Constraining The Mineralogy of Mercury Using Gamma Ray and Infrared Spectroscopy Maggie McAdam **Terps Conference** December 5, 2012

APOD 5/9/08, Nemiroff & Bonnell

### Mercury – A brief summary

# Small and difficult to observe with Space telescopes

- Difficult to get to with spacecraft
- Highly interesting scientifically
  - Highest uncompressed density in Solar System
  - Heterogeneous surface materials
  - Evidence for volcanic activity on surface
  - Thermal extremes on surface
  - Massive core and magneto sphere





#### **Observations of Mercury**

#### Infrared Spectroscopy

- Mineralogy of surface
- Grain size
- Thermal effects (longer wavelengths)
- Maturation of surface
- Gamma ray spectroscopy
  - Tells us what specific elements populate a surface
  - Probes top ~10 cm of surface

#### Gamma Ray Spectroscopy





#### **Reflectance Spectra of Mercury**



### Fayilitic Olivine



#### Infrared Spectroscopy



#### Mid-Infrared Observations



#### **Thermal Effects**





#### **Spectral Modeling**



#### **Current GRS Conclusions**

Confirmed abundances of Al, Ca, S and Fe
Comparisons with XRS and NS indicate vertical homogeneity on 10's of cm scales
Confirmation of low global Fe
Mineralogy constrained to albite, Na-rich plagioclase and Mg end member of pyroxene, enstatite.

#### Local abundances?

 GRS results for global counts
 MESSENGER Scientists working on mapping counts so that individual mineralogical units can be identified

### Questions?

#### Surface Maturation



#### Surface Maturation



#### Surface Maturation





## Mid-IR of Mercury

