ASTR 288C - Astronomy Research Techniques Fall 2019

Homework Assignment No. 6

- 1. Write an IDL procedure to do (crude) apeture photometry. This routine should (1) integrate all the flux within a circle of radius "rc" centered at (xc,yc). (2) integrate all the flux within a ring of inner radius "ri" and outer radius "ro", with the same center (xc,yc). (3) Scale the flux from the ring by the ratio of the (rc area)/(ring area) and subtract this "background" from the total flux within the "rc" circle to obtain the corrected flux.
- 2. Go to the Hubble Legacy website

http://hla.stsci.edu/hlaview.html

and search for images of the planetary nebula BD+30 3639 (also called PK64+5D1).

Last week you measured the flux from the narrow band H-beta image (f487n) taken in 1994: (proposal 05403 - HARRINGTON): hst_05403_02_wfpc2_f487n_pc_drz.fits

You will find many other narrow band filter observations taken as part of that observing proposal.

- (a) Download and examine files taken through other narrow filters. In particular, f656n (H α), f658n ([N II] 6584), f375n ([O II] 3727) and f953n ([S III] 9532). Use the IDL readfits procedure.
- (b) Look at the headers of each. Find the inverse sensitivity PHOTFLAM and the filter bandwidth PHOTBW and save the product of these values. (You will use these numbers to convert counts to flux in ergs/s.)
- (c) Use the "ring.pro" procedure to measure the counts in each filter, excluding the central star. Convert the counts to flux units (erg s⁻¹). Make a list of the ratios of each flux to $H\beta$.
- (d) Look back at the lecture notes Lect-2.pdf and note the Balmer decrement shown in Table 14.2. Compare the ratio of $H\alpha$ to $H\beta$ from your measurements with the theoretical ratio. Coment on the difference.

In the directory Astronomy 288C on my webpage,

http://www.astro.umd.edu/jph/A288C.html

there is a file with the predictions of a computer model of BD+30 3639. Compare the predictions with your measurements of the line ratios.

Due: 21 October 2019