

TODAY

FIRST HOMEWORK DUE

- ANCIENT ASTRONOMY

- COMPETING COSMOLOGIES
 - GEOCENTRIC VS. HELIOCENTRIC

 - PTOLEMY VS. COPERNICUS

 - RETROGRADE MOTION

 - PHASES OF VENUS

- GALILEO

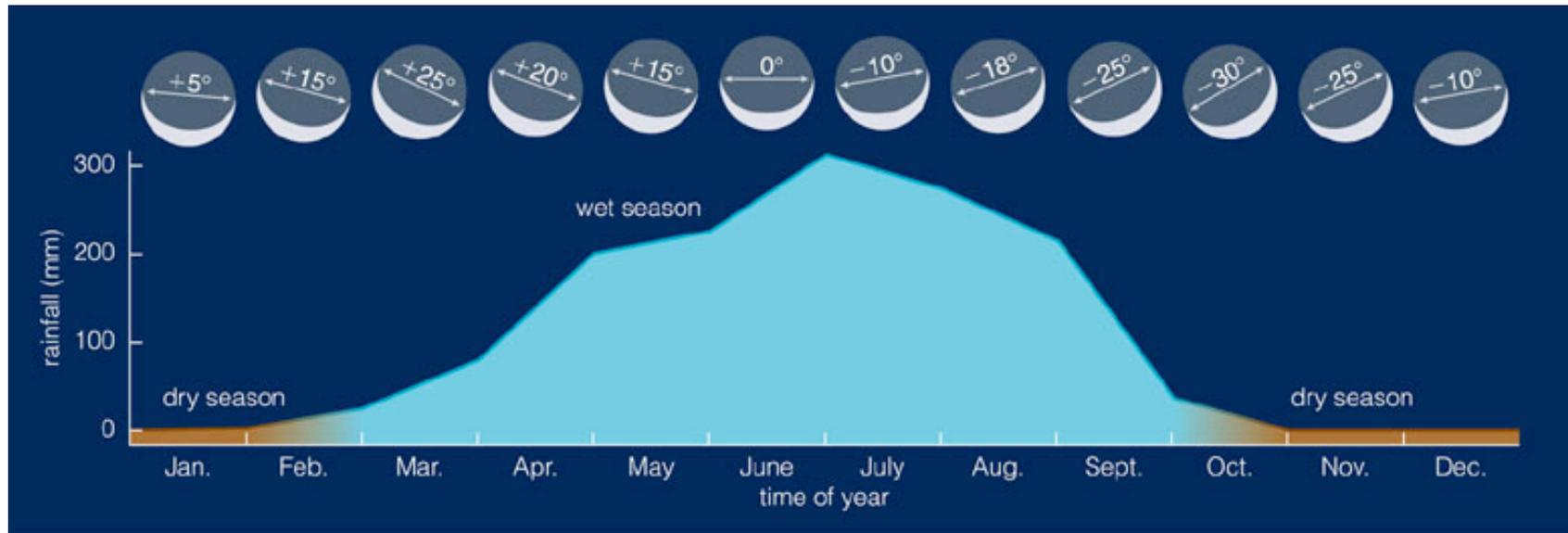
3.1 The Ancient Roots of Science

Our goals for learning:

- How did astronomical observations benefit ancient societies?
- What did ancient civilizations achieve in astronomy?

Astronomical observations were important to ancient societies

- In keeping track of time and seasons
 - for practical purposes, including agriculture
 - for religious and ceremonial purposes
- In aiding navigation



Ancient people of central Africa (6500 B.C.) could predict seasons from the orientation of the crescent moon.

Ancient achievements

- Daily timekeeping
- Tracking the seasons and calendar
- Monitoring lunar cycles
- Monitoring planets and stars
- Predicting eclipses
- Discovered precession
- And more...



Aztec calendar

- Egyptian obelisk:
Shadows tell time of
day.

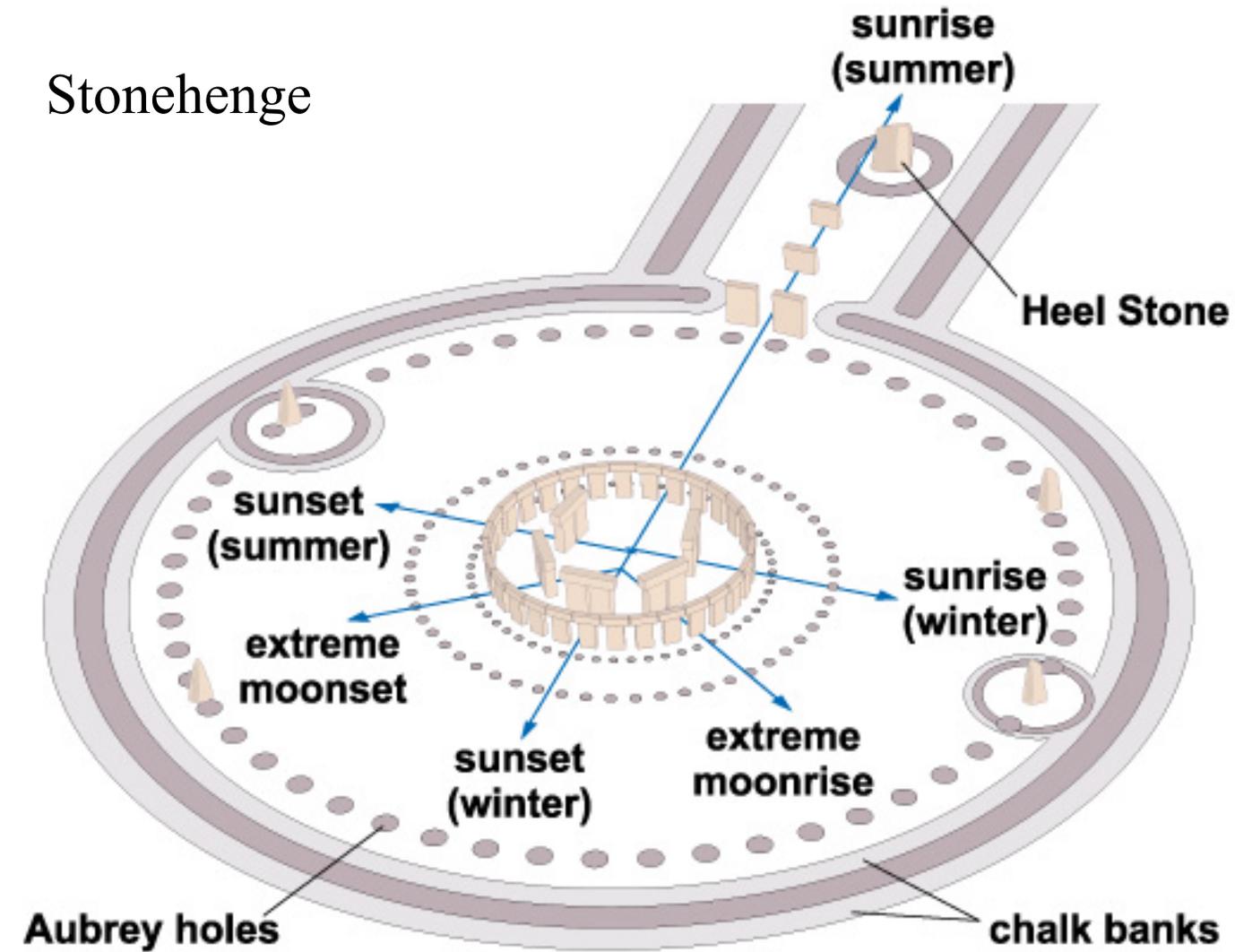




England: Stonehenge (completed around 1550 B.C.)

7

Stonehenge



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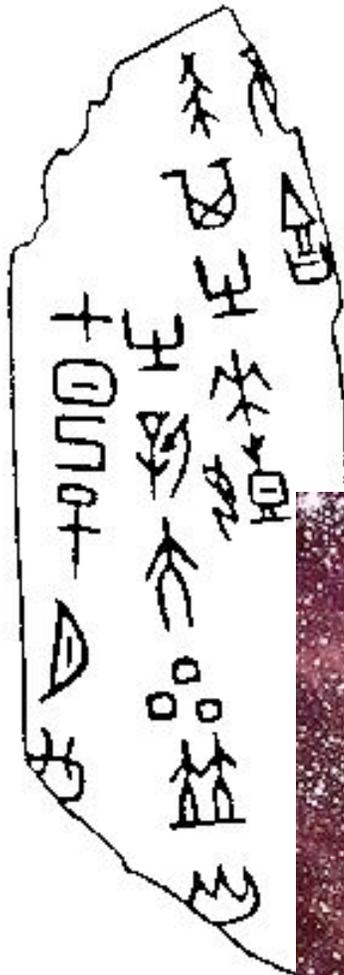
SW United States: “Sun Dagger” marks summer solstice



Scotland: 4,000-year-old stone circle; Moon rises as shown here every 18.6 years.



South Pacific: Polynesians were very skilled in the art of celestial navigation.



"On the Jisi day, the 7th day of the month, a big new star appeared in the company of the Ho star."

SN 1987A



...the new star dwindled."

China: earliest known records of supernova explosions (1400 B.C.)

Ancient Cosmology: A Flat Earth



Here there
be
dragons!



World Map of Hecataeus of Miletus (c. 500 BC)



Artist's reconstruction of the Library of Alexandria

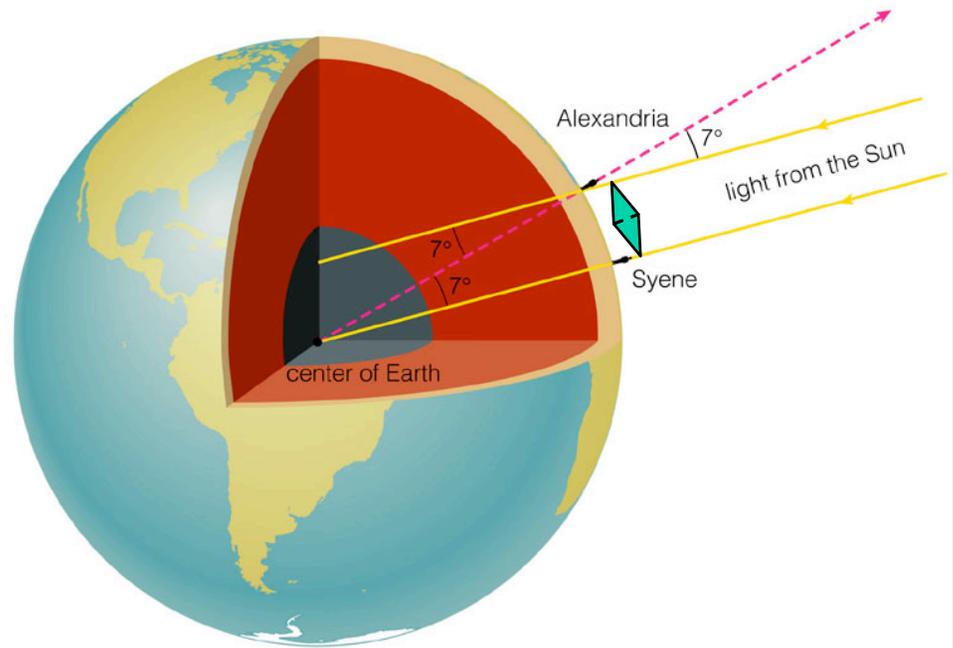
Eratosthenes became the third librarian at Alexandria under Ptolemy III in the Hellenistic period following the conquests of Alexander the Great. Ptolemy I had been one of Alexander's generals, and had taken Egypt as his own after Alexander's untimely death.

Eratosthenes measures the Earth (c. 240 B.C.)

Measurements:

Syene to Alexandria

- distance $\approx 5,000$ stadia
- angle = 7°
- i.e, $7/360$ of the circumference



Calculate circumference of Earth:

$$(7/360) \times (\text{circum. Earth}) = 5,000 \text{ stadia}$$

$$\Rightarrow \text{circum. Earth} = 5,000 \times 360/7 \text{ stadia} \approx 250,000 \text{ stadia}$$

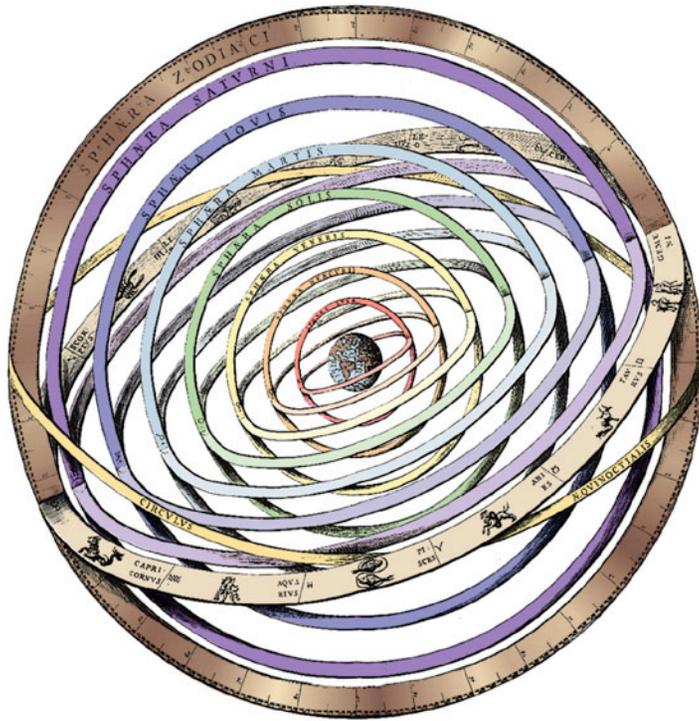
Compare to modern value ($\approx 40,100$ km):

$$\text{Greek stadium} \approx 1/6 \text{ km} \Rightarrow 250,000 \text{ stadia} \approx 42,000 \text{ km}$$

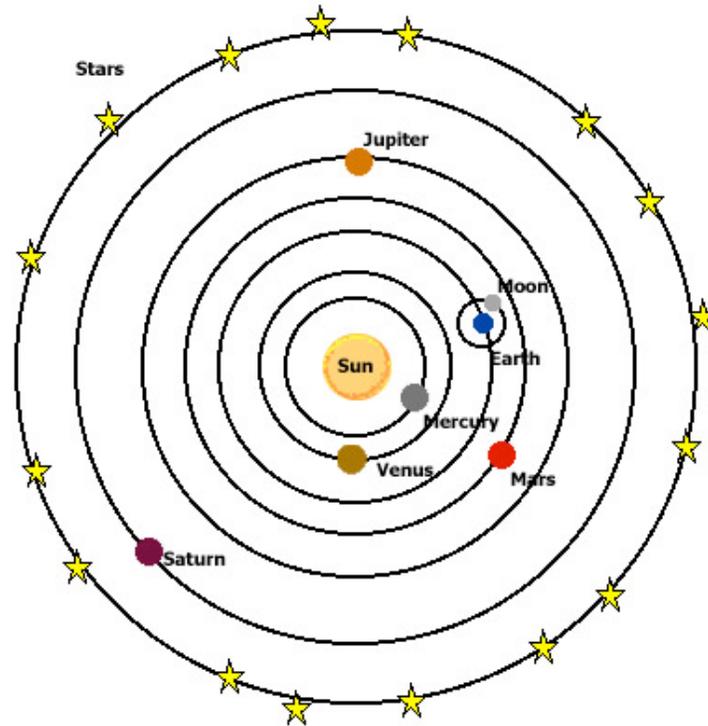
It was known long before Columbus that the Earth is not flat!

Competing Cosmologies

Geocentric
Ptolemaic
Earth at center



Heliocentric
Copernican
Sun at center



Earth- vs. Sun-Centered

To the ancients, which of the following would have been an argument in favor of a Sun-centered cosmology?

A. The Earth is clearly moving

B. The Sun is larger than the Earth, so it should be in the center

C. The Earth is corrupt, so it should move

D. None of the above

E. I don't know

Geocentric

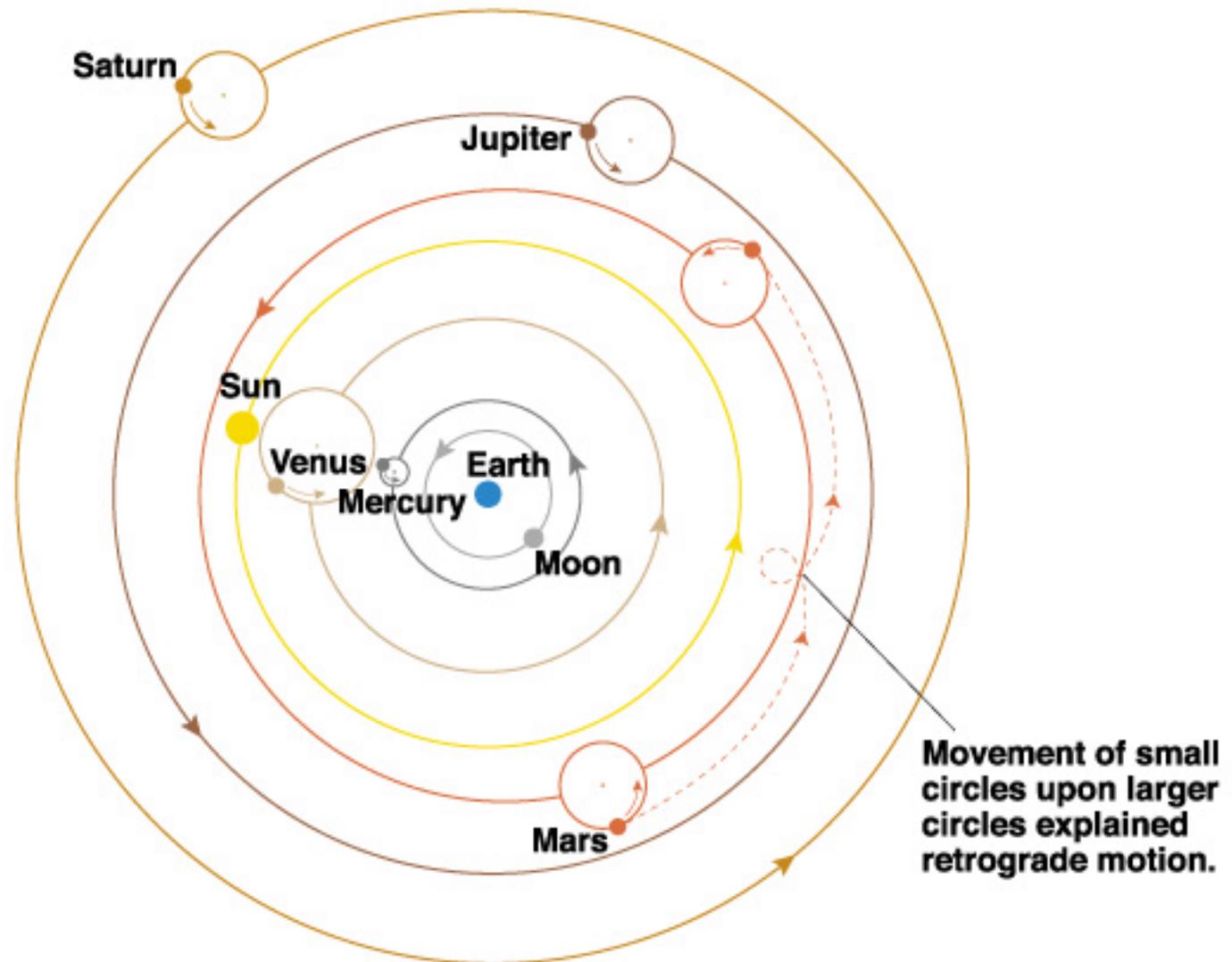


Ptolemy

The most sophisticated geocentric model was that of Ptolemy (A.D. 100–170) — the **Ptolemaic model**:

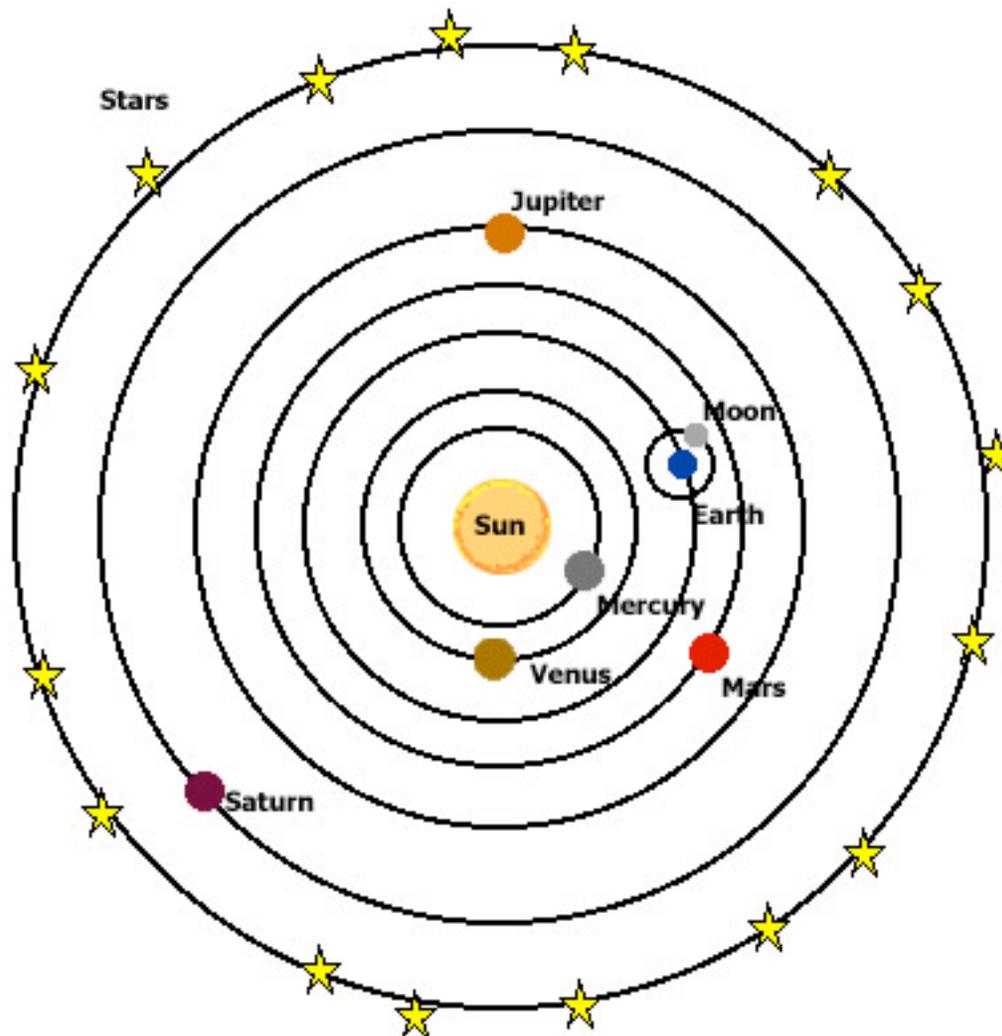
- Sufficiently accurate to remain in use for 1,500 years
 - i.e., predicted correct positions of planets for many centuries
- Arabic translation of Ptolemy's work named *Almagest* (“the greatest compilation”)

Geocentric Cosmology



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Heliocentric Cosmology



Heliocentric

Copernicus (1473–1543):



- He proposed the Sun-centered model (published 1543).
- He used the model to determine the layout of the solar system (planetary distances in AU).

But . . .

- The model was no more accurate than Ptolemaic model in predicting planetary positions, because it still used perfect circles.

Competing Cosmologies

Geocentric

Ptolemaic

Earth at center

Heliocentric

Copernican

Sun at center

The sun is the source of light in both models

Explains

- **Motion of Sun**
- **Motion of Moon**
- **Solar and Lunar Eclipses**
- **Phases of Moon**

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Hard to tell the difference!

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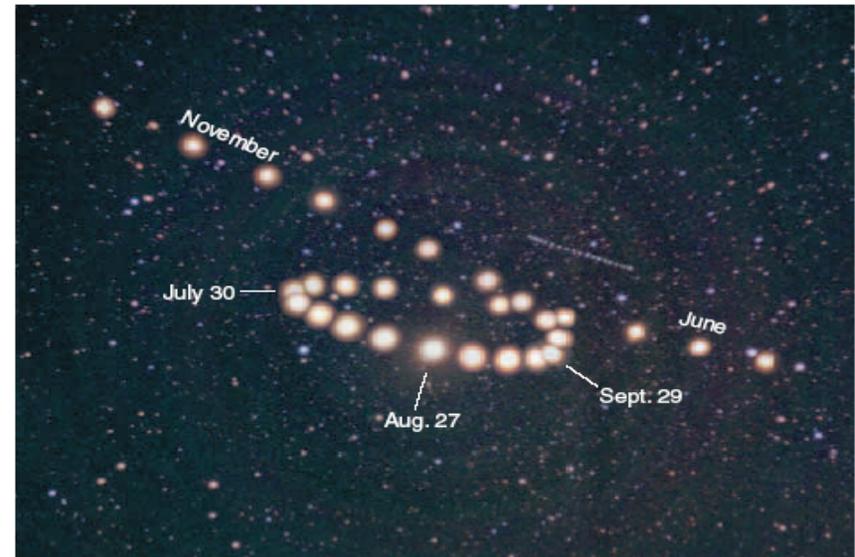
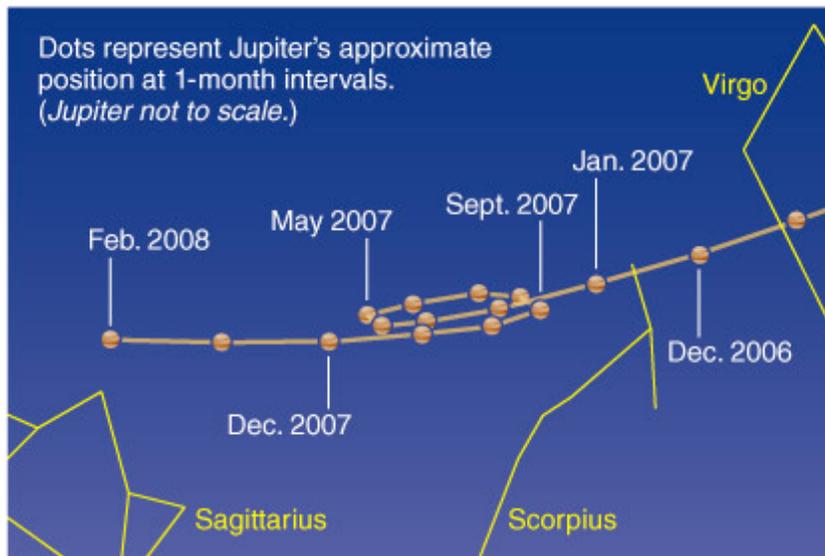
Retrograde Motion

Needs epicycles

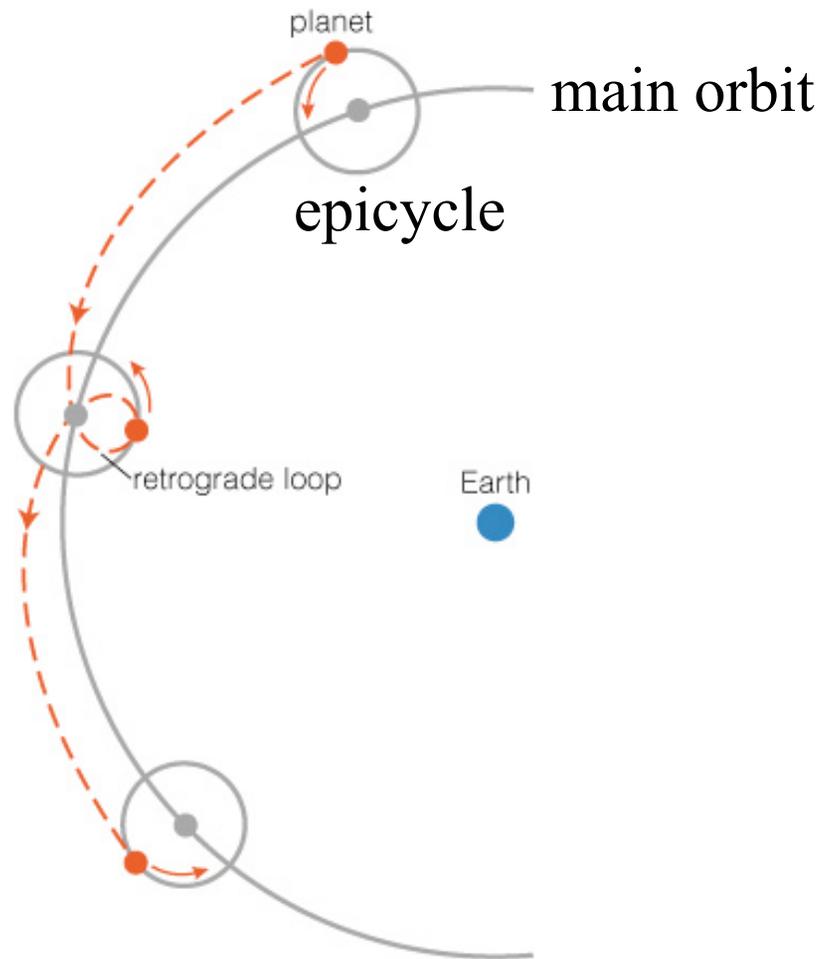
Consequence of Lapping

Retrograde motion

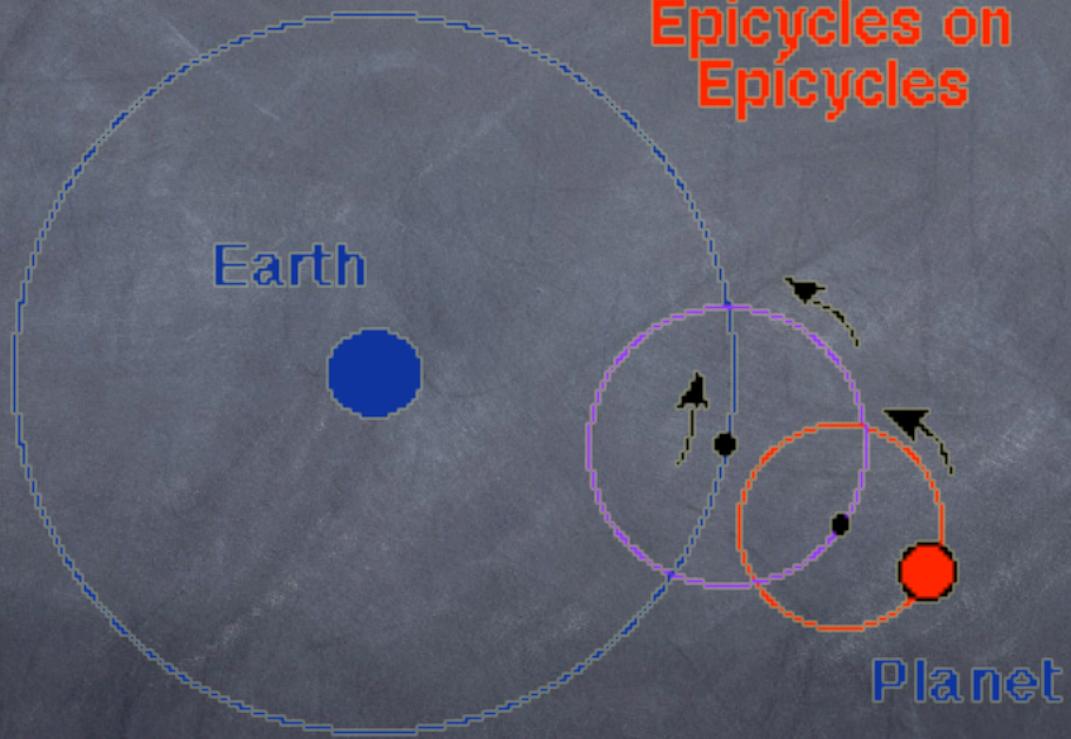
- Planets usually move slightly *eastward* from night to night relative to the stars.
- But, sometimes they go *westward* relative to the stars for a few weeks: **apparent retrograde motion.**



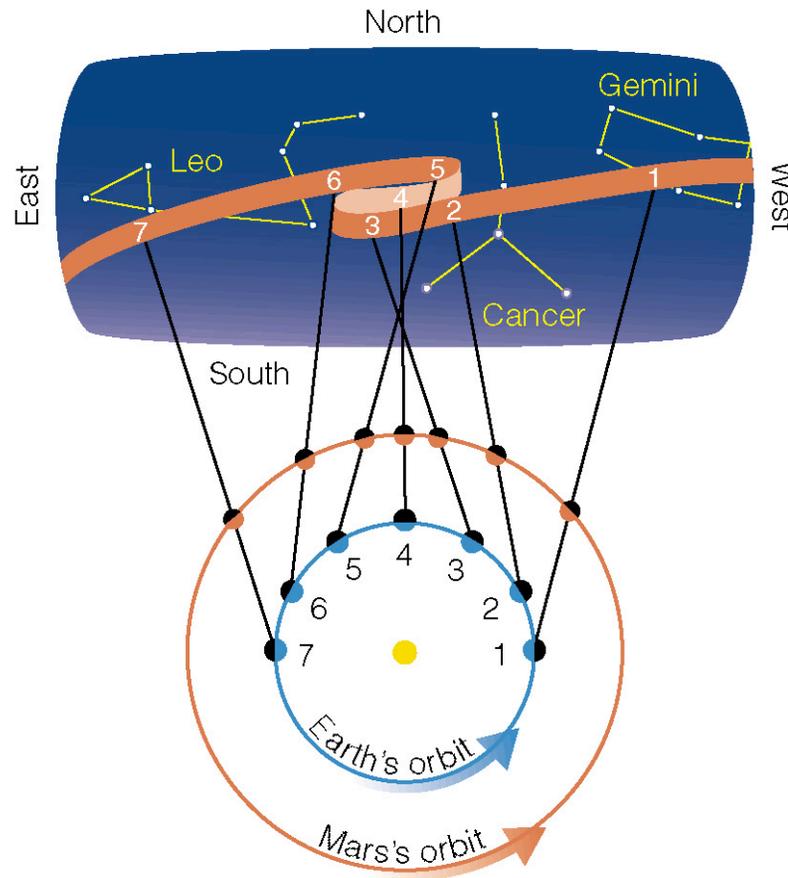
In the **Ptolemaic** model, planets *really do* go backwards.



Epicycles on Epicycles



In the **Copernican** model, retrograde motion is a consequence of one planet (Earth) “lapping” another in its orbit.



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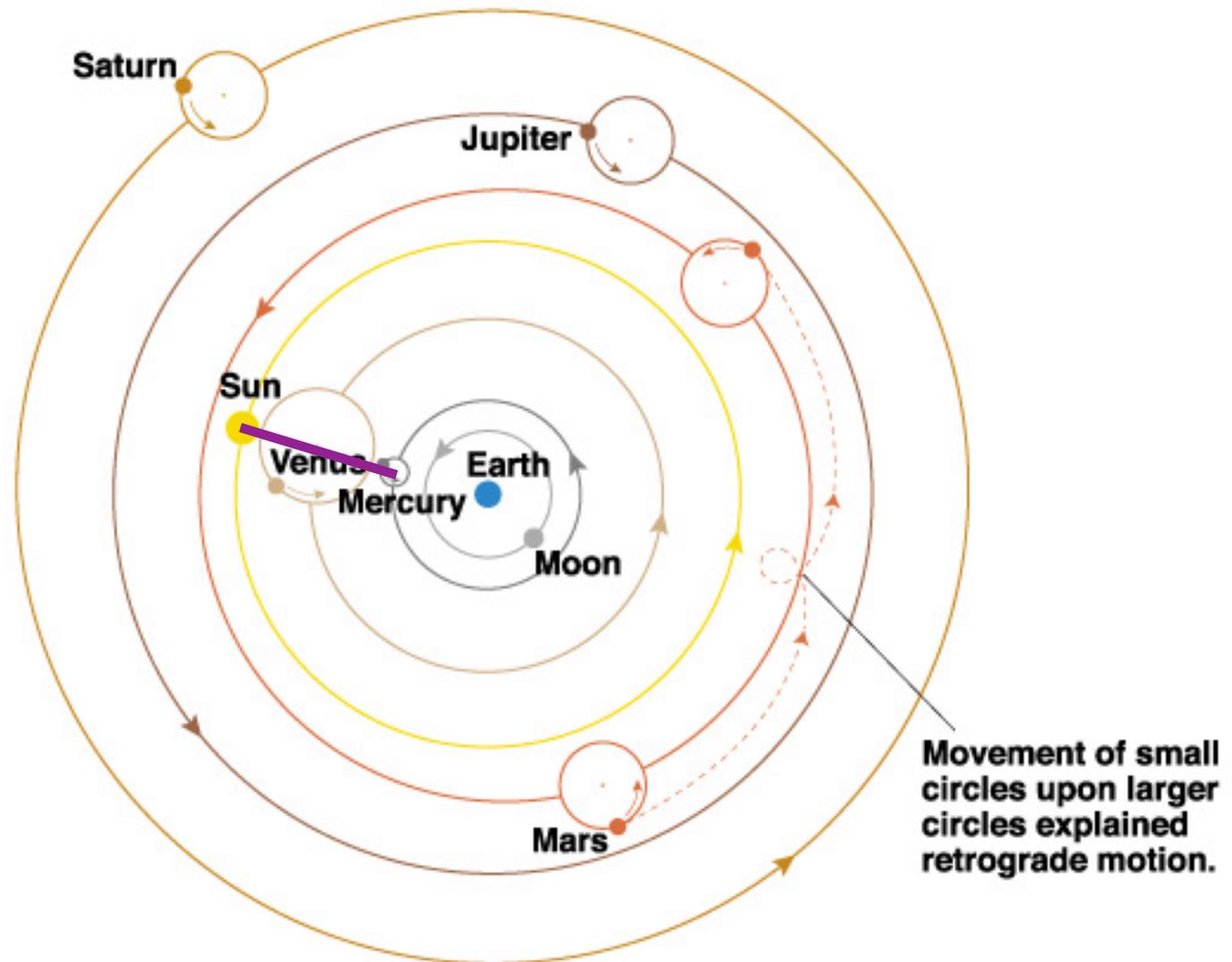
Consequence of Lapping

Inferiority of Mercury & Venus

Must tie to sun

Interior to Earth's Orbit

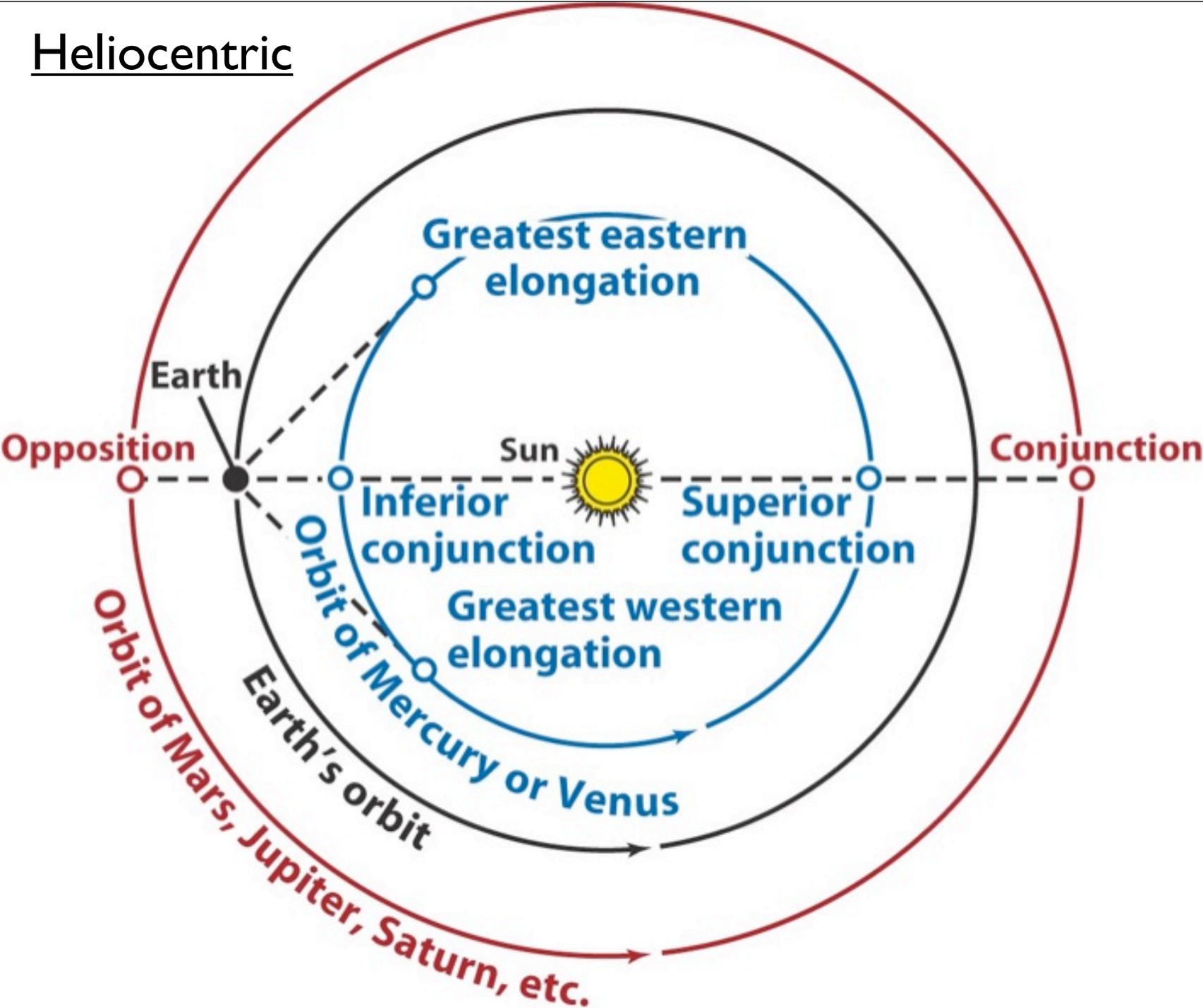
Geocentric Cosmology



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Mercury & Venus always close to sun on the sky

Heliocentric



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Copernican

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Retrograde Motion

Needs epicycles

Consequence of Lapping

Inferiority of Mercury & Venus **more natural**

Must tie to sun

Interior to Earth's Orbit

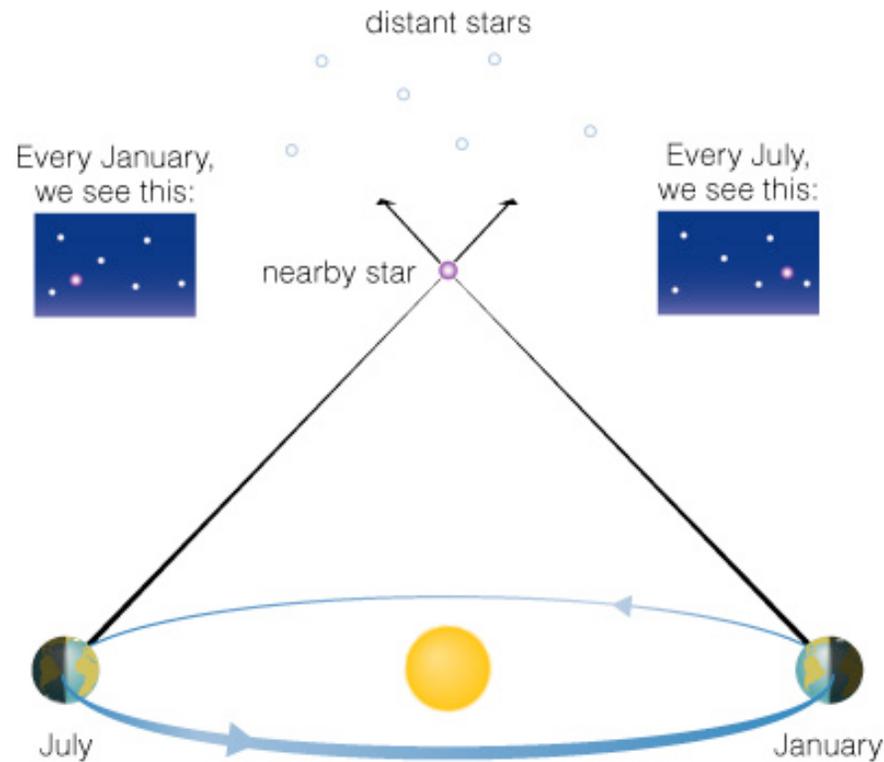
Predicts

- No parallax
- Venus: crescent phase only

- Parallax
- Venus: all phases

Parallax

If the Earth moves around the sun, the positions of stars should shift in reflex to that motion.



- The ancients could not detect **stellar parallax**.

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Retrograde Motion

Needs epicycles

Consequence of Lapping **nicer**

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Must tie to sun

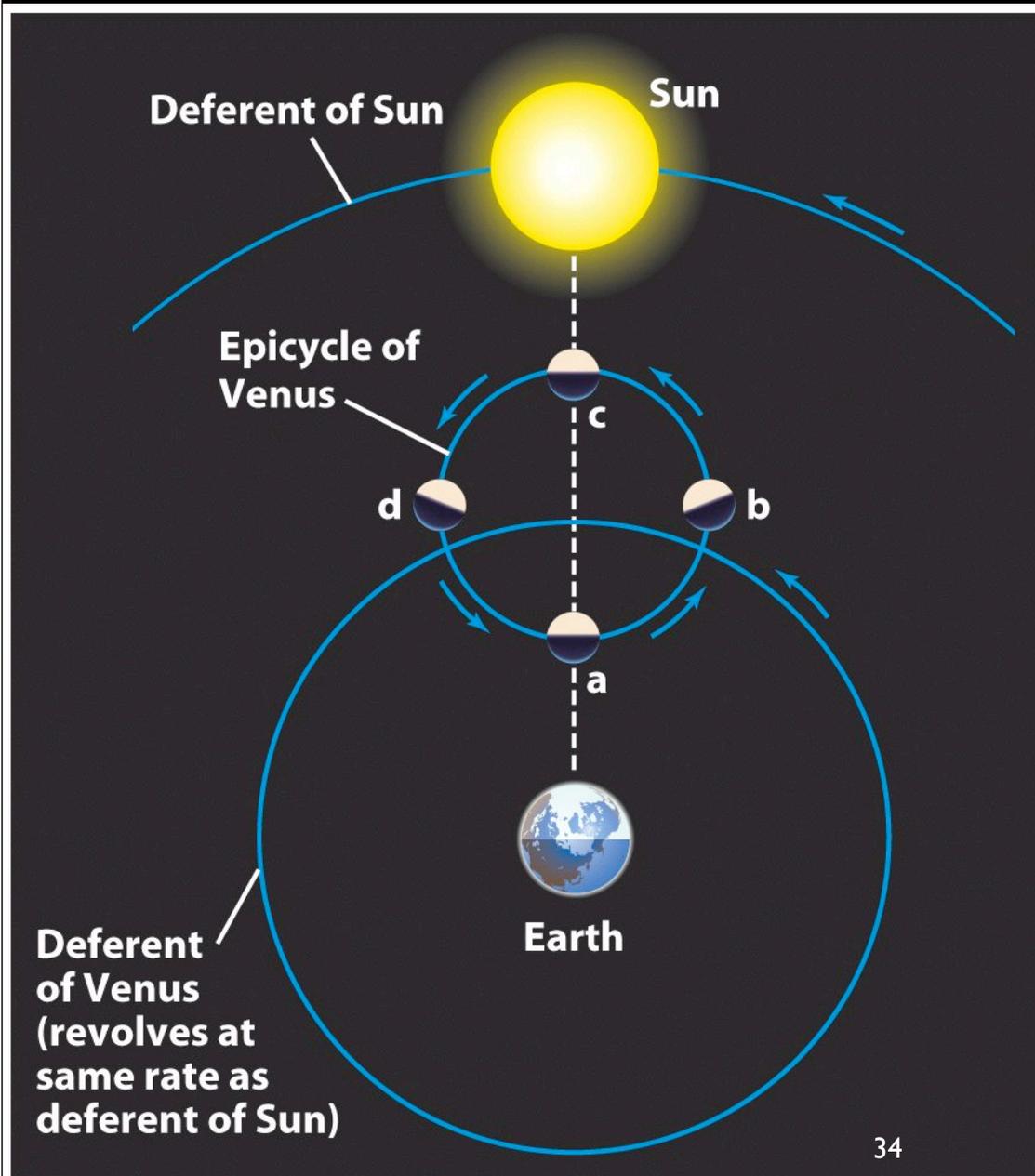
Interior to Earth's Orbit **nicer**

Predicts

- No parallax ✓
- Venus: crescent phase only

- Parallax **X**
- Venus: all phases
unknown to ancients

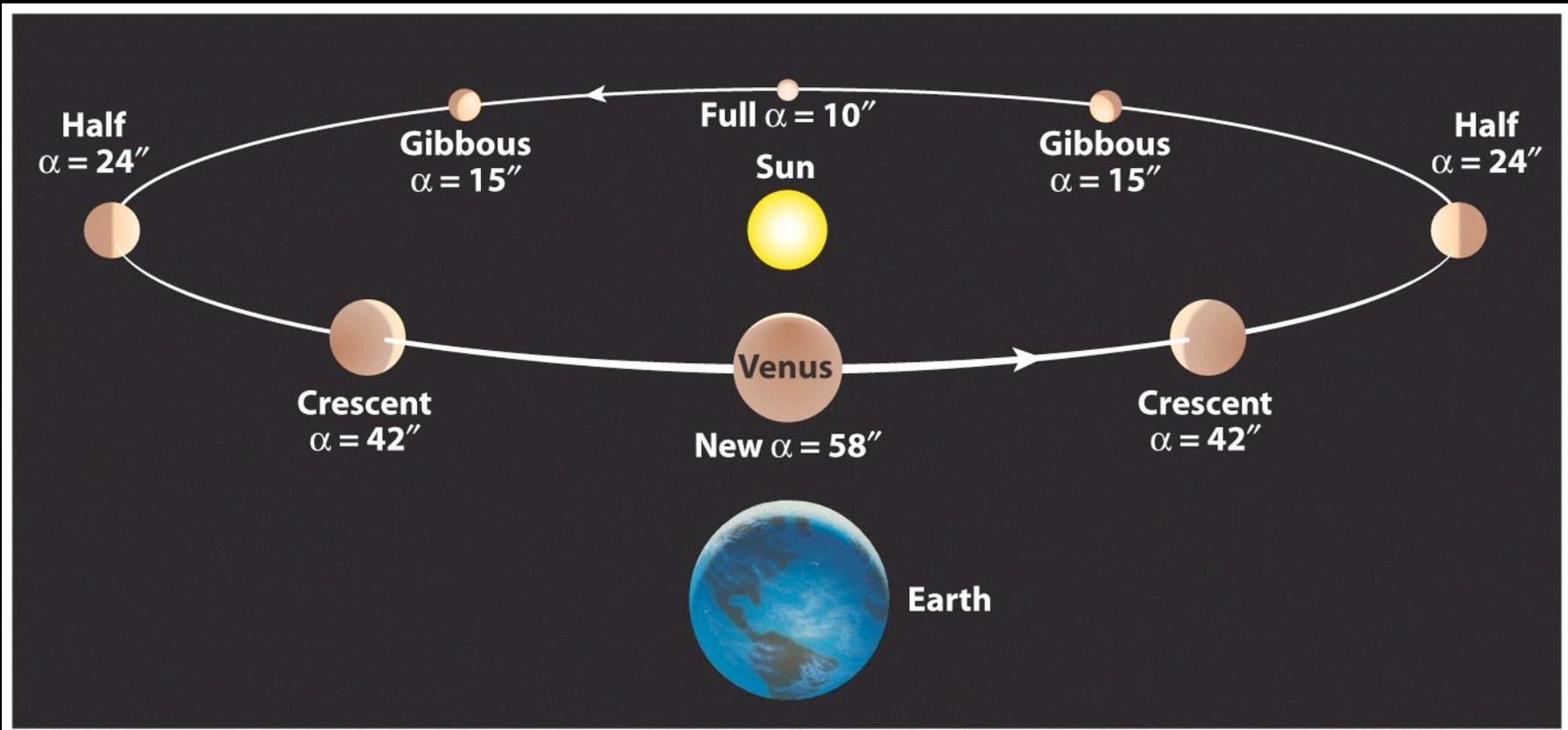
Geocentric



Only crescent phase can be observed - never full or even gibbous

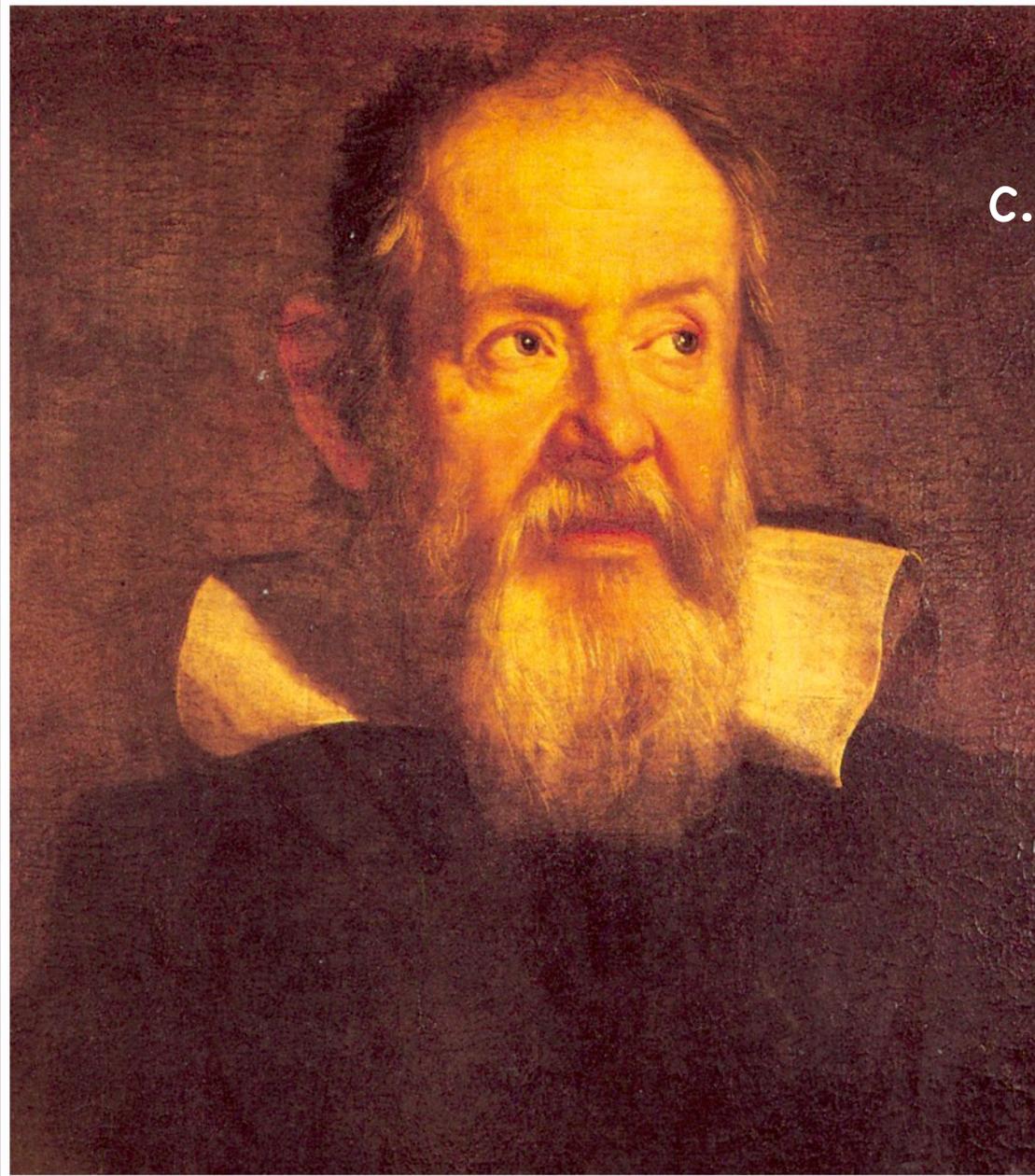
Heliocentric

The full range of phase can be observed
- from crescent to full



Phases of Venus

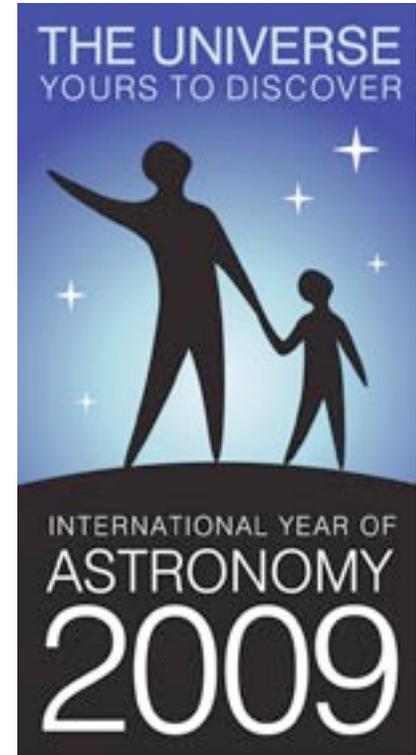
Galileo

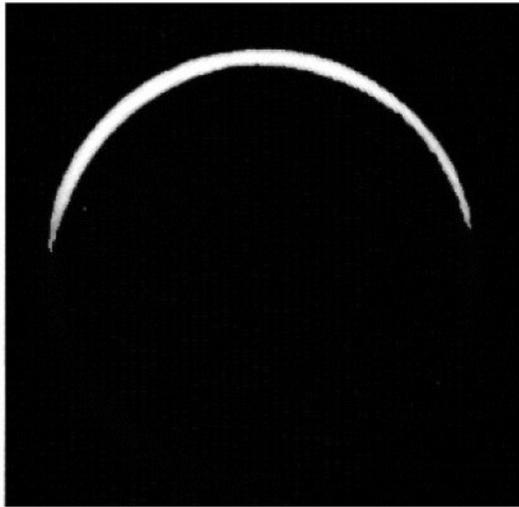


c. 1564–1640

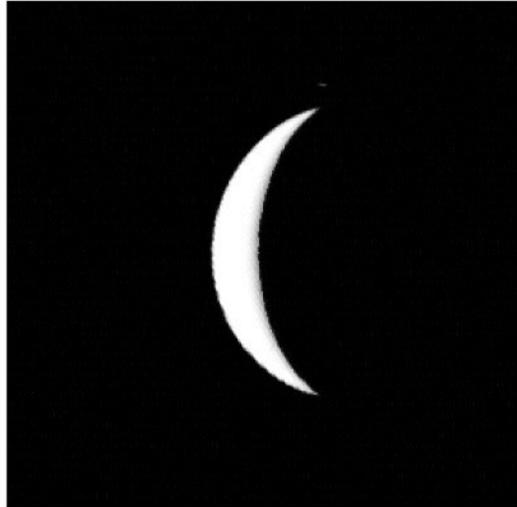
First telescopic astronomical observations

- First use of telescope for astronomy
- in 1609
- 400 years ago!



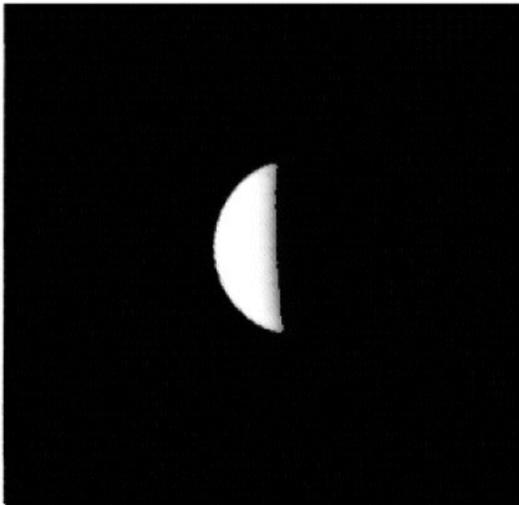


$\alpha = 58^\circ$

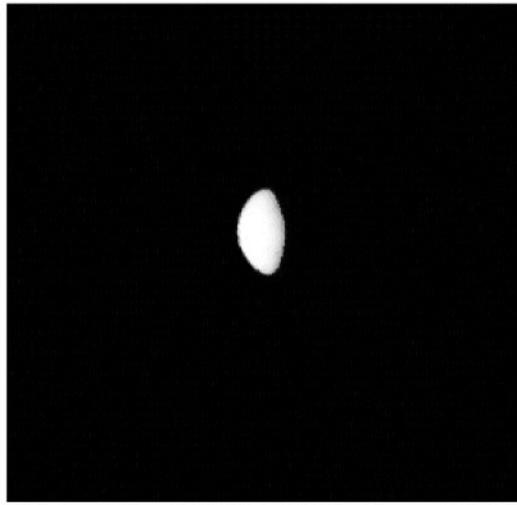


$\alpha = 42^\circ$

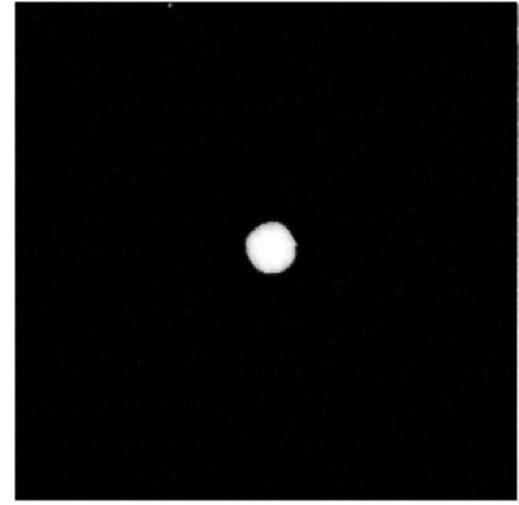
Phase and angular size of Venus depend on elongation



$\alpha = 24^\circ$



$\alpha = 15^\circ$



$\alpha = 10^\circ$

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