

# ASTR 320 - Theoretical Astrophysics

Spring 2013

## Homework Assignment No. 2

- At a time  $t_1 = 1$  sec after the big-bang, the temperature of the Universe is approximately  $T_1 = 10^{10}$  K.
  - At this time, is the Universe radiation dominated or mass dominated?
  - Free neutrons have a lifetime of about 880 sec. By what factor does the scale  $R$  of the Universe change during one neutron lifetime? I.e., compute the ratio  $R(t_2)/R(t_1)$ , where  $t_2 = 881$  sec.
  - What is the temperature  $T(t_2)$ ? By what factor does the energy density of the radiation,  $u_R$ , change during this time?
  - How does the number density of neutrons change during this interval?
- Titan, Saturn's largest satellite, has a radius of 2575 km and a mass of  $1.3455 \times 10^{23}$  kg.
  - What is the surface gravity of Titan?
  - Titan's atmosphere is mostly molecular nitrogen,  $N_2$ . If we take the temperature of the atmosphere to be 85 K (and constant), find the scale height in km of Titan's atmosphere.
  - How does this compare to the scale height of Earth's atmosphere?
- Consider the solar nebula, the disk of gas and dust orbiting the sun before the formation of the planets.
  - Use equation (21) of the notes

$$H = \left( \frac{kT}{m} \right)^{1/2} \left( \frac{GM_*}{R^3} \right)^{-1/2}, \quad (1)$$

to estimate the scale height,  $H$ , of the disk at the radius of the Earth's orbit. (Use the sun's present mass,  $M_\odot = 2 \times 10^{33}$  g, take  $T = 300$ K, and assume the composition is mostly molecular hydrogen,  $H_2$ .)

- How does the disk's scale height compare to its radius at this distance from the proto-sun?
- Repeat the calculation for Jupiter's orbital radius (5.2 AU), assuming  $T = 130$ K.

**Due: 21 February 2013**