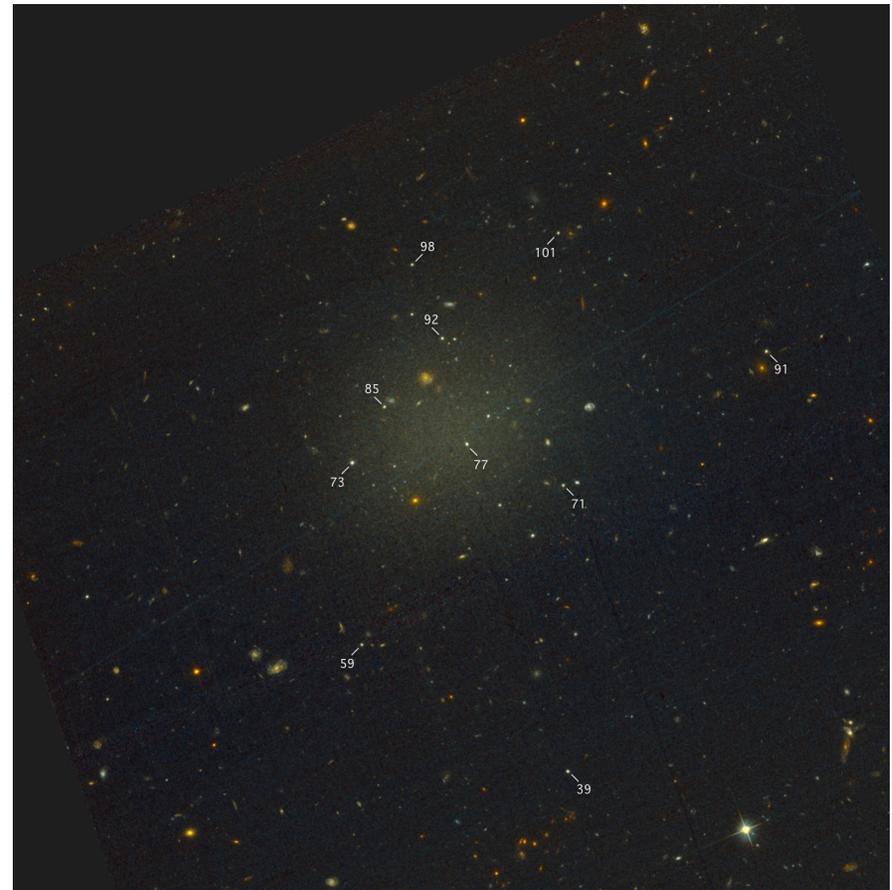


[16] Other Galaxies (3/29/18)

Upcoming Items

1. Homework #4 due on Tuesday.
2. Read Ch. 20.2 (and review the last part of Ch. 15.2 on variable stars) for next class and do the self-study quizzes.

NGC 1052-DF2

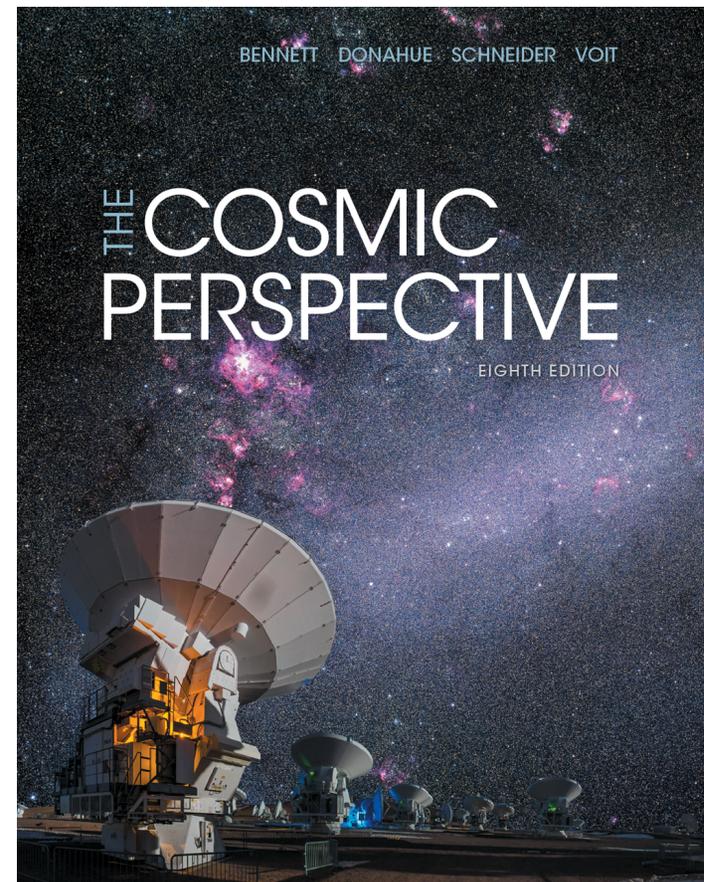


LEARNING GOALS

Ch. 20.1

For this class, you should be able to...

- ... classify a galaxy as a spiral, elliptical, or irregular based on morphology, and further estimate the subtype of spiral (barred or not, tightness of arms, etc.) or elliptical (degree of roundness);*
- ... contrast the relative proportion of gas, as well as stars of each spectral type, between galaxy types;*



Any astro questions?

In-Class Quiz

1. Why does ongoing star formation lead to a blue-white appearance?

- A. There aren't any red or yellow stars.
- B. Short-lived blue stars outshine others.
- C. Gas in the disk scatters blue light.



2. In which object would you expect to find K stars?

- A. Object 1 only.
- B. Object 2 only.
- C. Both object 1 and object 2.
- D. Neither object 1 nor object 2.



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Hubble
Ultra
Deep
Field



Elliptical Galaxy

A large, bright, yellowish elliptical galaxy with a diffuse, glowing core and a smooth, rounded shape. It is surrounded by a few smaller, fainter galaxies in a dark field.



Irregular Galaxies

A cluster of several small, blue, irregularly shaped galaxies with fragmented and clumpy structures. They are scattered across a dark field.



Spiral Galaxy

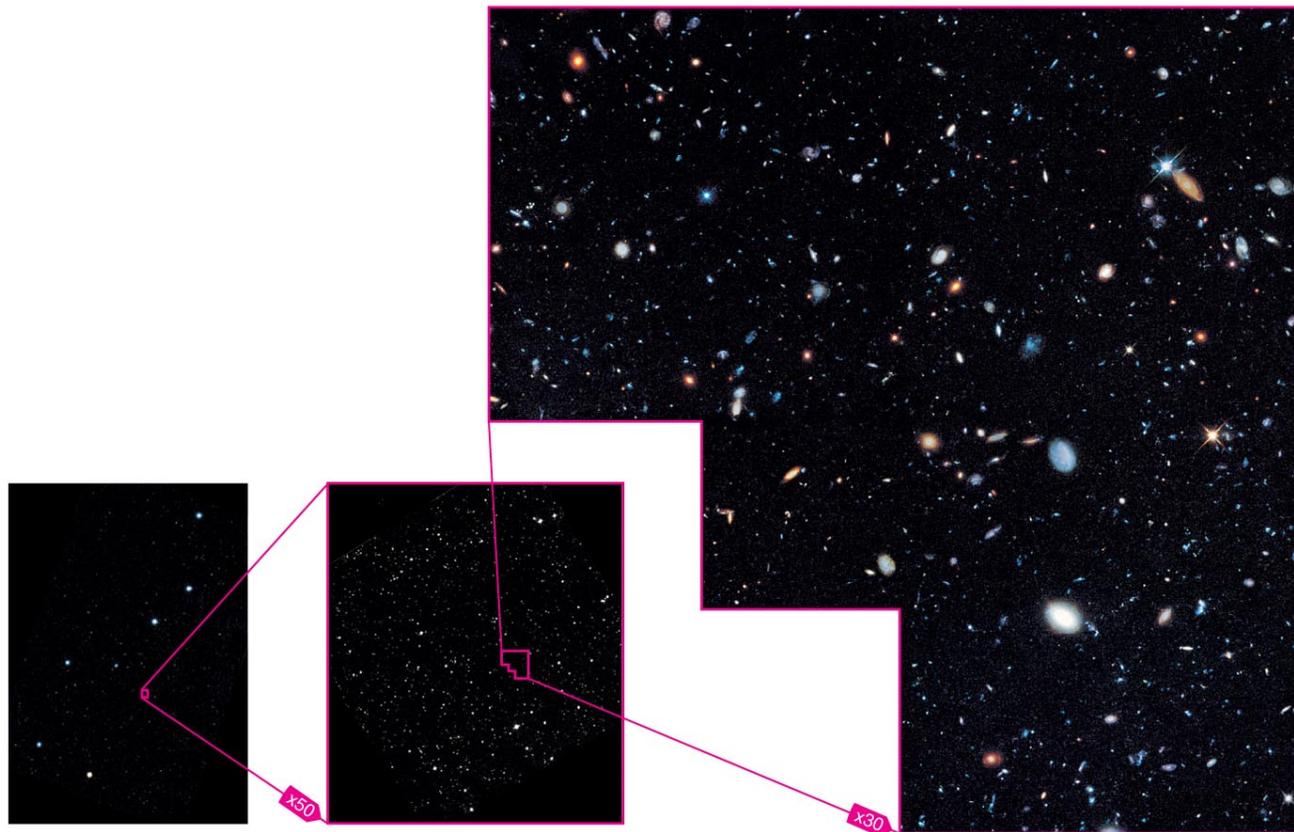
A large, bright, yellowish spiral galaxy with a prominent central bulge and several distinct, winding spiral arms. It is surrounded by a few smaller, fainter galaxies in a dark field.

Other Galaxies

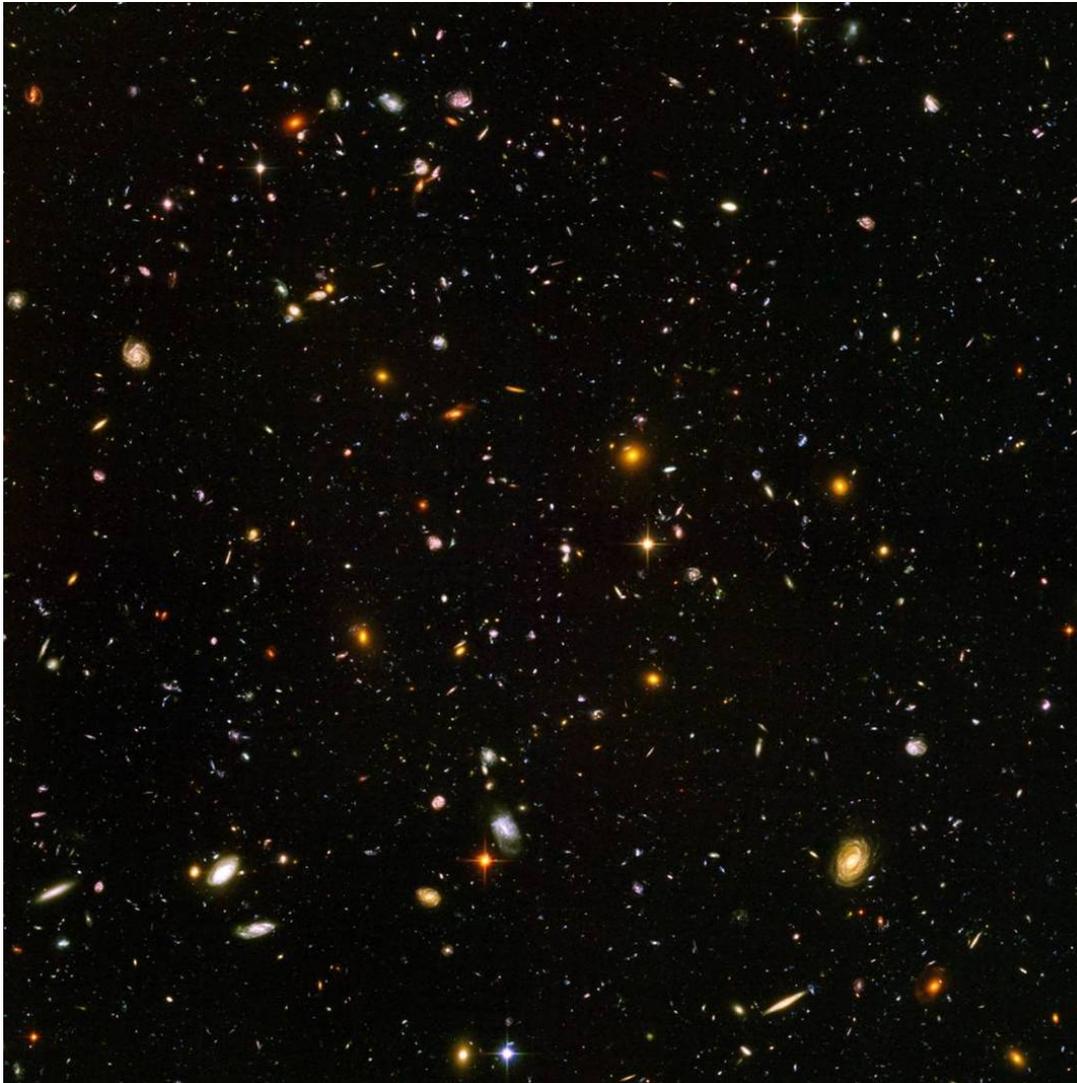
- Spirals (e.g., Milky Way).
 - Have disk and spheroid components, often a central bar.
 - Generally massive ($M \sim 10^{9-12} M_{\odot}$), have ISM in disk.
 - Contain both old and young stars: overall blue-white appearance.
- Ellipticals (e.g., M87).
 - All spheroid, no disk, huge range of masses ($M \sim 10^{5-13} M_{\odot}$).
 - Very little dust or cool gas, mostly old red stars: *red and dead*.
 - Biggest are usually in large clusters.
- Irregulars are in between (e.g., Small Magellanic Cloud).
 - Irregular shape, intermediate masses ($M \sim 10^{8-10} M_{\odot}$), Pop I stars.
- Use Hubble's classification scheme for galaxies.
- Galaxies often come in groups.

Hubble Deep Field

- Our deepest images of the universe show a great variety of galaxies, some of them billions of light-years away.



Galaxies and Cosmology



- A galaxy's age, its distance, and the age of the universe are all closely related.
- The study of galaxies is thus intimately connected with **cosmology**—the study of the structure and evolution of the universe.

Group Q: Why do Spirals have Gas?

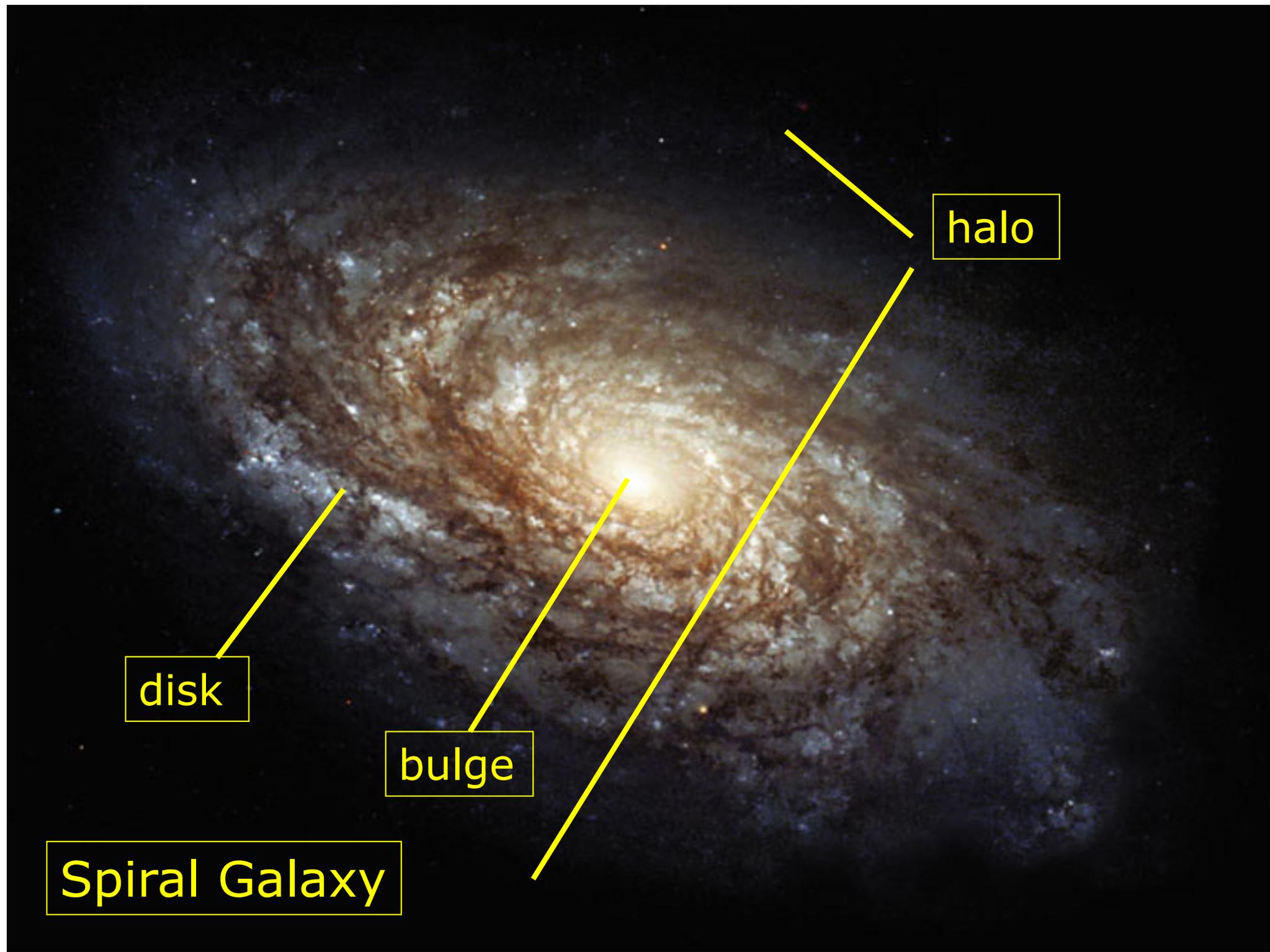
- You know that spirals have gas and significant ongoing star formation; ellipticals have little gas and little ongoing star formation
- But why do the spirals still have lots of gas?
- In particular, if we asked how long it would take for existing gas to cool, contract, and form stars, it is **way** shorter than the ages of the galaxies
- Thus almost all the gas in spirals should have formed stars, billions of years ago
- What might have prevented that?

Galaxy Types

- There are three basic varieties of galaxies:
 1. Disk/spiral galaxies (e.g., Milky Way).
 2. Elliptical galaxies (e.g., M87).
 3. Irregular galaxies (e.g., Small Magellanic Cloud, M82).
- We start by discussing the general properties of each type.

Properties of Disk (Spiral) Galaxies

- We have already discussed the Milky Way.
- More generally, spirals possess two main components:
 1. **Disk component:** Flattened structure orbiting center with organized rotation. Contains old and young stars, as well as cold gas and dust.
 - Coldest gas & youngest stars found near spiral arms (density waves).
 2. **Spheroidal component:** spheroidal distribution of old stars on random orbits with little dust or gas in between.
 - Includes the bulge and halo.
 - Like mini elliptical galaxies...
- Spiral galaxies can be *barred* or *unbarred*.



disk

bulge

halo

Spiral Galaxy

Disk component:

stars of all ages, many gas clouds.

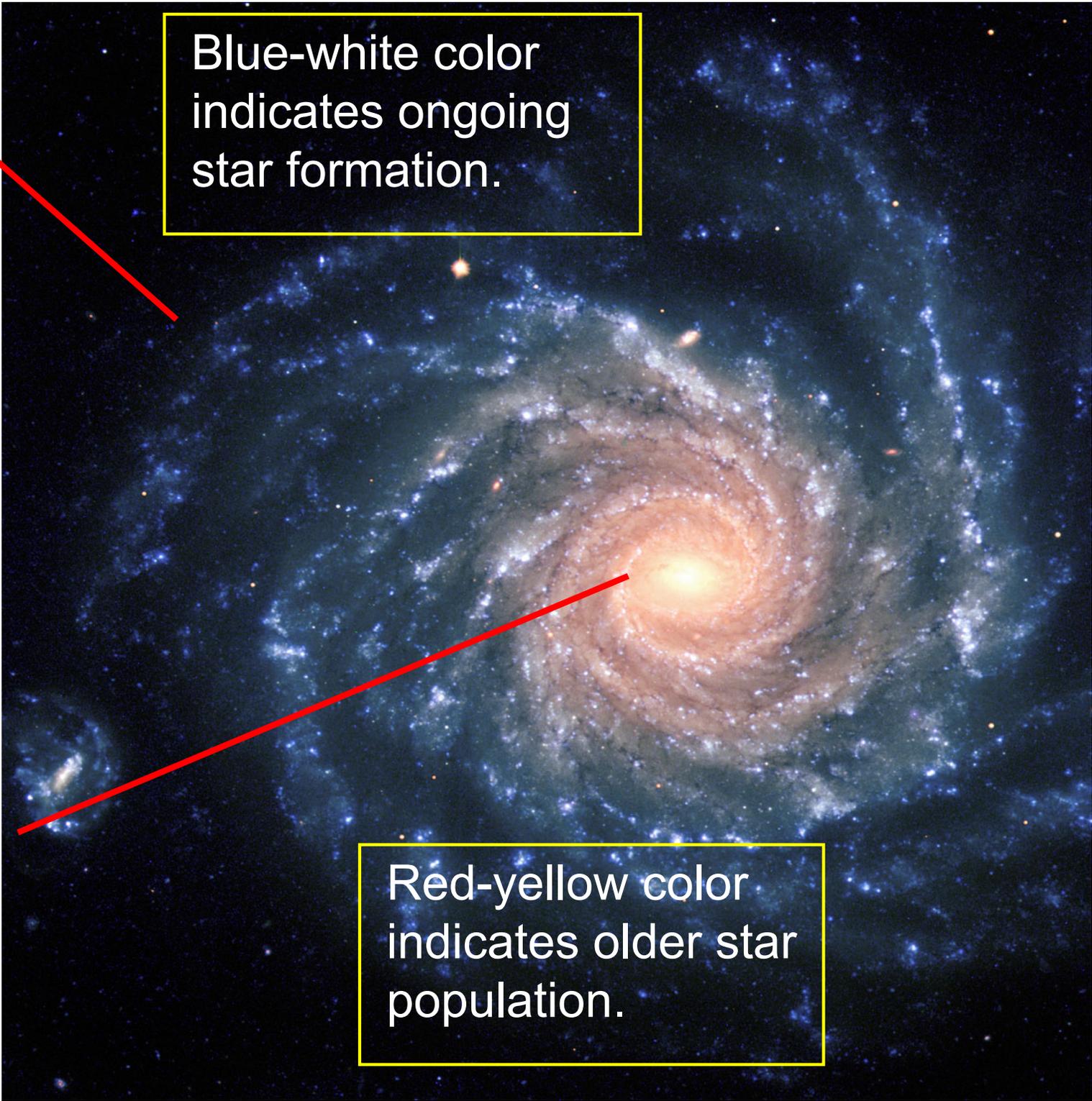


Spheroidal component:

bulge and halo, old stars, few gas clouds.

**Disk
Component:**
stars of all
ages,
many gas
clouds.

**Spheroidal
Component:**
bulge & halo,
old stars,
few gas
clouds.



Blue-white color
indicates ongoing
star formation.

Red-yellow color
indicates older star
population.



Barred Spiral Galaxy:

Has a bar of stars across the bulge.



Lenticular Galaxy:
Has a disk like a spiral galaxy but much less dusty gas (intermediate between spiral and elliptical).

Properties of Elliptical Galaxies

- Large “swarms” of stars orbiting with random orientations in the galaxy’s gravitational field: **spheroidal shape**.
- The randomness of the orbits means that the galaxy as a whole has **very little net rotation**.
- The **stellar population is old** (Pop. II): in general, there are no massive stars (they have all gone supernova).
- Usually **very little dust or cold gas** present... explains lack of star formation/young stars.
- Some **elliptical galaxies can be very massive** (the most massive galaxies in the universe are elliptical).
- There are also a lot of **dwarf ellipticals**... may have only a few million stars. Some of these may be the oldest galaxies (“fossils” of the first phase of galaxy formation).

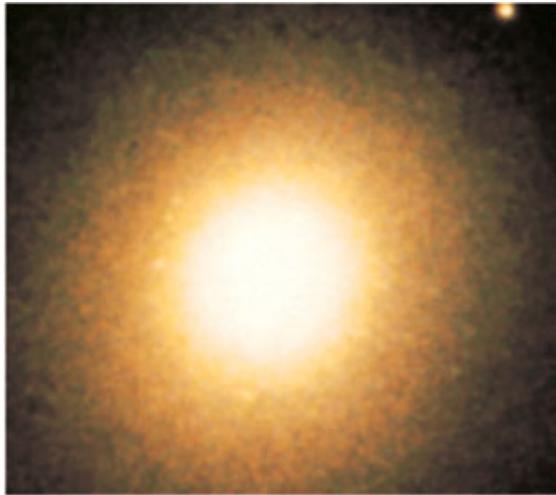


a M87, a giant elliptical galaxy in the Virgo Cluster, is one of the most massive galaxies in the universe. The region shown is more than 300,000 light-years across.

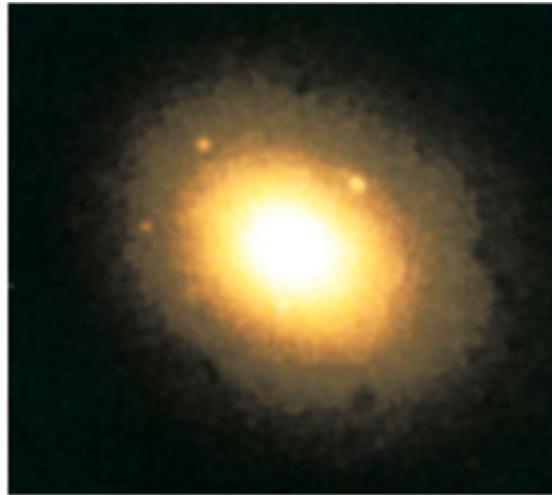
Elliptical Galaxy:

All spheroidal component, virtually no disk component.

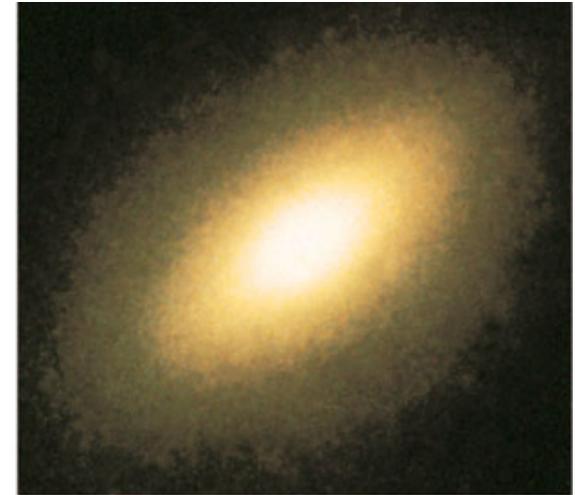
Red-yellow color indicates older star population:
RED and DEAD.



(a) E0 (M105)



(b) E3 (NGC 4365)

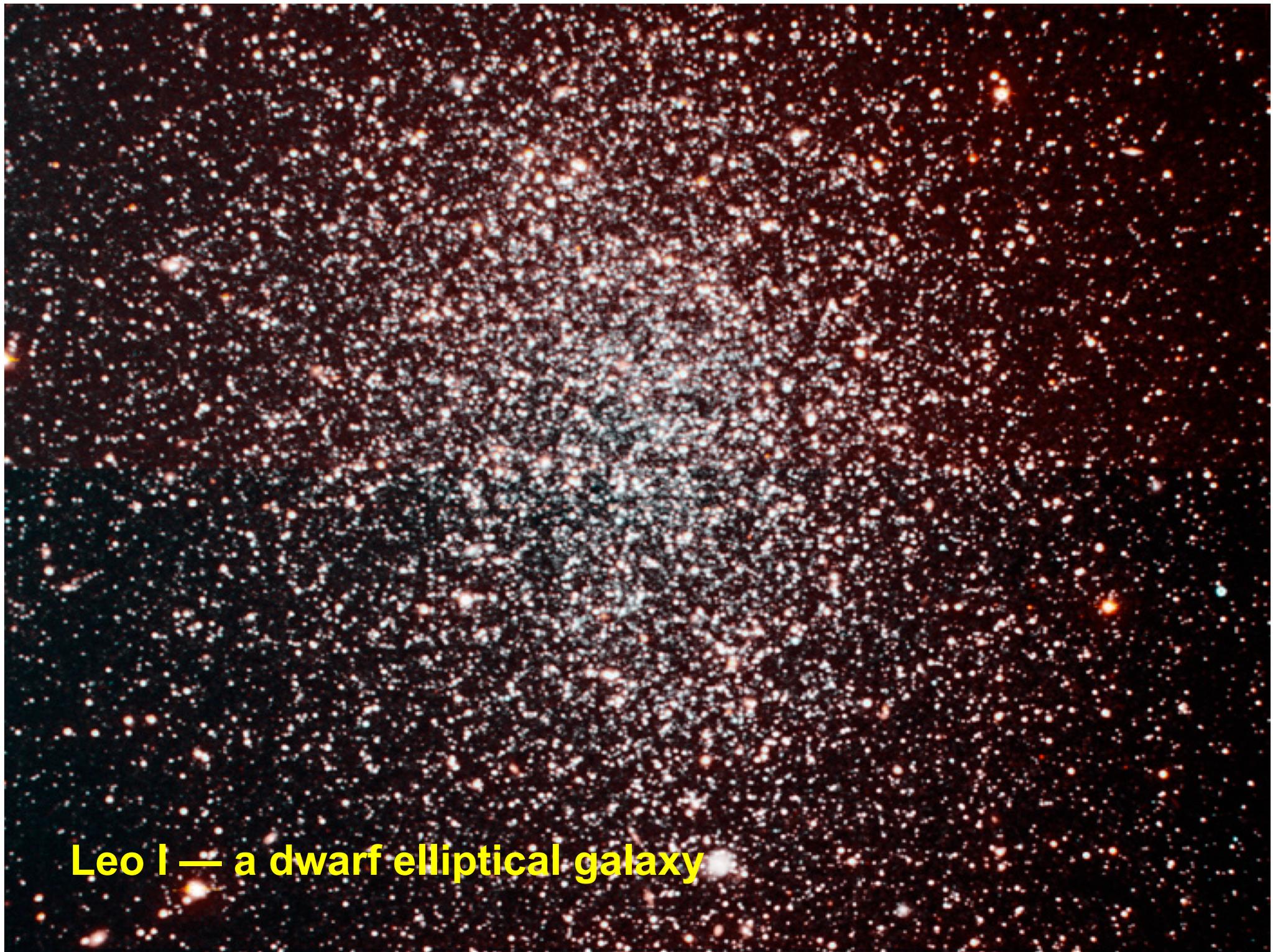


(c) E6 (NGC 3377)

Shape classification: $E\#$, where $\# = 10 \times (1 - b/a)$; a = major axis length of image, b = minor axis.

But note: we observe the projected shape!

Can we see an intrinsically elongated galaxy as nearly circular? Can we see an intrinsically spherical galaxy as highly elongated?



Leo I — a dwarf elliptical galaxy

Properties of Irregular Galaxies

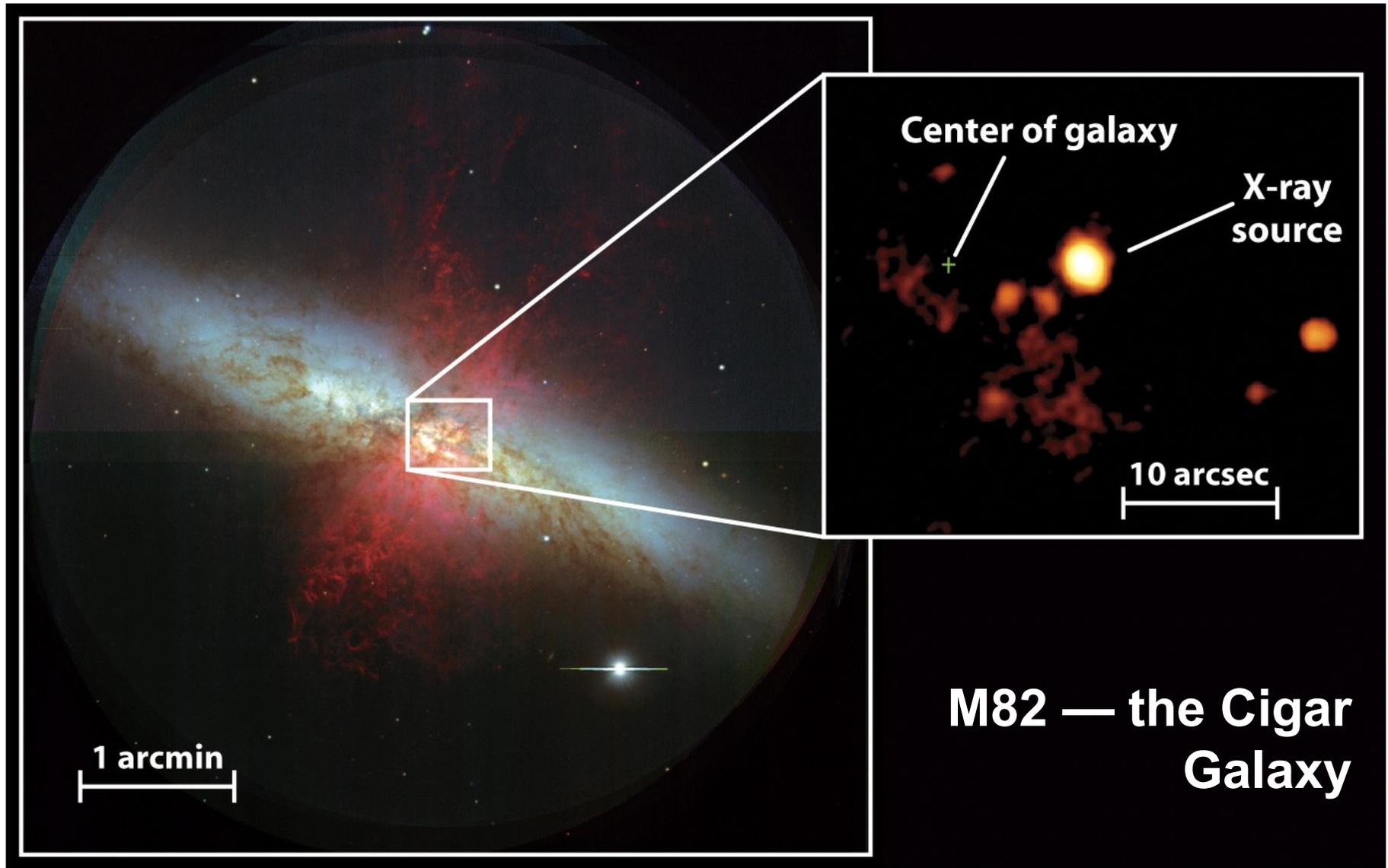
- Only hints of organized structure, or highly distorted objects caused by the collision of galaxies or violent activity in their centers.
- Often have many H II and star-forming regions.
- Examples: SMC, M82 (starburst galaxy).

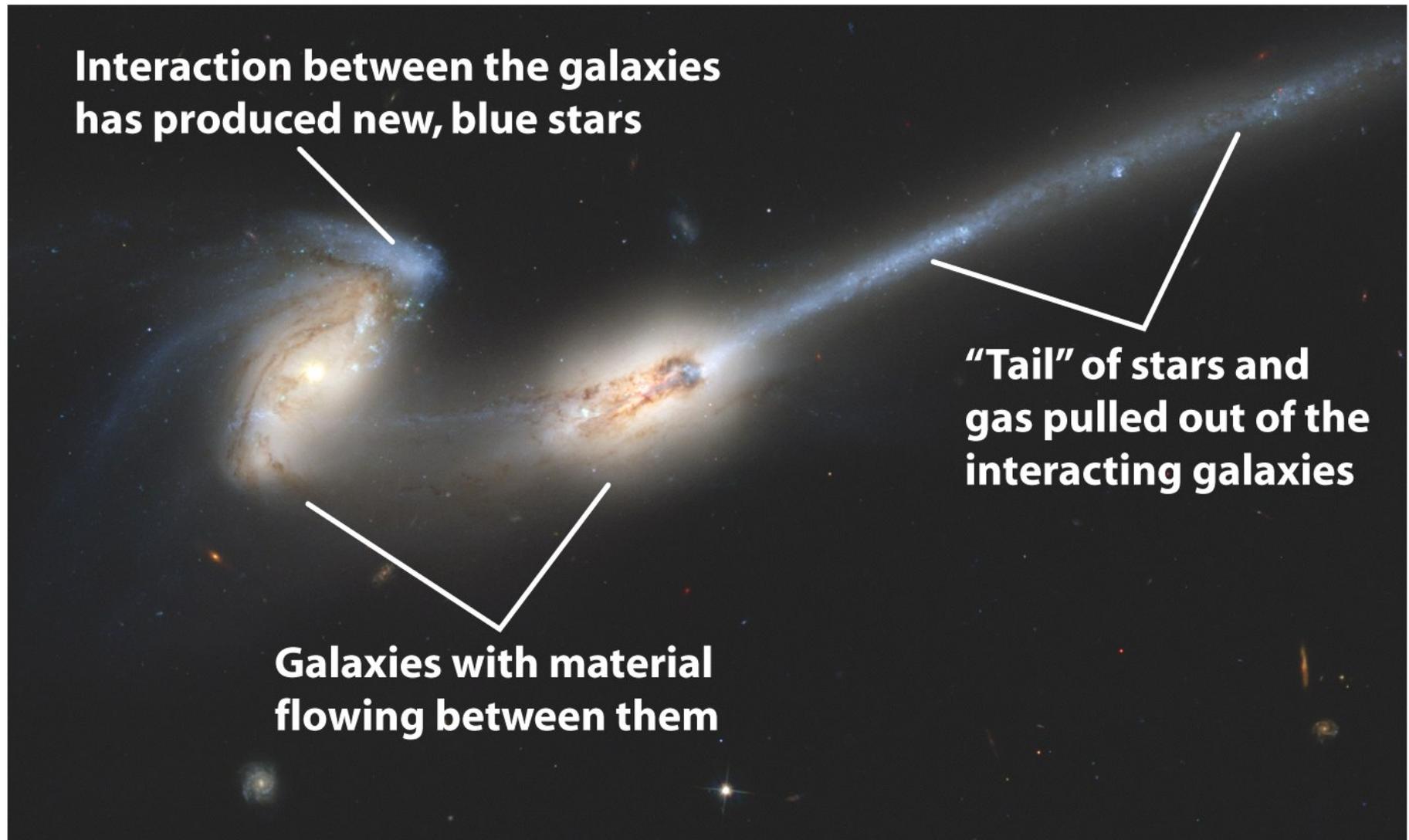


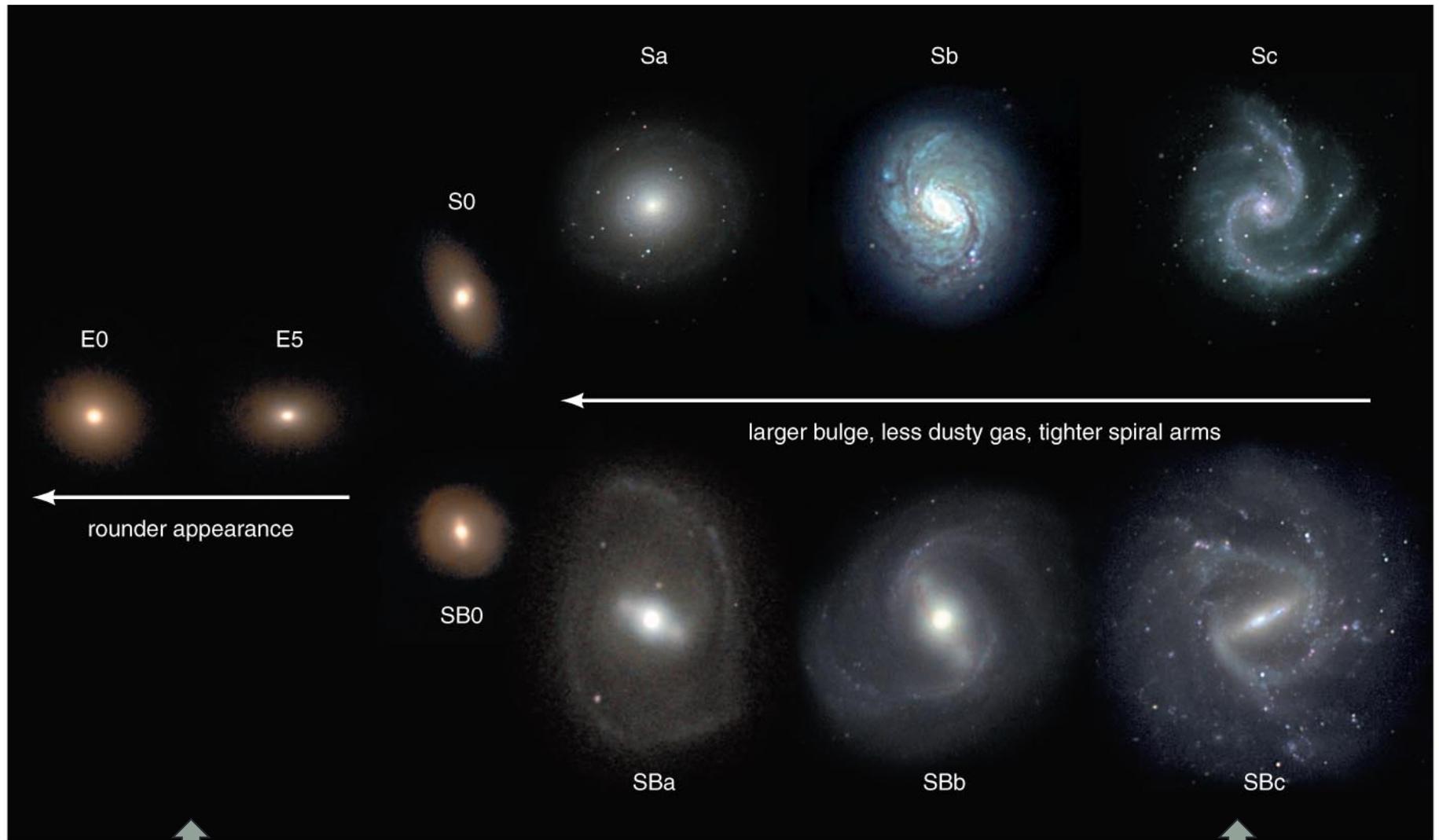
Irregular Galaxy:

Neither spiral nor elliptical.

Blue-white color indicates ongoing star formation.







Spheroid
Dominates

Hubble's Galaxy Classes
(irregulars not shown)

Disk
Dominates