

ASTR 380

## The History of the Earth



The real purpose of Stonehenge

# The History of the Earth

Earth's formation and bombardment

Formation of Moon and late heavy bombardment

Continental Motion

The Early Earth's Atmosphere

Life's interaction with the Atmosphere

What is important to life?



# Reminder

- First homework due at beginning of class this Thursday
- For essay-type questions, need to type or print out from computer to save grader's eyesight :)

# Composition of Terrestrials

- Terrestrial planets: those like Earth
- Earth by mass: 32% iron, 30% oxygen, then silicon, magnesium, ...
- Sun by mass: 74% hydrogen, 24% helium, 2% other
- What could explain differences?

Did heavy elements just  
sink closer to the Sun?



Did any elements just  
sink close to the Sun?

**No!**

# Grains and Ices

- Grains can stick together to make big things
- But grains need to be made of either heavy elements (silicon, iron, etc.) or ices
- Close to Sun, only heavies work  
Very little mass in them  
Can only make small planets
- Farther (beyond “frost line”), ices can form  
Incorporates hydrogen; lots of mass  
Later, gravity grabs H, He; big planets!

# Growth Through Collisions

- Grains come together, get bigger
- Collisions happen; sticking together eventually gets planetesimals
- Now gravitationally attracted, collisions escalate in strength
- Mopping-up process  
*A violent early solar system!*

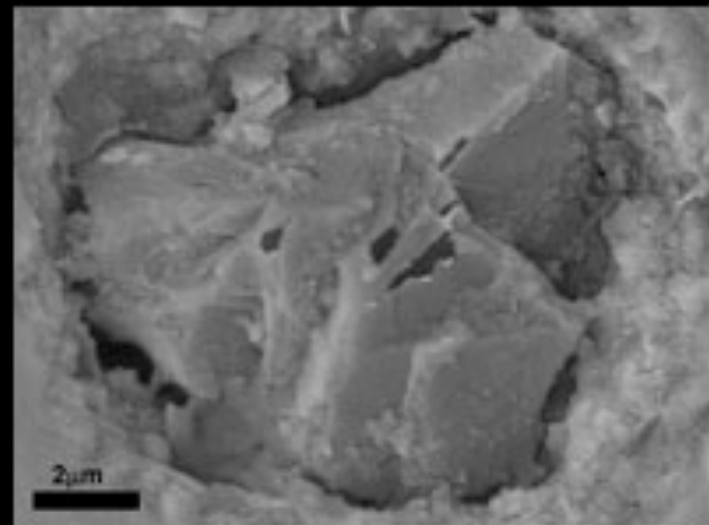


# Formation and Bombardment

Radioactive dating of meteorites finds that they date back to 4.57 billion years ago, with an uncertainty of 20 million years



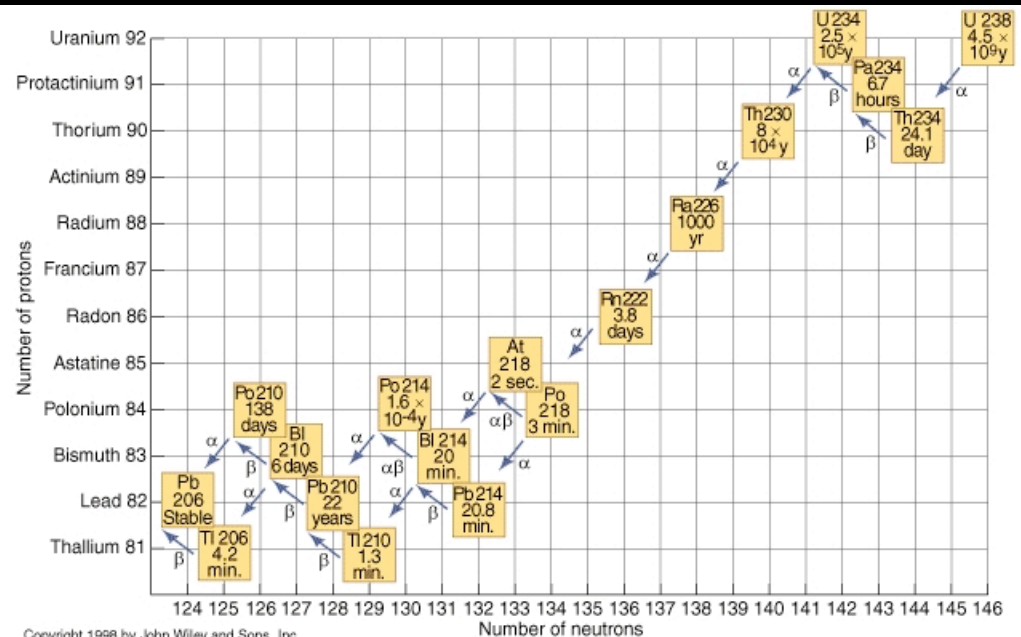
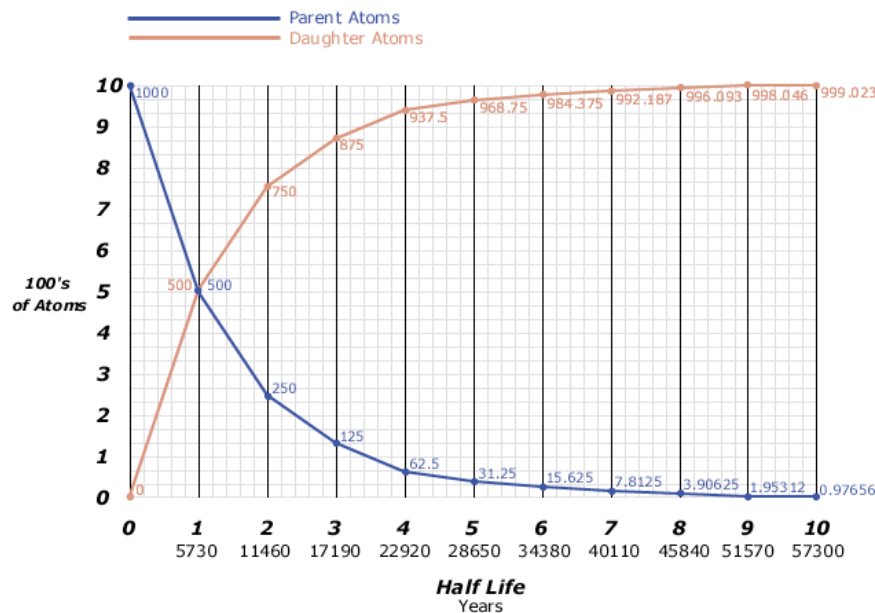
Tiny Mineral grains called Zircons have radioactive ages of 4.4 B yrs  
Oldest grains in Earth material



# Formation and Bombardment

Radiometric Dating: determine the age of something using the radioactive decay of an element.

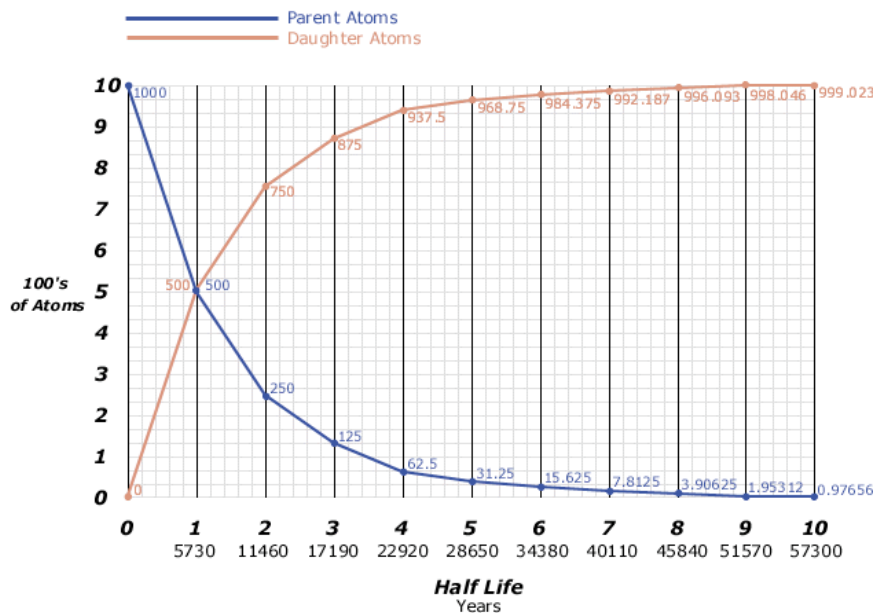
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# Formation and Bombardment

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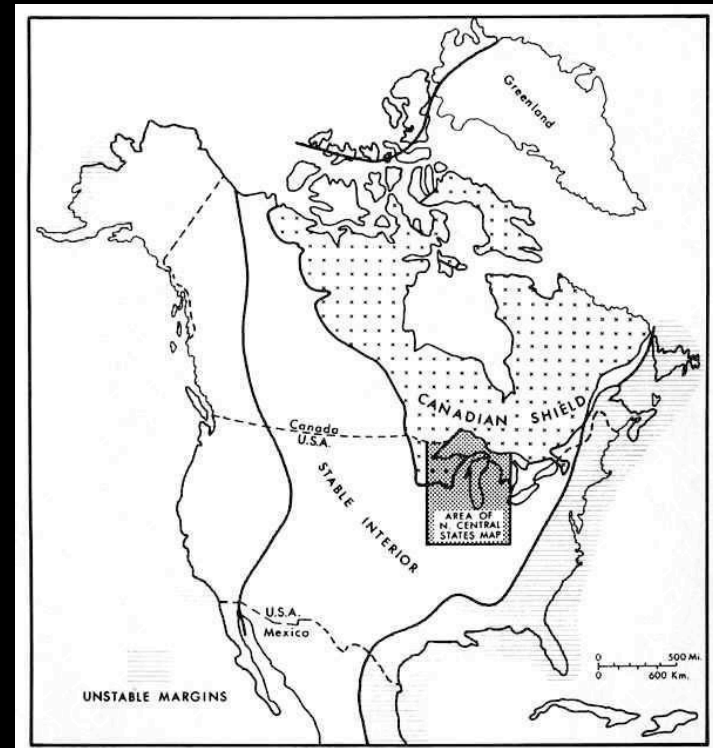


Eric B.

Isotope		Half - life
Carbon	$^{14}\text{C}$	5700 years
Iodine	$^{131}\text{I}$	8 days
Polonium	$^{214}\text{Po}$	$1,6 \times 10^{-4}$ seconds
Radium	$^{226}\text{Ra}$	1620 years
Uranium	$^{238}\text{U}$	$4,5 \times 10^9$ years

# Formation and Bombardment

The oldest intact rocks date from  
about 4.0 B yrs ago, found in the Canadian shield



Implications: Earth formed in less than 170 Million years.  
Earth had a stable surface after around 500 Million years.



# Formation and Bombardment

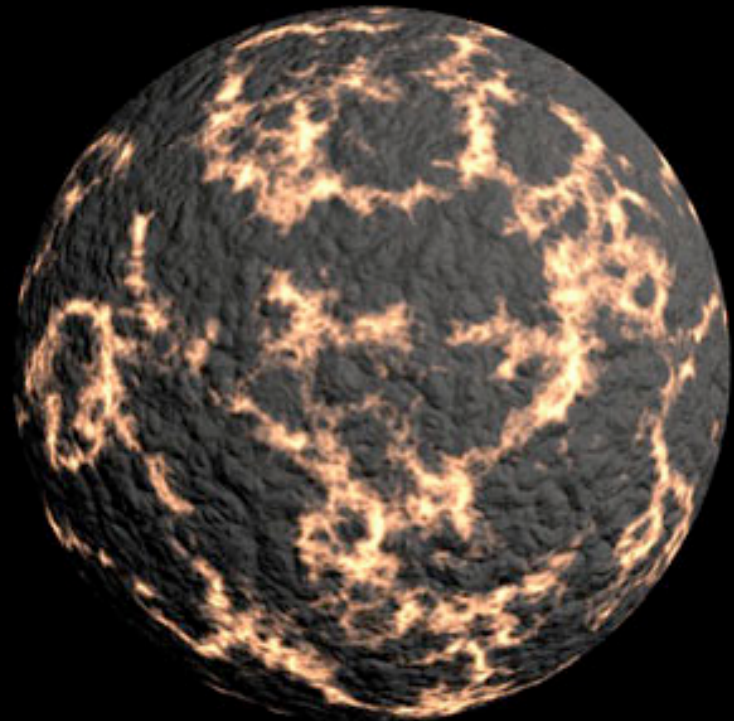
The Earth formed through the collection of planetesimals over a period of 50 - 100 Million years.



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The heat from impact of incoming bodies kept the entire Earth molten.

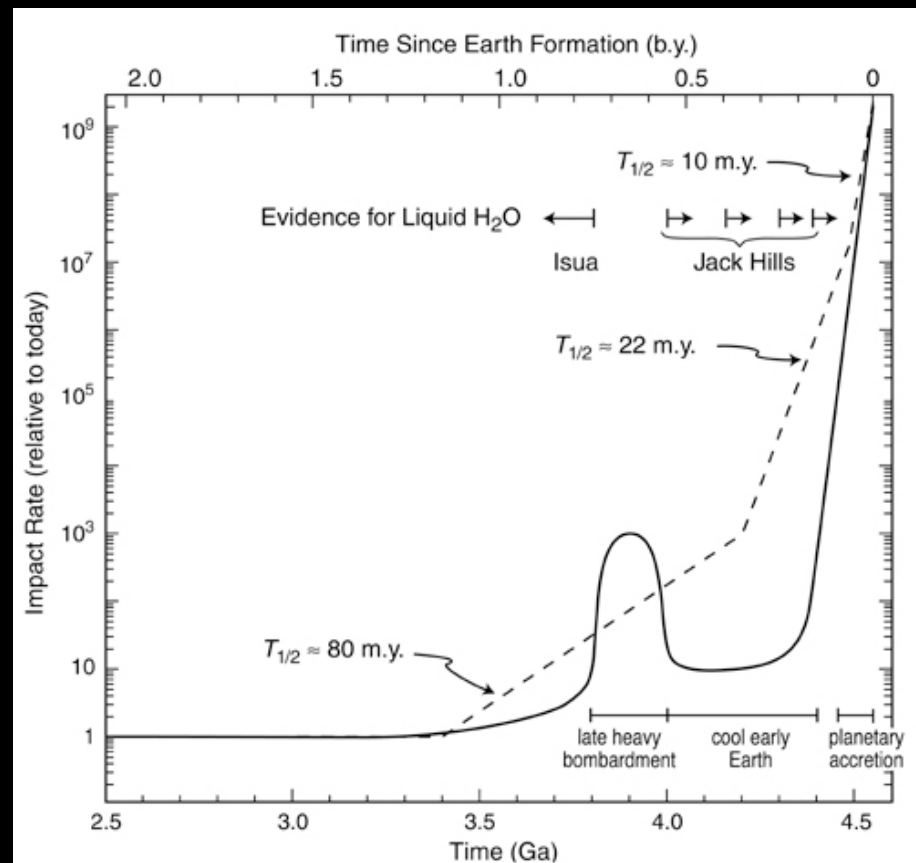


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These impact cleared small bodies out of the inner Solar System so the rate of impacts dropped off by 150 Million years.

By the end of this period the Earth was nearly its present mass and the surface was mostly solid.





# What about the Moon?

- Our moon is much larger relative to Earth than any other moon to a major planet
- Helps stabilize our rotation axis  
Therefore, seasons not extreme
- How did this happen? Would it be likely to happen elsewhere?

# Not Everyone Likes the Moon...

- Alexander Abian, late Iowa State math prof
- Said destroying the moon would give us eternal spring, remove hurricanes!
- Thoughts???



[http://wearsience.com/img450/destroy\\_the\\_moon.gif](http://wearsience.com/img450/destroy_the_moon.gif)

# Formation of the Moon and Late Heavy Bombardment

## Evidence:

Moon has less iron than Earth

Moon's orbit was much closer to Earth in past

Moon's rocks contain little gaseous material

What could have caused this?

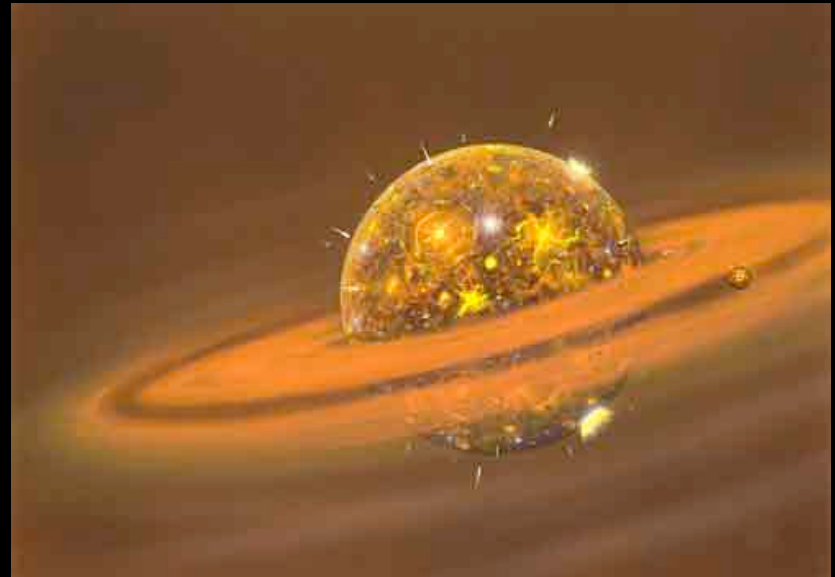
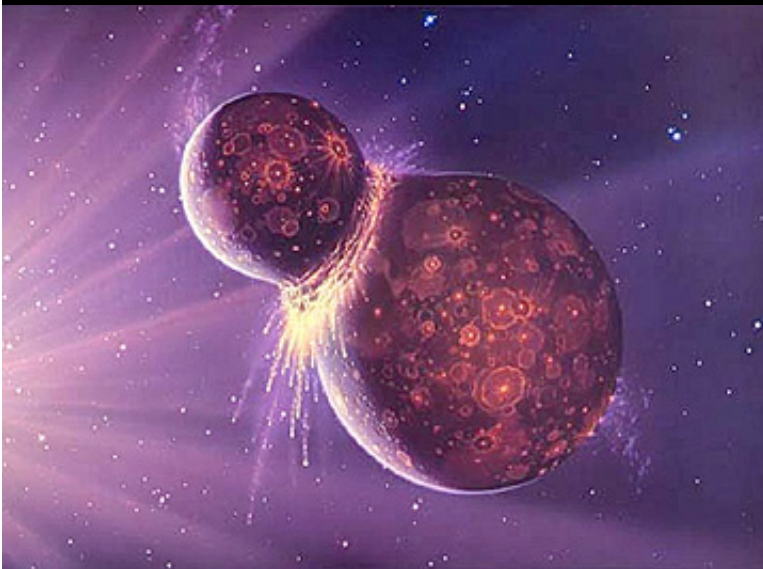


# Formation of the Moon and Late Heavy Bombardment

At around 70 Million years, the Moon was created by the impact of a Mars-sized body

It was a glancing impact which threw material into orbit around the Earth.

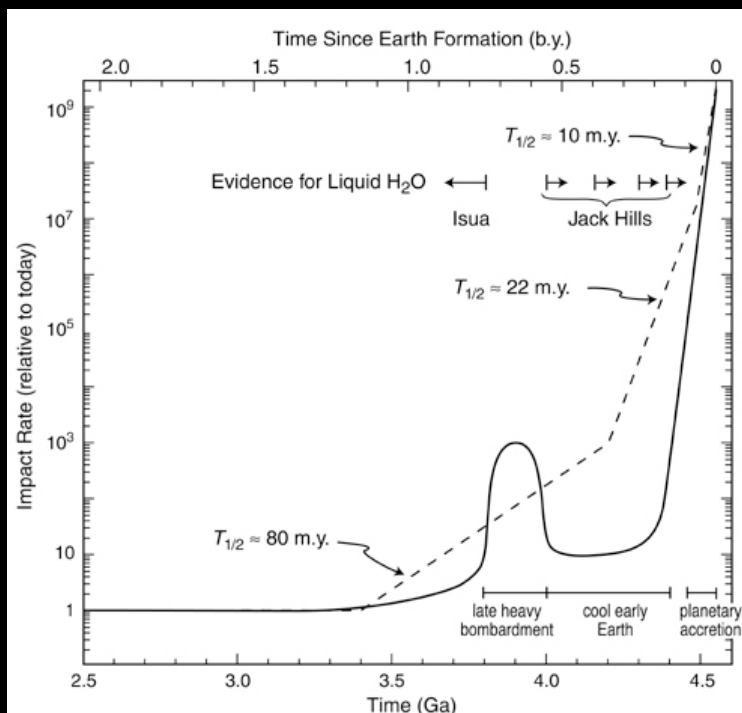
That material collected together to become the Moon



# Formation of the Moon and Late Heavy Bombardment

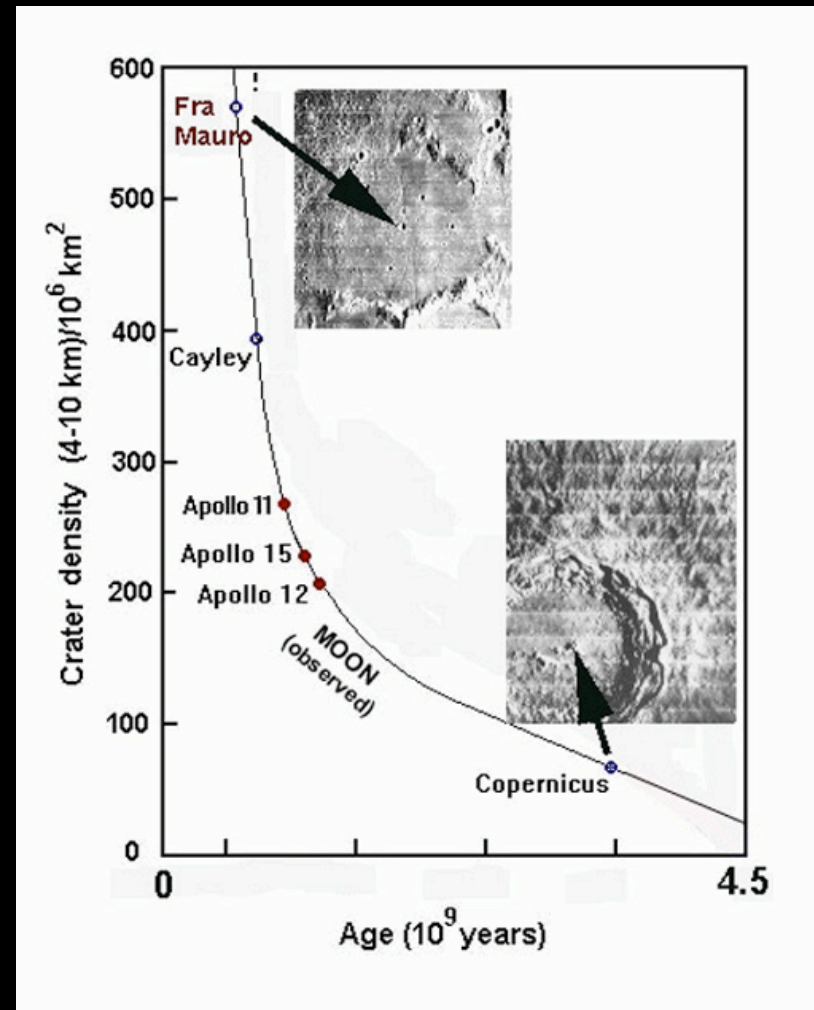
After its formation, the Moon's surface became a record of the continued bombardment of the Earth.

The big Maria on the Moon are evidence of a period of heavy impact activity around 3.9 Billion years ago



# Decreasing Rate of Impacts

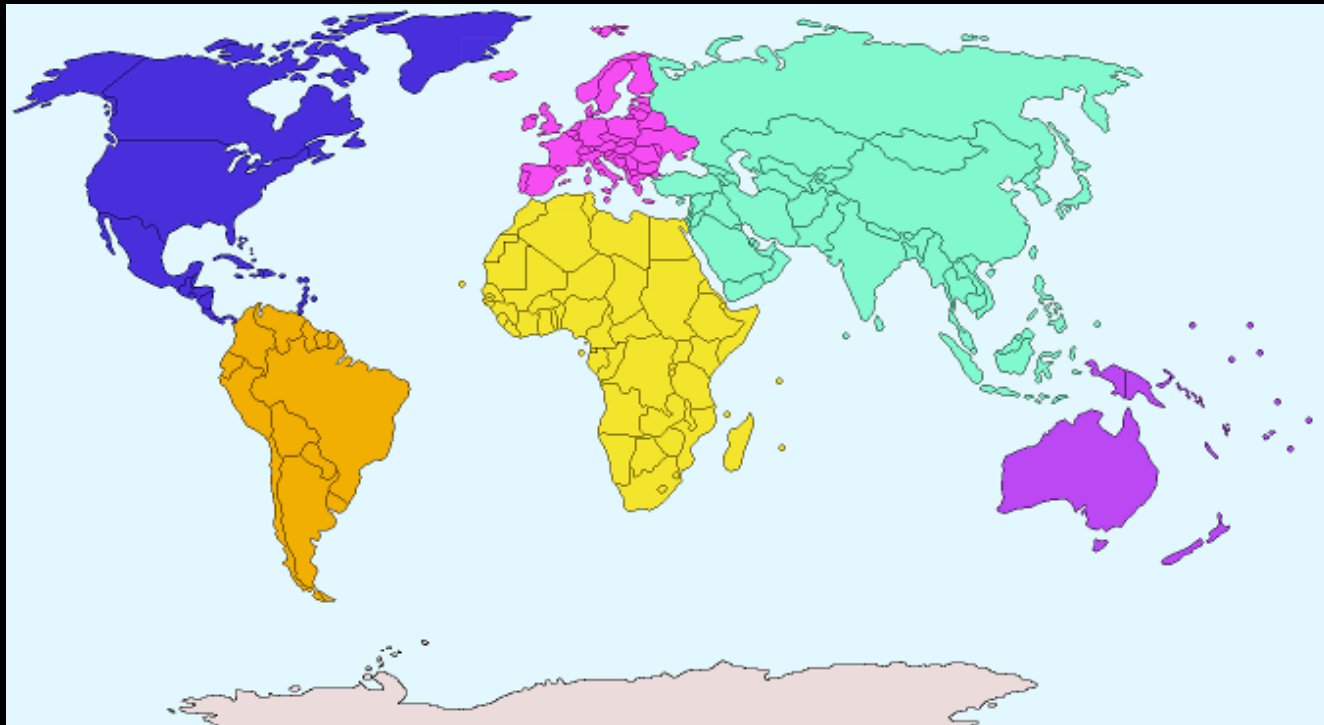
- Rate of impacts has gone down a lot
- Early Earth: no place for life
- Later, although impacts happened, not enough to wipe out all life
- Probably a general feature of planet formation





# The Crust of the Earth

- Continents fit together like jigsaw pieces
- Coincidence, or does it tell us about Earth's crust?

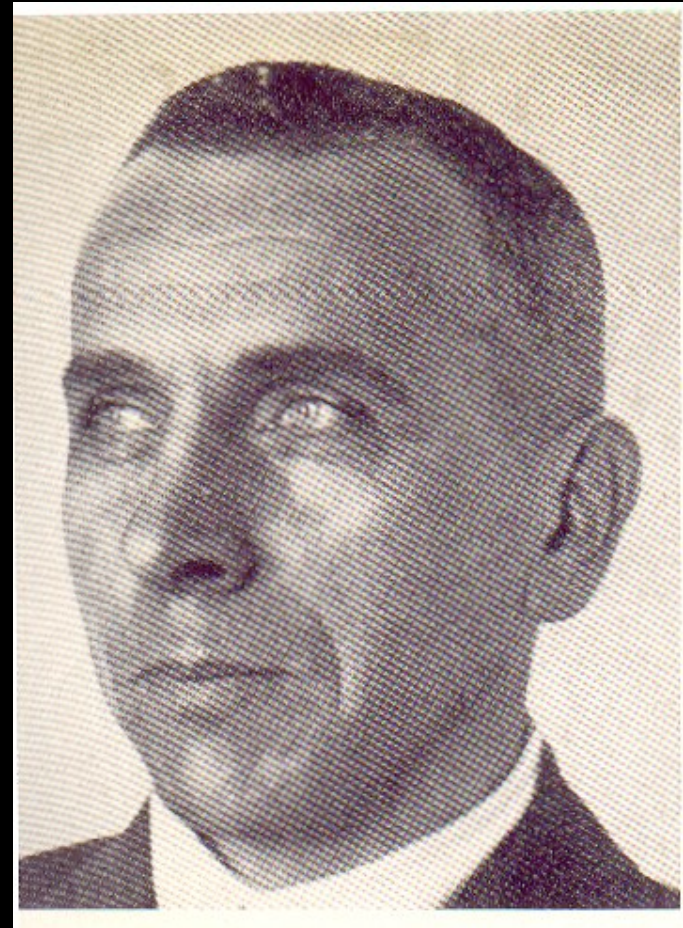


<http://www.naturalhistoryonthenet.com/Continents/images/continents.gif>

# Continental Drift

- Meteorologist
  - Continental shelves of Africa, South America fit together even better
  - Suggested continents drift all over Earth
- No physical mechanism**  
**How would light granite plow through dense basalt?**
- Not believed, but idea kept alive

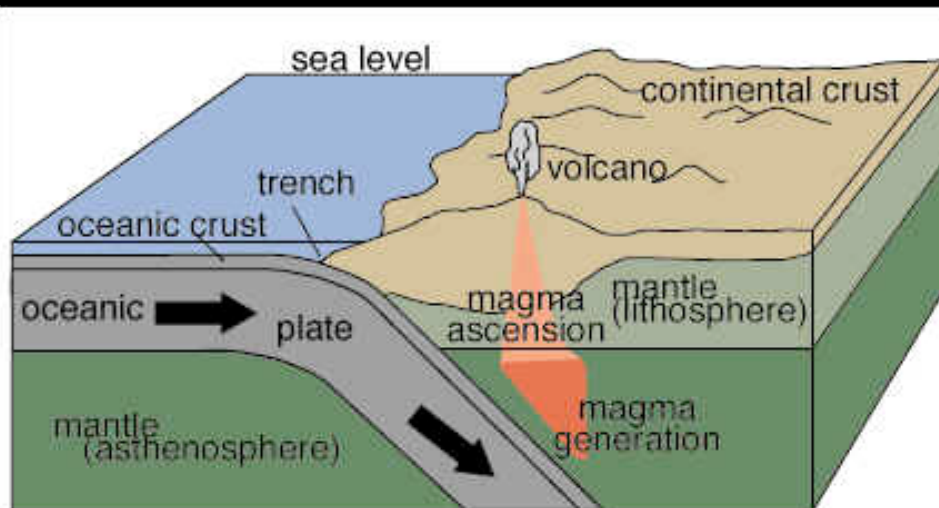
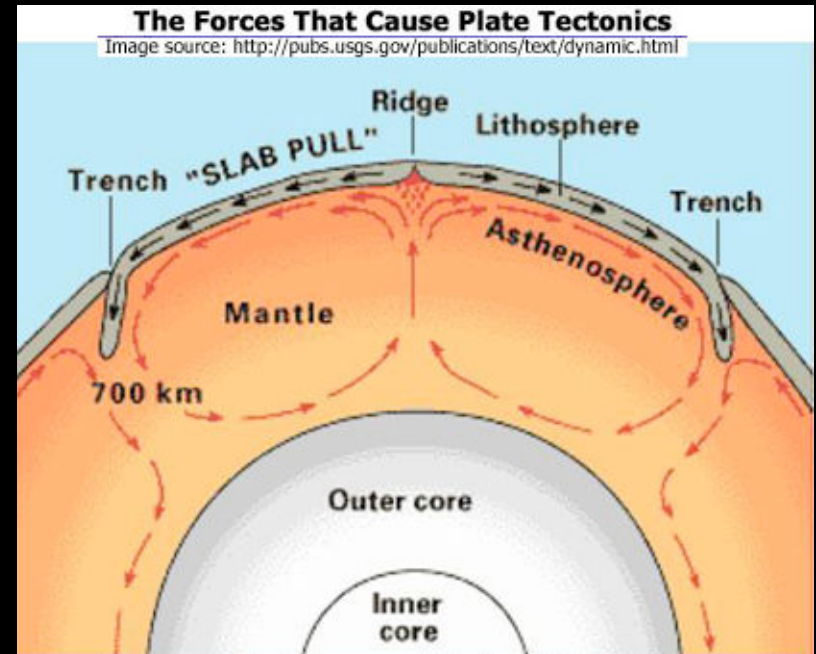
Alfred Wegener



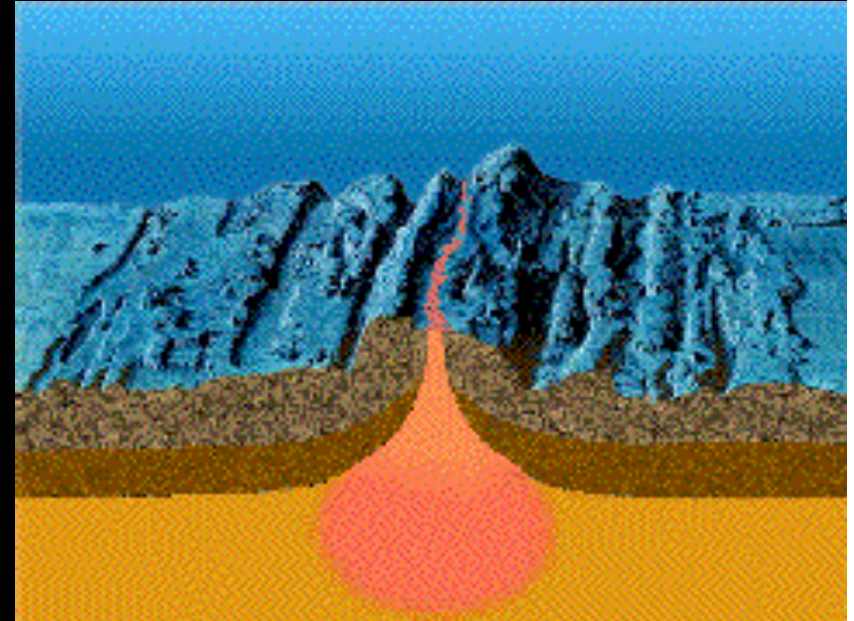


# Continental Formation and Motion

The crust of the Earth – the continents sit on a molten interior with strong convection which drives the continents around.



Magma is generated at subduction zones where dense oceanic plates are pushed under lighter continental plates.



# Continental Formation and Motion

## Active Volcanoes, Plate Tectonics, and the "Ring of Fire"



# Continental Formation and Motion

Because of plate tectonics, the continents are reformed on the scale of a few 100 Million years

250 Million years ago, all of the current continents were in one big super continent, Pangaea.

Before that there were many other arrangements of continents.

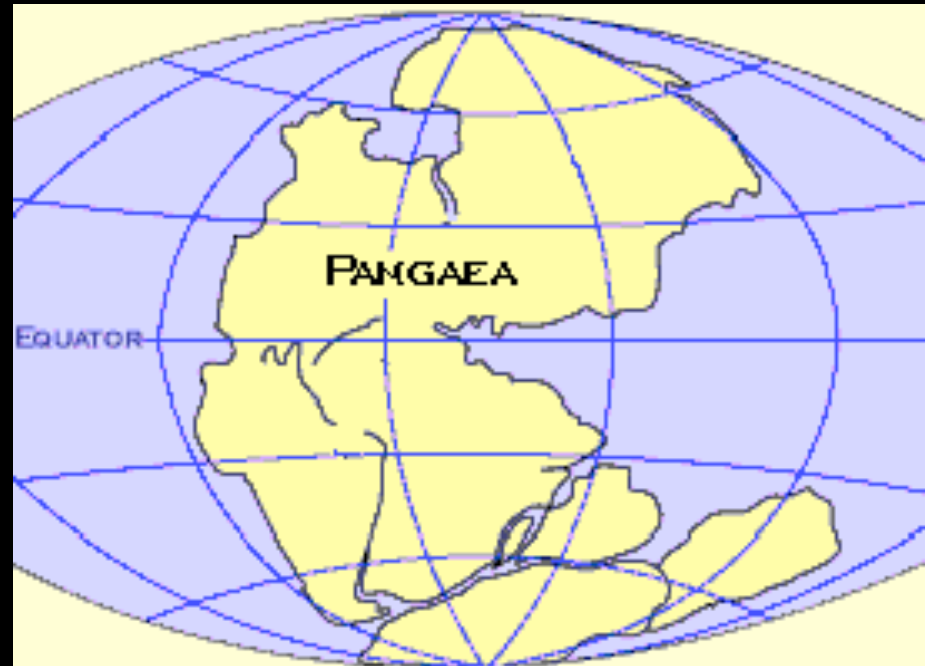


Plate tectonics drives change. Keeps mixture of land and water covering the Earth.

# Effects of Plate Tectonics

- Constantly changing environment may be an important driver for evolution
- Cycling of water may also play an important role

# Perspective: Other Terrestrials

- No other terrestrial planet (Mercury, Venus, Mars, even the Moon if you like) has active plate tectonics
- Why not?

# Perspective: Other Terrestrials

- No other terrestrial planet (Mercury, Venus, Mars, even the Moon if you like) has active plate tectonics
- Why not?
- Too small! Internal heat cools off too rapidly
- Is tectonics necessary for life?

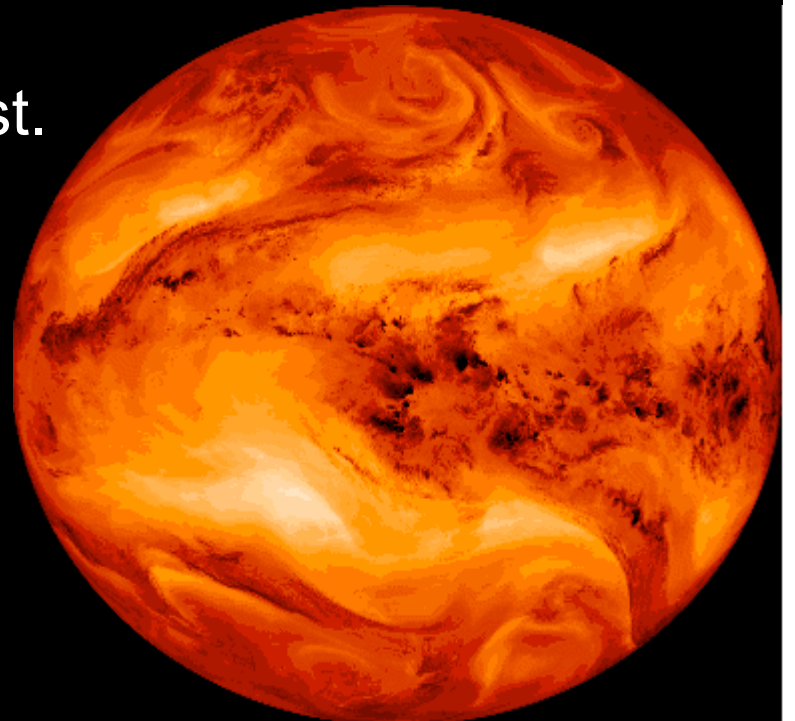
# The Early Earth's Atmosphere

The Earth may have started with a Hydrogen atmosphere which it could not hold on to for long.

The lava surface outgassed carbon dioxide, sulfur dioxide, water and ammonia.

During the collision that created the Moon all of this atmosphere was lost.

The lava surface continued outgassing – restoring the atmosphere.





# The Early Earth's Atmosphere

When the surface cooled sufficiently, water rained out to form seas, mountains grew.

Rainwater with dissolved  $\text{CO}_2$  is a mild acid which erodes rocks

In the oceans, the  $\text{CO}_2$  goes into carbonate minerals which form rock such as limestone over time.





# The Early Earth's Atmosphere

Amount of carbon dioxide tied up in rocks is comparable to the amount of CO<sub>2</sub> in the present Venus atmosphere!

The feedback loop of:

hotter => more water vapor => more rain => more  
acid rock erosion => more carbonate rock

may have regulated the  
amount of atmosphere  
in the early earth.

Until CO<sub>2</sub> was no longer  
dominant.



# The Early Earth's Atmosphere

When life came along, this cycle was sped up by cells joining in locking up  $\text{CO}_2$  into rocks

Then later turning  $\text{CO}_2$  into  $\text{O}_2$  when photosynthesis came along.

All along volcanoes dumped more  $\text{CO}_2$  into the atmosphere.

Life arose in this  $\text{CO}_2$  rich atmosphere --- not the present one



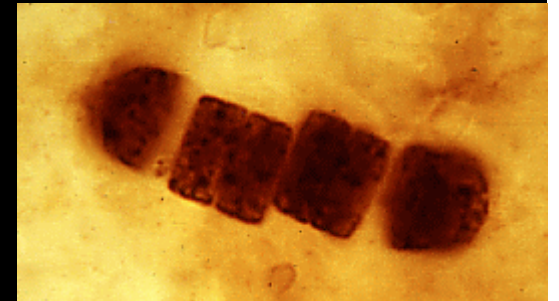
# Life's Interaction with the Atmosphere

The Archean, from 3.8 B yrs ago to 2.6 B yrs ago

Bacteria that lived in a world with CO<sub>2</sub>, methane, and ammonia atmosphere.

Limited fossil evidence.

Did their part by dying and locking up carbon into sedimentary rocks.



# Life's Interaction with the Atmosphere

During the Proterozoic, free O<sub>2</sub> in the atmosphere rose from 1 - 10%. Mostly released by the cyanobacteria – fossils in stromatolites. Stromatolites are rocks made from millions of microscopic layers comprised of the remains of bacteria.

2.5 billion years ago  
to 540 million years ago

By the end of this period  
there were oceans,  
continents, and surface  
temperature similar to  
today.





# Specifics of our Solar System

- Low-eccentricity planets; extrasolar systems often have high eccentricities
- Our large moon stabilizes our rotation axis  
Never have super-extreme seasons
- Jupiter: might protect us from asteroids  
Or, might be Mafia protection: without Jupiter, would be planet at asteroid belt!
- Do you think these are critical to life?

# Debate: Earth is Just Right

- Position 1: many aspects of Earth and the Solar System are crucial; without them, life would not be possible
- Position 2: not so! Life could originate in much different circumstances



# What is important to life?

Energy

Liquid water

Atmosphere

Plate tectonics?

Moon?



# What is important to intelligent life?

Photosynthesis?

A mix of land and oceans?

What else?



# Summary

- Formation of terrestrials appears to require a reasonable amount of heavy elements
- In early solar system, sweeping up of debris meant a high collision rate
- Our Earth is large enough for plate tectonics