Key points from Lecture 20 of ASTR 350

- 1. Stellar-mass black holes are born from core-collapse supernovae, and don't acquire much more mass later even if they accrete from a companion. In contrast, supermassive black holes start from much smaller "seeds" and grow a lot from there.
- 2. Supermassive black holes are tightly correlated with their host galaxies. This suggests that the two evolved in tandem. This is surprising given that SMBHs are so low-mass and small compared with the galaxies, but it suggests that feedback (probably from jets) from accreting SMBHs affects their galaxies.
- 3. There are some SMBHs that are quite high-mass (~ $10^9 M_{\odot}$) very early (some have been seen at redshift z = 7.5, which is only ~ 700 million years after the Big Bang). Where did they come from?
- 4. Not from primordial black holes shortly after the Big Bang; not enough mass. Not from dark matter accretion, although because that doesn't generate radiation it isn't limited by the Eddington luminosity; dark matter can't easily accrete because it can't easily lose angular momentum.
- 5. Moreover, if you start with a stellar-mass black hole (say, $\sim 10 100 \ M_{\odot}$) and accrete continuously at the Eddington rate, you can't get to $10^9 \ M_{\odot}$ by 700 million years after the Big Bang. That's because stars take at least 200 million years to form, and then the time to reach $10^9 \ M_{\odot}$ at a standard efficiency of $\eta = 0.1$ (recall that this is defined so that the luminosity from a mass accretion rate dM/dt is $L = \eta (dM/dt)c^2$) is more than 600 million years.
- 6. So there are two paths: (a) accrete with lower efficiency, because if the system is at the Eddington luminosity, the time to double the mass is proportional to the efficiency, or (b) start with a much higher mass (i.e., a more massive seed), more like $10^5 M_{\odot}$ than $10 M_{\odot}$. No one is sure of the answer.
- 7. Some specific possibilities include beginning with a massive seed from the collapse of a star cluster, or a supermassive star. Something Alex Dittmann (current UMd grad student) and I have suggested is that there may be many stellar-mass black holes in AGN disks; accreting those doesn't generate luminosity, so they lower the efficiency. Many people are working on ideas!