

Three years of Ulysses dust data: 2005 to 2007

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Abstract

The Ulysses spacecraft has been orbiting the Sun on a highly inclined ellipse ($i = 79^\circ$, perihelion distance 1.3 AU, aphelion distance 5.4 AU) since it encountered Jupiter in February 1992. Since then it made almost three revolutions about the Sun. Here we report on the final three years of data taken by the on-board dust detector. During this time, the dust detector recorded 609 dust impacts of particles with masses $10^{-16} \text{ g} \leq m \leq 10^{-7} \text{ g}$, bringing the mission total to 6719 dust data sets. The impact rate varied from a low value of 0.3 per day at high ecliptic latitudes to 1.5 per day in the inner solar system. The impact direction of the majority of impacts between 2005 and 2007 is compatible with particles of interstellar origin, the rest are most likely interplanetary particles. We compare the interstellar dust measurements from 2005/2006 with the data obtained during earlier periods (1993/1994) and (1999/2000) when Ulysses was traversing the same spatial region at southern ecliptic latitudes but the solar cycle was at a different phase. During these three intervals the impact rate of interstellar grains varied by more than a factor of two. Furthermore, in the two earlier periods the grain

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22 impact direction was in agreement with the flow direction of the interstellar helium
23 while in 2005/2006 we observed a shift in the approach direction of the grains by
24 approximately 30° away from the ecliptic plane. The reason for this shift remains
25 unclear but may be connected with the configuration of the interplanetary magnetic
26 field during solar maximum. We also find that the dust measurements are in agreement
27 with the interplanetary flux model of Staubach et al. (1997) which was developed to
28 fit a 5-year span of Ulysses data.

29 **1 Introduction**

30 Ulysses was the only space mission so far that left the ecliptic plane and flew over the poles
31 of the Sun. The spacecraft was launched in 1990 and was very successfully operated un-
32 til 30 June 2009, although individual instruments had to be turned off to conserve power.
33 Its orbital plane was almost perpendicular to the ecliptic plane (79° inclination) with an
34 aphelion at Jupiter. This special orbit orientation allowed Ulysses to unambiguously detect
35 interstellar dust grains entering the heliosphere because the spacecraft's orbital plane was
36 almost perpendicular to the flow direction of the interstellar dust. Ulysses had a highly
37 sensitive impact ionisation dust detector on board which measured impacts of micrometre
38 and sub-micrometre dust grains. The detector was practically identical with the dust in-
39 strument which flew on board the Galileo spaceprobe. Both instruments were described in
40 previous publications by Grün et al. (1992a,b, 1995c).

41 **1.1 Summary of results from the Ulysses dust investigations**

42 Comprehensive reviews of the scientific achievements of the Ulysses mission including
43 results from the dust investigation were given by Balogh et al. (2001) and Grün et al.
44 (2001). References to other works related to Ulysses and Galileo measurements on dust
45 in the planetary system were also given by Grün et al. (1995a,b); Krüger et al. (1999a,b,
46 2001a,b, 2006b,a); Krüger and Grün (2009). Some mission highlights are also summarized
47 in Table 1.

48 Various dust populations were investigated with the Ulysses and Galileo dust experiments
49 in interplanetary space: the interplanetary dust complex including β -meteoroids (i.e. dust
50 particles which leave the solar system on unbound orbits due to acceleration by radiation
51 pressure), interstellar grains sweeping through the heliosphere, and dust stream particles
52 expelled from the jovian system by electromagnetic forces, to name only the most signif-
53 icant dust types detected with Ulysses which have been analysed so far. In the following,
54 we summarise the most significant achievements of the Ulysses dust measurements.

55 Ulysses and Galileo dust measurements were used to study the 3-dimensional structure
56 of the interplanetary dust complex and its relation to the underlying populations of parent
57 bodies like asteroids and comets (Divine, 1993; Grün et al., 1997; Staubach et al., 1997).
58 Studies of asteroidal dust released from the IRAS dust bands show that they are not efficient

59 enough dust sources to maintain the observed interplanetary dust cloud (Mann et al., 1996).
60 The state of the inner solar system dust cloud within approximately 1 AU from the Sun in-
61 cluding dust destruction and ion formation processes in relation to so-called solar wind
62 pickup ions detected by the Solar Wind Ion Composition Spectrometer (SWICS) onboard
63 Ulysses was investigated (Mann et al., 2004; Mann and Czechowski, 2005). An improved
64 physical model was developed for the interplanetary meteoroid environment (Dikarev et al.,
65 2001, 2002, 2005) which uses long-term particle dynamics to define individual interplane-
66 tary dust populations. The Ulysses and Galileo in-situ dust data are an important data set
67 for the validation of this model. The properties of β -meteoroids were also studied with
68 the Ulysses data set (Wehry and Mann, 1999; Wehry et al., 2004). Finally, the potential
69 connection of cometary dust trails and enhancements of the interplanetary magnetic field
70 as measured by Ulysses was discussed by Jones and Balogh (2003).

71 During its first flyby at Jupiter in 1992, the Ulysses dust instrument discovered burst-like
72 intermittent streams of tiny dust grains in interplanetary space (Grün et al., 1993) which
73 had been emitted from the jovian system (Hamilton and Burns, 1993; Horányi et al., 1993;
74 Zook et al., 1996). This discovery was completely unexpected as no periodic phenomenon
75 for tiny dust grains in interplanetary space was previously known. These grains strongly
76 interacted with the interplanetary and the jovian magnetic fields (Horányi et al., 1997;
77 Grün et al., 1998) and the majority of them originated from Jupiter's moon Io (Graps et al.,
78 2000). In February 2004 Ulysses had its second Jupiter flyby (at 0.8 AU distance from the
79 planet) and again measured the jovian dust streams (Krüger et al., 2005a, 2006c; Flandes
80 and Krüger, 2007; Flandes et al., 2009).

81 Another important discovery made with Ulysses were interstellar dust particles sweeping
82 through the heliosphere (Grün et al., 1993). The grains which originated from the very
83 local interstellar environment of our solar system were identified by their impact direction
84 and impact velocities, the latter being compatible with particles moving on hyperbolic
85 heliocentric trajectories (Grün et al., 1994). Their dynamics depends on the grain size
86 and is strongly affected by the interaction with the interplanetary magnetic field and by
87 solar radiation pressure (Landgraf et al., 1999; Landgraf, 2000; Mann and Kimura, 2000;
88 Czechowski and Mann, 2003b,a; Landgraf et al., 2003). As a result, the size distribution
89 and fluxes of grains measured inside the heliosphere are strongly modified. Studies of
90 the dust impacts detected with both Ulysses and Galileo showed that the intrinsic size
91 distribution of interstellar grains in the local interstellar environment of our solar system
92 extends to grain sizes larger than those detectable by astronomical observations (Frisch
93 et al., 1999; Frisch and Slavin, 2003; Landgraf et al., 2000; Grün and Landgraf, 2000).
94 Observations of radar meteors entering the Earth's atmosphere at high speeds also indicate
95 the existence of even larger interstellar grains (Taylor et al., 1996; Baggaley and Neslušan,
96 2002).

97 The Ulysses and Galileo interstellar dust measurements showed that the dust-to-gas mass
98 ratio in the local interstellar cloud is higher than the standard interstellar value derived
99 from cosmic abundances (Landgraf, 1998; Frisch et al., 1999). This implied the existence
100 of inhomogeneities in the diffuse interstellar medium on relatively small length scales. In

101 2005/2006 the Ulysses measurements showed a 30° shift in the impact direction of inter-
102 stellar grains with respect to the interstellar helium flow (Krüger et al., 2007). The reason
103 of this shift is presently unclear. Finally, it turned out that the dust sensor side walls have a
104 similar sensitivity to dust impacts as the detector target itself (Altobelli et al., 2004; Willis
105 et al., 2005). This shows that earlier investigations which neglected the contributions of
106 the side wall overestimated the interstellar dust flux by about 20%. Since the contribution
107 of the sensor side wall increases the effective dust instrument field-of-view, the velocity
108 dispersion of the observed interstellar dust stream turned out to be smaller by about 30%
109 than previously thought (Altobelli et al., 2004).

110 Due to its unique highly inclined heliocentric trajectory Ulysses was able to monitor in-
111 terstellar dust at high ecliptic latitudes between 3 and 5 AU. Dust measurements between
112 0.3 and 3 AU in the ecliptic plane exist also from Helios, Galileo and Cassini. This data
113 shows evidence for distance-dependent alteration of the interstellar dust stream caused by
114 radiation pressure, gravitational focussing and electromagnetic interaction of the grains with
115 the time-varying interplanetary magnetic field (Altobelli et al., 2003, 2005b,a).

116 **1.2 The Ulysses and Galileo dust data papers**

117 The Ulysses dust detector obtained dust data for more than 17 years, making it the longest
118 continually-operating spaceborne dust detector to date. With the publication of this paper,
119 the full Ulysses data set, along with 13 years of Galileo data, will be fully described in the
120 scientific literature. The reduction process of Ulysses and Galileo dust data was described
121 by Grün et al. (1995c, hereafter Paper I). In the odd-numbered Papers III, V, VII and IX
122 (Grün et al., 1995a; Krüger et al., 1999b, 2001b, 2006a) we presented the Ulysses data set
123 spanning the time period from launch in October 1990 to December 2004. The companion
124 even-numbered Papers II, IV, VI and VIII (Grün et al., 1995b; Krüger et al., 1999a, 2001a,
125 2006b) discussed the ten years of Galileo data from October 1989 to December 1999. The
126 current paper (Paper XI) extends the Ulysses data set from January 2005 until the Ulysses
127 dust measurements ceased in November 2007, and a companion paper (Krüger et al., 2010,
128 Paper X) presents Galileo's final measurements at Jupiter from 2000 to 2003. A summary
129 of the temporal and spatial coverage of our Ulysses data papers is given in Table 1.

130 The main data products are a table of the impact rate of all impacts determined from the par-
131 ticle accumulators and a table of both raw and reduced data of all dust impacts for which the
132 full data set of measured impact parameters was transmitted to Earth. The information pre-
133 sented in these papers is similar to data which we have submitted to the various data archiving
134 centres (Planetary Data System, NSSDC, Ulysses Data Centre). Electronic access to
135 the data is also possible via the world wide web: <http://www.mpi-hd.mpg.de/dustgroup/>.

136 This paper is organised like our earlier Papers III, V, VII and IX. We begin with an overview
137 of important events of the Ulysses mission between 2005 and 2007 (Section 2). Sections 3
138 and 4 describe and analyse the Ulysses dust data set for this period. In Section 5 we discuss
139 the dust data set from the entire Ulysses mission, we analyse in particular the interstellar
140 dust measurements obtained in the 2005 to 2007 interval, and we compare these results to

141 previous ones. In Section 6 we summarise our conclusions.

142 **2 Mission and instrument operation**

143 **2.1 Ulysses mission and dust instrument characteristics**

144 The Ulysses spacecraft was launched on 6 October 1990. A swing-by manoeuvre at Jupiter
145 in February 1992 rotated the orbital plane 79° relative to the ecliptic plane. On the resulting
146 trajectory (Figure 1) Ulysses finished two full revolutions about the Sun. Passages over
147 the south pole of the Sun occurred in October 1994, November 2000 and February 2007,
148 passages through the ecliptic plane at a perihelion distance of 1.3 AU occurred in March
149 1995, May 2001 and August 2007, and passes over the Sun's north pole were in August
150 1995, October 2001 and January 2008. In April 1998 and July 2004 the spacecraft crossed
151 the ecliptic plane at an aphelion distance of 5.4 AU. Orbital elements for the out-of-ecliptic
152 part of the Ulysses trajectory are given in Paper VII. Figure 1 shows that during most of
153 the time interval considered in this paper Ulysses was at southern ecliptic latitudes.

154 Ulysses is a spinning spacecraft, and the dust sensor orientation at the time of a dust par-
155 ticle impact is recorded, allowing for an independent determination of the grain impact
156 direction. Ulysses spins at five revolutions per minute about the centre line of its high
157 gain antenna which normally points at Earth. Figure 2 shows the deviation of the spin
158 axis from the Earth direction for the period 2005 to 2007. Most of the time the spin axis
159 pointing was within 1° of the nominal Earth direction, similar to the mission before 2005.
160 This small deviation is usually negligible for the analysis of measurements with the dust
161 detector. The Ulysses spacecraft and mission are explained in more detail by Wenzel et al.
162 (1992). Details about the data transmission to Earth can also be found in Paper III.

163 The Ulysses dust detector (GRU) has a 140° wide field of view and is mounted at the
164 spacecraft nearly at right angles (85°) to the antenna axis (spacecraft spin axis). Due to
165 this mounting geometry, the dust sensor is most sensitive to particles approaching from
166 the plane perpendicular to the spacecraft-Earth direction. The impact direction of dust
167 particles is measured by the rotation angle which is the sensor viewing direction at the
168 time of a dust impact. During one spin revolution of the spacecraft the rotation angle
169 scans through a complete circle of 360° . Zero degrees rotation angle is defined to be the
170 direction closest to ecliptic north. At high ecliptic latitudes, however, the sensor pointing at
171 0° rotation angle significantly deviates from the actual north direction. During the passages
172 over the Sun's polar regions the sensor always scans through a plane tilted by about 30°
173 from the ecliptic plane and all rotation angles lie close to the ecliptic plane (cf. Figure 4
174 in Grün et al., 1997). A sketch of the viewing geometry around aphelion passage can be
175 found in Grün et al. (1993).

176 The available electrical power on board Ulysses became an issue beginning in 2001 due
177 to decreasing power generation of the radioisotope batteries (RTGs). Some instrument
178 heaters on board had to be switched off to save power which in turn also reduced the on

179 board temperature. Later, in 2002, power consumption could not be sufficiently reduced
180 anymore by simply switching off heaters, and a cycling instrument operation scheme had
181 to be implemented: one or more of the scientific instruments had to be switched off at a
182 time. In the 2005 to 2007 interval considered in this paper the dust instrument was switched
183 off during a total time period of about seven months, separated into three individual time
184 periods (see Sect. 2.2). After 30 November 2007 the dust instrument remained switched
185 off permanently. It was planned to switch the dust instrument on in January 2008 again,
186 but due to a failure on board the spacecraft, this did not happen. Hence, no dust data were
187 obtained after 30 November 2007 although operation of the Ulysses spacecraft continued
188 until 30 June 2009.

189 **2.2 Dust instrument operation**

190 Table 2 gives significant mission and dust instrument events from 2005 to 2007. Earlier
191 events are only listed if especially significant. A comprehensive list of events from launch
192 until the end of 2004 was given in Papers III, V, VII and IX.

193 During the earlier Ulysses mission phases until 2000, several spacecraft anomalies oc-
194 curred during which all scientific instruments on board were switched off automatically
195 (Disconnection of all Non-Essential Loads – DNELs for short). No such anomaly occurred
196 in the 2005 to 2007 interval. The dust instrument was switched off between 28 September
197 2006 and 9 March 2007, from 2 April 2007 to 30 April 2007 and after 30 November 2007
198 (*cf.* Table 1), respectively, because there was not enough electrical power generated by the
199 RTGs anymore to operate all instruments simultaneously.

200 The dust instrument has two heaters to allow for a relatively stable operating temperature
201 within the sensor. By heating one of the two or both heaters, three different heating power
202 levels can be achieved (0.4 W, 0.8 W or 1.2 W; for comparison, the total power consump-
203 tion of the instrument without heaters is 2.2 W). Sensor heating was necessary when the
204 spacecraft was outside about 2 AU because relatively little radiation was received from the
205 Sun.

206 Before 2001 one or both heaters were switched on beyond 2 AU, except close to the Sun
207 when both were switched off. Since November 2001 the maximum allowed heating power
208 for nominal dust instrument operation has been limited to 0.8 W to save power on board
209 Ulysses. This reduced the sensor temperature by about 10° C as compared to the config-
210 uration with 1.2 W heating power at similar heliocentric distance. The full heating power
211 was only allowed for short periods before instrument switch on to avoid damage of the
212 electronics. In 2005 and 2006 the instrument was operated with 0.8 W heating power, and
213 in 2007 the heating power was set to 0.4 W, except during switch-ons when it was raised
214 to 1.2 W.

215 Table 2 lists the total powers consumed by the heaters. From 2005 to 2007 the temperature
216 of the dust sensor was between -37°C and $+16^{\circ}\text{C}$. The lower limit for the specified oper-
217 ational range is -30°C . No major effect of the reduced temperature was recognised (for an

218 anomalous temperature-related flipping in the housekeeping value of the channeltron high
219 voltage in 2004 see Paper IX).

220 No reprogramming of the Ulysses dust instrument occurred in the 2005 to 2007 interval.
221 In the earlier mission the instrument was reprogrammed three times and the reader is re-
222 ferred to Papers V and IX for details. In particular, a new classification scheme for impact
223 events was implemented in April 2002 which is the same as the one installed in the Galileo
224 instrument in July 1994 (Paper IV).

225 **2.3 Instrument sensitivity and noise**

226 Analysis of the in-orbit noise characteristics of the dust instrument (Paper III) led to a
227 relatively noise-free configuration with which the instrument was normally operated until
228 August 2000: channeltron voltage 1140 V (HV = 3); event definition status such that
229 either the channeltron or the ion-collector channel could, independent of each other, start
230 a measurement cycle (EVD = C, I); detection thresholds for ion-collector, channeltron and
231 electron-channel set to the lowest levels and the detection threshold for the entrance grid
232 set to the first digital step (SSEN = 0, 0, 0, 1). See Paper I for a description of these terms.

233 During the entire Ulysses mission dedicated noise tests were performed at monthly in-
234 tervals in order to monitor instrument health and noise characteristics. During all these
235 tests the operational settings were changed in four steps at one-hour intervals, starting
236 from the nominal configuration described above: a) set the event definition status such
237 that the channeltron, the ion collector and the electron-channel can initiate a measure-
238 ment cycle (EVD = C, I, E); b) set the thresholds for all channels to their lowest lev-
239 els (SSEN = 0, 0, 0, 0); c) reset the event definition status to its nominal configuration
240 (EVD = C, I) and increase the channeltron high voltage by one step with respect to the
241 nominal configuration; d) reset the instrument to its nominal configuration (i.e. reduce the
242 channeltron high voltage by one step and set the detection thresholds to SSEN = 0, 0, 0, 1).

243 The noise tests performed in the earlier mission before 2000 revealed a long-term drop
244 in the noise sensitivity of the instrument which was most likely caused by a reduction in
245 the channeltron amplification due to electronics degradation (Paper IX and Krüger et al.,
246 2005b). To counterbalance the reduced amplification we increased the channeltron high
247 voltage by one digital step in August 2000 (HV = 4, 1250 V) which raised the instrument
248 sensitivity close to its original value from the earlier mission again (Paper IX). Although
249 the aging of the channeltron led to a drop in the number of class 3 impacts, the dust impacts
250 which have caused these events should have shown up in lower quality classes. The noise
251 response of the Ulysses dust detector was monitored since then and no significant drop
252 in the sensitivity was recognized. In the time interval considered here HV=4, EVD=C,I,
253 SSEN=0001 was the nominal configuration of the dust instrument (during noise tests the
254 channeltron voltage was always raised by one digital step, i.e. to HV = 5, 1370 V).

255 Figure 3 shows the noise rate of the dust instrument for the 2005 to 2007 period. The upper
256 panel shows the daily maxima of the noise rate. In the earlier mission before November
257 2001 the daily maxima were dominated by noise due to interference with the sounder of

258 the Unified RADio and Plasma wave instrument (URAP) on board Ulysses (Stone et al.,
259 1992). Since November 2001 the sounder was switched off permanently for power saving
260 and no sounder noise occurred anymore (see Papers III, V, VII and IX for details of the
261 sounder operation and sounder-related noise). Individual sharp spikes in the upper panel
262 of Figure 3 are caused by noise tests which occurred at approximately monthly intervals,
263 They are best seen in 2007 when Ulysses was close to the Sun which is known to be a
264 strong source of noise.

265 The bottom panel of Figure 3 shows the daily averages in the noise rate. The average
266 was about 10 events per day at random times and it shows that dead time is negligible.
267 These noise rates are very similar to those measured since August 2000 (Paper IX) when
268 the instrument was operated with the same channeltron voltage as in the 2005-07 interval,
269 implying that no significant channeltron degradation has occurred since 2000.

270 **3 Impact events**

271 The dust instrument classifies all impact events into four classes and six ion charge ampli-
272 tude ranges which leads to 24 individual categories, with one accumulator belonging to one
273 individual category. Class 3, our highest class, are real dust impacts and class 0 are mostly
274 noise events. Depending upon the noise of the charge measurements, classes 1 and 2 can be
275 true dust impacts or noise events. Two classification schemes were used in the Ulysses dust
276 instrument: since 26 March 2002, when the dust instrument was reprogrammed, the clas-
277 sification scheme has been the same as the one used in the Galileo instrument (described
278 in Paper IV). A different scheme was used before which is described in Paper I.

279 Between 1 January 2005 and 30 November 2007 the complete data sets (sensor orientation,
280 charge amplitudes, charge rise times, etc.) of 6970 events including 609 dust impacts were
281 transmitted to Earth. Table 3 lists the number of all dust impacts counted with the 24
282 accumulators of the instrument. ‘AC xy ’ refers to class number ‘ x ’ and amplitude range
283 ‘ y ’ (for a detailed description of the accumulator categories see Paper I). As discussed in
284 the previous section, most noise events were recorded during the time periods when the
285 dust instrument was configured to its high sensitive state for noise tests. During these
286 periods many events were only counted by one of the 24 accumulators because their full
287 information was overwritten before the data could be transmitted to Earth. Since the dust
288 impact rate was low during times outside these periods, it is expected that only the data
289 sets of very few true dust impacts were lost.

290 Table 4 lists all 609 particles detected between January 2005 and November 2007 for which
291 the complete information exists. Note that approximately 45 jovian stream particles were
292 detected in six dust streams in 2005 (AR1, Krüger et al., 2006c) which are about 10 nm
293 in size and their velocities exceed 200km s^{-1} (Zook et al., 1996). Their mass and speed
294 calibration is unreliable because their masses and speeds are outside of the calibration
295 range of the dust instrument.

296 In Table 4 dust particles are identified by their sequence number and their impact time.

297 The event category – class (CLN) and amplitude range (AR) – are given. Raw data as
298 transmitted to Earth are displayed in the next columns: sector value (SEC) which is the
299 spacecraft spin orientation at the time of impact, impact charge numbers (IA, EA, CA) and
300 rise times (IT, ET), time difference and coincidence of electron and ion signals (EIT, EIC),
301 coincidence of ion and channeltron signal (IIC), charge reading at the entrance grid (PA)
302 and time (PET) between this signal and the impact. Then the instrument configuration
303 is given: event definition (EVD), charge sensing thresholds (ICP, ECP, CCP, PCP) and
304 channeltron high voltage step (HV). See Paper I for further explanation of the instrument
305 parameters.

306 The next four columns in Table 4 give information about Ulysses’ orbit: heliocentric dis-
307 tance (R), ecliptic longitude and latitude (LON, LAT) and distance from Jupiter (D_{Jup} , in
308 astronomical units). The next column gives the rotation angle (ROT) as described in Sec-
309 tion 2. Then follows the pointing direction of the dust instrument at the time of particle
310 impact in ecliptic longitude and latitude (S_{LON} , S_{LAT}). Mean impact velocity (v , in km s^{-1})
311 and velocity error factor (VEF, i.e. multiply or divide stated velocity by VEF to obtain
312 upper or lower limits) as well as mean particle mass (m , in grams) and mass error factor
313 (MEF) are given in the last columns. For $\text{VEF} > 6$, both velocity and mass values should
314 be discarded. This occurs for 72 impacts. No intrinsic dust charge values are given (see
315 Svestka et al., 1996, for a detailed analysis). Recently, reliable charge measurements for
316 interplanetary dust grains were reported for the Cassini dust detector (Kempf et al., 2004).
317 These measurements may lead to an improved understanding of the charge measurements
318 of Ulysses and Galileo in the future.

319 4 Analysis

320 The most important impact parameter determined by the dust instrument is the positive
321 charge measured on the ion collector, Q_{I} , because it is relatively insensitive to noise. Fig-
322 ure 4 shows the distribution of Q_{I} for all dust particles detected from 2005 to 2007. Ion
323 impact charges have been detected over the entire range of six orders of magnitude in im-
324 pact charge that the dust instrument can measure. One impact is close to the saturation
325 limit of $\sim 10^{-8} \text{ C}$ and may thus constitute a lower limit of the actual impact charge. The
326 impact charge distribution of the big particles ($Q_{\text{I}} > 10^{-13} \text{ C}$) follows a power law distri-
327 bution with index -0.45 and is shown as a dashed line. This value is very similar to earlier
328 Ulysses measurements of 2000 to 2004 (-0.40 ; cf. Paper IX).

329 In the earlier 1993 to 2004 data set (Papers V, VII and IX) the impact charge distribution
330 was reminiscent of three individual populations: small particles with impact charges $Q_{\text{I}} <$
331 10^{-13} C (AR1), intermediate size particles with $10^{-13} \text{ C} \leq Q_{\text{I}} \leq 10^{-11} \text{ C}$ (AR2 and AR3)
332 and big particles with $Q_{\text{I}} > 10^{-11} \text{ C}$ (AR4 to AR6). This is also visible in the present
333 data set, although less pronounced. The intermediate particles are mostly of interstellar
334 origin and the big particles are attributed to interplanetary grains (Grün et al., 1997, see
335 also Section 5). The small particle impacts (AR1) detected over the polar regions of the

336 Sun are candidates for being interplanetary β -meteoroids (Hamilton et al., 1996; Wehry
337 and Mann, 1999; Wehry et al., 2004).

338 It should be noted that the charge distribution shown in Figure 4 is very similar to the
339 one measured with Galileo in interplanetary space between 1993 and 1995 (i.e. between
340 1 and 5 AU; Paper IV). In particular, the power law index of -0.43 for the big particles
341 was practically identical. This indicates that both dust instruments basically detected the
342 same dust populations in interplanetary space and that their responses to dust impacts are
343 very similar. The only significant difference is a dip in the Ulysses charge distribution at
344 $2 \cdot 10^{-10} \text{ C}$ (Figure 4) which is also evident in all earlier Ulysses data (Papers III, V, VII
345 and IX). In particular, it is visible in both the ecliptic and out-of-ecliptic phases of Ulysses
346 which indicates that it is not sensitive to the impact speed or the dust population of the
347 grains. A small but much weaker dip is also seen in some but not all Galileo data sets. We
348 therefore conclude that the dip is most likely due to an artefact in the Ulysses instrument
349 electronics in AR5.

350 The ratio of the channeltron charge Q_C and the ion collector charge Q_I is a measure of
351 the channeltron amplification A , which in turn is an important parameter for dust impact
352 identification (Paper I). In Figure 5 we show the charge ratio Q_C/Q_I as a function of Q_I for
353 the 2005 to 2007 dust impacts with the channeltron high voltage set to 1250 V ($HV = 4$).
354 This diagram is directly comparable with similar diagrams in the previous Papers III, V
355 and VII for $HV = 3$ and Paper IX for $HV = 4$, respectively.

356 The mean amplification determined from particles with $10^{-12} \text{ C} \leq Q_I \leq 10^{-11} \text{ C}$ and $HV = 4$
357 in the 2005 to 2007 interval is $A \simeq 1.57$. This value is somewhat lower than the values
358 derived from the first five years of the mission (1990 to 1995; Papers III and V) and com-
359 parable to the more recent determinations (1996 to 2004; Papers VII and IX). It indicates
360 that a sufficiently high channeltron amplification and stable instrument operation could be
361 maintained with this higher voltage ($HV = 4$) since 2000. It shows in particular that the
362 channeltron degradation could be mostly counterbalanced by increasing the high voltage
363 by one digital step. Much more severe electronics degradation was found for the Galileo
364 dust detector during Galileo's orbital tour in the jovian system. It was likely related to
365 the harsh radiation environment in the magnetosphere of the giant planet (Krüger et al.,
366 2005b).

367 In Figure 6 we show the masses and velocities of all dust particles detected between 2005
368 and 2007. As in the earlier periods before 2005, velocities occur over the entire calibrated
369 range from 2 to 70 km s^{-1} . The masses vary over almost nine orders of magnitude from
370 $\sim 10^{-7} \text{ g}$ to 10^{-16} g . The mean errors are a factor of 2 for the velocity and a factor of
371 10 for the mass. The clustering of the velocity values is due to discrete steps in the rise
372 time measurement but this quantisation is much smaller than the velocity uncertainty. For
373 many particles in the lowest two amplitude ranges (AR1 and AR2) the velocity had to be
374 computed from the ion charge signal alone which leads to the vertical striping in the lower
375 mass range in Figure 6 (most prominent above 10 km s^{-1}). In the higher amplitude ranges
376 the velocity could normally be calculated from both the target and the ion charge signal,
377 resulting in a more continuous distribution in the mass-velocity plane. Impact velocities

378 below about 3 km s^{-1} should be treated with caution because anomalous impacts onto the
379 sensor grids or structures other than the target generally lead to prolonged rise times and
380 hence to unnaturally low impact velocities.

381 5 Discussion

382 In Figure 7 we show the dust impact rate detected in various amplitude ranges together
383 with the total impact rate summed over all amplitude ranges. The highest overall impact
384 rate was recorded in 2007 when Ulysses was in the inner solar system and relatively close
385 to the ecliptic plane. It coincides with a peak in the rate of bigger particles in the three
386 highest ion amplitude ranges (AR4 – AR6). These impacts are attributed to interplanetary
387 particles on low inclination orbits (Grün et al., 1997). The majority of them are the impacts
388 with $Q_I > 10^{-10} \text{ C}$ shown in Figure 4. The impact rate of intermediate sized particles in
389 AR2 and AR3 showed relatively little variation. This size range is dominated by interstellar
390 impactors. The details of the various dust populations are discussed further below.

391 Figure 8 shows the sensor orientation at the time of a particle impact (rotation angle). The
392 big particles (diamonds, impact charge $Q_I \geq 8 \cdot 10^{-14} \text{ C}$ which roughly corresponds to
393 AR2-6) are concentrated towards the upstream direction of interstellar helium (cf. Fig-
394 ure 10, bottom panel; Witte et al., 1996; Witte, 2004; Witte et al., 2004). They have been
395 detected with a relatively constant rate during the entire three-year period (Figure 7). The
396 particles with the highest ion amplitude ranges (AR4 to AR6) are not distinguished in this
397 diagram because they cannot be separated from interstellar particles by directional argu-
398 ments alone. They have to be distinguished by other means (e. g. mass and speed). In
399 addition, their total number is so small that they constitute only a small ”contamination“ of
400 the interstellar particles in Figure 8. In the ecliptic plane at 1.3 AU, however, interplanetary
401 particle flux dominates over interstellar flux by a factor of about 3 (in number).

402 5.1 Interstellar dust

403 Interstellar particles move on hyperbolic trajectories through the solar system and approach
404 Ulysses from the same direction as the interstellar gas (Grün et al., 1994; Baguhl et al.,
405 1995a; Witte et al., 1996; Witte, 2004; Witte et al., 2004). They can therefore be identified
406 by their impact direction and their impact speed. Earlier investigations showed that in the
407 Ulysses and Galileo dust data sets the interstellar impactors are mostly found in amplitude
408 ranges AR2 and AR3.

409 In the earlier mission from 1993 to 2004 the impact rate of interstellar grains (as derived
410 from the AR2 and AR3 accumulators) varied by about a factor of 2.5: between 1993 and
411 1995 the rate was $\sim 2 \cdot 10^{-6} \text{ s}^{-1}$ (Paper V) while in the 1996 to 1999 interval it dropped
412 to $\sim 8 \cdot 10^{-7} \text{ s}^{-1}$ (Paper VII) and from 2000 to 2004 the average impact rate was again
413 $\sim 2 \cdot 10^{-6} \text{ s}^{-1}$ (Paper IX). In 2005/2006 the average impact rate of interstellar impactors
414 was somewhat higher: $\sim 3 \cdot 10^{-6} \text{ s}^{-1}$ (bottom panel of Figure 7). In 2007, when Ulysses

415 was in the inner solar system again, the impact rate derived from AR2 and AR3 was still
416 higher. Here, however, the impact rate is not due to interstellar particles alone because of a
417 strong contribution from interplanetary impactors.

418 The dashed curve in Figure 7 (bottom panel) shows the expected impact rate of interstel-
419 lar particles assuming that they approach from the direction of interstellar helium (Witte
420 et al., 1996; Witte, 2004; Witte et al., 2004) and that they move through the solar system
421 on straight trajectories with a relative velocity of 26 km s^{-1} . This assumption means dy-
422 namically that radiation pressure cancels gravity for these particles ($\beta = 1$) and that their
423 Larmor radii are large compared with the dimension of the solar system. Both assumptions
424 are reasonable for particles with masses between 10^{-13} and 10^{-12} g which is the dominant
425 size range measured for interstellar grains (Grün et al., 1997). The variation predicted by
426 the model is caused by changes in the instrument’s viewing direction with respect to the
427 approach direction of the particles and changes in the relative velocity between the space-
428 craft and the particles. The dust particle flux is independent of heliocentric distance in this
429 simple model, which gives relatively good agreement with the observed impact rate.

430 Ulysses has monitored the interstellar dust flow through the solar system for more than 15
431 years. This time period covers more than two and a half revolutions of the spacecraft about
432 the Sun through more than 2/3 of a complete 22-year solar cycle. Thus, Ulysses measured
433 interstellar dust during solar minimum and solar maximum conditions of the interplanetary
434 magnetic field (IMF). The interstellar dust flux modulation due to grain interaction with
435 the magnetic field during solar minimum could be well explained (Landgraf, 1998, 2000;
436 Landgraf et al., 2003). By taking into account the sensor side wall in the instrument field
437 of view we could recently improve the flux determination (Altobelli et al., 2004).

438 Ulysses provides the unique opportunity to compare repeated dust measurements at the
439 same locations in the solar system for different phases of the solar cycle and the inter-
440 planetary magnetic field. In Figure 9 we show the approach direction of interstellar grains
441 for three selected periods when Ulysses was between approximately -8 and -56° ecliptic
442 latitude during approach to the inner solar system. During these three intervals the dust de-
443 tection geometry as indicated by the contour lines was very similar so that we can compare
444 them directly. At least three differences are obvious between the three panels:

- 445 • The approach direction of the majority of grains was compatible with the flow di-
446 rection of the interstellar helium gas as measured with Ulysses (Witte, 2004; Witte
447 et al., 2004) during all three time intervals.
- 448 • The dust flux varied by a factor of about two during the three intervals, the lowest
449 flux being measured in the period 1999 to 2000, while the highest flux occurred in
450 2005/2006 (Krüger et al., 2007).
- 451 • A subset of the grains detected in 2005 shows a clear shift in the detected impact
452 direction away from the ecliptic plane towards southern ecliptic latitudes.

453 Our preliminary analysis indicates that this shift is about 30° away from the ecliptic plane
454 towards southern ecliptic latitudes (Krüger et al., 2007). The reason for this shift remains

455 mysterious. Whether it is connected to a secondary stream of interstellar neutral atoms
456 shifted from the main neutral gas flow (Collier et al., 2004; Wurz et al., 2004; Nakagawa
457 et al., 2006) is presently unclear. Given, however, that the neutral gas stream is shifted
458 along the ecliptic plane while the shift in the dust flow is offset from the ecliptic, a connec-
459 tion between both phenomena seems unlikely.

460 Even though Ulysses' position in the heliosphere and the dust detection conditions were
461 very similar during all three time intervals considered in Figure 9, the configurations of the
462 solar wind driven interplanetary magnetic field (IMF), which strongly affects the dynamics
463 of the smallest grains, were completely different. We have to consider that the interstellar
464 grains need approximately twenty years to travel from the heliospheric boundary to the
465 inner solar system where they are detected by Ulysses. Thus, the effect of the IMF on
466 the grain dynamics is the accumulated effect caused by the interaction with the IMF over
467 several years: In the earlier time intervals (1993/1994 and 1999/2000) the grains had a
468 recent dynamic history dominated by solar minimum conditions (Landgraf, 2000), while
469 the grains detected during the third interval (2005/2006) had a recent history dominated by
470 the much more disturbed solar maximum conditions of the IMF. During the solar maximum
471 conditions the overall magnetic dipole field changed polarity. Morfill and Grün (1979)
472 predicted that due to this effect in a 22-year cycle, small interstellar grains experience either
473 focussing or defocusing conditions. During these times they are systematically deflected
474 by the solar wind magnetic field either towards or away from the solar magnetic equator
475 plane (close to the ecliptic plane). This latter configuration likely has a strong influence on
476 the dust dynamics and the total interstellar flux in the inner heliosphere but it has not been
477 modelled in detail. An explanation of the grain interaction with the IMF at the recent solar
478 maximum conditions is still pending.

479 Detailed modelling of the dynamics of the electrically charged dust grains in the helio-
480 sphere can give us information about the local interstellar environment of the solar system
481 where the particles originate from. The models developed by Landgraf et al. (2003) fit the
482 observed flux variation by assuming a constant dust concentration in the local interstellar
483 environment of our solar system. It implies that the local interstellar dust phase must be
484 homogeneously distributed over length scales of 50 AU, which is the distance travelled by
485 the Sun during the measurement period of Ulysses from 1992 to 2002. This conclusion
486 is supported by the more recent Ulysses data until the end of 2004 (Krüger et al., 2006a).
487 Our latest Ulysses dust data of 2005/2006, on the other hand, put a question mark onto
488 this conclusion because if the observed shift in impact direction turns out to be intrinsic, it
489 would imply that this homogeneity breaks down on larger length scales.

490 **5.2 Interplanetary dust**

491 Figure 10 shows the data (AR3 to AR6) from launch in 1990 until the end of the Ulysses
492 mission in 2007 and compares them with two meteoroid environment models by Divine
493 (1993) and Staubach et al. (1997). For most of the mission time, excluding the ecliptic
494 plane crossings in 1995, 2001 and 2007 the Divine model predicts impacts from a very

495 broad range of directions, with spin angles from 45° to 300° . The impactors belong mostly
496 to the so-called "halo" population which was introduced in the model to explain the Pioneer
497 data (Divine, 1993). The Ulysses directionality of the impacts had not been available at the
498 time of construction of the model, and they are not well fit by the "halo" population. In
499 contrast, the Staubach et al. (1997) model was fit to 5 years of Ulysses data from the craft's
500 first orbit about the Sun, taking the crucial directional information into account. It is in
501 better agreement with the data and has been confirmed by Ulysses' second and third orbits.
502 It places most of the impacts into the spin angle range from 30° to 120° . These are due to
503 interstellar dust flowing through the solar system (Grün et al., 1994).

504 One more observation from these plots is that both the Divine and Staubach models predict
505 higher flux during the ecliptic plane crossings than the data permit. The time dependence
506 of the interstellar dust flux was asserted after the meteoroid models under review had been
507 constructed (Landgraf, 2000). However, the disagreement at the ecliptic plane crossings
508 was not anticipated, since almost all data incorporated in the models were taken from the
509 ecliptic plane. Possible explanations of the discrepancy are the roughness of model fits as
510 well as different representations of the data taken for model adjustments and displayed in
511 Figure 10. While Figure 10 selects all impacts above a fixed threshold of charge released,
512 the models were fitted using more uncertain inferred mass thresholds. The inference of
513 mass is based on speed determination that is uncertain by a factor of 2.

514 **5.3 β -meteoroids**

515 When Ulysses was in the inner solar system in 1995 and in 2001, maxima were evident in
516 the impact rate of the smallest particles (AR1; Papers V and IX). A similar maximum in the
517 inner solar system occurred again in 2007 (Fig. 7; top panel). Although in all three cases the
518 maximum was reached during a short period around ecliptic plane crossing, many particles
519 were also detected at high ecliptic latitudes. These grains are attributed to a population of
520 submicron-sized interplanetary particles whose dynamics is dominated by solar radiation
521 pressure. They move on escape trajectories from the solar system (β -meteoroids, Baguhl
522 et al., 1995b; Hamilton et al., 1996). β -meteoroids were identified with Ulysses over the
523 Sun's poles in 1994/95 (Wehry and Mann, 1999) and 2000/01 (Wehry et al., 2004). Due
524 to the detection geometry, however, they were undetectable outside these periods, explain-
525 ing the lower impact rates in AR1. β -meteoroids were detectable again from mid-2006
526 until early-2008 (Wehry et al., 2004). In independent potential detection of β -meteoroids
527 was recently reported from the plasma wave instrument on board the STEREO spacecraft
528 (Meyer-Vernet et al., 2009).

529 Wehry and Mann (1999) and Wehry et al. (2004) identified a significant asymmetry in the
530 flux of these particles between the northern and the southern hemisphere from the first two
531 heliocentric orbits of Ulysses, the reason of which is presently unknown. A comprehensive
532 analysis of these three data sets will be the subject of a future investigation and may reveal
533 whether the asymmetry is connected with the solar cycle variation of the interplanetary
534 magnetic field.

535 6 Conclusions

536 In this paper, the eleventh and final in a series of Ulysses and Galileo dust data papers, we
537 present data from the Ulysses dust instrument for the period January 2005 to November
538 2007. In this time interval, starting from a heliocentric distance of 5.3 AU close to aphelion,
539 the spacecraft approached the Sun, flew over the Sun's south pole, crossed the ecliptic plane
540 at 1.4 AU heliocentric distance and reached a northern ecliptic latitude of 60° .

541 A total number of 609 dust impacts were recorded during this period. Together with 6110
542 impacts recorded in interplanetary space and near Jupiter between Ulysses' launch in Oc-
543 tober 1990 and December 2004 (Grün et al., 1995a; Krüger et al., 1999b, 2001b, 2006a),
544 the complete dust data set measured during the entire Ulysses mission consists of 6719
545 impacts. Given its temporal coverage and Ulysses' unique orbital orientation, the Ulysses
546 dust data set will be a treasure for decades to come.

547 The total recorded dust impact rate dropped from an initial value of 0.7 impacts per day in
548 2005 when Ulysses was at low ecliptic latitudes to a value of 0.3 impacts at higher latitudes.
549 Most of these particles were of interstellar origin, in particular at higher ecliptic latitudes,
550 a minor fraction being interplanetary particles. A maximum of 1.5 per day was measured
551 in 2007 in the inner solar system; here the majority of the grains were of interplanetary
552 origin.

553 The measurements from the entire Ulysses mission since launch in 1990 are in disagree-
554 ment with the interplanetary dust flux model by Divine (1993). Instead, they are well
555 matched by the model of Staubach et al. (1997) which was originally developed with a
556 shorter data set from the first Ulysses orbit and with the dust measurements from Galileo's
557 interplanetary cruise.

558 Noise tests performed regularly during the three years period revealed no degradation in
559 the noise sensitivity of the dust instrument, and the nominal instrument operational con-
560 figuration remained unchanged during the entire period. In particular, no change in the
561 channeltron high voltage setting was required.

562 With Ulysses we had the unprecedented opportunity to measure dust in interplanetary space
563 for approximately 17 years. In particular, due to the unique orientation of Ulysses' orbital
564 plane approximately perpendicular to the flow direction of the interstellar dust through the
565 solar system and the spacecraft's 6-year revolution period about the Sun, we obtained dust
566 measurements from three traverses of the same spatial region at southern ecliptic latitudes.
567 These passes were separated by six years in time and were obtained at different phases of
568 the solar cycle. Variations in the interstellar dust flux by a factor of three are evident during
569 the entire mission. Flux variations until 2002 have been explained by the interaction of the
570 dust grains with the time-varying interplanetary magnetic field, while detailed modelling
571 of the later data is still pending. The data obtained in 2005/2006 reveal an approximately
572 30° shift in the approach direction of the grains away from the approach direction of the
573 interstellar helium gas. The reason for this shift remains mysterious and will be the subject
574 of a future investigation.

575 Even though this is the final paper in our series of Ulysses dust data papers published

576 during the last 15 years, the evaluation of this unique data set is continuing. A list of
577 specific open questions raised in this and earlier data papers includes:

- 578 • *β -meteoroids*: The detailed evaluation of the measurements from Ulysses' third so-
579 lar orbit is still pending. Open questions include the measured north-south asym-
580 metry in the measured flux and the heliocentric distance range where these grains
581 are generated. More measurements from other spacecraft (e.g. STEREO) would be
582 advantageous to answer these questions.

- 583 • *Comparison with Pioneer 10 and 11 measurements*: The flux of particles with $m \geq$
584 10^{-9} g as measured with Ulysses is about a factor of five lower than expected from
585 the Pioneer 10/11 measurements. Potential reasons for this discrepancy are errors
586 on the mass calibrations of the Ulysses and/or Pioneer detectors in this mass range,
587 many of the Pioneer detections may not be due to actual meteoroid impacts, or the
588 dust grains by Pioneer belong to a population of dust which could not or only parti-
589 tially be detected with Ulysses (Krüger et al., 1999b). Measurements with New
590 Horizons in the outer solar system may shed new light onto this question.

- 591 • *Interstellar dust*: Earlier comprehensive investigations of the interstellar impactors
592 were mostly performed in the late 1990s and relied upon the significantly smaller
593 data set available at the time. In the meantime, until the end of the Ulysses mission,
594 the interstellar dust data set has grown by at least a factor of two so that a complete
595 re-analysis of the entire data set is worthwhile and can give new insights into, e.g., the
596 grain dynamics inside the heliosphere and into the conditions in the local interstellar
597 environment where these grains originate from. In particular, the reason for the
598 observed 30° shift remains an open question.

- 599 • *Interplanetary dust*: The interplanetary dust model by Divine (1993) was developed
600 before the Ulysses data became available and the model by Staubach et al. (1997)
601 used only data from within the ecliptic plane and from a fraction of Ulysses' first
602 heliocentric orbit. A new model is presently being developed by Dikarev et al. (2005)
603 which will incorporate infrared observations of the zodiacal cloud by the COBE
604 DIRBE instrument, in-situ flux measurements by the dust detectors on board Galileo
605 and Ulysses, and the crater size distributions on lunar rock samples retrieved by the
606 Apollo missions.

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Table 1: Summary of Ulysses data papers, significant mission events and dust detector switch-off times. Jupiter distance $R_J = 71492$ km.

Time Interval	Significant Mission Events	Dust Detector off	Paper Number
1990 – 1992	Ulysses launch (6 Oct 1990), Jupiter flyby (8 Feb 1992, distance $6.3 R_J$)	Before 27 Oct 1990, 14 Jun 1991 - 18 Jun 1991	III (Grün et al., 1995a)
1993 – 1995	Maximum southern latitude -79° (3 Oct 1994), Perihelion (12 Mar 1995), Maximum northern latitude 79° (19 Aug 1995)	8 Aug 1993 - 11 Aug 1993, 27 Nov 1993 - 28 Nov 1993, 10 Oct 1994 - 11 Oct 1994, 10 Dec 1995	V (Krüger et al., 1999b)
1996 – 1999	Aphelion (20 Apr 1998),	17 Aug 1996 - 18 Aug 1996, 1 Apr 1997 - 2 Apr 1997, 15 Feb 1999 - 16 Feb 1999	VII (Krüger et al., 2001b)
2000 – 2004	Maximum southern latitude -80° (27 Nov 2000), Perihelion (23 May 2001), Maximum northern latitude 80° (13 Oct 2001), Jupiter flyby (4 Feb 2004, distance $0.8 AU$), Aphelion (30 Jun 2004)	26 Mar 2002 - 8 Apr 2002, 1 Dec 2002 - 3 Jun 2003, 28 Jun 2003 - 22 Aug 2003, 30 Nov 2003 - 2 Dec 2003	IX (Krüger et al., 2006a)
2005 – 2007	Maximum southern latitude -80° (7 Feb 2007), Perihelion (18 Aug 2007)	28 Sep 2006 - 9 Mar 2007, 2 Apr 2007 - 30 Apr 2007, After 30 Nov 2007	XI (this paper)

Table 2: Ulysses mission and dust detector (GRU) configuration, tests and other events. Only selected events are given before 2005. See Section 2 for details.

Yr-day	Date	Time	Event
90-279	06.10.90		Ulysses launch
03-154	03.06.03	03:38	GRU 0.8 W heater on
04-182	30.06.04		Ulysses aphelion passage (5.4 AU)
04-337	02.12.04	11:00	GRU nominal configuration: HV=4, EVD=C,I, SSEN=0001
05-006	06.01.05	07:00	GRU noise test
05-034	03.02.05	04:20	GRU noise test
05-063	04.03.05	03:00	GRU noise test
05-090	31.03.05	01:00	GRU noise test
05-120	30.04.05	00:00	GRU noise test
05-146	26.05.05	04:00	GRU noise test
05-174	23.06.05	20:00	GRU noise test
05-201	20.07.05	17:27	GRU noise test
05-229	17.08.05	11:49	GRU noise test
05-259	16.09.05	19:00	GRU noise test
05-285	12.10.05	09:00	GRU noise test
05-314	10.11.05	13:14	GRU noise test
05-342	08.12.05	11:00	GRU noise test
06-005	05.01.06	08:30	GRU noise test
06-033	02.02.06	09:16	GRU noise test
06-061	02.03.06	04:00	GRU noise test
06-089	30.03.06	02:00	GRU noise test
06-118	28.04.06	00:30	GRU noise test
06-148	28.05.06	06:28	GRU noise test
06-173	22.06.06	07:33	GRU noise test
06-202	21.07.06	19:00	GRU noise test
06-232	20.08.06	17:00	GRU noise test
06-256	13.09.06	20:11	GRU noise test
06-271	28.09.06	16:28	GRU off (both heaters off)
07-038	07.02.07		Ulysses maximum south solar latitude (-79.7°)
07-067	08.03.07	18:08	GRU 0.8 W heater on
07-067	08.03.07	18:28	GRU 0.4 W heater on
07-068	09.03.07	16:44	GRU 0.8 W heater off
07-068	09.03.07	17:31	GRU on
07-069	10.03.07	07:54	GRU nominal configuration
07-079	20.03.07	08:40	GRU noise test
07-092	02.04.07	18:11	GRU off (both heaters off)
07-120	30.04.07	01:23	GRU 0.8 W heater on
07-120	30.04.07	01:33	GRU 0.4 W heater on
07-120	30.04.07	21:37	GRU 0.8 W heater off
07-120	30.04.07	21:38	GRU on
07-121	01.05.07	04:07	GRU nominal configuration
07-130	10.05.07	20:00	GRU noise test
07-158	07.06.07	01:00	GRU noise test
07-187	06.07.07	16:15	GRU noise test
07-214	02.08.07	01:53	GRU noise test
07-231	18.08.07		Ulysses perihelion passage (1.39 AU)
07-232	19.08.07		Ulysses ecliptic plane crossing
07-242	30.08.07	00:17	GRU noise test
07-270	27.09.07	17:00	GRU noise test
07-298	25.10.07	13:00	GRU noise test
07-326	22.11.07	15:00	GRU noise test
07-334	30.11.07	16:20	GRU off (both heaters off)
08-014	14.01.08		Ulysses maximum north solar latitude (79.8°)
09-181	30.06.09		Ulysses end of mission

Abbreviations used: HV: channeltron high voltage step; EVD: event definition, ion- (I), channeltron- (C), or electron-channel (E); SSEN: detection thresholds, ICP, CCP, ECP and PCP. Times when no data could be collected with the dust instrument: 28 September 2006 to 9 March 2007, 2 April to 30 April 2007 and after 30 November 2007.

Table 3. Overview of dust impacts detected with the Ulysses dust detector between 1 January 2005 and 30 November 2007 as derived from the accumulators[†]. Switch-on of the instrument is indicated by horizontal lines. The heliocentric distance R , the lengths of the time interval Δt (days) from the previous table entry, and the corresponding numbers of impacts are given for the 24 accumulators. The accumulators are arranged with increasing signal amplitude ranges (AR), with four event classes for each amplitude range (CLN = 0,1,2,3); e.g. AC31 means counter for AR = 1 and CLN = 3. The Δt in the first line (05-006) is the time interval counted from the last entry in Table 2 in Paper IX. The totals of counted impacts[†], of impacts with complete data, and of all events (noise plus impact events) for the entire period are given as well.

Date	Time	R [AU]	Δt [d]	AC 01 [†]	AC 11 [†]	AC 21	AC 31	AC 02 [†]	AC 12	AC 22	AC 32	AC 03	AC 13	AC 23	AC 33	AC 04	AC 14	AC 24	AC 34	AC 05	AC 15	AC 25	AC 35	AC 06	AC 16	AC 26	AC 36
05-006	11:45	5.300	7.561	-	-	-	-	-	-	-	1	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
05-039	23:36	5.258	33.49	-	-	1	2	-	-	5	2	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
05-061	19:36	5.227	21.83	-	-	3	8	-	-	2	2	-	-	-	-	-	-	1	-	*	-	-	-	*	-	-	-
05-089	09:46	5.184	27.59	-	-	1	11	-	-	5	2	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
05-104	17:01	5.158	15.30	-	-	4	2	-	-	-	-	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
05-140	22:27	5.089	36.22	-	-	5	6	-	-	5	2	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
05-177	01:56	5.013	36.14	-	-	5	4	-	-	9	9	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
05-204	13:40	4.949	27.48	-	-	2	5	-	-	2	2	-	-	-	-	-	-	-	2	*	-	-	-	*	-	-	-
05-227	18:13	4.891	23.18	-	-	4	5	-	-	3	7	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
05-250	17:23	4.831	22.96	-	-	1	6	-	-	1	7	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
05-284	14:06	4.734	33.86	-	-	2	3	-	-	2	4	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
05-310	21:37	4.654	26.31	-	-	1	3	-	-	5	5	-	-	-	-	-	-	-	3	*	-	-	-	*	-	-	-
05-337	10:38	4.568	26.54	-	-	3	-	-	-	2	2	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
05-364	06:07	4.475	26.81	-	-	1	2	-	-	1	1	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
06-028	18:17	4.368	29.50	-	-	2	2	-	-	1	3	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
06-052	22:35	4.274	24.17	-	-	2	3	-	-	2	4	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
06-078	05:59	4.172	25.30	-	-	2	2	-	-	1	2	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
06-100	06:53	4.078	22.03	-	-	1	-	-	-	3	1	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
06-132	07:12	3.935	32.01	-	-	2	-	-	-	2	1	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
06-165	22:45	3.776	33.64	-	-	1	2	-	-	1	2	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
06-192	21:07	3.642	26.93	-	-	1	-	-	-	1	1	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
06-217	08:19	3.514	24.46	-	-	1	-	-	-	-	1	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
06-238	08:40	3.400	21.01	-	-	2	-	-	-	-	1	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
06-271	16:30	3.211	33.32	-	-	1	-	-	-	1	2	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
07-069	07:54	2.163	162.6	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
07-092	18:15	2.004	23.43	-	-	1	-	-	-	1	2	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
07-121	04:07	1.818	28.41	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
07-152	04:50	1.634	31.02	-	-	1	6	-	-	1	1	-	-	-	-	-	-	-	-	*	-	-	-	*	-	-	-
07-182	21:34	1.489	30.69	-	-	13	10	-	-	4	2	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
07-200	08:56	1.433	17.47	2	12	6	1	-	-	8	5	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
07-222	17:12	1.396	22.34	2	25	17	3	-	-	1	3	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
07-243	23:21	1.402	21.25	3	25	8	1	-	-	3	9	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
07-266	13:56	1.452	22.60	-	5	13	4	-	-	2	6	-	-	-	-	-	-	-	2	*	-	-	-	*	-	-	-
07-288	22:33	1.538	22.35	-	8	9	-	-	-	1	1	-	-	-	-	-	-	-	1	*	-	-	-	*	-	-	-
07-309	22:37	1.644	21.00	-	1	5	-	-	-	2	2	-	-	-	-	-	-	-	2	*	-	-	-	*	-	-	-

Table 3 continued.

Date	Time	R [AU]	Δt [d]	AC 01 [†]	AC 11 [†]	AC 21	AC 31	AC 02 [†]	AC 12	AC 22	AC 32	AC 03	AC 13	AC 23	AC 33	AC 04	AC 14	AC 24	AC 34	AC 05	AC 15	AC 25	AC 35	AC 06	AC 16	AC 26	AC 36	
07-334	16:30	1.790	24.74	-	3	2	-	-	-	-	-	-	-	1	-	-	-	-	-	-	*	-	-	*	-	-	-	-
				7 [†]	131 [†]	159	31	0 [‡]	12	88	76	0	0	17	61	0	0	4	19	*	0	2	2	*	0	1	0	0
				7	131	158	31	0	12	88	76	0	0	17	61	0	0	4	19	*	0	2	2	*	0	1	0	0
				6243	132	159	31	120	12	88	76	2	0	17	61	1	0	4	19	*	0	2	2	*	0	1	0	0

‡: Entries for AC01, AC11 and AC02 are the number of impacts with complete data. Due to the noise contamination of these three categories the number of impacts cannot be determined from the accumulators. The method to separate dust impacts from noise events in these three categories has been given by Baguhl et al. (1993).

*: No entries are given for AC05 and AC06 because they count the overflows of accumulators AC21 and AC31 since the reprogramming in December 2004 (Paper IX).

Table 4. Raw data: No., impact time, CLN, AR, SEC, IA, EA, CA, IT, ET, EIT, EIC, ICC, PA, PET, EVD, ICP, ECP, CCP, PCP, HV, and evaluated data: R, LON, LAT, D_{dup} , rotation angle (ROT), instr. pointing (S_{LOW} , S_{LAT}), impact speed (v , in km s^{-1}), speed error factor (VEF), mass (m , in grams) and mass error factor (MEF).

No.	IMP.	DATE	C	AR	S	IA	EA	CA	IT	ET	E	E	I	E	I	PA	P	E	E	I	E	C	P	C	P	C	P	HV	R	LON	LAT	D_{dup}	ROT	S_{LON}	S_{LAT}	V	VEF	M	MEF
6111	05-006	03:30	3	4	17	29	49	13	9	5	7	7	0	0	1	46	0	1	0	0	0	0	0	0	0	0	4	5.29763	159.2	-8.8	2.70789	280	76	11	7.8	1.9	5.4	10^{-10}	10.5
6112	05-006	11:45	3	2	156	10	20	9	9	7	6	7	0	0	1	36	0	1	0	0	0	0	0	0	0	0	4	5.29735	159.2	-8.8	2.70959	116	269	-23	14.1	1.9	8.9	10^{-13}	10.5
6113	05-010	12:16	3	2	166	10	20	1	8	6	7	0	1	0	0	37	0	1	0	0	0	0	0	0	0	0	4	5.29256	159.2	-9.0	2.73766	132	274	-39	21.4	1.9	1.9	10^{-13}	10.5
6114	05-010	15:01	2	1	87	2	6	0	9	10	11	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.29242	159.2	-9.0	2.73851	21	245	69	26.5	2.0	3.3	10^{-15}	12.5	
6115	05-011	21:22	2	1	105	5	2	0	8	8	7	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.29094	159.2	-9.1	2.74700	46	255	44	28.0	1.6	7.4	10^{-15}	6.0	
6116	05-014	23:24	2	1	193	6	10	0	7	8	8	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.28737	159.3	-9.2	2.76737	169	311	-70	35.4	1.6	4.8	10^{-15}	6.0	
6117	05-017	00:12	2	1	102	5	10	0	8	8	8	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.28481	159.3	-9.4	2.78178	43	254	47	28.0	1.6	8.8	10^{-15}	6.0	
6118	05-018	08:39	2	1	51	3	7	0	9	9	0	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.28329	159.3	-9.4	2.79025	331	87	61	21.0	1.6	1.1	10^{-14}	6.0	
6119	05-019	14:03	3	1	87	2	6	10	15	8	10	0	1	0	0	0	0	1	0	0	0	0	0	0	0	4	5.28176	159.3	-9.5	2.79871	21	244	68	36.7	2.0	1.1	10^{-15}	12.5	
6120	05-021	13:52	1	1	98	5	2	0	8	15	0	1	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.27929	159.3	-9.6	2.81224	40	251	50	21.4	1.9	6.4	10^{-15}	10.5	
6121	05-022	22:45	2	1	204	7	12	0	7	7	6	0	0	0	0	0	34	0	1	0	0	0	0	0	0	4	5.27758	159.4	-9.7	2.82153	189	15	-71	43.7	1.6	3.3	10^{-15}	6.0	
6122	05-023	08:32	2	2	186	8	13	0	7	8	6	0	0	0	0	0	34	0	1	0	0	0	0	0	0	4	5.27711	159.4	-9.7	2.82406	163	300	-66	35.4	1.6	1.1	10^{-14}	6.0	
6123	05-023	10:04	2	2	248	13	12	1	6	15	0	1	1	0	0	27	31	1	0	0	0	0	0	0	0	4	5.27695	159.4	-9.7	2.82491	250	68	-17	43.5	1.9	8.9	10^{-15}	10.5	
6124	05-024	05:20	2	1	241	6	11	0	8	8	7	0	0	0	0	0	35	0	1	0	0	0	0	0	0	4	5.27601	159.4	-9.7	2.82997	240	65	-27	28.0	1.6	1.2	10^{-14}	6.0	
6125	05-025	11:04	2	1	217	5	10	0	8	8	7	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.27444	159.4	-9.8	2.83841	210	49	-55	22.7	1.6	1.8	10^{-14}	6.0	
6126	05-030	11:00	3	2	179	12	21	2	8	5	6	0	1	0	0	1	39	0	1	0	0	0	0	0	0	4	5.26804	159.4	-10.0	2.87210	158	293	-62	21.4	1.9	3.1	10^{-13}	10.5	
6127	05-032	05:43	2	2	157	9	19	0	8	5	8	0	0	0	0	0	37	0	1	0	0	0	0	0	0	4	5.25915	159.5	-10.4	2.91743	141	277	-47	25.9	1.6	6.9	10^{-15}	6.0	
6128	05-032	09:14	2	1	187	3	8	0	9	8	12	0	0	0	0	42	0	1	0	0	0	0	0	0	0	4	5.25545	159.5	-10.5	2.95585	43	248	47	34.1	1.9	6.0	10^{-14}	10.5	
6129	05-034	00:12	2	2	76	8	13	0	7	8	6	0	0	0	0	6	21	1	0	0	0	0	0	0	0	4	5.25307	159.6	-10.6	2.94755	98	260	-6	26.5	2.0	3.3	10^{-15}	12.5	
6130	05-034	23:18	2	1	21	5	10	0	8	8	7	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.25290	159.6	-10.6	2.94839	216	49	-49	14.1	1.9	2.8	10^{-14}	10.5	
6131	05-035	00:03	3	1	115	7	12	0	8	8	6	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.25273	159.6	-10.6	2.94922	50	249	40	31.3	2.0	1.9	10^{-15}	12.5	
6132	05-037	05:39	2	1	161	4	8	13	8	9	7	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.25187	159.6	-10.7	2.95340	78	256	13	28.0	1.6	1.0	10^{-14}	6.0	
6133	05-039	23:36	2	2	92	12	21	0	7	4	6	0	0	0	0	42	0	1	0	0	0	0	0	0	0	4	5.25136	159.6	-10.7	2.95591	182	349	-71	7.8	1.9	9.3	10^{-14}	10.5	
6134	05-041	16:13	2	1	125	2	6	0	9	10	10	0	0	0	0	6	21	1	0	0	0	0	0	0	0	4	5.25119	159.6	-10.7	2.95674	255	65	-13	22.7	1.6	7.8	10^{-15}	6.0	
6135	05-041	18:10	1	1	209	4	3	0	9	15	0	1	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.25050	159.6	-10.7	2.96008	56	251	34	22.7	1.6	9.4	10^{-15}	6.0	
6136	05-041	21:24	3	1	91	2	6	3	9	9	9	0	1	0	0	0	0	1	0	0	0	0	0	0	0	4	5.24878	159.6	-10.8	2.96842	196	26	-65	28.0	1.6	6.2	10^{-15}	6.0	
6137	05-042	12:20	2	1	111	6	10	1	8	8	6	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.24808	159.6	-10.8	2.97176	149	282	-53	34.6	1.6	2.7	10^{-15}	6.0	
6138	05-042	22:42	2	1	185	4	1	0	11	12	0	1	0	0	0	37	31	1	0	0	0	0	0	0	0	4	5.24791	159.6	-10.8	2.97259	127	268	-33	43.7	1.6	2.9	10^{-15}	6.0	
6139	05-043	00:56	2	1	237	3	7	0	9	8	9	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.24617	159.6	-10.9	2.98092	26	237	63	35.4	1.6	8.1	10^{-15}	6.0	
6140	05-043	12:24	2	1	96	3	8	0	9	8	9	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.24494	159.6	-10.9	2.98676	164	300	-64	34.6	1.6	2.1	10^{-12}	6.0	
6141	05-044	19:41	3	1	196	4	9	5	8	8	7	0	1	0	0	0	0	1	0	0	0	0	0	0	0	4	5.24442	159.6	-11.0	2.98925	123	267	-29	28.0	1.6	6.2	10^{-15}	6.0	
6142	05-045	06:28	2	1	155	3	9	0	8	7	6	0	0	0	0	45	0	1	0	0	0	0	0	0	0	4	5.24283	159.6	-11.0	2.99674	142	276	-47	35.4	1.6	3.9	10^{-13}	6.0	
6143	05-045	10:34	3	1	159	7	11	1	7	7	8	0	1	0	0	1	38	0	1	0	0	0	0	0	0	4	5.23624	159.7	-11.3	3.02747	169	310	-67	34.1	1.9	5.2	10^{-14}	10.5	
6144	05-046	17:38	2	1	68	7	12	0	7	8	7	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.23317	159.7	-11.4	3.04156	227	52	-38	11.8	11.8	2.4	10^{-14}	5858.3	
6145	05-047	12:38	2	4	166	27	30	8	6	5	4	1	1	0	0	1	40	0	1	0	0	0	0	0	0	4	5.23208	159.7	-11.4	3.04653	144	276	-49	21.4	1.9	5.0	10^{-13}	10.5	
6146	05-047	23:00	2	1	137	4	9	0	8	8	8	0	0	0	0	0	34	0	1	0	0	0	0	0	0	4	5.22434	159.8	-11.7	3.08123	318	78	48	28.0	1.6	1.7	10^{-14}	6.0	
6147	05-049	02:18	3	3	151	21	26	6	5	6	6	0	1	0	0	45	0	1	0	0	0	0	0	0	0	4	5.22359	159.8	-11.7	3.08453	45	242	44	31.3	2.0	1.5	10^{-15}	12.5	
6148	05-053	13:55	2	2	152	12	20	0	8	5	6	0	0	0	0	40	0	1	0	0	0	0	0	0	0	4	5.21870	159.8	-11.9	3.10594	226	48	-39	21.4	1.9	2.6	10^{-13}	10.5	
6154	05-066	11:46	2	1	73	6	11	0	8	7	0	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.21737	159.8	-11.9	3.11170	116	260	-22	28.0	1.6	1.2	10^{-14}	6.0	
6155	05-069	03:40	2	1	95	1	6	0	15	9	13	0	0	0	0	0	0	1	0	0	0	0	0	0	0	4	5.21317	159.9	-12.1	3.12977	148	278	-52	31.3	2.0	1.5	10^{-15}	12.5	
6156	05-069	05:19	2</																																				

No.	IMP.	DATE	C	AR	S	IA	EA	CA	IT	ET	E	I	E	I	PA	P	E	I	E	C	P	C	P	C	P	R	LV	HV	R	LV	HV	R	LV	HV	ROT	S _{LV}	S _{LV}	V	VEF	M	MEF
6161	05-073	02:49	3	1	187	7	12	1	7	7	6	0	1	4	21	1	0	0	0	0	0	0	0	0	0	4	5.20717	159.9	-12.3	3.15517	288	67	19	43.7	1.6	3.3	10 ⁻¹⁵	6.0			
6162	05-074	22:31	2	1	71	5	5	1	10	15	0	1	1	0	0	0	0	0	0	0	0	0	0	0	4	5.20423	159.9	-12.4	3.16744	133	266	-38	10.4	1.9	9.3	10 ⁻¹⁴	10.5				
6163	05-076	13:07	3	2	52	8	12	10	7	8	6	0	1	0	0	0	0	0	0	0	0	0	0	4	5.20167	159.9	-12.5	3.17805	107	256	-14	35.4	1.6	9.7	10 ⁻¹⁵	6.0					
6164	05-078	06:56	3	3	92	19	22	7	6	5	6	0	1	41	0	1	0	0	0	0	0	0	0	4	5.19889	160.0	-12.6	3.18947	169	307	-66	34.6	1.6	1.4	10 ⁻¹³	6.0					
6165	05-078	08:08	2	2	91	8	13	0	8	8	6	0	0	34	0	1	0	0	0	0	0	0	0	4	5.19889	160.0	-12.6	3.18947	168	304	-65	28.0	1.6	2.4	10 ⁻¹⁴	6.0					
6166	05-078	18:10	2	2	101	2	6	0	10	9	6	0	0	0	0	0	0	0	0	0	0	0	0	4	5.19809	160.0	-12.6	3.19273	182	342	-69	31.3	2.0	1.9	10 ⁻¹⁵	12.5					
6167	05-081	13:15	2	2	114	13	21	0	8	6	6	0	0	39	0	1	0	0	0	0	0	0	0	4	5.19367	160.0	-12.7	3.21063	202	24	-59	21.4	1.9	3.6	10 ⁻¹³	10.5					
6168	05-081	13:19	3	2	101	12	20	8	7	5	6	0	1	38	0	1	0	0	0	0	0	0	0	4	5.19367	160.0	-12.7	3.21063	184	346	-68	34.1	1.9	5.2	10 ⁻¹⁴	10.5					
6169	05-082	00:30	2	1	122	5	4	1	9	15	0	1	0	0	0	0	0	0	0	0	0	0	0	4	5.19286	160.0	-12.8	3.21389	214	36	-50	14.1	1.9	3.8	10 ⁻¹⁴	10.5					
6170	05-082	16:49	1	1	168	3	1	0	9	15	0	1	0	2	21	1	0	0	0	0	0	0	0	4	5.19185	160.0	-12.8	3.21795	279	63	10	14.1	1.9	1.8	10 ⁻¹⁴	10.5					
6171	05-082	17:46	1	2	41	10	1	0	6	15	0	1	0	7	31	1	0	0	0	0	0	0	0	4	5.19185	160.0	-12.8	3.21795	100	253	-7	43.5	1.9	9.0	10 ⁻¹⁶	10.5					
6172	05-083	08:36	2	1	113	4	9	0	8	7	8	0	1	3	21	1	0	0	0	0	0	0	0	4	5.19084	160.0	-12.8	3.22201	205	27	-56	34.6	1.6	3.2	10 ⁻¹⁵	6.0					
6173	05-086	08:55	3	1	127	6	11	3	8	8	6	0	1	0	0	0	0	0	0	0	0	0	0	4	5.18593	160.0	-13.0	3.24147	227	44	-39	28.0	1.6	1.2	10 ⁻¹⁴	6.0					
6174	05-086	13:47	2	1	90	3	8	0	8	8	0	0	0	0	0	0	0	0	0	0	0	0	0	4	5.18552	160.0	-13.0	3.24309	175	319	-68	28.0	1.6	4.5	10 ⁻¹⁵	6.0					
6175	05-089	09:46	2	2	72	8	12	0	9	10	9	0	0	0	0	0	0	0	0	0	0	0	0	4	5.18076	160.0	-13.1	3.26170	153	278	-55	19.3	1.6	8.8	10 ⁻¹⁴	6.0					
6176	05-091	14:38	2	2	253	11	19	0	8	5	7	0	0	37	0	1	0	0	0	0	0	0	0	4	5.17721	160.1	-13.3	3.27543	48	235	41	21.4	1.9	1.7	10 ⁻¹³	10.5					
6177	05-093	22:13	3	1	96	6	11	1	8	8	0	1	0	1	0	0	0	0	0	0	0	0	0	4	5.17321	160.1	-13.4	3.29074	188	355	-67	28.0	1.6	1.2	10 ⁻¹⁴	6.0					
6178	05-094	00:05	2	1	160	4	8	0	8	8	0	0	0	0	0	0	0	0	0	0	0	0	0	4	5.17299	160.1	-13.4	3.29154	278	60	10	28.0	1.6	5.2	10 ⁻¹⁵	6.0					
6179	05-094	20:54	2	1	185	5	10	0	7	7	0	0	0	0	0	0	0	0	0	0	0	0	0	4	5.17172	160.1	-13.4	3.29637	314	71	44	43.7	1.6	1.7	10 ⁻¹⁵	6.0					
6180	05-096	05:25	2	1	201	1	5	0	15	9	10	0	0	8	22	1	0	0	0	0	0	0	0	4	5.16938	160.2	-13.5	3.30522	337	86	65	31.3	2.0	1.3	10 ⁻¹⁵	12.5					
6181	05-097	04:05	2	2	7	9	4	9	9	13	0	1	1	11	21	1	0	0	0	0	0	0	0	4	5.16767	160.2	-13.6	3.31164	66	240	25	14.1	1.9	7.5	10 ⁻¹⁴	10.5					
6182	05-097	14:47	3	1	150	7	12	8	7	7	6	0	1	3	21	1	0	0	0	0	0	0	0	4	5.16703	160.2	-13.6	3.31405	267	56	0	43.7	1.6	3.3	10 ⁻¹⁵	6.0					
6183	05-100	15:28	2	1	24	2	7	0	15	8	10	0	0	0	0	0	0	0	0	0	0	0	0	4	5.16163	160.2	-13.8	3.34009	91	247	0	36.7	2.0	1.3	10 ⁻¹⁵	12.5					
6184	05-100	20:16	2	2	234	10	9	6	8	8	0	1	1	0	0	0	0	0	0	0	0	0	0	4	5.16142	160.2	-13.8	3.34889	26	222	62	21.4	1.9	4.7	10 ⁻¹⁴	10.5					
6185	05-104	17:01	3	4	40	30	49	14	7	5	5	0	1	46	0	1	0	0	0	0	0	0	0	4	5.15464	160.3	-14.0	3.35967	115	254	-21	14.1	1.9	1.4	10 ⁻¹⁰	10.5					
6186	05-105	05:18	2	1	231	6	11	0	9	8	8	0	0	0	0	0	0	0	0	0	0	0	0	4	5.15376	160.3	-14.0	3.36286	25	218	63	22.7	1.6	2.5	10 ⁻¹⁴	6.0					
6187	05-114	14:37	1	1	50	1	5	0	15	0	1	0	1	11	31	1	0	0	0	0	0	0	0	4	5.13694	160.4	-14.5	3.42242	133	261	-38	2.3	2.0	6.1	10 ⁻¹²	12.5					
6188	05-116	12:48	3	2	96	12	21	1	8	6	6	0	1	39	0	1	0	0	0	0	0	0	0	4	5.13328	160.4	-14.6	3.43506	199	13	-61	21.4	1.9	3.1	10 ⁻¹³	10.5					
6189	05-118	08:10	2	2	84	13	9	3	9	15	0	1	1	42	31	1	0	0	0	0	0	0	0	4	5.13005	160.4	-14.7	3.44611	183	337	-68	14.1	1.9	3.5	10 ⁻¹³	10.5					
6190	05-119	14:49	1	1	161	5	3	0	8	15	0	1	0	0	0	0	0	0	0	0	0	0	0	4	5.12774	160.4	-14.8	3.45399	291	60	22	21.4	1.9	7.3	10 ⁻¹⁵	10.5					
6191	05-122	09:46	2	1	185	4	9	0	8	8	0	0	0	0	0	0	0	0	0	0	0	0	0	4	5.12237	160.5	-14.9	3.47206	326	74	55	28.0	1.6	6.2	10 ⁻¹⁵	6.0					
6192	05-123	05:19	3	4	7	30	49	17	8	5	5	0	1	46	0	1	0	0	0	0	0	0	0	4	5.12097	160.5	-15.0	3.47677	136	240	15	10.4	1.9	3.0	10 ⁻¹⁰	10.5					
6193	05-123	07:10	2	1	50	5	0	1	9	15	0	1	1	0	0	0	0	0	0	0	0	0	0	4	5.12073	160.5	-15.0	3.47755	76	262	-41	14.1	1.9	3.8	10 ⁻¹⁴	10.5					
6194	05-123	14:27	2	1	29	7	12	0	8	8	7	0	0	35	0	1	0	0	0	0	0	0	0	4	5.12026	160.5	-15.0	3.47912	107	249	-14	28.0	1.6	1.7	10 ⁻¹⁴	6.0					
6195	05-123	18:00	3	1	245	7	12	11	7	8	6	0	1	0	0	0	0	0	0	0	0	0	0	4	5.11979	160.5	-15.0	3.48069	51	232	39	35.4	1.6	8.1	10 ⁻¹⁵	6.0					
6196	05-123	20:44	1	1	85	3	2	0	9	15	0	1	0	0	0	0	0	0	0	0	0	0	0	4	5.11979	160.5	-15.0	3.48069	186	345	-67	14.1	1.9	2.1	10 ⁻¹⁴	10.5					
6197	05-124	06:04	2	2	105	9	20	0	8	6	7	0	0	37	0	1	0	0	0	0	0	0	0	4	5.11885	160.5	-15.0	3.48383	214	30	-49	21.4	1.9	1.6	10 ⁻¹³	10.5					
6198	05-124	20:40	3	2	26	14	21	19	8	7	6	0	1	40	0	1	0	0	0	0	0	0	0	4	5.11791	160.5	-15.1	3.48696	103	248	-10	21.4	1.9	4.2	10 ⁻¹³	10.5					
6199	05-127	19:53	1	1	106	4	3	0	9	15	0	1	0	0	0	0	0	0	0	0	0	0	0	4	5.11221	160.5	-15.2	3.50574	217	32	-47	14.1	1.9	2.8	10 ⁻¹⁴	10.5					
6200	05-128	02:41	1	1	67	2	2	0	9	15	0	1	0	0	0	0	0	0	0	0	0	0	0	4	5.11173	160.5	-15.2	3.50730	162	287	-61	11.8	11.8	2.8	10 ⁻¹⁴	5858.3					
6201	05-128	14:07	2	1	72	5	3	4	9	15	0	1	1	0	0	0	0	0	0	0	0	0	0	4	5.11077	160.5	-15.3	3.51042	169	301	-65	14.1	1.9	3.3	10 ⁻¹⁴	10.5					
6202	05-129	22:22	2	2	64	12	5	4	9	15	0	1	1	36	31	1	0	0	0	0	0	0	0	4	5.10814	160.5	-15.3	3.51900	158	282	-59	14.1	1.9	1.4	10 ⁻¹³	10.5					
6203	05-130	07:30	2	2	6	9	13	0	8	6	0	0	34	0	1	0	0																								

No.	IMP DATE	C L N	AR S E C	IA	EA	CA	IT	ET	E I T	E I C	I I C	PA	P E T	E V D	I C P	E C P	C C P	P P	HV	R	LON	LAT	D _{rup}	ROT	S _{LON}	S _{LAT}	V	V _{EF}	M	MEF	
6216	05-152 08:34	2	2	97	9	19	0	8	6	7	0	0	37	0	1	0	0	0	1	4	5.06350	160.8	-16.6	3.65701	211	27	-52	21.4	1.9	1.2·10 ⁻¹³	10.5
6217	05-152 19:34	3	2	52	12	20	2	8	5	6	0	1	38	0	1	0	0	0	1	4	5.06221	160.8	-16.6	3.66006	148	270	-51	21.4	1.9	2.6·10 ⁻¹³	10.5
6218	05-152 23:51	3	2	50	12	20	3	9	7	6	0	1	38	0	1	0	0	0	1	4	5.06227	160.8	-16.6	3.66082	145	268	-49	14.1	1.9	1.2·10 ⁻¹²	10.5
6219	05-157 23:57	3	1	66	5	9	13	7	8	7	0	1	0	0	0	0	0	0	1	4	5.05177	160.9	-16.9	3.69120	169	301	-65	35.4	1.6	3.5·10 ⁻¹⁵	6.0
6220	05-160 17:43	2	2	93	10	19	0	8	6	7	0	0	36	0	1	0	0	0	1	4	5.04596	160.9	-17.1	3.70785	208	24	-54	21.4	1.9	1.5·10 ⁻¹³	10.5
6221	05-164 23:25	2	2	68	14	21	0	8	5	6	0	0	40	0	1	0	0	0	1	4	5.03688	161.0	-17.3	3.73346	174	313	-67	21.4	1.9	4.2·10 ⁻¹³	10.5
6222	05-164 23:43	2	1	21	6	11	0	9	8	0	0	0	0	0	0	0	0	0	1	4	5.03688	161.0	-17.3	3.73346	108	249	-15	21.0	1.6	3.6·10 ⁻¹⁴	6.0
6223	05-169 11:56	3	2	52	10	14	1	9	10	8	0	0	1	38	0	1	0	0	1	4	5.02714	161.0	-17.6	3.76046	152	276	-54	19.3	1.6	1.7·10 ⁻¹³	6.0
6224	05-173 14:39	2	1	38	1	1	0	15	1	11	0	0	5	24	1	0	0	0	1	4	5.01809	161.1	-17.8	3.78508	134	261	-39	11.8	1.1	1.9·10 ⁻¹⁴	58858.3
6225	05-173 19:34	3	2	59	8	12	10	7	8	7	0	1	0	0	0	0	0	0	1	4	5.01754	161.1	-17.8	3.78657	163	291	-62	35.4	1.6	9.7·10 ⁻¹⁵	6.0
6226	05-173 23:29	2	2	19	8	12	0	7	8	7	0	0	34	0	1	0	0	0	1	4	5.01727	161.1	-17.8	3.78731	107	250	-14	35.4	1.6	9.7·10 ⁻¹⁵	6.0
6227	05-174 12:06	2	2	59	8	13	0	8	6	6	0	0	39	0	1	0	0	0	1	4	5.01588	161.1	-17.8	3.79103	163	292	-62	28.0	1.6	2.4·10 ⁻¹⁴	6.0
6228	05-175 11:05	3	2	48	13	21	1	8	6	6	0	1	35	9	1	0	0	0	1	4	5.01395	161.1	-17.9	3.79623	148	273	-51	21.4	1.9	3.6·10 ⁻¹³	10.5
6229	05-175 17:44	3	1	112	6	10	3	7	8	6	0	1	0	0	0	0	0	0	1	4	5.01339	161.1	-17.9	3.79772	238	46	-28	35.4	1.6	4.8·10 ⁻¹⁵	6.0
6230	05-175 22:10	3	2	72	12	20	0	7	5	6	0	0	38	0	1	0	0	0	1	4	5.01284	161.1	-17.9	3.79920	182	337	-67	34.1	1.9	5.2·10 ⁻¹⁴	10.5
6231	05-176 16:57	2	2	44	10	20	12	7	5	6	0	1	36	0	1	0	0	0	1	4	5.01117	161.1	-18.0	3.80365	143	268	-47	34.1	1.9	3.8·10 ⁻¹⁴	10.5
6232	05-176 17:16	2	1	54	3	8	0	9	8	10	0	0	21	1	0	0	0	0	1	4	5.01117	161.1	-18.0	3.80365	157	283	-58	22.7	1.6	9.4·10 ⁻¹⁵	6.0
6233	05-176 18:49	2	2	56	13	8	9	10	11	0	1	1	35	9	1	0	0	0	1	4	5.01089	161.1	-18.0	3.80440	160	287	-60	10.4	1.9	6.0·10 ⁻¹³	10.5
6234	05-177 01:56	3	2	4	13	21	21	8	7	6	0	1	39	0	1	0	0	0	1	4	5.01033	161.1	-18.0	3.80588	87	245	4	21.4	1.9	3.6·10 ⁻¹³	10.5
6235	05-177 06:08	3	1	63	5	9	4	7	8	8	0	1	0	0	0	0	0	0	1	4	5.00977	161.1	-18.0	3.80736	170	305	-65	35.4	1.6	3.5·10 ⁻¹⁵	6.0
6236	05-178 21:22	3	2	39	9	13	4	9	9	7	0	1	0	0	0	0	0	0	1	4	5.00614	161.2	-18.1	3.81699	137	264	-41	21.0	1.6	8.3·10 ⁻¹⁴	6.0
6237	05-184 03:54	2	1	238	3	9	0	9	8	9	0	0	0	0	0	0	0	0	1	4	4.99428	161.2	-18.4	3.84796	59	238	31	22.7	1.6	1.1·10 ⁻¹⁴	6.0
6238	05-186 01:59	1	1	65	7	4	0	11	15	0	1	0	40	31	1	0	0	0	1	4	4.99000	161.2	-18.5	3.85897	176	321	-67	7.8	1.9	2.4·10 ⁻¹³	10.5
6239	05-186 17:18	2	1	250	5	3	8	9	15	0	1	0	0	0	0	0	0	0	1	4	4.98857	161.3	-18.5	3.86264	76	243	14	14.1	1.9	3.3·10 ⁻¹⁴	10.5
6240	05-189 06:06	2	1	167	7	14	0	7	7	7	0	1	0	36	0	1	0	0	1	4	4.98253	161.3	-18.7	3.87800	321	75	50	34.1	1.9	1.3·10 ⁻¹⁴	10.5
6241	05-189 17:19	2	1	40	5	8	3	11	15	0	1	1	0	0	0	0	0	0	1	4	4.96265	161.4	-19.2	3.92742	177	326	-67	14.1	1.9	3.8·10 ⁻¹⁴	10.5
6242	05-190 20:37	3	2	29	11	20	1	8	5	6	0	1	31	1	0	0	0	0	1	4	4.98166	161.3	-18.7	3.88019	143	270	-46	4.2	1.6	2.4·10 ⁻¹²	6.0
6243	05-192 03:33	3	3	63	19	24	1	6	6	6	0	1	43	0	1	0	0	0	1	4	4.97905	161.3	-18.8	3.88676	128	261	-33	21.4	1.9	2.2·10 ⁻¹³	10.5
6244	05-192 12:28	3	3	40	23	28	8	6	5	5	0	1	46	0	1	0	0	0	1	4	4.97586	161.3	-18.9	3.89477	176	322	-67	28.0	1.6	3.8·10 ⁻¹³	6.0
6245	05-197 19:36	2	1	62	5	4	2	9	15	0	1	1	0	0	0	0	0	0	1	4	4.96324	161.4	-19.2	3.92597	146	274	-49	34.6	1.6	8.3·10 ⁻¹³	6.0
6246	05-199 20:21	2	2	73	8	20	0	9	4	9	0	0	39	0	1	0	0	0	1	4	4.95790	161.4	-19.3	3.93897	193	6	-63	14.1	1.9	6.5·10 ⁻¹³	10.5
6247	05-201 14:07	3	3	234	19	23	3	6	6	0	1	42	0	0	0	0	0	0	1	4	4.95373	161.4	-19.4	3.94905	61	241	29	28.0	1.6	3.3·10 ⁻¹³	6.0
6248	05-201 18:03	2	2	84	11	20	0	7	5	6	0	1	37	0	1	0	0	0	1	4	4.95313	161.4	-19.4	3.95048	211	31	-52	34.1	1.9	4.4·10 ⁻¹⁴	10.5
6249	05-204 13:40	1	1	36	3	1	0	9	15	0	1	0	0	0	0	0	0	0	1	4	4.94652	161.5	-19.6	3.96628	144	273	-47	14.1	1.9	1.8·10 ⁻¹⁴	10.5
6250	05-206 08:15	3	2	15	8	13	20	7	8	6	0	1	0	0	0	0	0	0	1	4	4.94229	161.5	-19.7	3.97630	117	259	-23	35.4	1.6	1.1·10 ⁻¹⁴	6.0
6251	05-208 10:23	2	2	27	7	12	0	8	8	0	0	0	35	0	1	0	0	0	1	4	4.93712	161.5	-19.8	3.98843	134	267	-39	28.0	1.6	2.8·10 ⁻¹⁴	6.0
6252	05-208 20:15	2	1	37	7	12	0	7	8	7	0	0	0	0	0	0	0	0	1	4	4.93621	161.5	-19.8	3.99057	149	278	-51	35.4	1.6	8.1·10 ⁻¹⁵	6.0
6253	05-209 02:16	3	2	4	10	20	12	8	6	6	0	1	36	0	1	0	0	0	1	4	4.93560	161.5	-19.8	3.99199	102	254	-9	21.4	1.9	1.9·10 ⁻¹³	10.5
6254	05-209 08:06	1	1	124	3	3	0	9	15	0	1	0	35	31	1	0	0	0	1	4	4.93499	161.5	-19.9	3.99342	271	61	3	14.1	1.9	2.4·10 ⁻¹⁴	10.5
6255	05-209 15:45	2	1	31	7	3	2	8	15	0	1	1	0	0	0	0	0	0	1	4	4.93407	161.6	-19.9	3.99555	141	271	-44	21.4	1.9	1.0·10 ⁻¹⁴	10.5
6256	05-209 22:02	2	2	224	8	5	7	10	13	0	1	1	10	18	1	0	0	0	1	4	4.93346	161.6	-19.9	3.99698	52	238	38	10.4	1.9	1.6·10 ⁻¹³	10.5
6257	05-210 08:54	3	2	23	11	20	11	8	6	6	0	1	37	0	1	0	0	0	1	4	4.93254	161.6	-19.9	3.99911	130	264	-35	21.4	1.9	2.2·10 ⁻¹³	10.5
6258	05-210 11:08	3	2	34	14	22	9	7	5	6	0	1	40	0	1	0	0	0	1	4	4.93224	161.6	-19.9	3.99982	145	275	-48	34.1	1.9	9.7·10 ⁻¹⁴	10.5
6259	05-211 09:50	3	3	21	21	25	18	7	8	6	0	1	43	0	1	0	0	0	1	4	4.92978	161.6	-20.0	4.00551	129	265	-34	19.3	1.6	3.0·10 ⁻¹²	6.0
6260	05-211 17:32	2	1	26	7	5	10	9	13	0	1	1	0	0	0	0	0	0	1	4	4.92917	161.6	-20.0	4.00693	136	269	-40	14.1	1.9	6.4·10 ⁻¹⁴	10.5
6261	05-211 22:26	2	1	186	6	11	0	8	7	0	0	0	0	0	0	0	0	0	1	4	4.92855	161.6	-20.0	4.00835	1	162	78	28.0	1.6	1.2·10 ⁻¹⁴	6.0
6262	05-212 16:59	3	1	119	7	11	3	7	8	6	0	1	0	0	0	0	0	0	1	4	4.92670	161.6	-20.1	4.01260	267	60	0	35.4	1.6	6	

No.	IMP.	DATE	C	AR	S	IA	EA	CA	IT	ET	E	I	C	E	I	PA	P	E	V	D	P	C	C	P	P	R	LON	LAT	D _{rup}	ROT	S _{LON}	S _{LAT}	V	VEFF	M	MEF	
6271	05-224	03-53	1	1	76	2	4	0	11	15	0	1	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.89789	161.8	-20.7	4.07730	221	43	-43	11.8	11.8	3.6·10 ⁻¹⁴	5858.3
6272	05-225	20-41	3	3	27	20	24	1	5	5	6	0	1	44	0	1	44	0	1	0	0	0	0	0	0	4	4.89375	161.8	-20.8	4.08636	153	285	-54	43.7	1.6	9.8·10 ⁻¹⁴	6.0
6273	05-226	03-58	3	3	25	19	23	5	6	7	6	0	1	41	0	1	41	0	1	0	0	0	0	0	4	4.89279	161.8	-20.9	4.08845	150	283	-52	25.9	1.6	4.3·10 ⁻¹³	6.0	
6274	05-226	13-00	3	3	230	22	29	11	6	3	5	0	1	46	0	1	46	0	1	0	0	0	0	0	4	4.89183	161.8	-20.9	4.09054	78	249	12	21.4	1.9	4.4·10 ⁻¹²	10.5	
6275	05-226	18-20	1	1	206	5	5	0	10	15	0	1	0	13	31	1	31	1	0	0	0	0	0	0	4	4.89119	161.8	-20.9	4.09193	45	238	44	10.4	1.9	9.3·10 ⁻¹⁴	10.5	
6276	05-226	19-20	2	2	3	9	3	5	9	15	0	1	0	0	0	0	0	0	0	0	0	0	0	0	4	4.89119	161.8	-20.9	4.09193	119	263	-25	14.1	1.9	6.6·10 ⁻¹⁴	10.5	
6277	05-227	18-13	2	1	52	5	1	6	9	13	0	1	1	0	0	0	0	0	0	0	0	0	0	0	4	4.88863	161.8	-21.0	4.09748	189	360	-65	14.1	1.9	2.5·10 ⁻¹⁴	10.5	
6278	05-228	19-33	2	1	246	4	1	3	9	15	0	1	0	0	0	0	0	0	0	0	0	0	0	0	4	4.88606	161.8	-21.0	4.10303	107	259	-13	14.1	1.9	2.1·10 ⁻¹⁴	10.5	
6279	05-228	22-34	2	1	53	2	5	0	15	9	10	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.88573	161.8	-21.0	4.10372	195	14	-62	31.3	2.0	1.6·10 ⁻¹⁵	12.5	
6280	05-229	08-28	1	2	239	13	14	0	12	15	0	1	0	24	31	1	31	1	0	0	0	0	0	0	4	4.88477	161.8	-21.1	4.10580	97	256	-4	5.0	1.9	1.4·10 ⁻¹¹	10.5	
6281	05-229	15-08	3	2	247	14	21	14	9	9	7	0	1	39	0	1	39	0	1	0	0	0	0	0	4	4.88380	161.8	-21.1	4.10788	108	259	-15	14.1	1.9	1.9·10 ⁻¹²	10.5	
6282	05-229	15-23	3	2	245	10	20	1	7	6	6	0	1	36	0	1	36	0	1	0	0	0	0	0	4	4.88380	161.8	-21.1	4.10788	106	258	-12	34.1	1.9	3.8·10 ⁻¹⁴	10.5	
6283	05-231	03-05	3	2	251	9	14	4	8	5	7	0	1	37	0	1	37	0	1	0	0	0	0	0	4	4.87992	161.8	-21.2	4.11618	115	262	-21	28.0	1.6	3.3·10 ⁻¹⁴	6.0	
6284	05-232	23-06	3	2	249	14	22	4	7	6	6	0	1	41	0	1	41	0	1	0	0	0	0	0	4	4.87537	161.9	-21.3	4.12584	119	264	-25	34.1	1.9	9.7·10 ⁻¹⁴	10.5	
6285	05-233	17-50	3	2	43	9	13	2	9	8	0	1	0	0	0	0	0	0	0	0	0	0	0	0	4	4.87342	161.9	-21.3	4.12997	190	3	-64	21.0	1.6	8.3·10 ⁻¹⁴	6.0	
6286	05-235	04-56	2	1	57	4	8	0	8	8	7	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.86950	161.9	-21.4	4.13822	137	274	-41	28.0	1.6	1.5·10 ⁻¹⁴	6.0	
6287	05-235	20-36	1	1	58	4	1	0	10	15	0	1	0	0	0	0	0	0	0	0	0	0	0	0	4	4.86786	161.9	-21.4	4.14165	211	36	-51	10.4	1.9	4.3·10 ⁻¹⁴	10.5	
6288	05-238	01-14	3	2	253	11	20	1	8	5	6	0	1	44	0	1	44	0	1	0	0	0	0	0	4	4.86194	161.9	-21.6	4.15398	133	272	-38	21.4	1.9	2.2·10 ⁻¹³	10.5	
6289	05-240	14-27	3	3	28	19	23	1	6	5	6	0	1	41	0	1	41	0	1	0	0	0	0	0	4	4.85533	162.0	-21.7	4.16763	177	334	-66	34.6	1.6	1.7·10 ⁻¹³	6.0	
6290	05-243	09-49	2	1	233	6	11	0	7	7	8	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.84766	162.0	-21.9	4.18327	116	265	-22	43.7	1.6	2.3·10 ⁻¹⁵	6.0	
6291	05-244	00-30	2	1	157	4	8	0	8	8	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.84599	162.0	-21.9	4.18666	9	196	74	28.0	1.6	5.2·10 ⁻¹⁵	6.0	
6292	05-244	07-08	3	4	25	27	49	28	8	8	5	0	1	44	0	1	44	0	1	0	0	0	0	0	4	4.84532	162.0	-22.0	4.18801	183	349	-66	10.4	1.9	1.9·10 ⁻¹⁰	10.5	
6293	05-246	17-55	3	3	24	19	25	1	6	5	6	0	1	46	0	1	46	0	1	0	0	0	0	0	4	4.83893	162.1	-22.1	4.20086	191	8	-63	34.6	1.6	2.3·10 ⁻¹³	6.0	
6294	05-246	20-27	2	1	191	2	6	0	9	7	10	0	0	3	21	1	21	1	0	0	0	0	0	0	4	4.83860	162.1	-22.1	4.20153	66	249	24	56.0	2.0	1.8·10 ⁻¹⁶	12.5	
6295	05-249	07-11	3	2	137	9	14	1	8	8	6	0	1	36	0	1	36	0	1	0	0	0	0	0	4	4.83183	162.1	-22.3	4.21500	1	165	76	28.0	1.6	3.3·10 ⁻¹⁴	6.0	
6296	05-250	17-23	2	2	0	8	6	5	8	15	0	1	1	0	0	0	0	0	0	0	0	0	0	0	4	4.82809	162.1	-22.3	4.22239	167	314	-62	21.4	1.9	2.0·10 ⁻¹⁴	10.5	
6297	05-253	04-35	2	2	252	11	20	0	7	5	7	0	0	38	0	1	38	0	1	0	0	0	0	0	4	4.82126	162.2	-22.5	4.23578	171	323	-64	34.1	1.9	4.4·10 ⁻¹⁴	10.5	
6298	05-253	12-40	3	2	240	14	21	7	8	6	6	0	1	39	0	1	39	0	1	0	0	0	0	0	4	4.82023	162.2	-22.5	4.23778	154	294	-54	21.4	1.9	4.2·10 ⁻¹³	10.5	
6299	05-254	21-23	3	3	194	19	22	7	7	6	6	0	1	41	0	1	41	0	1	0	0	0	0	0	4	4.81645	162.2	-22.6	4.24512	89	258	2	22.7	1.6	5.8·10 ⁻¹³	6.0	
6300	05-255	14-08	3	2	5	14	22	22	7	5	6	0	1	40	0	1	40	0	1	0	0	0	0	0	4	4.81472	162.2	-22.7	4.24845	183	351	-65	34.1	1.9	9.7·10 ⁻¹⁴	10.5	
6301	05-255	19-29	3	2	209	14	21	2	7	5	6	0	1	39	0	1	39	0	1	0	0	0	0	0	4	4.81403	162.2	-22.7	4.24978	110	265	-16	34.1	1.9	8.3·10 ⁻¹⁴	10.5	
6302	05-255	21-02	1	1	163	4	2	0	9	15	0	1	0	21	1	21	1	0	0	0	0	0	0	0	4	4.81369	162.2	-22.7	4.25045	45	302	44	14.1	1.9	2.5·10 ⁻¹⁴	10.5	
6303	05-255	22-30	1	1	245	3	10	0	14	15	0	1	0	58	31	1	58	31	1	0	0	0	0	0	4	4.81369	162.2	-22.7	4.25045	160	302	-58	2.0	1.9	3.9·10 ⁻¹¹	10.5	
6304	05-256	11-26	3	3	227	19	23	23	7	6	6	0	1	41	0	1	41	0	1	0	0	0	0	0	4	4.81231	162.2	-22.7	4.25311	135	277	-39	22.7	1.6	6.8·10 ⁻¹³	6.0	
6305	05-256	23-37	3	3	218	19	25	8	7	7	6	0	1	43	0	1	43	0	1	0	0	0	0	0	4	4.81093	162.2	-22.7	4.25576	130	275	-34	21.0	1.6	1.3·10 ⁻¹²	6.0	
6306	05-260	20-31	3	3	208	20	26	13	5	6	5	0	1	44	0	1	44	0	1	0	0	0	0	0	4	4.80015	162.3	-23.0	4.27029	122	271	-27	35.4	1.6	3.3·10 ⁻¹³	6.0	
6307	05-267	02-36	2	3	230	21	26	12	5	5	5	0	1	44	0	1	44	0	1	0	0	0	0	0	4	4.78255	162.4	-23.4	4.30913	164	311	-60	43.7	1.6	1.6·10 ⁻¹³	6.0	
6308	05-270	06-15	2	1	161	6	11	0	8	8	7	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.77329	162.4	-23.6	4.32607	72	255	19	28.0	1.6	1.2·10 ⁻¹⁴	6.0	
6309	05-272	18-16	3	3	199	19	23	3	5	6	6	0	1	42	0	1	42	0	1	0	0	0	0	0	4	4.76611	162.5	-23.7	4.33904	124	275	-29	35.4	1.6	1.5·10 ⁻¹³	6.0	
6310	05-274	00-02	2	1	162	3	7	1	13	15	0	1	1	42	31	1	42	31	1	0	0	0	0	0	4	4.76251	162.5	-23.8	4.34551	77	257	14	2.4	1.6	1.1·10 ⁻¹¹	6.0	
6311	05-276	19-00	3	3	185	20	25	13	7	8	6	0	1	43	0	1	43	0	1	0	0	0	0	0	4	4.75454	162.5	-24.0	4.35968	113	271	-18	19.3	1.6	2.5·10 ⁻¹²	6.0	
6312	05-276	23-19	3	2	220	15	23	7	8	7	6	0	1	41	0	1	41	0	1	0	0	0	0	0	4	4.75418	162.5	-24.0	4.36032	162	311	-58	21.4	1.9	6.4·10 ⁻¹³	10.5	

No.	IMP.	DATE	C	AR	S	IA	EA	CA	IT	ET	E	I	E	E	I	PA	P	E	E	I	E	C	P	C	P	C	P	R	LON	LAT	D _{rup}	ROT	S _{LON}	S _{LAT}	V	VEF	M	MEF
6326	05-301	06:10	3	3	187	21	25	8	5	5	6	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.68123	162.9	-25.6	4.48295	132	286	-34	43.7	1.6	1.4-10 ⁻¹³	6.0
6327	05-310	14:22	2	1	7	3	8	0	9	8	9	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.66046	162.9	-25.6	4.48418	238	64	-26	22.7	1.6	9.4-10 ⁻¹⁵	6.0	
6328	05-307	08:03	2	1	155	3	7	0	9	8	9	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.64263	163.0	-26.0	4.51229	84	264	7	22.7	1.6	7.8-10 ⁻¹⁵	6.0		
6329	05-308	03:39	3	2	210	13	21	20	7	5	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.65990	163.0	-26.0	4.51655	168	330	-58	34.1	1.9	7.1-10 ⁻¹⁴	10.5		
6330	05-308	14:42	3	2	218	12	21	1	8	7	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.65872	163.0	-26.0	4.51837	179	352	-61	21.4	1.9	3.1-10 ⁻¹³	10.5		
6331	05-310	21:22	3	2	191	13	21	10	9	9	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.65127	163.1	-26.2	4.52986	140	295	-41	14.1	1.9	1.7-10 ⁻¹²	10.5		
6332	05-310	21:37	2	2	94	9	14	0	8	8	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.65127	163.1	-26.2	4.52986	4	186	71	28.0	1.6	3.3-10 ⁻¹⁴	6.0		
6333	05-312	05:11	2	2	216	10	8	9	15	0	1	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.64733	163.1	-26.3	4.53589	179	352	-60	14.1	1.9	1.8-10 ⁻¹³	10.5		
6334	05-314	15:17	3	2	156	13	22	5	8	6	7	0	0	0	0	0	0	0	0	0	0	0	0	0	5	4.63941	163.1	-26.4	4.54790	94	270	0	21.4	1.9	4.2-10 ⁻¹³	10.5		
6335	05-318	21:22	3	3	150	19	23	12	7	9	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.62586	163.2	-26.7	4.56819	88	269	4	16.9	1.6	2.3-10 ⁻¹²	6.0		
6336	05-323	02:13	1	1	206	3	4	0	11	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.61257	163.3	-27.0	4.58770	169	335	-57	7.8	1.9	1.2-10 ⁻¹³	10.5		
6337	05-323	08:49	1	1	104	1	2	0	15	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.61176	163.3	-27.0	4.58888	26	230	58	11.8	1.1	2.3-10 ⁻¹⁴	5858.3		
6338	05-330	12:19	1	1	215	6	3	0	9	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.58811	163.3	-27.5	4.62277	183	2	-58	14.1	1.9	3.9-10 ⁻¹⁴	10.5		
6339	05-332	22:39	2	2	128	11	9	7	6	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.58027	163.4	-27.7	4.63376	60	257	29	43.5	1.9	3.9-10 ⁻¹⁵	10.5		
6340	05-336	10:51	3	4	161	24	31	12	8	6	5	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.56865	163.5	-27.9	4.64985	109	280	-14	10.4	1.9	7.5-10 ⁻¹¹	10.5		
6341	05-337	10:38	3	2	168	11	20	9	9	8	7	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.56532	163.5	-28.0	4.65443	119	286	-22	14.1	1.9	1.0-10 ⁻¹²	10.5		
6342	05-337	21:40	3	2	154	14	20	14	9	11	7	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.56365	163.5	-28.0	4.65671	99	275	-5	10.4	1.6	3.4-10 ⁻¹²	6.0		
6343	05-352	18:12	1	1	189	1	1	0	15	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.51310	163.8	-29.0	4.72350	154	317	-47	11.8	1.1	1.9-10 ⁻¹⁴	5858.3		
6344	05-354	00:23	3	3	156	19	23	8	6	6	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.50878	163.8	-29.1	4.72901	107	281	-11	28.0	1.6	3.3-10 ⁻¹³	6.0		
6345	05-364	06:07	2	2	200	9	6	2	9	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.47290	164.0	-29.9	4.77357	178	355	-52	14.1	1.9	1.1-10 ⁻¹³	10.5		
6346	06-001	17:36	3	2	152	13	20	14	9	8	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.46448	164.1	-30.0	4.78374	110	283	-13	14.1	1.9	1.4-10 ⁻¹²	10.5		
6347	06-002	08:42	3	1	118	1	6	0	15	9	11	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.46225	164.1	-30.1	4.78641	62	257	27	31.3	2.0	1.5-10 ⁻¹⁵	12.5		
6348	06-005	07:11	3	3	136	19	23	6	6	6	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.45154	164.1	-30.3	4.79915	91	273	2	28.0	1.6	3.3-10 ⁻¹³	6.0		
6349	06-005	08:33	2	1	156	4	6	4	11	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.45154	164.1	-30.3	4.79915	119	290	-20	4.2	1.6	1.4-10 ⁻¹²	6.0		
6350	06-007	05:26	2	2	123	10	14	9	10	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.44480	164.2	-30.4	4.80707	72	263	18	4.9	1.7	9.6-10 ⁻¹²	7.7		
6351	06-009	06:00	3	2	181	13	20	11	7	6	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.43714	164.2	-30.6	4.81599	153	320	-43	21.4	1.9	3.1-10 ⁻¹³	10.5		
6352	06-013	19:03	3	2	164	13	19	10	9	9	8	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.42080	164.3	-30.9	4.83473	134	301	-31	16.9	1.6	6.1-10 ⁻¹³	6.0		
6353	06-016	15:04	1	2	160	8	5	0	13	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.41029	164.4	-31.1	4.84658	134	301	-30	21.4	1.9	1.7-10 ⁻¹⁴	10.5		
6354	06-017	06:04	3	1	254	3	8	13	8	13	8	9	0	0	0	0	0	0	0	0	0	0	0	0	4	4.40799	164.4	-31.2	4.84915	266	81	0	28.0	1.6	4.5-10 ⁻¹⁵	6.0		
6355	06-018	04:42	2	2	182	8	8	4	8	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.40477	164.4	-31.2	4.85273	165	335	-46	21.4	1.9	2.9-10 ⁻¹⁴	10.5		
6356	06-019	07:32	3	2	118	13	20	11	7	6	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.40063	164.4	-31.3	4.85733	74	262	16	34.1	1.9	6.1-10 ⁻¹⁴	10.5		
6357	06-020	03:05	1	1	206	5	2	0	10	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.39740	164.4	-31.4	4.86090	198	24	-45	10.4	1.9	6.1-10 ⁻¹⁴	10.5		
6358	06-026	14:41	1	1	114	3	11	0	13	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.37368	164.6	-31.8	4.88662	80	264	11	2.5	1.9	1.8-10 ⁻¹¹	10.5		
6359	06-028	18:17	2	2	120	8	3	6	10	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.36524	164.6	-32.0	4.89559	88	269	5	10.4	1.9	1.2-10 ⁻¹³	10.5		
6360	06-030	01:25	3	3	147	23	30	26	8	5	5	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.36053	164.6	-32.1	4.90055	126	294	-23	10.4	1.9	5.8-10 ⁻¹¹	10.5		
6361	06-030	23:04	1	1	98	6	9	0	11	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.35723	164.6	-32.2	4.90402	57	248	29	4.2	1.6	3.3-10 ⁻¹²	6.0		
6362	06-032	08:21	3	2	136	13	21	11	7	6	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.35202	164.7	-32.3	4.90944	110	282	-11	21.4	1.9	3.6-10 ⁻¹³	10.5		
6363	06-033	13:06	2	2	114	9	14	0	8	8	7	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.34728	164.7	-32.4	4.91435	87	267	5	28.0	1.6	3.3-10 ⁻¹⁴	6.0		
6364	06-033	15:24	2	1	176	3	8	0	8	8	9	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.34681	164.7	-32.4	4.91484	175	347	-45	28.0	1.6	4.5-10 ⁻¹⁵	6.0		
6365	06-034	15:58	3	2	138	15	22	1	8	6	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.34301	164.7	-32.5	4.91875	121	289	-19	21.4	1.9	5.4-10 ⁻¹³	10.5		
6366	06-035	15:36	2	2	120	10	14	8	8	8	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.33920	164.7	-32.5	4.92266	95	272	0	28.0	1.6	3.9-10 ⁻¹⁴	6.0		
6367	06-041	14:21	3	2	135	13	21	11	7	6	6	0	0	0	0	0	0	0	0	0	0	0	0	0	4	4.31665	164.9	-33.0	4.94537	122	290	-20	34.1	1.9	7.1-10 ⁻¹⁴	10.5		
6368	06-043	03:59	1																																			

No.	IMP DATE	C L N	AR	S E C	IA	EA	CA	IT	ET	E I T	E I C	I I C	I I C	PA	P E T	E V D	I C P	E C C	C C P	P P	HV	R	LON	LAT	D _{rup}	ROT	S _{LON}	S _{LAT}	V	V _{EF}	M	MEF	
6381	06-085 16:50	3	3	97	19	23	18	6	6	5	0	1	42	0	1	0	0	0	0	0	1	4	4.13870	165.9	-36.5	5.10310	156	308	-34	28.0	1.6	3.3-10 ⁻¹³	6.0
6382	06-090 17:18	2	3	96	22	28	29	5	5	5	0	1	46	0	1	0	0	0	0	0	1	4	4.11756	166.1	-36.9	5.11949	157	308	-34	43.7	1.6	2.7-10 ⁻¹³	6.0
6383	06-091 02:03	2	2	71	14	12	6	9	15	0	0	1	36	31	1	0	0	0	0	0	1	4	4.11597	166.1	-36.9	5.12071	122	274	-17	14.1	1.9	6.7-10 ⁻¹³	10.5
6384	06-092 10:11	2	2	58	9	10	7	9	9	0	1	1	0	0	0	1	0	0	0	0	1	4	4.11012	166.1	-37.1	5.12516	112	264	-10	14.1	1.9	2.2-10 ⁻¹³	10.5
6385	06-094 02:05	2	1	45	5	10	0	8	8	7	0	0	0	0	0	1	0	0	0	0	1	4	4.10318	166.2	-37.2	5.13039	94	251	1	28.0	1.6	8.8-10 ⁻¹⁵	6.0
6386	06-099 10:30	3	2	60	12	20	5	8	6	8	0	1	36	0	1	0	0	0	0	0	1	4	4.08008	166.3	-37.7	5.14745	123	271	-17	21.4	1.9	2.6-10 ⁻¹³	10.5
6387	06-100 06:53	2	2	64	9	14	0	8	7	0	0	1	36	0	1	0	0	0	0	0	1	4	4.07630	166.3	-37.7	5.15019	129	276	-21	28.0	1.6	3.3-10 ⁻¹⁴	6.0
6388	06-100 11:10	2	2	43	14	5	12	9	15	0	0	1	37	31	1	0	0	0	0	0	1	4	4.07576	166.3	-37.7	5.15058	99	253	-2	14.1	1.9	2.0-10 ⁻¹³	10.5
6389	06-102 07:49	3	3	107	20	25	6	6	6	6	0	1	44	0	1	0	0	0	0	0	1	4	4.06763	166.4	-37.9	5.15642	190	345	-38	28.0	1.6	5.9-10 ⁻¹³	6.0
6390	06-104 23:53	2	1	239	5	10	0	7	7	8	0	0	0	0	0	1	0	0	0	0	1	4	4.05621	166.5	-38.1	5.16453	17	177	47	43.7	1.6	1.7-10 ⁻¹⁵	6.0
6391	06-109 23:14	2	2	111	10	4	4	9	15	0	1	1	34	31	1	0	0	0	0	0	1	4	4.03428	166.6	-38.6	5.17972	204	1	-34	14.1	1.9	8.9-10 ⁻¹⁴	10.5
6392	06-112 08:49	2	1	111	2	7	0	15	8	11	0	0	0	0	1	0	0	0	0	0	1	4	4.02380	166.7	-38.8	5.18683	205	2	-34	36.7	2.0	1.3-10 ⁻¹⁵	12.5
6393	06-116 21:11	3	3	39	19	23	3	6	5	6	0	1	42	0	1	0	0	0	0	0	1	4	4.00325	166.8	-39.2	5.20046	112	258	-11	34.6	1.6	1.7-10 ⁻¹³	6.0
6394	06-118 16:34	3	3	42	19	24	10	7	6	6	0	1	42	0	1	0	0	0	0	0	1	4	3.99543	166.9	-39.3	5.20555	116	261	-14	22.7	1.6	7.9-10 ⁻¹³	6.0
6395	06-118 19:52	3	3	51	21	27	6	5	5	6	0	1	46	0	1	0	0	0	0	0	1	4	3.99487	166.9	-39.3	5.20591	129	272	-22	43.7	1.6	1.9-10 ⁻¹³	6.0
6396	06-120 07:42	3	3	39	19	24	5	6	5	6	0	1	44	0	1	0	0	0	0	0	1	4	3.98814	166.9	-39.5	5.21023	113	259	-12	34.6	1.6	1.9-10 ⁻¹³	6.0
6397	06-132 07:12	3	2	79	12	20	7	9	6	7	0	1	37	0	1	0	0	0	0	0	1	4	3.93366	167.3	-40.6	5.24378	179	326	-41	14.1	1.9	1.2-10 ⁻¹²	10.5
6398	06-137 10:15	2	1	206	5	9	0	8	8	7	0	0	0	0	1	0	0	0	0	0	1	4	3.91002	167.5	-41.0	5.25753	359	146	52	28.0	1.6	7.4-10 ⁻¹⁵	6.0
6399	06-137 19:25	3	2	58	14	21	8	8	6	6	0	1	39	0	1	0	0	0	0	0	1	4	3.90828	167.5	-41.1	5.25852	151	292	-34	21.4	1.9	4.2-10 ⁻¹³	10.5
6400	06-139 02:15	3	3	35	19	22	5	7	9	6	0	1	39	0	1	0	0	0	0	0	1	4	3.90248	167.5	-41.2	5.26182	120	262	-16	16.9	1.6	2.0-10 ⁻¹²	6.0
6401	06-141 00:15	2	2	15	12	6	3	8	15	0	1	1	39	31	1	0	0	0	0	0	1	4	3.89317	167.6	-41.4	5.26704	92	242	2	21.4	1.9	3.7-10 ⁻¹⁴	10.5
6402	06-141 17:10	3	1	108	7	10	1	8	6	6	0	1	0	0	1	0	0	0	0	0	1	4	3.89025	167.6	-41.4	5.26866	223	18	-27	28.0	1.6	1.2-10 ⁻¹⁴	6.0
6403	06-141 20:06	3	2	129	9	13	1	7	8	6	0	1	0	0	1	0	0	0	0	0	1	4	3.88967	167.6	-41.4	5.26899	253	42	-8	35.4	1.6	1.3-10 ⁻¹⁴	6.0
6404	06-157 01:25	1	1	79	2	2	0	11	15	0	1	0	39	31	1	0	0	0	0	0	1	4	3.81747	168.1	-42.9	5.30691	194	347	-41	11.8	11.8	2.8-10 ⁻¹⁴	5888.3
6405	06-158 11:56	3	3	53	21	27	12	5	5	5	0	1	46	0	1	0	0	0	0	0	1	4	3.81086	168.2	-43.0	5.31017	158	299	-39	43.7	1.6	1.9-10 ⁻¹³	6.0
6406	06-165 22:45	3	4	32	27	49	25	4	5	0	1	46	0	1	0	0	0	0	0	0	1	4	3.77454	168.5	-43.8	5.32751	134	276	-27	21.4	1.9	2.0-10 ⁻¹¹	10.5
6407	06-172 18:16	3	3	35	23	29	8	5	4	5	0	1	46	0	1	0	0	0	0	0	1	4	3.74082	168.8	-44.4	5.34272	141	282	-31	48.9	1.6	2.3-10 ⁻¹³	6.0
6408	06-176 21:48	2	1	12	5	9	0	8	7	0	0	4	22	1	0	0	0	0	0	0	1	4	3.72039	168.9	-44.9	5.35153	110	257	-11	28.0	1.6	7.4-10 ⁻¹⁵	6.0
6409	06-184 16:06	2	2	239	9	4	10	15	0	1	0	1	38	31	1	0	0	0	0	0	1	4	3.68162	169.3	-45.7	5.36741	75	236	14	10.4	1.9	3.8-10 ⁻¹³	10.5
6410	06-192 21:07	3	2	32	13	21	3	8	6	6	0	1	40	0	1	0	0	0	0	0	1	4	3.63976	169.6	-46.5	5.38339	149	294	-36	21.4	1.9	3.6-10 ⁻¹³	10.5
6411	06-206 22:08	3	3	14	22	27	23	5	5	5	0	1	46	0	1	0	0	0	0	0	1	4	3.56738	170.3	-48.0	5.40823	134	283	-28	43.7	1.6	2.2-10 ⁻¹³	6.0
6412	06-210 16:33	3	4	1	25	49	12	9	5	5	0	1	46	0	1	0	0	0	0	0	1	4	3.54439	170.5	-48.5	5.41439	117	269	-16	7.8	1.9	2.9-10 ⁻¹⁰	10.5
6413	06-211 06:40	2	1	182	3	2	0	8	7	0	1	0	41	31	1	0	0	0	0	0	1	4	3.54440	170.5	-48.5	5.41540	12	177	53	14.1	1.9	2.1-10 ⁻¹⁴	10.5
6414	06-215 23:39	3	2	56	10	19	2	7	7	7	0	1	36	0	1	0	0	0	0	0	1	4	3.51993	170.8	-49.0	5.42267	199	5	-40	34.1	1.9	2.9-10 ⁻¹⁴	10.5
6415	06-217 08:19	2	1	2	7	11	0	8	8	6	0	0	0	0	1	0	0	0	0	0	1	4	3.51262	170.8	-49.2	5.42477	124	276	-21	28.0	1.6	1.5-10 ⁻¹⁴	6.0
6416	06-223 21:49	2	1	208	6	5	6	10	15	0	1	43	31	1	0	0	0	0	0	0	1	4	3.47715	171.2	-50.0	5.43450	59	233	26	10.4	1.9	1.1-10 ⁻¹³	10.5
6417	06-228 19:44	3	3	205	19	24	11	10	9	6	0	1	42	0	1	0	0	0	0	0	1	4	3.45081	171.5	-50.5	5.44123	60	235	26	10.0	1.6	1.1-10 ⁻¹¹	6.1
6418	06-228 22:13	2	1	4	5	9	0	8	8	7	0	0	0	0	1	0	0	0	0	0	1	4	3.45013	171.5	-50.5	5.44140	137	291	-29	28.0	1.6	7.4-10 ⁻¹⁵	6.0
6419	06-238 08:40	3	2	254	13	21	1	8	5	6	0	1	39	0	1	0	0	0	0	0	1	4	3.39886	172.1	-51.7	5.45331	139	296	-30	21.4	1.9	3.6-10 ⁻¹³	10.5
6420	06-254 09:14	3	2	217	12	19	1	8	6	7	0	1	37	0	1	0	0	0	0	0	1	4	3.30884	173.2	-53.6	5.47063	100	272	-3	21.4	1.9	2.0-10 ⁻¹³	10.5
6421	06-263 18:27	3	3	199	20	24	11	5	6	6	0	1	43	0	1	0	0	0	0	0	1	4	3.25543	174.0	-54.8	5.47883	83	263	9	35.4	1.6	2.4-10 ⁻¹³	6.0
6422	06-266 05:23	2	1	4	2	6	1	12	15	0	1	1	31	1	0	0	0	0	0	0	1	4	3.24178	174.2	-55.2	5.48069	173	348	-40	2.3	2.0	8.9-10 ⁻¹²	12.5
6423	06-266 06:35	3	4	230	29	49	10	11	9	6	0	1	31	0	1	0	0	0	0	0	1	4	3.24106	174.2	-55.2	5.48079	131	301	-24	2.5	1.9	2.3-10 ⁻⁰⁸	10.5
6424	06-267 08:17	2	2	165	11	10	8	10	15	0	1	39	31	1	0	0	0	0	0	0	1	4	3.23529	174.3	-55.3	5.48154	39	230	38	10.4	1.9	6.1-10 ⁻¹³	10.5
6425	06-268 21:29	3	2	222	15	23	9	9	7	6	0	1	41	0	1	0	0	0	0	0	1	4	3.22591	174.4									

No.	IMP.DATE	C L N	AR	S E C	IA	EA	CA	IT	ET	E T	E I C	I I C	E I C	P E T	PA	P E T	C C P	C C P	HV	R	LON	LAT	D _{rup}	ROT	S _{LON}	S _{LAT}	V	VEF	M	MEF
6436	07-130 08:34	3	2	137	11	19	1	7	5	7	0	0	0	0	1	37	0	1	4	1.76099	314.6	-62.5	5.26218	41	59	32	34.1	1.9	3.3·10 ⁻¹⁴	10.5
6437	07-132 00:14	2	1	123	4	2	0	7	15	12	0	0	0	0	1	31	1	4	1.75028	315.4	-61.8	5.25860	23	99	40	34.1	1.9	1.1·10 ⁻¹⁵	10.5	
6438	07-132 11:44	2	1	162	4	8	0	8	8	8	0	0	0	0	2	24	1	4	1.74800	315.6	-61.6	5.25784	78	92	12	28.0	1.6	5.2·10 ⁻¹⁵	6.0	
6439	07-136 03:34	2	1	154	5	8	0	8	7	0	0	0	0	0	1	25	1	4	1.72537	317.1	-59.9	5.25023	69	87	17	28.0	1.6	6.2·10 ⁻¹⁵	6.0	
6440	07-146 04:00	2	1	165	6	1	0	0	15	12	0	0	0	0	2	25	1	4	1.66706	320.6	-55.2	5.23056	91	108	3	11.8	11.8	4.4·10 ⁻¹⁴	5858.3	
6441	07-147 06:12	2	1	141	5	9	0	8	8	7	0	0	0	0	3	25	1	4	1.66071	320.9	-54.6	5.22842	58	83	25	28.0	1.6	7.4·10 ⁻¹⁵	6.0	
6442	07-149 00:16	1	2	106	10	7	0	9	15	0	1	0	0	0	4	31	1	4	1.65091	321.5	-53.8	5.22511	9	29	46	14.1	1.9	1.5·10 ⁻¹³	10.5	
6443	07-149 22:51	2	3	100	20	7	2	10	15	0	1	0	0	0	1	47	29	1	4	1.64606	321.7	-53.3	5.22348	2	18	47	5.0	1.9	1.1·10 ⁻¹¹	10.5
6444	07-152 04:50	1	1	79	2	3	0	15	2	0	1	0	0	0	4	46	13	1	4	1.63439	322.4	-52.2	5.21955	333	339	42	11.8	11.8	3.2·10 ⁻¹⁴	5858.3
6445	07-154 06:49	1	1	62	5	1	0	7	15	0	1	0	0	0	3	31	1	4	1.62218	323.0	-51.0	5.21545	311	316	31	34.1	1.9	1.1·10 ⁻¹⁵	10.5	
6446	07-157 06:41	1	1	173	2	4	0	15	15	0	1	0	0	0	38	30	1	4	1.60630	323.8	-49.4	5.21014	109	125	-8	11.8	11.8	3.6·10 ⁻¹⁴	5858.3	
6447	07-157 13:06	3	1	88	6	10	12	7	7	6	0	1	0	0	3	25	1	4	1.60499	323.9	-49.3	5.20970	350	2	48	43.7	1.6	2.0·10 ⁻¹⁵	6.0	
6448	07-157 20:10	2	1	141	2	6	0	15	8	10	0	0	0	0	3	31	1	4	1.60369	324.4	-48.1	5.20605	74	92	21	36.7	2.0	1.1·10 ⁻¹⁵	12.5	
6449	07-157 21:26	2	1	52	1	1	0	15	11	0	0	0	0	0	3	31	1	4	1.60304	324.0	-49.1	5.20905	299	307	24	11.8	11.8	1.9·10 ⁻¹⁴	5858.3	
6450	07-158 03:43	2	1	11	5	9	0	8	7	7	0	0	0	0	34	0	1	5	1.60174	324.0	-49.0	5.20862	242	264	-14	34.6	1.6	3.8·10 ⁻¹⁵	6.0	
6451	07-158 09:31	1	1	115	2	1	0	9	15	0	1	0	0	0	3	31	1	4	1.60044	324.1	-48.8	5.20819	28	58	42	11.8	11.8	2.4·10 ⁻¹⁴	5858.3	
6452	07-159 01:50	2	1	134	5	10	0	7	7	8	0	0	0	0	3	31	1	4	1.59721	324.3	-48.5	5.20712	55	85	27	43.7	1.6	1.7·10 ⁻¹⁵	6.0	
6453	07-159 17:01	1	1	147	6	7	1	15	15	1	0	0	0	0	4	35	1	4	1.59400	324.4	-48.1	5.20605	74	100	15	34.1	1.9	1.5·10 ⁻¹⁵	10.5	
6454	07-161 05:17	2	1	110	7	7	0	15	15	0	1	1	0	0	37	31	1	4	1.58636	324.8	-47.3	5.20353	23	52	45	21.4	1.9	2.0·10 ⁻¹⁴	10.5	
6455	07-162 17:23	2	4	165	26	22	3	8	12	0	1	1	0	0	63	29	1	4	1.57883	325.2	-46.5	5.20105	101	121	-3	10.4	1.9	2.5·10 ⁻¹¹	10.5	
6456	07-164 19:50	1	1	177	2	15	0	9	15	0	1	0	0	0	47	31	1	4	1.56834	325.7	-45.3	5.19762	119	135	-16	11.8	11.8	2.2·10 ⁻¹³	5858.3	
6457	07-165 00:47	3	1	247	7	12	5	7	7	6	0	1	0	0	3	31	1	4	1.56712	325.8	-45.1	5.19723	218	245	-29	43.7	1.6	3.3·10 ⁻¹⁵	6.0	
6458	07-165 12:02	3	3	11	19	24	14	6	7	6	0	1	0	0	4	43	0	1	4	1.56469	325.9	-44.8	5.19644	246	269	-12	25.9	1.6	5.0·10 ⁻¹³	6.0
6459	07-165 14:29	3	1	8	5	9	1	7	7	7	0	1	0	0	3	31	1	4	1.56469	325.9	-44.8	5.19644	242	266	-15	43.7	1.6	1.4·10 ⁻¹⁵	6.0	
6460	07-165 18:37	2	1	224	1	6	0	15	9	12	0	0	0	0	3	31	1	4	1.56348	325.9	-44.7	5.19605	186	208	-39	31.3	2.0	1.5·10 ⁻¹⁵	12.5	
6461	07-165 19:56	1	1	133	5	13	0	9	15	15	1	0	0	0	58	28	1	4	1.56348	325.9	-44.7	5.19605	58	90	26	14.1	1.9	1.8·10 ⁻¹³	10.5	
6462	07-165 22:29	2	1	169	5	1	0	6	2	12	0	0	0	0	3	31	1	4	1.56288	326.0	-44.6	5.19585	109	127	-9	43.5	1.9	3.9·10 ⁻¹⁶	10.5	
6463	07-169 23:20	1	1	81	3	10	0	7	15	0	1	0	0	0	4	46	31	1	4	1.54399	326.9	-42.3	5.18979	348	2	50	34.1	1.9	3.4·10 ⁻¹⁵	10.5
6464	07-170 07:57	1	1	196	3	9	0	9	15	15	1	0	1	0	1	31	1	4	1.54226	326.9	-42.0	5.18925	150	163	-33	5.7	2.1	8.2·10 ⁻¹³	15.7	
6465	07-171 00:28	3	2	134	10	19	8	7	5	6	0	1	0	0	3	31	1	4	1.53883	327.1	-41.6	5.18816	63	95	23	34.1	1.9	2.9·10 ⁻¹⁴	10.5	
6466	07-173 19:19	2	1	206	3	8	0	8	8	0	0	0	0	0	2	31	1	4	1.52651	327.7	-39.9	5.18432	166	183	-40	28.0	1.6	4.5·10 ⁻¹⁵	6.0	
6467	07-175 13:26	3	1	249	5	10	3	7	7	6	0	1	0	0	2	25	1	4	1.51890	328.1	-38.9	5.18198	227	257	-25	43.7	1.6	1.7·10 ⁻¹⁵	6.0	
6468	07-176 23:13	1	1	176	2	2	0	9	15	0	1	0	0	0	3	31	1	4	1.51304	328.3	-38.0	5.18021	126	141	-21	11.8	11.8	2.8·10 ⁻¹⁴	5858.3	
6469	07-177 09:20	2	1	102	5	14	0	15	12	12	0	0	0	0	15	1	1	4	1.51094	328.5	-37.7	5.17958	22	56	48	2.0	1.9	1.0·10 ⁻¹⁰	10.5	
6470	07-177 16:13	1	1	179	4	8	0	5	7	0	1	0	0	0	45	31	1	4	1.50989	328.5	-37.5	5.17926	130	145	-24	62.6	1.6	2.3·10 ⁻¹⁶	6.0	
6471	07-177 19:34	3	3	49	19	24	4	5	5	5	0	1	0	0	1	43	0	1	4	1.50937	328.5	-37.5	5.17911	308	316	31	43.7	1.6	7.3·10 ⁻¹⁴	6.0
6472	07-179 07:57	2	1	9	6	11	0	7	7	6	0	0	0	0	2	31	1	4	1.50319	328.8	-36.5	5.17728	252	277	-9	43.7	1.6	2.3·10 ⁻¹⁵	6.0	
6473	07-181 16:16	1	1	48	2	7	0	9	15	15	1	0	0	0	4	46	30	1	4	1.49370	329.3	-35.0	5.17453	309	316	32	2.3	2.0	1.1·10 ⁻¹¹	12.5
6474	07-182 14:46	1	1	47	3	1	0	7	0	0	1	0	0	0	7	1	1	4	1.49031	329.5	-34.5	5.17358	308	316	31	34.1	1.9	7.7·10 ⁻¹⁶	10.5	
6475	07-182 21:34	1	1	195	4	11	0	5	15	0	1	0	0	0	47	31	1	4	1.48886	329.5	-34.2	5.17317	156	171	-38	70.0	1.9	2.5·10 ⁻¹⁶	10.5	
6476	07-183 14:00	3	4	67	24	30	17	8	6	5	0	1	0	0	1	47	0	1	4	1.48647	329.6	-33.8	5.17251	337	346	49	19.5	1.6	1.1·10 ⁻¹¹	6.0
6477	07-184 04:39	1	1	118	3	10	0	9	15	15	1	0	0	0	19	31	1	4	1.48411	329.8	-33.4	5.17185	49	87	34	14.1	1.9	8.0·10 ⁻¹⁴	10.5	
6478	07-184 13:14	1	1	120	3	8	0	9	15	0	1	0	0	0	4	31	1	4	1.48270	329.8	-33.2	5.17147	52	90	32	5.7	2.1	6.9·10 ⁻¹³	15.7	
6479	07-184 13:46	3	2	75	8	12	5	7	7	6	0	1	0	0	3	31	1	4	1.48270	329.8	-33.2	5.17147	348	4	53	43.7	1.6	4.0·10 ⁻¹⁵	6.0	
6480	07-185 02:29	2	1	126	6	5	1	7	15	15	1	1	0	0	3	31	1	4	1.48085	329.9	-32.9	5.17096	61	97	26	34.1	1.9	2.3·10 ⁻¹⁵	10.5	
6481	07-185 03:01	2	1	110	4	15	0	7	0	8	0	0	0	0	3	31	1	4	1.48038	330.0	-32.8	5.17084	38	78	41	34.1	1.9	8.4·10 ⁻¹⁵	10.5	
6482	07-185 03:19	1	1	111	5	1	0	11	15	0	1	0	0	0	41	28	1	4	1.48038	330.0	-32.8	5.17084	40	79	40	7.8	1.9	1.1·10 ⁻¹³	10.5	
6483	07-185 05:28	3	1	153	1	5	1	15	7	11	0	1	0	0	3	31	1	4	1.48038	330.0	-32.8	5.17084	99	122	-2	56.0	2.0	1.3·10 ⁻¹⁶	12.5	
6484	07-185 08:52	2	2	28	9	13	0	7	7	6	0	0	0	0	36	0	1	4	1.47992	330.0	-32.7	5.17071	283							

No.	IMP DATE	C L N	AR S E C	IA	EA	CA	IT	ET	E I T	E I C	I I C	PA	P E T	E V D	I C P	E C P	C C P	P P	HV	R	LON	LAT	D _{dup}	ROT	S _{LON}	S _{LAT}	V	V _{EF}	M	MEF
6491	07-187 16:13	2	221	8	13	0	7	7	6	0	0	35	0	1	0	0	0	1	4	1.47136	330.4	-31.1	5.16845	196	225	-43	43.7	1.6	4.7-10 ⁻¹⁵	6.0
6492	07-189 03:45	3	115	20	26	3	7	6	6	0	1	45	0	1	0	0	0	1	4	1.46615	330.7	-30.1	5.16713	48	88	35	22.7	1.6	1.5-10 ⁻¹²	6.0
6493	07-189 16:45	1	181	2	7	0	15	15	15	1	0	31	1	0	0	0	0	1	4	1.46445	330.8	-29.8	5.16671	141	154	-32	7.7	2.0	1.9-10 ⁻¹³	12.5
6494	07-190 00:00	1	154	1	1	0	15	14	15	1	0	3	31	1	0	0	0	1	4	1.46319	330.8	-29.6	5.16640	103	125	-6	11.8	11.8	1.9-10 ⁻¹⁴	5858.3
6495	07-190 04:34	2	66	6	9	1	8	15	0	1	1	44	30	1	0	0	0	1	4	1.46277	330.9	-29.5	5.16630	339	349	52	7.0	2.9	6.8-10 ⁻¹³	46.7
6496	07-191 01:13	2	46	14	21	0	7	7	7	0	0	41	0	1	0	0	0	1	4	1.45987	331.0	-28.9	5.16560	312	318	35	34.1	1.9	8.3-10 ⁻¹⁴	10.5
6497	07-191 15:49	1	156	1	1	0	15	14	15	1	0	31	1	0	0	0	0	1	4	1.45783	331.1	-28.5	5.16512	107	128	-9	11.8	11.8	1.9-10 ⁻¹⁴	5858.3
6498	07-191 20:50	2	159	8	13	0	7	7	6	0	1	36	0	1	0	0	0	1	4	1.45726	331.2	-28.4	5.16502	111	130	-12	43.7	1.6	4.7-10 ⁻¹⁵	6.0
6499	07-193 08:01	2	183	8	15	9	7	0	6	0	1	3	31	1	0	0	0	1	4	1.45266	331.4	-27.4	5.16394	146	159	-36	34.1	1.9	1.7-10 ⁻¹⁴	10.5
6500	07-195 12:39	1	235	1	12	0	15	15	15	1	0	45	30	1	0	0	0	1	4	1.44582	331.8	-25.8	5.16249	220	254	-33	11.8	11.8	1.2-10 ⁻¹³	5858.3
6501	07-195 16:32	2	27	12	0	7	7	7	7	0	0	2	31	1	0	0	0	1	4	1.44545	331.8	-25.8	5.16242	288	300	17	43.7	1.6	3.3-10 ⁻¹⁵	6.0
6502	07-196 10:09	2	96	12	20	0	7	6	6	0	1	38	0	1	0	0	0	1	4	1.44326	332.0	-25.2	5.16198	25	66	51	34.1	1.9	5.2-10 ⁻¹⁴	10.5
6503	07-196 22:21	2	50	11	7	12	9	15	15	1	1	26	1	0	0	0	0	1	4	1.44183	332.0	-24.9	5.16171	321	325	43	14.1	1.9	1.7-10 ⁻¹³	10.5
6504	07-197 03:32	1	166	1	11	0	15	10	15	1	0	13	31	1	0	0	0	1	4	1.44112	332.1	-24.7	5.16157	124	139	-23	11.8	11.8	1.0-10 ⁻¹³	5858.3
6505	07-197 12:26	0	29	3	7	0	10	6	10	0	1	18	21	1	0	0	0	1	4	1.44006	332.1	-24.5	5.16138	292	302	20	27.0	2.3	4.3-10 ⁻¹⁵	19.8
6506	07-197 14:55	2	255	7	12	0	7	7	7	0	0	3	26	1	0	0	0	1	4	1.44006	332.1	-24.5	5.16138	250	277	-12	43.7	1.6	3.3-10 ⁻¹⁵	6.0
6507	07-197 17:23	3	40	9	14	1	7	7	6	0	1	36	0	1	0	0	0	1	4	1.43971	332.2	-24.4	5.16131	307	313	33	43.7	1.6	6.4-10 ⁻¹⁵	6.0
6508	07-198 11:59	1	183	3	9	0	10	15	15	1	0	19	30	1	0	0	0	1	4	1.43764	332.3	-23.9	5.16094	149	161	-39	4.9	1.7	1.3-10 ⁻¹²	7.7
6509	07-198 14:15	3	40	8	13	6	7	7	7	0	1	35	0	1	0	0	0	1	4	1.43461	332.5	-23.1	5.16043	308	313	34	43.7	1.6	4.7-10 ⁻¹⁵	6.0
6510	07-199 22:18	0	61	2	1	0	15	0	1	1	0	21	2	1	0	0	0	1	4	1.43363	332.5	-22.8	5.16028	338	344	54	11.8	11.8	2.4-10 ⁻¹⁴	5858.3
6511	07-200 08:56	2	74	9	11	11	11	15	0	1	1	6	30	1	0	0	0	1	4	1.43265	332.6	-22.6	5.16012	357	16	60	4.2	1.6	7.9-10 ⁻¹²	6.0
6512	07-200 15:12	1	174	2	8	0	15	15	0	1	0	47	31	1	0	0	0	1	4	1.43169	332.7	-22.3	5.15998	137	149	-33	2.3	2.0	1.3-10 ⁻¹¹	12.5
6513	07-200 16:33	2	37	20	15	5	7	6	0	1	0	27	1	0	0	0	0	1	4	1.43169	332.7	-22.3	5.15998	305	310	31	14.1	1.9	2.1-10 ⁻¹²	10.5
6514	07-201 23:34	2	65	10	9	5	9	15	0	1	1	39	30	1	0	0	0	1	4	1.42855	332.9	-21.4	5.15954	345	354	58	14.1	1.9	2.2-10 ⁻¹³	10.5
6515	07-203 22:58	1	41	5	3	0	8	15	0	1	0	36	30	1	0	0	0	1	4	1.42379	333.2	-20.0	5.15899	312	314	38	21.4	1.9	7.3-10 ⁻¹⁵	10.5
6516	07-205 08:36	2	117	5	10	0	8	7	7	0	0	3	31	1	0	0	0	1	4	1.42071	333.4	-19.0	5.15872	60	100	28	34.6	1.6	4.5-10 ⁻¹⁵	6.0
6517	07-205 19:39	1	127	1	1	0	15	0	1	0	0	3	31	1	0	0	0	1	4	1.41962	333.5	-18.7	5.15864	74	108	16	11.8	11.8	1.9-10 ⁻¹⁴	5858.3
6518	07-205 20:27	2	43	5	9	0	20	15	15	0	3	26	1	0	0	0	0	1	4	1.41962	333.5	-18.7	5.15864	316	316	41	43.7	1.6	1.4-10 ⁻¹⁵	6.0
6519	07-206 11:10	1	121	2	10	0	8	15	15	1	0	47	31	1	0	0	0	1	4	1.41830	333.6	-18.2	5.15856	66	103	23	11.8	11.8	7.3-10 ⁻¹⁴	5858.3
6520	07-206 23:24	1	198	4	13	0	6	14	0	1	0	47	31	1	0	0	0	1	4	1.41726	333.7	-17.9	5.15851	175	193	-53	43.5	1.9	2.4-10 ⁻¹⁵	10.5
6521	07-207 07:24	2	60	5	10	0	8	7	7	0	0	2	26	1	0	0	0	1	4	1.41650	333.7	-17.6	5.15848	341	343	59	34.6	1.6	4.5-10 ⁻¹⁵	6.0
6522	07-207 08:16	1	172	1	8	0	15	14	15	1	0	19	31	1	0	0	0	1	4	1.41650	333.7	-17.6	5.15848	138	146	-36	11.8	11.8	6.1-10 ⁻¹⁴	5858.3
6523	07-209 18:08	1	174	1	9	0	15	15	15	1	0	47	31	1	0	0	0	1	4	1.41171	334.1	-15.8	5.15846	143	148	-39	11.8	11.8	7.3-10 ⁻¹⁴	5858.3
6524	07-209 22:45	2	57	2	7	0	15	7	10	0	1	3	31	1	0	0	0	1	4	1.41148	334.1	-15.7	5.15846	338	336	58	56.0	2.0	2.2-10 ⁻¹⁶	12.5
6525	07-210 09:13	3	65	5	9	3	7	7	6	0	1	2	26	1	0	0	0	1	4	1.41059	334.2	-15.4	5.15850	350	356	64	43.7	1.6	1.4-10 ⁻¹⁵	6.0
6526	07-210 20:30	2	62	3	11	0	8	5	13	0	0	24	0	1	0	0	0	1	4	1.40994	334.3	-15.1	5.15853	346	348	63	21.4	1.9	2.1-10 ⁻¹⁴	10.5
6527	07-211 07:09	1	200	2	11	0	15	15	0	1	0	63	31	1	0	0	0	1	4	1.40909	334.4	-14.7	5.15859	180	200	-56	11.8	11.8	1.2-10 ⁻¹³	5858.3
6528	07-211 07:11	1	205	1	10	0	15	15	15	1	0	46	31	1	0	0	0	1	4	1.40909	334.4	-14.7	5.15859	187	213	-55	11.8	11.8	8.6-10 ⁻¹⁴	5858.3
6529	07-212 03:24	0	42	2	7	0	15	10	14	1	0	45	28	1	0	0	0	1	4	1.40765	334.5	-14.1	5.15872	318	312	44	26.5	2.0	3.9-10 ⁻¹⁵	12.5
6530	07-212 11:35	2	123	1	5	0	15	10	10	0	0	4	31	1	0	0	0	1	4	1.40725	334.5	-13.9	5.15876	72	106	18	26.5	2.0	2.2-10 ⁻¹⁵	12.5
6531	07-212 21:29	1	53	2	11	0	15	15	0	1	0	46	31	1	0	0	0	1	4	1.40646	334.6	-13.5	5.15886	334	326	57	11.8	11.8	1.2-10 ⁻¹³	5858.3
6532	07-213 15:25	1	249	1	11	0	15	15	0	1	0	46	31	1	0	0	0	1	4	1.40535	334.7	-13.0	5.15905	250	276	-15	11.8	11.8	1.0-10 ⁻¹³	5858.3
6533	07-213 19:17	2	23	6	11	0	7	7	7	0	0	3	26	1	0	0	0	1	4	1.40516	334.7	-12.9	5.15908	292	295	22	43.7	1.6	2.3-10 ⁻¹⁵	6.0
6534	07-213 19:30	1	25	2	7	0	15	13	15	1	0	45	31	1	0	0	0	1	4	1.40516	334.7	-12.9	5.15908	295	296	24	7.7	2.0	1.9-10 ⁻¹³	12.5
6535	07-213 22:26	2	146	22	1	17	9	15	15	0	1	47	0	1	0	0	0	1	4	1.40498	334.8	-12.8	5.15911	105	121	-10	7.8	1.9	1.4-10 ⁻¹²	10.5
6536	07-214 06:23	1	14	2	11	0	15	15	0	1	0	31	1	0	0	0	0	1	4	1.40444	334.8	-12.5	5.15921	280	289	11	11.8	11.8	1.2-10 ⁻¹³	5858.3
6537	07-214 09:44	3	85	27	49	18	6	5	5	0	1	47	0	1	0	0	0	1	4	1.40426	334.8	-12.5	5.15925	20	65	62	21.4	1.9	2.0-10 ⁻¹¹	10.5
6538	07-214 11:07	2	123	6	11	0	10	3	14	0	0	56	7	1	0	0	0													

No.	IMP DATE	C L N	AR	S E C	IA	EA	CA	IT	ET	E I T	E I C	E I C	P E T	P E T	E V D	E I C P	E I C P	C C P	C C P	H V	R	LON	LAT	D _{dup}	ROT	S _{LON}	S _{LAT}	V	V _{EF}	M	MEF	
6546	07-216 01:58	1	1	123	2	2	0	15	15	15	1	0	3	31	1	0	0	0	0	1	4	1.40208	335.1	-11.3	5.15978	73	106	17	11.8	11.8	2.8-10 ⁻¹⁴	5858.3
6547	07-216 07:21	1	1	44	2	1	0	15	8	0	0	23	31	1	0	0	0	0	1	4	1.40176	335.1	-11.1	5.15987	322	310	49	11.8	11.8	2.4-10 ⁻¹⁴	5858.3	
6548	07-216 08:51	2	1	161	1	5	0	15	7	11	0	0	3	31	1	0	0	0	1	4	1.40176	335.1	-11.1	5.15987	127	330	-30	56.0	2.0	1.3-10 ⁻¹⁶	12.5	
6549	07-216 12:22	1	1	241	1	9	0	15	15	0	1	0	46	31	1	0	0	0	1	4	1.40145	335.2	-10.9	5.15997	240	270	-24	11.8	11.8	7.3-10 ⁻¹⁴	5858.3	
6550	07-216 16:35	1	1	92	2	10	0	15	15	0	1	0	47	31	1	0	0	0	1	4	1.40130	335.2	-10.8	5.16002	30	80	56	11.8	11.8	1.1-10 ⁻¹³	5858.3	
6551	07-216 16:59	3	4	43	28	49	26	6	5	0	1	47	0	1	0	0	0	0	1	4	1.40130	335.2	-10.8	5.16002	321	308	48	10.4	1.9	2.2-10 ⁻¹⁰	10.5	
6552	07-216 22:10	3	4	83	27	49	10	6	5	0	1	47	0	1	0	0	0	0	1	4	1.40100	335.2	-10.6	5.16012	17	63	65	21.4	1.9	2.0-10 ⁻¹¹	10.5	
6553	07-216 23:15	2	1	117	5	9	0	7	7	8	0	0	31	1	0	0	0	0	1	4	1.40100	335.2	-10.6	5.16012	65	103	25	43.7	1.6	1.4-10 ⁻¹⁵	6.0	
6554	07-217 01:22	2	1	19	6	12	0	7	7	7	0	0	34	0	1	0	0	0	1	4	1.40085	335.2	-10.5	5.16018	287	290	18	43.7	1.6	2.7-10 ⁻¹⁵	6.0	
6555	07-217 03:29	1	1	173	2	9	0	9	15	0	1	0	47	31	1	0	0	0	1	4	1.40070	335.3	-10.4	5.16023	144	142	-44	11.8	11.8	8.8-10 ⁻¹⁴	5858.3	
6556	07-217 09:12	2	1	148	2	7	0	15	8	10	0	0	4	31	1	0	0	0	1	4	1.40041	335.3	-10.3	5.16034	109	120	-14	36.7	2.0	1.3-10 ⁻¹⁵	12.5	
6557	07-217 12:15	2	1	155	3	7	0	9	8	9	0	0	3	31	1	0	0	0	1	4	1.40026	335.3	-10.2	5.16039	119	124	-23	22.7	1.6	7.8-10 ⁻¹⁵	6.0	
6558	07-218 06:46	2	1	47	4	10	0	8	7	8	0	0	2	26	1	0	0	0	1	4	1.39943	335.4	-9.6	5.16074	327	310	54	34.6	1.6	3.8-10 ⁻¹⁵	6.0	
6559	07-218 12:54	3	2	96	12	20	1	7	7	6	0	1	39	0	1	0	0	0	1	4	1.39917	335.5	-9.4	5.16087	36	87	51	34.1	1.9	5.2-10 ⁻¹⁴	10.5	
6560	07-218 21:05	1	1	104	2	15	0	9	0	0	1	0	47	31	1	0	0	0	1	4	1.39878	335.5	-9.2	5.16106	47	94	41	11.8	11.8	2.2-10 ⁻¹³	5858.3	
6561	07-218 21:31	1	1	182	2	15	0	9	0	0	1	0	47	31	1	0	0	0	1	4	1.39878	335.5	-9.2	5.16106	157	153	-55	11.8	11.8	2.2-10 ⁻¹³	5858.3	
6562	07-219 14:12	2	1	3	2	1	0	15	0	13	0	0	46	0	1	0	0	0	1	4	1.39816	335.6	-8.7	5.16140	266	280	-1	11.8	11.8	2.4-10 ⁻¹⁴	5858.3	
6563	07-220 01:27	1	1	62	3	13	0	7	15	0	1	0	3	31	1	0	0	0	1	4	1.39768	335.7	-8.3	5.16168	348	337	72	34.1	1.9	5.5-10 ⁻¹⁵	10.5	
6564	07-220 19:30	2	1	155	5	10	0	7	7	8	0	0	2	31	1	0	0	0	1	4	1.39702	335.8	-7.8	5.16213	119	121	-25	43.7	1.6	1.7-10 ⁻¹⁵	6.0	
6565	07-221 07:13	2	1	109	4	9	0	7	7	7	0	0	3	31	1	0	0	0	1	4	1.39660	335.9	-7.4	5.16245	55	98	35	43.7	1.6	1.2-10 ⁻¹⁵	6.0	
6566	07-221 14:32	1	2	11	11	3	0	12	7	15	1	0	57	30	1	0	0	0	1	4	1.39640	335.9	-7.2	5.16261	277	282	8	5.0	1.9	1.7-10 ⁻¹²	10.5	
6567	07-221 16:36	3	1	143	3	7	1	7	7	8	0	1	2	31	1	0	0	0	1	4	1.39630	335.9	-7.1	5.16269	102	114	-9	43.7	1.6	7.2-10 ⁻¹⁶	6.0	
6568	07-221 22:53	3	2	60	13	21	9	8	9	6	0	1	39	0	1	0	0	0	1	4	1.39611	336.0	-7.0	5.16286	346	324	71	21.4	1.9	3.6-10 ⁻¹³	10.5	
6569	07-222 01:05	2	2	169	8	13	0	7	7	8	0	0	3	31	1	0	0	0	1	4	1.39602	336.0	-6.9	5.16295	139	129	-43	43.7	1.6	4.7-10 ⁻¹⁵	6.0	
6570	07-222 06:24	1	1	145	1	10	0	15	15	0	1	0	47	31	1	0	0	0	1	4	1.39583	336.0	-6.7	5.16312	105	114	-12	11.8	11.8	8.6-10 ⁻¹⁴	5858.3	
6571	07-222 17:12	1	1	199	1	3	0	15	15	0	1	0	5	30	1	0	0	0	1	4	1.39557	336.1	-6.4	5.16339	181	197	-68	11.8	11.8	8.6-10 ⁻¹⁴	5858.3	
6572	07-222 20:45	1	1	70	1	10	0	15	9	0	1	0	46	28	1	0	0	0	1	4	1.39548	336.1	-6.3	5.16348	360	11	79	11.8	11.8	2.6-10 ⁻¹⁴	5858.3	
6573	07-223 05:27	1	1	98	2	10	0	15	15	0	1	0	47	31	1	0	0	0	1	4	1.39524	336.2	-6.0	5.16376	39	93	50	11.8	11.8	1.1-10 ⁻¹³	5858.3	
6574	07-223 22:44	1	1	145	3	11	0	8	15	0	1	0	47	31	1	0	0	0	1	4	1.39478	336.3	-5.5	5.16433	105	112	-12	21.4	1.9	2.1-10 ⁻¹⁴	10.5	
6575	07-224 03:05	1	2	179	9	7	0	7	15	15	1	0	2	31	1	0	0	0	1	4	1.39464	336.3	-5.3	5.16453	152	136	-56	34.1	1.9	5.4-10 ⁻¹⁵	10.5	
6576	07-224 18:40	2	1	202	2	1	0	15	2	12	0	0	3	26	1	0	0	0	1	4	1.39431	336.4	-4.8	5.16505	185	207	-71	11.8	11.8	2.4-10 ⁻¹⁴	5858.3	
6577	07-224 19:25	2	2	179	11	20	0	7	7	6	0	0	40	0	1	0	0	0	1	4	1.39425	336.4	-4.8	5.16505	152	134	-56	34.1	1.9	4.4-10 ⁻¹⁴	10.5	
6578	07-224 21:20	2	1	133	3	7	0	8	7	9	0	0	3	31	1	0	0	0	1	4	1.39425	336.4	-4.7	5.16515	87	107	4	34.6	1.6	1.9-10 ⁻¹⁵	6.0	
6579	07-225 01:58	3	4	80	24	29	3	8	6	6	0	1	47	0	1	0	0	0	1	4	1.39419	336.5	-4.6	5.16526	13	73	75	19.5	1.6	9.0-10 ⁻¹²	6.0	
6580	07-225 11:42	2	2	70	10	14	2	10	15	15	1	1	60	31	1	0	0	0	1	4	1.39401	336.5	-4.4	5.16558	358	359	83	4.9	1.7	9.6-10 ⁻¹²	7.7	
6581	07-225 15:12	1	1	24	7	8	0	9	15	15	1	0	45	30	1	0	0	0	1	4	1.39391	336.5	-4.2	5.16580	294	281	24	14.1	1.9	1.1-10 ⁻¹³	10.5	
6582	07-225 15:20	2	2	69	10	10	7	8	15	0	1	1	6	30	1	0	0	0	1	4	1.39391	336.5	-4.2	5.16580	357	346	83	21.4	1.9	5.6-10 ⁻¹⁴	10.5	
6583	07-225 15:25	1	1	93	3	10	0	9	15	15	1	0	47	31	1	0	0	0	1	4	1.39391	336.5	-4.2	5.16580	31	92	59	14.1	1.9	8.0-10 ⁻¹⁴	10.5	
6584	07-225 21:24	1	1	159	2	11	0	15	15	0	1	0	47	31	1	0	0	0	1	4	1.39380	336.6	-4.0	5.16603	123	114	-30	11.8	11.8	1.2-10 ⁻¹³	5858.3	
6585	07-226 07:40	2	6	79	59	28	31	4	0	4	1	1	31	0	1	0	0	0	1	4	1.39366	336.6	-3.7	5.16637	10	75	78	43.5	1.9	8.1-10 ⁻¹²	10.5	
6586	07-226 07:42	1	1	84	5	6	0	6	15	15	1	0	63	31	1	0	0	0	1	4	1.39366	336.6	-3.7	5.16637	17	85	72	43.5	1.9	8.5-10 ⁻¹⁶	10.5	
6587	07-226 22:39	1	1	110	3	10	0	9	15	15	1	0	19	30	1	0	0	0	1	4	1.39345	336.7	-3.3	5.16696	54	100	37	14.1	1.9	8.0-10 ⁻¹⁴	10.5	
6588	07-227 07:52	2	2	100	13	11	13	10	15	0	1	1	2	26	1	0	0	0	1	4	1.39334	336.8	-3.0	5.16732	39	98	51	10.4	1.9	1.0-10 ⁻¹²	10.5	
6589	07-227 15:16	3	4	8	26	1	28	11	15	6	0	1	22	0	1	0	0	0	1	4	1.39325	336.8	-2.7	5.16769	269	275	0	2.5	1.9	1.1-10 ⁻¹⁰	10.5	
6590	07-227 19:25	1	1	180	2	9	0	15	15	15	1	0	19	30	1	0	0	0	1	4	1.39322	336.9	-2.6	5.16782	151	122	-57	11.8	11.8	8.8-10 ⁻¹⁴	5858.3	
6591	07-227 22:51	2	1	215	7	11	0	7	6	0	1	0	2	26	1	0	0	0	1	4	1.39319	336.9	-2.5	5.16794	200	248	-65	43.7	1.6	2.9-10 ⁻¹⁵	6.0	
6592	07-228 05:46	1	1	112	3	12	0	10	15	15	1	0	20	31	1	0	0	0	1	4</												

No.	IMP DATE	C L N	AR S E C	IA	EA	CA	IT	ET	E I T	E I C	E I C	PA	P E T	E V D	E I C P	E I C P	C C P	C C P	HV	R	LON	LAT	D _{rup}	ROT	S _{Lon}	S _{Lat}	V	VEF	M	MEF		
6601	07-229 20:42	1	1	91	5	12	0	6	15	15	1	0	47	31	1	0	0	0	1	4	1.39294	337.2	-1.1	5.16994	23	102	67	43.5	1.9	2.4	10 ⁻¹⁵	10.5
6602	07-229 22:52	1	1	105	2	11	0	15	15	15	1	0	46	31	1	0	0	0	1	4	1.39293	337.2	-1.0	5.17008	42	101	48	11.8	11.8	1.2	10 ⁻¹³	5858.3
6603	07-229 23:43	1	1	154	3	2	0	10	15	15	1	0	43	30	1	0	0	0	1	4	1.39293	337.2	-1.0	5.17008	111	104	-20	10.4	1.9	4.4	10 ⁻¹⁴	10.5
6604	07-230 22:00	1	1	209	1	10	0	15	14	15	1	0	46	31	1	0	0	0	1	4	1.39294	337.3	-0.3	5.17123	186	232	-80	11.8	11.8	8.6	10 ⁻¹⁴	5858.3
6605	07-231 01:40	1	1	103	4	3	0	9	15	15	1	0	45	30	1	0	0	0	1	4	1.39295	337.3	-0.2	5.17138	37	103	53	14.1	1.9	2.8	10 ⁻¹⁴	10.5
6606	07-231 11:29	3	2	72	13	21	3	7	6	0	1	0	40	0	1	0	0	0	1	4	1.39298	337.4	0.1	5.17183	352	242	81	34.1	1.9	7.1	10 ⁻¹⁴	10.5
6607	07-231 12:51	2	1	94	2	7	0	15	7	11	0	0	0	0	1	0	0	0	1	4	1.39299	337.4	0.2	5.17198	23	109	66	56.0	2.0	2.2	10 ⁻¹⁶	12.5
6608	07-232 05:46	1	1	154	4	2	0	8	15	15	1	0	4	26	1	0	0	0	1	4	1.39308	337.5	0.6	5.17276	270	270	0	21.4	1.9	5.4	10 ⁻¹⁵	10.5
6609	07-232 23:46	2	1	54	7	7	5	10	15	15	1	1	0	0	1	0	0	0	1	4	1.39322	337.6	1.2	5.17371	322	261	51	10.4	1.9	1.9	10 ⁻¹³	10.5
6610	07-233 05:45	1	1	60	2	10	0	15	14	15	1	0	23	30	1	0	0	0	1	4	1.39329	337.7	1.4	5.17404	330	257	59	11.8	11.8	1.1	10 ⁻¹³	5858.3
6611	07-233 08:18	1	1	135	2	10	0	15	15	15	1	0	19	31	1	0	0	0	1	4	1.39332	337.7	1.5	5.17420	75	99	14	11.8	11.8	1.1	10 ⁻¹³	5858.3
6612	07-233 18:33	3	5	53	55	30	30	13	6	5	0	1	63	0	1	0	0	0	1	4	1.39346	337.8	1.8	5.17487	318	259	46	2.0	1.9	4.1	10 ⁻¹⁷	10.5
6613	07-234 02:33	1	1	148	2	13	0	4	15	15	1	0	63	31	1	0	0	0	1	4	1.39354	337.8	2.0	5.17521	90	97	0	11.8	11.8	1.7	10 ⁻¹³	5858.3
6614	07-234 04:49	2	2	44	13	24	6	10	7	15	0	1	27	0	1	0	0	0	1	4	1.39359	337.8	2.1	5.17538	303	262	32	10.4	1.9	5.5	10 ⁻¹²	10.5
6615	07-234 11:22	2	2	66	10	19	0	7	7	7	0	0	36	0	1	0	0	0	1	4	1.39367	337.8	2.3	5.17572	333	248	61	34.1	1.9	2.9	10 ⁻¹⁴	10.5
6616	07-234 18:50	3	3	84	19	24	10	6	6	6	0	1	43	0	1	0	0	0	1	4	1.39382	337.9	2.6	5.17624	357	195	76	28.0	1.6	3.8	10 ⁻¹³	6.0
6617	07-236 00:41	2	1	166	4	8	0	8	7	8	0	0	6	26	1	0	0	0	1	4	1.39439	338.1	3.5	5.17804	105	91	-15	34.6	1.6	2.7	10 ⁻¹⁵	6.0
6618	07-236 01:26	1	1	124	2	8	0	15	15	15	1	0	47	31	1	0	0	0	1	4	1.39439	338.1	3.5	5.17804	46	106	41	2.3	2.0	1.3	10 ⁻¹¹	12.5
6619	07-236 14:31	1	1	103	2	10	0	11	15	15	1	0	47	31	1	0	0	0	1	4	1.39466	338.2	3.9	5.17879	14	139	68	11.8	11.8	1.1	10 ⁻¹³	5858.3
6620	07-237 03:07	2	3	80	23	8	3	7	15	15	1	1	47	30	1	0	0	0	1	4	1.39502	338.3	4.3	5.17974	338	228	61	14.1	1.9	1.1	10 ⁻¹²	10.5
6621	07-237 04:44	1	1	153	1	9	0	15	15	15	1	0	19	31	1	0	0	0	1	4	1.39502	338.3	4.3	5.17974	80	95	8	11.8	11.8	7.3	10 ⁻¹⁴	5858.3
6622	07-237 12:37	2	5	30	49	30	9	9	15	15	1	1	24	10	1	0	0	0	1	4	1.39526	338.3	4.6	5.18032	265	263	-5	11.8	11.8	3.4	10 ⁻¹⁰	5858.3
6623	07-238 04:37	1	1	156	3	11	0	9	15	15	1	0	47	31	1	0	0	0	1	4	1.39568	338.4	5.1	5.18130	78	94	10	14.1	1.9	9.5	10 ⁻¹⁴	10.5
6624	07-238 13:55	2	1	130	3	7	0	9	10	10	0	0	0	0	1	0	0	0	1	4	1.39595	338.5	5.4	5.18191	39	112	46	21.0	1.6	1.1	10 ⁻¹⁴	6.0
6625	07-239 02:06	1	2	84	12	10	0	14	13	15	1	0	63	30	1	0	0	0	1	4	1.39633	338.5	5.7	5.18272	330	228	52	2.0	1.9	1.7	10 ⁻¹⁰	10.5
6626	07-239 04:23	0	1	80	4	10	0	8	13	13	1	0	59	26	1	0	0	0	1	4	1.39643	338.6	5.8	5.18293	324	234	47	28.0	1.6	7.4	10 ⁻¹⁵	6.0
6627	07-239 07:31	2	2	93	11	13	12	10	15	15	1	1	7	30	1	0	0	0	1	4	1.39653	338.6	5.9	5.18314	341	213	59	4.9	1.7	9.4	10 ⁻¹²	7.7
6628	07-239 10:53	1	2	146	10	11	0	9	15	15	1	0	1	30	1	0	0	0	1	4	1.39663	338.6	6.0	5.18335	55	102	31	14.1	1.9	3.0	10 ⁻¹³	10.5
6629	07-239 20:37	1	1	119	2	9	0	10	15	15	1	0	47	31	1	0	0	0	1	4	1.39694	338.7	6.3	5.18397	13	145	61	11.8	11.8	8.8	10 ⁻¹⁴	5858.3
6630	07-241 05:19	3	2	89	12	20	9	7	8	6	0	1	39	0	1	0	0	0	1	4	1.39819	338.9	7.3	5.18636	321	228	42	34.1	1.9	5.2	10 ⁻¹⁴	10.5
6631	07-242 12:36	0	1	42	3	6	0	8	5	9	1	0	46	20	1	0	0	0	1	4	1.39961	339.1	8.3	5.18884	244	266	-24	38.7	1.6	1.0	10 ⁻¹⁵	6.0
6632	07-242 19:32	3	3	127	21	27	15	5	6	0	1	46	31	1	0	0	0	0	1	4	1.39988	339.1	8.5	5.18931	2	163	57	43.7	1.6	1.9	10 ⁻¹³	6.0
6633	07-243 00:09	3	4	93	28	22	18	11	9	15	1	1	24	30	1	0	0	0	1	4	1.40017	339.1	8.7	5.19365	318	217	35	10.4	1.9	2.2	10 ⁻¹¹	10.5
6634	07-243 01:22	2	2	122	8	12	0	7	6	0	0	4	31	1	0	0	0	0	1	4	1.40017	339.1	8.7	5.19365	353	179	55	43.7	1.6	4.0	10 ⁻¹⁵	6.0
6635	07-243 09:57	2	3	104	20	10	14	14	14	15	1	1	25	31	1	0	0	0	1	4	1.40060	339.2	8.9	5.19049	325	215	42	2.0	1.9	4.8	10 ⁻¹⁰	10.5
6636	07-243 10:37	3	3	159	22	27	6	7	5	6	0	1	47	0	1	0	0	0	1	4	1.40060	339.2	8.9	5.19049	43	109	36	28.1	1.6	1.1	10 ⁻¹²	6.0
6637	07-243 23:21	1	1	244	2	13	0	15	15	15	1	0	46	31	1	0	0	0	1	4	1.40119	339.3	9.3	5.19144	159	26	-57	11.8	11.8	1.7	10 ⁻¹³	5858.3
6638	07-245 00:09	3	3	106	20	27	8	8	5	6	0	1	46	0	1	0	0	0	1	4	1.40261	339.4	10.1	5.19365	318	217	35	10.4	1.9	2.2	10 ⁻¹¹	10.5
6639	07-245 03:06	2	3	155	17	9	2	15	15	15	1	1	47	30	1	0	0	0	1	4	1.40278	339.5	10.2	5.19390	26	124	44	2.0	1.9	2.2	10 ⁻¹⁰	10.5
6640	07-245 07:18	3	3	111	19	24	8	8	6	6	0	1	43	0	1	0	0	0	1	4	1.40294	339.5	10.3	5.19415	324	211	38	19.5	1.6	1.5	10 ⁻¹²	6.0
6641	07-245 11:02	2	2	159	12	14	13	11	11	15	1	1	36	9	1	0	0	0	1	4	1.40311	339.5	10.4	5.19440	31	119	41	12.6	1.6	8.9	10 ⁻¹³	6.0
6642	07-245 18:19	3	1	107	4	9	4	7	7	8	0	1	0	0	1	0	0	0	1	4	1.40362	339.6	10.7	5.19516	315	217	32	43.7	1.6	1.2	10 ⁻¹⁵	6.0
6643	07-246 12:56	1	1	145	1	11	0	15	14	15	1	0	47	31	1	0	0	0	1	4	1.40468	339.7	11.2	5.19671	5	153	48	11.8	11.8	1.0	10 ⁻¹³	5858.3
6644	07-247 07:06	2	1	211	7	12	0	7	7	8	0	0	35	0	1	0	0	0	1	4	1.40578	339.8	11.8	5.19829	94	70	-5	43.7	1.6	3.3	10 ⁻¹⁵	6.0
6645	07-247 09:04	1	2	188	9	2	0	13	15	15	1	0	44	30	1	0	0	0	1	4	1.40597	339.8	11.9	5.19855	61	91	19	2.5	1.9	1.1	10 ⁻¹¹	10.5
6646	07-247 09:21	3	3	170	22	25	6	7	6	0	1	47	0	1	0	0	0	0	1	4	1.40597	339.8	11.9	5.19855	36	112	36					

No.	IMP. DATE	C L N	AR S E C	IA	EA	CA	IT	ET	E I T	E I C	I I C	PA	P E T	E V D	I C P	E C P	C C P	P P	HV	R	LON	LAT	D _{dup}	ROT	S _{Lon}	S _{Lat}	V	VEF	M	MEF
6656	07-254 04:44	2	1 148	2	7	0	15	7	10	0	0	2	26	1	0	0	0	0	4	1.41814	340.8	16.7	5.21426	339	170	34	56.0	2.0	2.2·10 ⁻¹⁶	12.5
6657	07-254 06:33	1	1 157	1	1	0	15	15	15	1	0	2	31	1	0	0	0	1	4	1.41840	340.9	16.8	5.21458	351	155	37	11.8	1.9·10 ⁻¹⁴	5858.3	
6658	07-254 08:21	2	3 178	18	13	7	10	15	15	1	1	47	30	1	0	0	0	1	4	1.41840	340.9	16.8	5.21458	21	118	34	5.0	1.9	1.7·10 ⁻¹¹	10.5
6659	07-254 11:05	2	1 163	5	10	0	7	7	8	0	0	2	26	1	0	0	0	1	4	1.41866	340.9	16.9	5.21489	359	144	37	43.7	1.6	1.7·10 ⁻¹⁵	6.0
6660	07-254 22:47	2	1 152	1	6	0	15	8	11	0	0	2	26	1	0	0	0	1	4	1.41973	341.0	17.3	5.21617	342	164	35	36.7	2.0	8.9·10 ⁻¹⁶	12.5
6661	07-255 08:39	2	1 134	4	9	0	7	8	0	0	0	2	26	1	0	0	0	1	4	1.42054	341.0	17.5	5.21713	316	191	24	43.7	1.6	1.2·10 ⁻¹⁵	6.0
6662	07-255 12:32	3	1 185	5	9	1	7	7	7	0	1	2	31	1	0	0	0	1	4	1.42109	341.1	17.7	5.21778	27	109	31	43.7	1.6	1.4·10 ⁻¹⁵	6.0
6663	07-255 14:05	2	1 179	1	7	0	15	9	11	0	0	2	26	1	0	0	0	1	4	1.42109	341.1	17.7	5.21778	18	119	34	31.3	2.0	1.8·10 ⁻¹⁵	12.5
6664	07-255 15:28	3	3 201	23	26	6	7	7	6	0	1	47	0	1	0	0	0	1	4	1.42136	341.1	17.8	5.21811	49	87	21	21.0	1.6	3.4·10 ⁻¹²	6.0
6665	07-255 19:41	3	1 106	5	8	7	7	8	8	0	1	2	26	1	0	0	0	1	4	1.42164	341.1	17.9	5.21843	275	224	0	35.4	1.6	2.9·10 ⁻¹⁵	6.0
6666	07-255 20:04	2	1 161	4	9	0	8	7	8	0	0	3	26	1	0	0	0	1	4	1.42164	341.1	17.9	5.21843	352	150	36	34.6	1.6	3.2·10 ⁻¹⁵	6.0
6667	07-255 20:39	2	1 119	1	5	0	15	9	10	0	0	2	26	1	0	0	0	1	4	1.42164	341.1	17.9	5.21843	293	210	11	31.3	2.0	1.3·10 ⁻¹⁵	12.5
6668	07-256 11:43	2	2 189	8	8	7	9	15	15	1	1	35	30	1	0	0	0	1	4	1.42304	341.2	18.3	5.22008	30	105	30	14.1	1.9	1.3·10 ⁻¹³	10.5
6669	07-257 00:34	2	2 111	8	13	0	7	7	7	0	0	36	0	1	0	0	0	1	4	1.42448	341.3	18.8	5.22175	278	219	2	43.7	1.6	4.7·10 ⁻¹⁵	6.0
6670	07-258 09:31	2	2 62	8	12	0	8	7	6	0	0	2	26	1	0	0	0	1	4	1.42775	341.5	19.7	5.22549	206	282	-38	34.6	1.6	1.1·10 ⁻¹⁴	6.0
6671	07-261 08:19	1	2 117	8	0	12	15	0	1	0	0	12	30	1	0	0	0	1	4	1.43507	342.0	21.7	5.23367	276	213	0	5.0	1.9	2.4·10 ⁻¹²	10.5
6672	07-261 18:01	1	1 160	2	1	0	15	15	15	1	0	3	31	1	0	0	0	1	4	1.43641	342.1	22.1	5.23514	335	159	29	11.8	11.8	2.4·10 ⁻¹⁴	5858.3
6673	07-262 13:16	2	3 145	23	20	13	7	15	15	1	1	63	29	1	0	0	0	1	4	1.43846	342.2	22.6	5.23737	312	181	20	14.1	1.9	5.6·10 ⁻¹²	10.5
6674	07-262 14:14	3	4 142	30	51	29	6	13	5	0	1	47	0	1	0	0	0	1	4	1.43846	342.2	22.6	5.23737	308	184	18	21.4	1.9	5.0·10 ⁻¹¹	10.5
6675	07-265 04:43	2	1 7	2	1	0	9	15	11	0	0	2	25	1	0	0	0	1	4	1.44596	342.6	24.4	5.24542	113	20	-16	11.8	11.8	2.4·10 ⁻¹⁴	5858.3
6676	07-265 12:09	1	1 73	4	1	0	8	15	0	1	0	3	31	1	0	0	0	1	4	1.44707	342.7	24.7	5.24660	205	273	-36	21.4	1.9	4.5·10 ⁻¹⁵	10.5
6677	07-265 12:21	2	3 170	23	14	11	12	14	15	1	0	11	31	1	0	0	0	1	4	1.44707	342.7	24.7	5.24660	342	147	29	2.0	1.9	1.5·10 ⁻⁰⁹	10.5
6678	07-266 13:56	2	1 129	5	9	0	7	7	7	0	0	3	31	1	0	0	0	1	4	1.45009	342.8	25.4	5.24979	282	201	3	43.7	1.6	1.4·10 ⁻¹⁵	6.0
6679	07-267 22:27	2	2 109	15	15	10	10	0	0	1	1	46	30	1	0	0	0	1	4	1.45437	343.1	26.3	5.25426	251	225	-14	10.4	1.9	2.3·10 ⁻¹²	10.5
6680	07-268 05:28	2	3 209	19	5	4	3	2	3	1	1	47	0	1	0	0	0	1	4	1.45517	343.1	26.5	5.25508	31	87	25	70.0	1.9	6.9·10 ⁻¹⁶	10.5
6681	07-268 05:52	1	1 236	7	2	0	10	14	0	1	0	18	29	1	0	0	0	1	4	1.45517	343.1	26.5	5.25508	69	52	8	10.4	1.9	8.6·10 ⁻¹⁴	10.5
6682	07-268 07:16	1	1 152	2	7	0	15	15	15	1	0	15	31	1	0	0	0	1	4	1.45556	343.1	26.6	5.25500	311	174	18	2.3	2.0	1.1·10 ⁻¹¹	12.5
6683	07-268 09:29	2	1 153	1	4	1	15	15	0	1	1	9	30	1	0	0	0	1	4	1.45596	343.2	26.6	5.25591	312	173	18	11.8	11.8	3.0·10 ⁻¹⁴	5858.3
6684	07-268 13:09	2	1 162	4	7	0	8	7	8	0	0	3	31	1	0	0	0	1	4	1.45636	343.2	26.7	5.25633	325	161	24	34.6	1.6	2.2·10 ⁻¹⁵	6.0
6685	07-270 15:52	1	1 107	5	10	0	5	10	0	1	0	47	31	1	0	0	0	1	4	1.46334	343.6	28.1	5.26351	244	229	-18	43.1	1.6	1.8·10 ⁻¹⁵	6.0
6686	07-272 22:38	1	2 174	8	5	0	10	15	0	1	0	40	30	1	0	0	0	1	4	1.47108	343.9	29.6	5.27138	334	147	26	10.4	1.9	1.6·10 ⁻¹³	10.5
6687	07-274 07:27	2	1 122	1	4	0	15	8	12	0	0	3	31	1	0	0	0	1	4	1.47598	344.2	30.5	5.27632	259	213	-9	11.8	11.8	3.0·10 ⁻¹⁴	5858.3
6688	07-276 04:27	1	1 176	4	4	0	13	15	0	1	0	45	30	1	0	0	0	1	4	1.48287	344.5	31.8	5.28322	332	146	25	2.5	1.9	6.1·10 ⁻¹²	10.5
6689	07-277 17:34	2	1 120	2	7	0	15	8	10	0	0	2	31	1	0	0	0	1	4	1.48855	344.8	32.7	5.28887	251	217	-14	36.7	2.0	1.3·10 ⁻¹⁵	12.5
6690	07-277 18:10	1	1 128	4	3	0	8	15	0	1	0	3	31	1	0	0	0	1	4	1.48903	344.8	32.8	5.28935	263	207	-7	21.4	1.9	6.1·10 ⁻¹⁵	10.5
6691	07-277 22:33	1	1 109	2	1	0	15	14	0	1	0	3	25	1	0	0	0	1	4	1.48951	344.9	32.9	5.28983	236	231	-22	11.8	11.8	2.4·10 ⁻¹⁴	5858.3
6692	07-278 14:31	2	1 190	3	5	0	12	3	9	0	0	17	1	1	0	0	0	1	4	1.49194	345.0	33.3	5.29222	349	127	29	18.7	4.1	1.3·10 ⁻¹⁴	151.2
6693	07-279 14:27	2	1 167	3	6	0	8	9	8	0	0	4	31	1	0	0	0	1	4	1.49586	345.2	33.9	5.29610	315	161	19	25.9	1.6	4.1·10 ⁻¹⁵	6.0
6694	07-280 00:46	2	1 198	2	1	0	15	4	12	0	0	3	31	1	0	0	0	1	4	1.49785	345.2	34.3	5.29806	358	115	29	11.8	11.8	2.4·10 ⁻¹⁴	5858.3
6695	07-281 00:14	3	4 172	26	49	14	7	5	5	0	1	47	0	1	0	0	0	1	4	1.50187	345.4	34.9	5.30201	320	155	21	14.1	1.9	7.7·10 ⁻¹¹	10.5
6696	07-282 23:46	2	1 146	5	4	0	8	1	9	0	0	8	1	1	0	0	0	1	4	1.50958	345.8	36.1	5.30955	281	189	2	21.4	1.9	8.3·10 ⁻¹⁵	10.5
6697	07-283 00:22	3	3 190	21	27	11	6	5	5	0	1	46	0	1	0	0	0	1	4	1.51010	345.8	36.1	5.31006	343	130	28	34.6	1.6	5.0·10 ⁻¹³	6.0
6698	07-284 09:41	1	1 67	3	1	0	8	15	0	1	0	3	31	1	0	0	0	1	4	1.51590	346.1	37.0	5.31571	169	305	-37	21.4	1.9	3.9·10 ⁻¹⁵	10.5
6699	07-288 08:40	3	3 171	22	28	28	6	5	6	0	1	34	31	1	0	0	0	1	4	1.53283	346.9	39.4	5.33213	311	162	17	34.6	1.6	7.0·10 ⁻¹³	6.0
6700	07-288 14:40	1	1 190	3	9	0	10	15	0	1	0	47	31	1	0	0	0	1	4	1.53395	346.9	39.5	5.33322	337	135	26	4.9	1.7	1.3·10 ⁻¹²	7.7
6701	07-288 22:33	2	1 13	3	7	0	8	8	9	0	0	2	25	1	0	0	0	1	4	1.53564	347.0	39.7	5.33485	88	24	-2	28.0	1.6	3.7·10 ⁻¹⁵	6.0
6702	07-294 09:24	1	1 229	5	11	0	11	15	0	1	0	46	31	1	0	0	0	1	4	1.56130	348.2	43.0	5.35953	26	79	25	4.2	1.6	4.0·10 ⁻¹²	6.0
6703	07-295 14:24	2	1 86	5	9	0	7	7	7	0																				

No.	IMP.DATE	C I N	AR	S E C	IA	EA	CA	IT	ET	E I T	E I C	I I C	PA	P E T	E V D	I C P	E C P	C C P	P C P	HV	R	LON	LAT	D _{up}	ROT	S _{Lon}	S _{Lat}	V	VEF	M	MEF
6711	07-308 19:03	3	4	191	24	28	10	8	6	5	0	1	47	0	1	0	0	0	0	4	1.63524	351.7	50.9	5.43033	321	151	21	19.5	1.6	7.6·10 ⁻¹²	6.0
6712	07-309 11:53	3	3	196	23	29	6	5	5	0	1	47	0	1	0	0	0	0	4	1.63865	351.9	51.2	5.43360	327	145	23	34.6	1.6	9.8·10 ⁻¹³	6.0	
6713	07-309 22:37	2	2	149	10	19	0	8	6	6	0	38	0	1	0	0	0	0	4	1.64140	352.0	51.5	5.43622	261	204	-8	21.4	1.9	1.5·10 ⁻¹³	10.5	
6714	07-313 19:02	2	1	238	1	1	0	15	15	12	0	0	2	31	1	0	0	0	4	1.66299	353.1	53.5	5.45691	23	84	26	11.8	11.8	1.9·10 ⁻¹⁴	5858.3	
6715	07-315 23:14	2	1	62	7	11	0	8	7	7	0	3	31	1	0	0	0	0	4	1.67505	353.8	54.5	5.46850	134	344	-26	34.6	1.6	7.6·10 ⁻¹⁵	6.0	
6716	07-322 23:06	1	1	152	4	3	0	9	15	0	1	0	2	24	1	0	0	0	4	1.71586	356.0	57.9	5.50781	255	211	-11	14.1	1.9	2.8·10 ⁻¹⁴	10.5	
6717	07-327 11:18	1	1	181	5	7	0	9	15	0	1	0	2	24	1	0	0	0	4	1.74286	357.7	59.9	5.53397	293	181	8	5.7	2.1	7.9·10 ⁻¹³	15.7	
6718	07-327 23:30	1	1	150	2	9	0	9	15	0	1	0	37	31	1	0	0	0	4	1.74589	357.8	60.1	5.53692	249	218	-14	11.8	11.8	8.8·10 ⁻¹⁴	5858.3	
6719	07-334 08:17	2	3	151	19	7	1	7	15	0	1	1	46	31	1	0	0	0	4	1.78513	0.4	62.9	5.57519	246	223	-16	14.1	1.9	4.3·10 ⁻¹³	10.5	

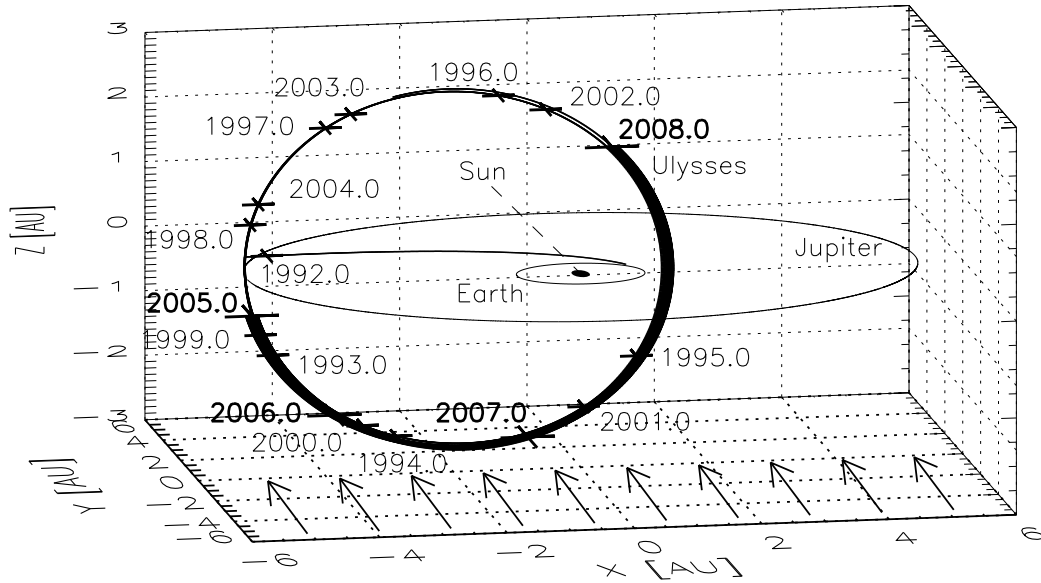


Figure 1: The trajectory of Ulysses in ecliptic coordinates with the Sun at the centre. The orbits of Earth and Jupiter indicate the ecliptic plane, and the initial trajectory of Ulysses was in this plane. Since Jupiter flyby in early 1992 the orbit has been almost perpendicular to the ecliptic plane (79° inclination). Crosses mark the spacecraft position at the beginning of each year. The 2005 to 2007 part of the trajectory is shown as a thick line. Vernal equinox is to the right (positive x axis). Arrows indicate the undisturbed interstellar dust flow direction which is within the measurement accuracy co-aligned with the direction of the interstellar helium gas flow. It is almost perpendicular to the orbital plane of Ulysses.

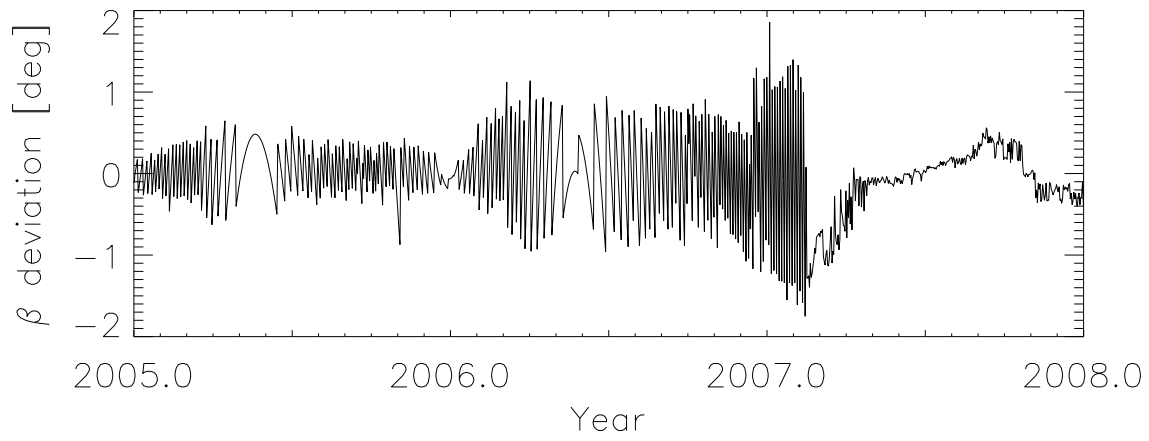
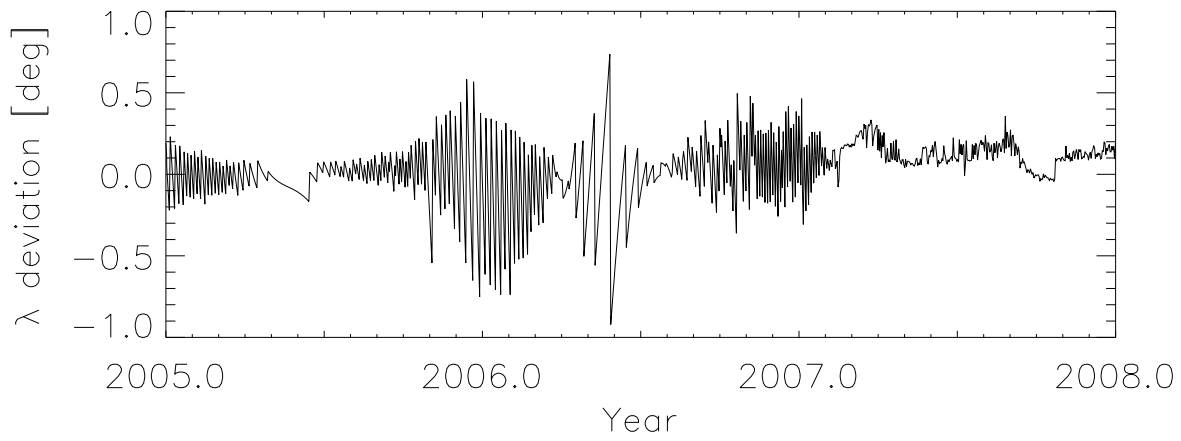


Figure 2: Spacecraft attitude: deviation of the antenna pointing direction (i.e. positive spin axis) from the nominal Earth direction. The angles are given in ecliptic longitude (top) and latitude (bottom, equinox 1950.0).

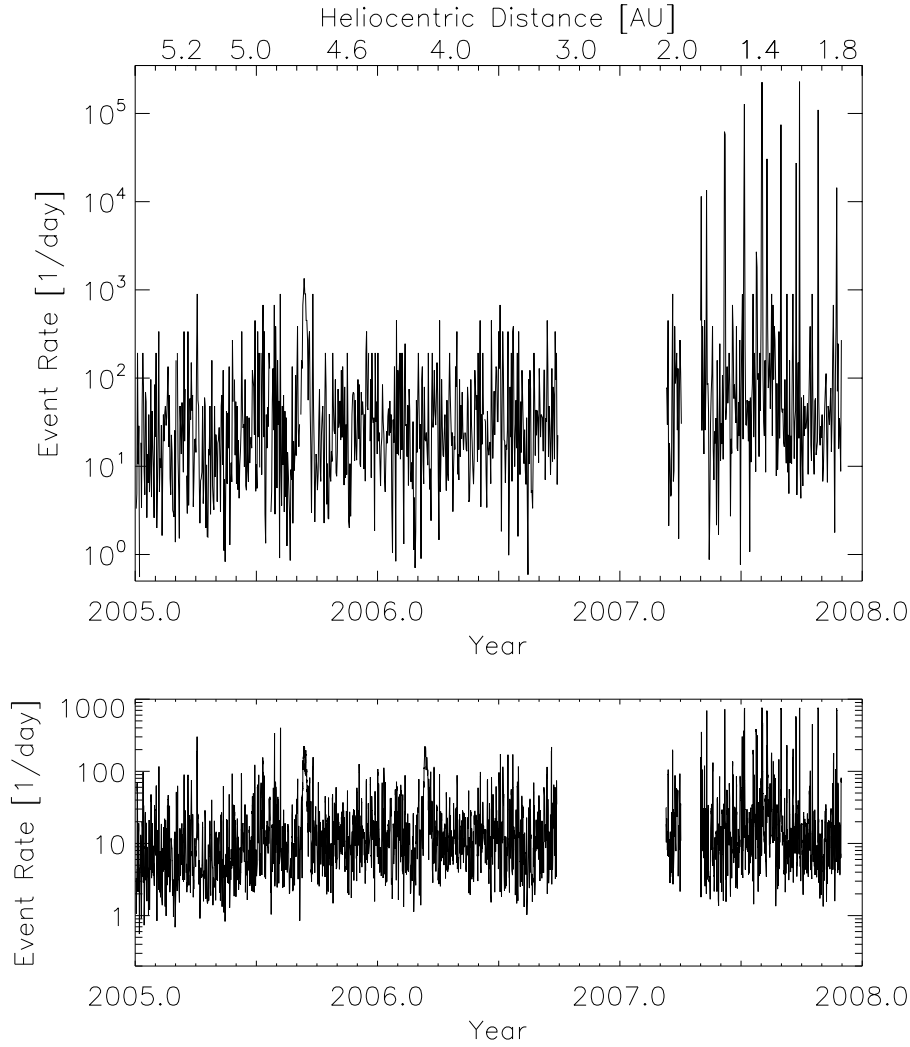


Figure 3: Noise rate (class 0 events) detected with the dust instrument. The heliocentric distance of Ulysses is indicated at the top. Upper panel: Daily maxima in the noise rate (determined from the AC01 accumulator). Sharp spikes which are strongest after May 2008 when Ulysses was in the inner solar system are caused by periodic noise tests. Lower panel: One-day average of the noise rate calculated from the number of AC01 events for which the complete information was transmitted to Earth.

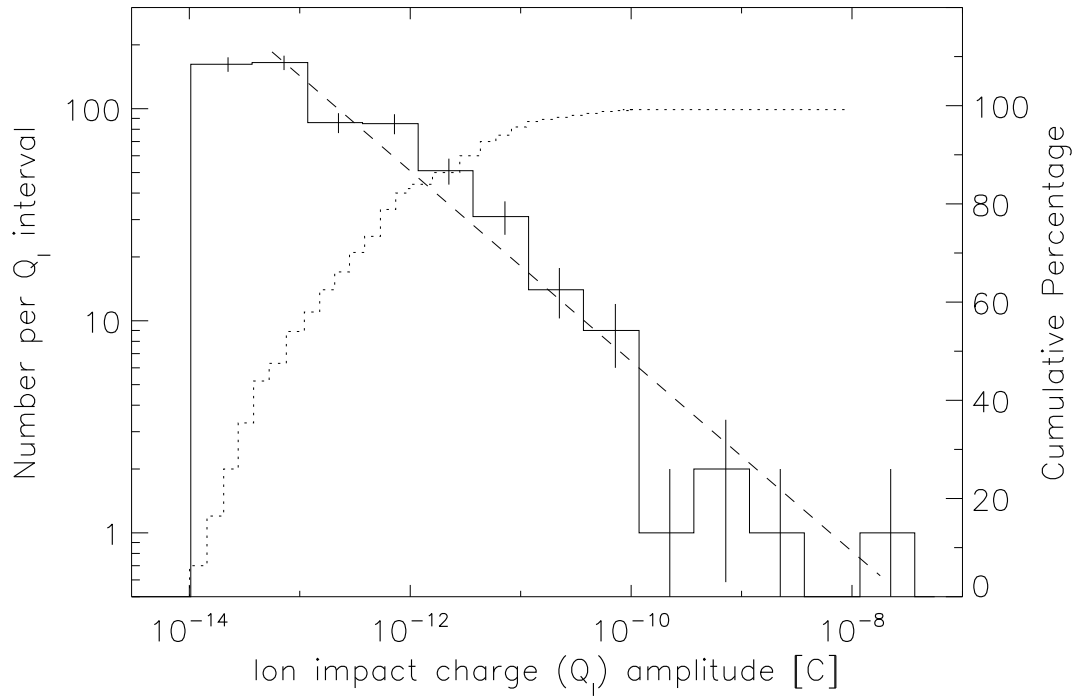


Figure 4: Distribution of the impact charge amplitude Q_I for all dust particles detected from 2005 to 2007. The solid line indicates the number of impacts per charge interval, and the dotted line shows the cumulative percentage. Vertical bars indicate the \sqrt{n} statistical fluctuation. A power law fit to the data with $Q_I > 3 \cdot 10^{-14} \text{ C}$ is shown as a dashed line (Number $N \sim Q_I^{-0.45}$).

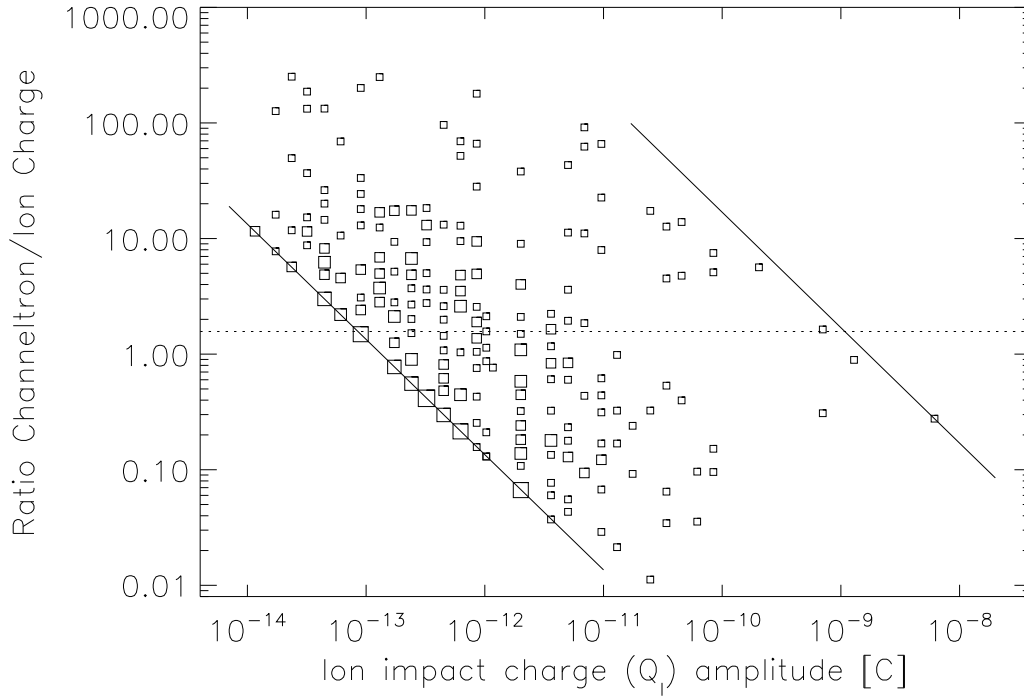


Figure 5: Channeltron amplification factor $A = Q_C/Q_I$ as a function of impact charge Q_I for all dust impacts detected between 2005 and 2007 with channeltron voltage set to $HV = 4$. The solid lines denote the sensitivity threshold (lower left) and the saturation limit (upper right) of the channeltron. Squares indicate dust particle impacts. The area of each square is proportional to the number of events included (the scaling of the squares is not the same as in earlier papers). The dotted horizontal line shows the mean value of the channeltron amplification $A = 1.57$ for ion impact charges $10^{-12} \text{ C} < Q_I < 10^{-11} \text{ C}$ (82 particles).

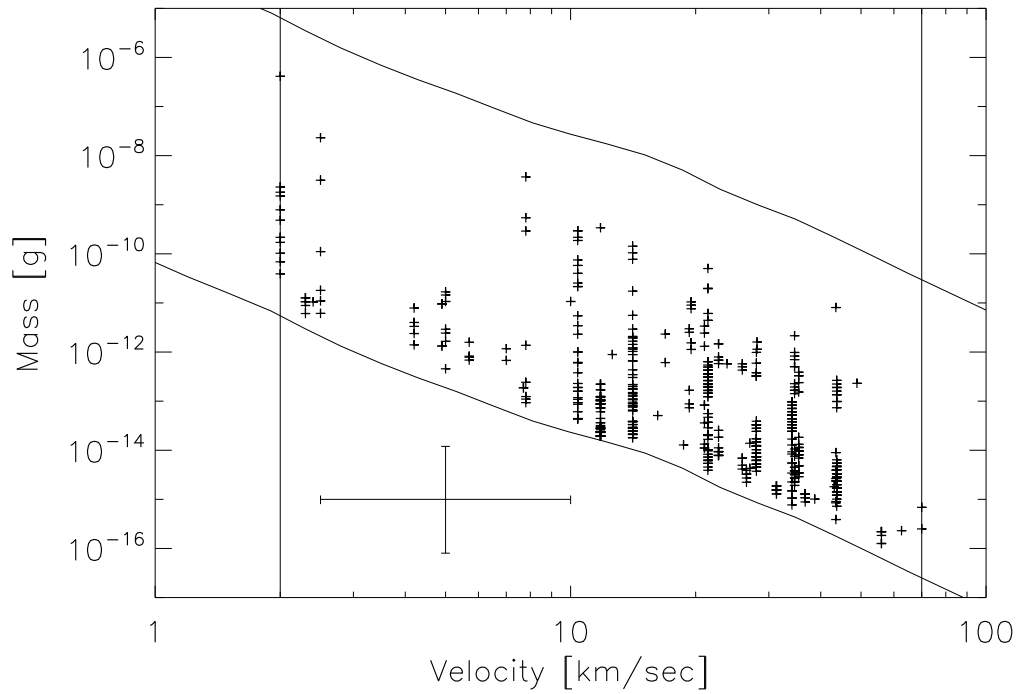


Figure 6: Masses and impact velocities of all impacts recorded with the Ulysses sensor from 2005 to 2007. The lower and upper solid lines indicate the threshold and the saturation limit of the detector, respectively, and the vertical lines indicate the calibrated velocity range. A sample error bar is shown that indicates a factor of 2 uncertainty for the velocity and a factor of 10 for the mass determination.

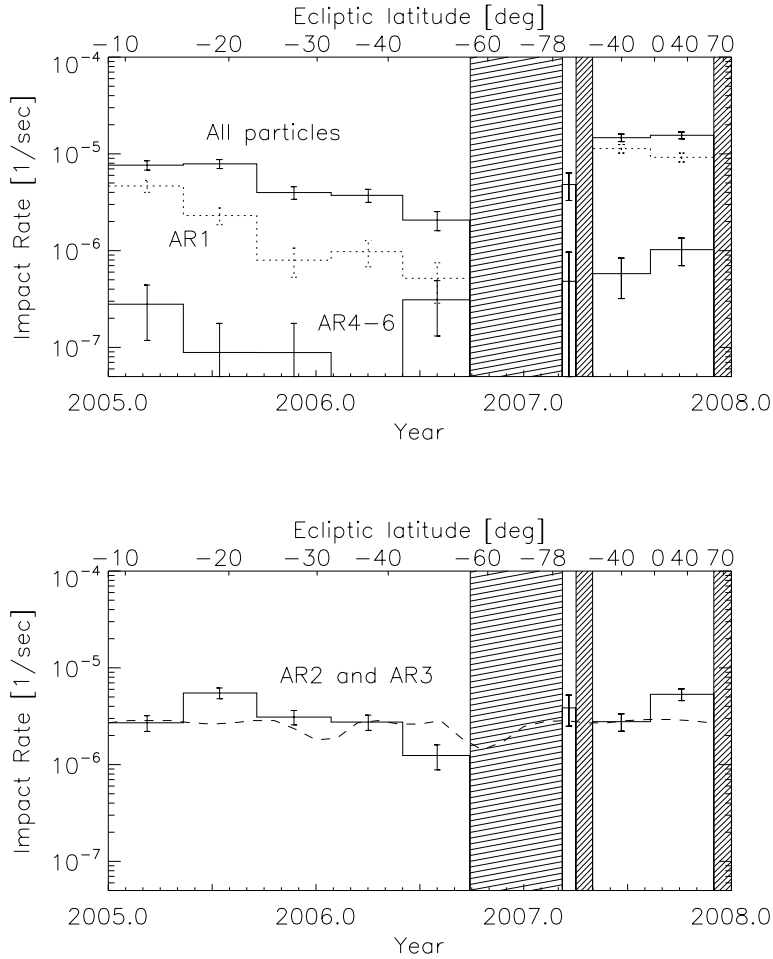


Figure 7: Impact rate of dust particles detected with the Ulysses dust sensor as a function of time. The ecliptic latitude of the spacecraft is indicated at the top. Shaded areas indicate periods when the dust instrument was switched off. Upper panel: total impact rate (upper solid histogram), impact rate of small particles (AR1, dotted histogram), and impact rate of big particles (AR4 – AR6, lower solid histogram). Note that a rate of about 10^{-7} impacts per second is caused by a single dust impact in the averaging interval of about 130 days. An averaging interval of only 24 days had to be used for March 2007 because the instrument was switched off before and after this period. Lower panel: impact rate of intermediate size particles (AR2 and AR3, solid histogram). A model for the rate of interstellar particles assuming a constant flux was fit to the data and is superimposed as a dashed line. Vertical bars indicate the \sqrt{n} statistical fluctuation.

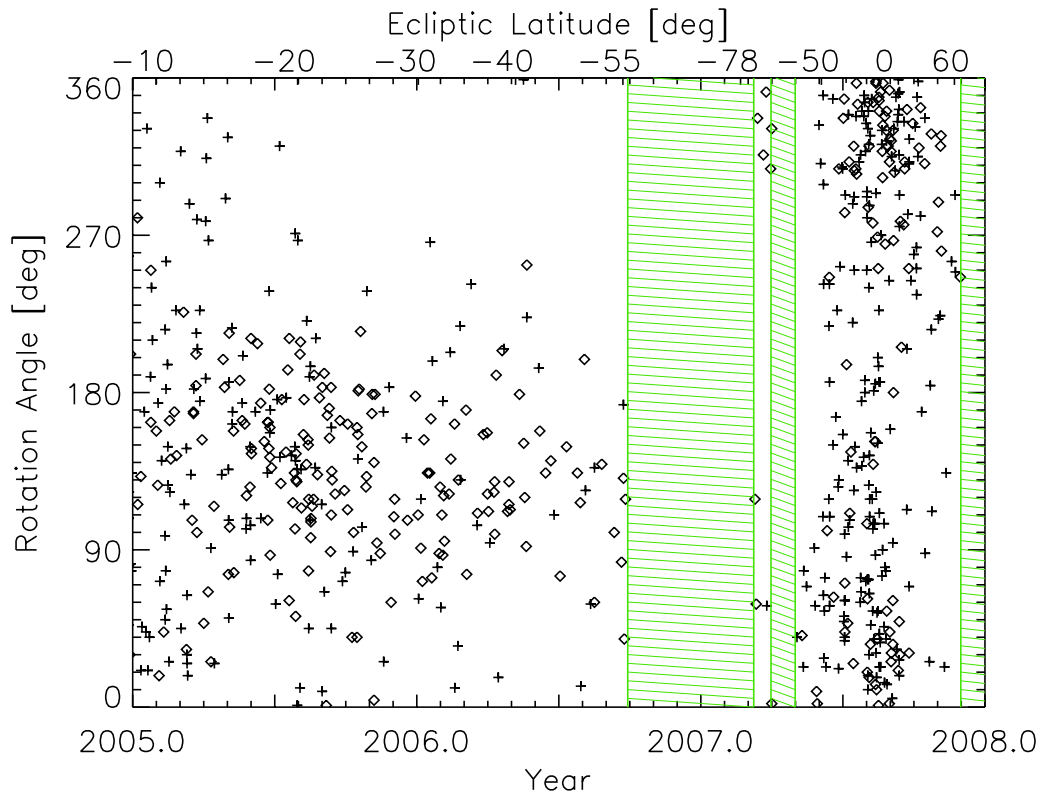


Figure 8: Rotation angle vs. time for all particles detected in the 2005-07 interval. Plus signs indicate particles with impact charge $Q_I < 10^{-13}$ C (AR1), diamonds those with $Q_I > 10^{-13}$ C (AR2 to AR6). Ulysses' ecliptic latitude is indicated at the top.

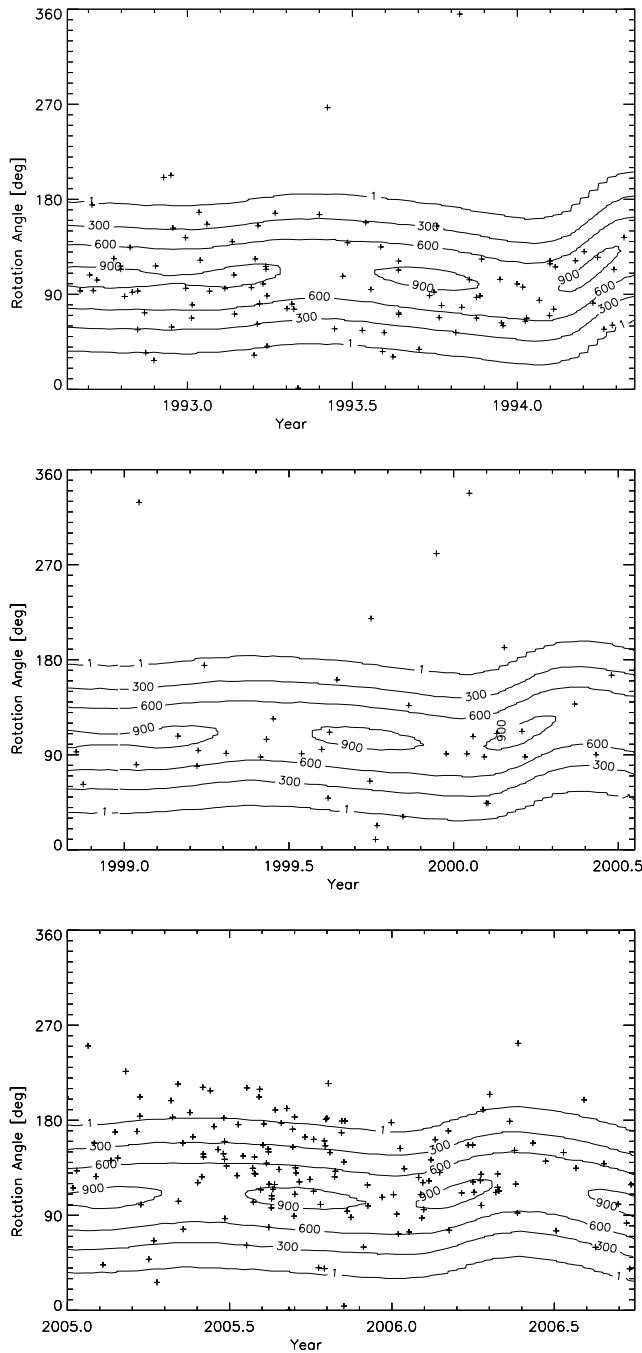


Figure 9: Rotation angle vs. time for dust impacts detected during three time intervals when Ulysses was traversing the same portion of its trajectory between -8.5° and -56° ecliptic latitude. Contour lines show the effective sensor area for dust particles approaching from the upstream direction of interstellar helium (Witte, 2004; Witte et al., 2004). Here we show particles with impact charges $1.5 \cdot 10^{-13} \text{ C} < Q_I < 10^{-11} \text{ C}$ which approximately coincides with AR2-3.

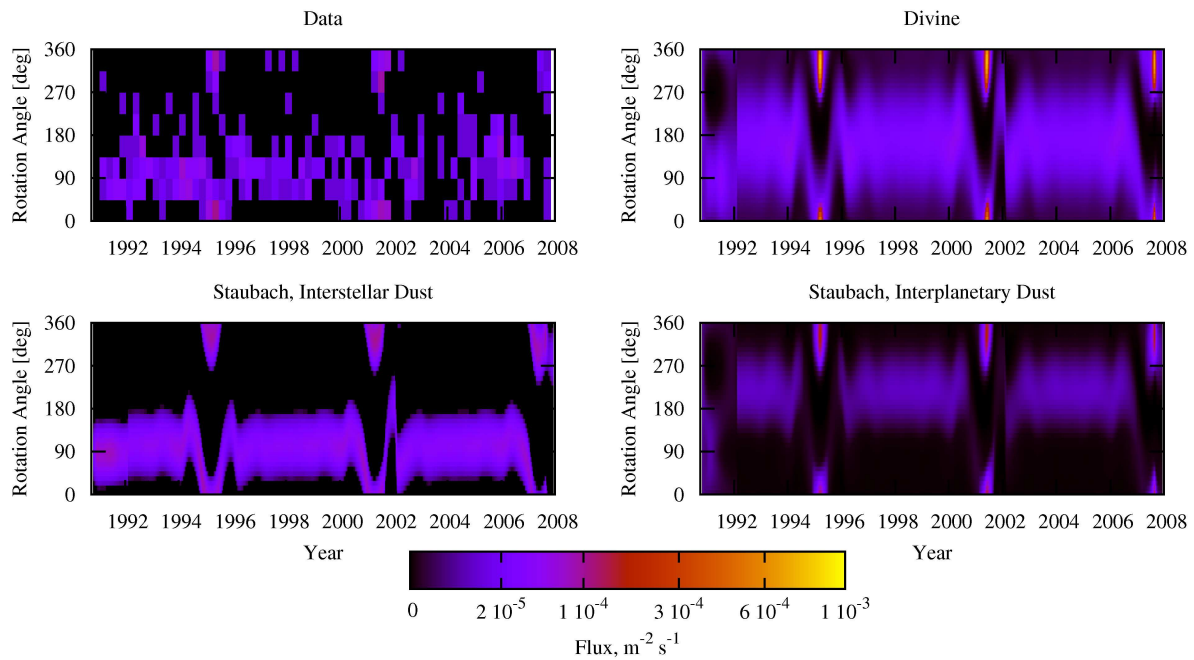


Figure 10: Particle fluxes onto the Ulysses dust detector as seen in the data (AR3 to AR6) and predicted by the meteoroid environment models by Divine (1993) and Staubach et al. (1997). The Jupiter encounter in 1992 is best seen on the model plots as a sharp swing in directionality of impacts caused by the rotation of spacecraft velocity relative to the giant planet. The ecliptic plane crossings near the perihelia in 1995, 2001 and 2007 are marked by the high impact rates due to the number density of dust increase near the Sun and high speed of Ulysses relative to this dust. As the spacecraft moved north at the times of the crossings, the meteoroids came from the ecliptic north as well (rotation angle 0).