CALCULATED TO EXCEED THE ESCAPE VELOCITY OF THE MILKY WAY GALAXY
- HVSs SPEEDS CONFIRMED WITH SPECTROSCOPY

PRODUCTION OF A HYPERVELOCITY STAR



SIMULATION RESULTS

(SAMPLE = 1000)



SIMPLE CALCULATIONS

$$V_{\rm esc} = \sqrt{2\,G\,M/r} \implies G = 6.674E-11\,m^3/kg/sec^2$$
 $M = 5E11~(2E30), mass~MW~halo$
 $r = 50~kpc = 50(3.1E19)$
 $V_{\rm esc} = 292.88~E3 = 293~km/sec$
 $V_{\rm (HVS1)} = spectroscopy \implies radial~velocity = 853~km/sec$
 $V_{\rm (HVS1)} = calculated \implies v = \sqrt{V_0}V \quad ; \delta v = \sqrt{Gm}*/\alpha \; ;$
 $v \approx \sqrt{GM(bh)/r}; \sqrt{v_0}v = 1,000~km/sec$

THE FATE OF COMPANIONS TO HVSs?

- IF A HVS HAS PLANETS, CAN A PLANET STILL ORBIT A HVS? SIMULATIONS SHOW UP TO 10% OF PLANETS TIGHTLY ORBITING A HVS COULD STAY BOUND AS THE STAR IS FLUNG OUTWARD.
- WHAT HAPPENS TO THE STAR CAPTURED BY THE BLACK HOLE? THE FORMER COMPANION STAR TO A HVS IS TYPICALLY KICKED INTO A HIGHLY ECCENTRIC ORBIT AROUND THE BH; WITH SHORT PERIODS OF 36 100+YEARS.
- CAN PLANETS BECOME HYPER-V PLANETS? YES: EITHER FLUNG OFF THE HVS; OR RIPPED FROM THE COMPANION, CAPTURED STAR (ORBITING THE BLACK HOLE).

WHY I CAN'T CONCLUSIVELY PROVE CONSERVATION OF ENERGY

- I CAN ROUGHLY ESTIMATE THE CHANGE IN VELOCITY OF THE EJECTED STAR AND THE ASSOCIATED DELTA ENERGY OF THE HVS.
- HOWEVER, THERE IS INSUFFICIENT INFORMATION ON THE RESULTING ORBIT OF THE STAR CAPTURED BY THE BLACK HOLE, FOR GIVEN HVS.
- BOTH THE EJECTION SPEED OF THE HVS AND THE SIZE/ PERIOD OF THE CAPTURED STAR ARE HIGHLY DEPENDENT ON ORIENTATION OF THE BINARY WHEN IT GETS DISRUPTED, AS WELL BINARY SEPARATION.
- SIMULATION IS REALLY A BETTER TOOL!