ASTR 380 The Requirements for Life



"Pretty cool, Dewey. Hey! Shake the jar and see if they'll fight!"

Outline

- Chemical requirements
- Is water necessary?
- Type of star?
- Nature of solar system?
- Location in galaxy?
- An evaluation of other spots in our solar system

Midterm Results

- Average: 81% Nicely done!
- Numerical problem gave many trouble Remember, I link to websites that will help
- No letter grade assigned to just this exam Only class as a whole, at end

The Requirements for Life

Environmental Requirements for Originating Life:

Chemical building blocks Energy a liquid medium stability

Environmental Requirements for Sustaining Life lowering the bar....

Looking at the Moon, Mercury and the Moons of Mars

Chemical building blocks:

carbon, oxygen, nitrogen, hydrogen make up 96% of the mass of organisms on Earth

Why?

reasonable cosmic abundance

strong chemical bonds

complex chemistry possible





Complex chemistry possible: Hydrogen forms 1 chemical bond Oxygen can form 2 chemical bonds Nitrogen can form 3 bonds Carbon can form 4 bonds



Carbon Dioxide



Methanol

Methyl Cyanide CH₃CN

H

Ñ

н

Н



CH₃OH

Complex chemistry possible with ability to form a variety of bonds.

Carbon can form chains and rings

Other atoms like sulfur, calcium, iron, phosphorus play important roles in many essential bio-molecules



What other chemistry might be good for life?

Bonding structure suggest silicon, phosphorus, and sulfur. silicon for carbon, phosphorus for N and sulfur for oxygen —

S

N

	1	New												pnd	pno	pno	on	S 18
	IA	Original		Alkali	i metals		Ac	tinide serie	es	С	Solid			S	S S	S	Q	VIIIA
1	1 ¹ H Hydrogen	2 11A		Alkali	ine earth n	netals	Po	or metals		Br	Liquid		13 IIIA	14 IVA	15	16 VIA	17 VIIA	2 ² K He Helium
2	3 2 Li Lithium 6.941	4 2 Be Beryllium 9.012182		Trans Lantr	sition meta nanide seri	ls es	No No	nmetals ble gases		H Tc	Gas Synthetic		5 B Boron 10.811	6 ² C Carbon	7 25 N Nitrogen 14.00674	8 2 0 Oxygen 15 PT 4	9 27 F Fluorine 18.9984032	10 28 K Neon 20.1797
3	11 28 Na ¹ Sodium 22.989770	12 28 Mg Magnesium 24.3050	3 IIIB	4 IVB	5 VB	6 VIB	7 VIIB	8	9 - VIIIB	10	11 IB	12 IIB	13 28 3 Aluminum 26.981538	14 28 Si Silicon 28.0855	15 28 P Phosphorus 30.973761	16 28 S 5 Sulfur 32.066	17 28 CI Chlorine 35.453	18 28 KL Argon 39.948
4	19 28 K 1 Potassium 39.0983	20 28 Ca 2 Calcium 40.078	21 28 Sc 22 Scandium 44.955910	22 28 Ti 10 2 Titanium 47.867	23 28 V 11 Vanadium 50.9415	24 28 Cr 13 Chromium 51.9961	25 28 Mn 13 Manganese 54.938049	26 28 Fe 14 Iron 55.8457	27 28 4 Co 15 Cobalt 58.933200	28 28 Ni 16 Nickel 58.6934	29 28 Cu 18 Copper 63.546	30 28 Zn 18 Zinc 65.409	31 28 Gallium 69.723	32 28 Ge 4 Germanium 72.64	33 28 As 18 Arsenic 74.92160	34 28 Se 18 Selenium 78.96	35 28 Br 18 7 Bromine 79.904	36 28 L Kr 18 8 N Krypton 83.798
5	37 2 Rb 18 Rubidium 85.4678	38 28 Sr Strontium 87.62	39 28 Y 18 92 Yttrium 88.90585	40 28 Zr 18 2/r 91.224	41 28 Nb 18 Niobium 92.90638	42 28 Mo Molybdenum 95.94	43 28 Tc 18 13 2 Technetium (98)	44 19 Ru Ruthenium 101.07	45 28 Rh 18 Rhodium 102.90550	46 28 Pd 18 Palladium 106.42	47 28 Ag 18 Silver 107.8682	48 28 Cd 18 Cadmium 112.411	49 28 In 18 18 18 18 18 18 18 18 18 18	50 28 Sn 18 18 18 18 18 18 18 18 18 18	51 28 5b 18 Antimony 121.760	52 28 Te 18 18 18 18 18 18 18 18 18 18 18 18 18	53 28 18 18 18 18 18 18 18 18 18 18 18 18 18	54 28 L Xe 18 N Xenon 131.293
6	55 28 Cs 18 Cesium 1 132.90545	56 28 Ba 18 Barium 2 137.327	57 to 71	72 28 Hf 18 32 Hafnium 2 178.49	73 28 Ta 18 32 Tantalum 12 180.9479	74 28 W 18 22 Tungsten 2 183.84	75 28 Re 18 Rhenium 21 186.207	76 26 Os 16 Osmium 190.23	77 8 17 18 18 32 17 18 18 32 18 32 192.217	78 28 Pt 18 92 Platinum 1 195.078	79 28 Au 18 Gold 1 196.96655	80 28 Hg 18 Mercury 200.59	81 28 TI 18 32 204.3833	82 8 Pb 18 32 Lead 4 207.2	83 28 Bi 18 Bismuth 5 208.98038	84 28 Po 18 Polonium 20 (209)	85 28 At 18 Astatine 7 (210)	86 2 K L Rn 322 N Radon 8 (222)
7	87 2 8 Fr 18 5 Francium 8 (223) 1	88 28 Ra 18 32 Radium 8 (226) 2	89 to 103	104 28 Rf 32 Rutherfordium 10 (261) 2	105 28 Db 322 Dubnium 11 (262) 2	106 28 Sg 18 Seaborgium 12 (266) 2	107 2 8 Bh 32 80hrium 13 (264) 2	108 Hs Hassium (269)	2 109 2 8 Mt 18 32 2 32 Meitnerium 15 (268) 2	110 28 Ds 18 Ds 322 Darmstadtium 17 (271) 1	111 28 Rg 32 Roentgenium 18 (272) 1	112 28 Uub 32 Ununbium 18 (285) 2	113 Uut Ununtrium (284)	114 Uuq ^{Ununguadium} (289)	115 Uup ^{Ununpentium} (288)	116 Uuh ^{Ununhexium} (292)	117 Uus ^{Ununseptium}	118 Uuo Ununactium

What other elements might be good for life?

Bonding structure suggests: silicon for carbon phosphorus for nitrogen sulfur for oxygen

But...





 $SiO_2 = glass!$ Silicon important in plants.

phosphorus is a solid at room temperature but is self-igniting when in contact with oxygen It becomes a liquid at 317K = 111 F



What other chemistry might be good for life?



Sulfur is solid at room temperature and becomes a liquid at 388K = 239 F

All three are insoluble in water!



But both sulfur and phosphorus are soluble in carbon disulfide which is a liquid between -112 C and 46C.

But carbon disulfide is flammable in the presence of oxygen and very reactive with other molecules in solution.

Energy Requirements for Life:

The ultimate source of energy for much life on Earth now is the Sun.

- Plants harvest energy from sunlight
- Animals eat plants sugars and starches
- Bacteria eat the complex molecules created by plants and animals.

The secondary source is geothermal activity which releases highly reactive molecules.

Table 1. Planetary Data

	Energy	Requirem	nents for	Life:
--	--------	----------	-----------	-------

Planet	Distance from Sun (AU)	Radius (km)	Albedo	Est. Temp (K)
Mercury	0.387	2439	0.06	100-700
Venus	0,723	6052	0.76	700
Earth	1,000	6371	0.30	288
Mars	1.524	3393	0.16	210-300
Jupiter	5.203	71,398	0.51	110-150
Saturn	9.54	60,000	0.50	95
Uranus	19.18	25,559	0.66	58
Neptune	30.06	24,800	0.62	56
Pluto	39.44	1140	0.4-0.6	40

Being farther from a star decreases the energy from sunlight – colder temperatures – less energy – less stored energy – chemical reaction, evolution, motion will go slower.

Tidal heating and geothermal activity might provide for some energy.... But would it be enough?

Also consider that evolution to complex multi-celled organisms requires even more energy!

Does life require liquid water?

Life requires a way for molecules to be transported to the organism: gas or liquid.

Life requires a way to move molecules within the cells. liquid

Life requires a way to mediate/enable chemical reactions liquid

In all life on Earth, water does all of these things....

What are the alternatives to water?

In standard Earth organic chemistry, ammonia methane, and ethane might be possible... but the liquid temperatures are cold.
All chemical reactions go slower at these cold temperatures.
Methane clathrates (methane and water) found on ocean floors suggest that alternative situations are possible....



Substance	Freezing	Boiling	Range	
	Temperature	Temperature		
Water	0 C	100 C	100 C	
Ammonia	-78 C	-33 C	45 C	
Methane	-182 C	-164 C	18 C	
Ethane	-183 C	-89 C	94 C	

Big advantages to water:

- (1) Common molecule found in comets
- (2) Liquid over broad range of temperature
- (3) Ice floats
- (4) Water is a polar (= negative-positive charge separation) molecule which helps in dissolving things
- (5) Water is a stable molecule so it does not enter into most reactions that occur in water.



Stability:

It takes 10's to 100's of millions of years for life to happen. The environment must be stable on that timescale.

Implications:

- Day-night variations are not too extreme.
 --- no swings which can destroy molecules
- Seasonable variations are not too extreme.
 --- Planet is in reasonably circular orbit.
- Star is not too variable or short-lived.

Or are we being closed-minded?

- Organisms exist that can survive remarkable swings in environment Deinococcus radiodurans, for example
- Some large organisms require it E.g., plants needing fire to reproduce
- Could life evolve in such circumstances?
- How extreme is too extreme?

The Environmental Requirements for Life Origin versus survivability:

There might be bacterial life on Earth which could survive on Mars on the polar caps....

There could be bacterial life on Earth which could survive on Mars 10 meters down in the soil.

If life exists on a moon of Jupiter, it is possible that life might survive if it were placed in the Jovian atmosphere. Maybe....

Endospores are dormant non-reproductive states entered by some bacteria to survive extreme heat, dry, poisons, and even in space for hours. The Environmental Requirements for Life Origin versus survivability:

The major change is that stability is likely less important!

In summary, life requires:

chemical building blocks – good arguments that C, N, O, and H are the best.

energy – reasonable arguments that the Sun is the best stable, long term source of energy

liquid – good arguments that water is best; ammonia and methane are possibilities.

stability – at least to get started (maybe!)

Precession of Equinoxes

- Over 26,000 yrs, Earth's axis moves
- But tilt remains 23.5 degrees
 Stabilized by Moon
- Therefore, seasons are never too extreme



What About Mars?

- Currently, obliquity close to that of Earth
- But over millions of years, varies from 10 to 50 deg
- At large tilts, would have extreme seasons
- Is this a disqualifier?



http://updatecenter.britannica.com/eb/image?binaryId=73524&rendTypeId=4

Pros and Cons of Climate Change

- Suppose we had a more variable climate
- Disadvantage
 Might be tougher on complex organisms,
 which tend to specialize to environment
- Advantage Changes over millions of years might stimulate evolution
- Which do you think would win out?

Eccentric Orbits: Ejection

- Planets in Solar System have nearly circular orbits
- What if they didn't?
- If two orbits cross, one planet will be ejected eventually
- Therefore, planetary systems can't form this way



http://www.oklo.org/wp-content/images/analoghydro1.gif

Eccentric Orbits: Climate

- Nested eccentric orbits can be stable
- But distance variation means wide swings in temperature, extreme climate variation
- What are consequences?



star systems in comparison with each other, and alzo with our own inner solar system. © Govin Symil 05

http://www.gavinrymill.com/dinosaurs/extra-solar-planets/extra-solar-planets-reticulum.gif

Eccentricity and Climate

- At poles, go from zero solar illumination to illumination comparable to our winter
- But temperature doesn't change by nearly that much! Why not? Winds, reradiated heat
- High eccentricity could change illumination by much larger factor
- Stronger winds? Might compensate, though

The Example of Venus

- Venus has thickest atmosphere of terrestrials (90x Earth's pressure)
- "Day" is very long; about 120 days, noon to noon Without atmosphere, huge day/night change
- But with atmosphere, temperature is nearly constant over whole planet; day, night, equator, pole
- No atmosphere (Moon) means much more extreme swings
- Lessons?

Our Friend, the Sun

- Long-lived Now 4.6 Gyr old Time for evolution
- Very stable in short term
- But over its history, has increased its brightness by 30-40%
- How do other stars compare?



http://z.about.com/d/space/1/5/Y/Q/sun_tour.jpg

Other Stars: Preview

- We'll discuss this more later, but...
- Big stars live a short time
- Small stars live a long time but have larger flares, hence short-term changes in brightness
- How critical do you think it is to have a star much like our Sun?

Planets Around a Pulsar

- First extrasolar planets discovered!
- Lots of energy, but in form of high-energy particles
- Subsurface life?



http://www.daviddarling.info/images/pulsar_planets.gif

Our Sun in Milky Way

- Sun is close enough to our galactic center that heavy elements are numerous
- But not close enough that nearby supernovae happen frequently
- How critical a balance do you think has to be struck?



The Moon, Mercury, and the Moons of Mars



The Moon – Been there. Done that.







NY Sun, August 1835



The Moon – Brought back rocks ©

No atmosphere No liquid on surface Rocks show no evidence for life or complex organics

Ice might have been found in shadows of craters at pole – where sunlight never hits.

Ice is probably collected from comet hits on Moon.



The Moon – left seismometers

Moon quakes exist!

Caused by:

- impacts
- tides from Sun
- heating of surface

Moon is no longer molten.



The Moon versus our checklist:

chemical building blocks: light on amounts of C, N, and O

energy: lots of sunlight

liquid: No. And no atmosphere



Mercury: Mariner 10 spacecraft took pics in 1974

A flyby with the cameras rolling.





Mercury:

Old, very cratered surface No evidence of volcanic activity Has very thin atmosphere which is constantly escaping

No evidence for water

Core may still be molten but inactive mantle.



Mercury versus our checklist:

chemical building blocks: 70% metallic and 30% silicate may have lost much C, N, O in a late large collision.

energy: lots and lots of sunlight

liquid: No. Nearly no atmosphere

stability: Due to 59 day long rotation (Mercury day) and very slight atmosphere... night time lows = -183 C daytime highs = 427 C

Moons of Mars: 10-30 km in size, irregular shaped.

Deimos

Phobos

Moons of Mars: Both moons appear very similar to carbonaceous asteroids.

They are likely captured asteroids.

Surface temperature is around -40 C on average



The Moons of Mars versus our checklist:

chemical building blocks: Carbonaceous asteroids so good C,N,O

energy: reasonable sunlight

liquid: No. No ices. No atmosphere

stability: Probably reasonable but no data on temperature variations at specific locations on moons

Summary

- As far as we can tell, carbon chemistry and liquid water are really good for life
- We are indeed in a good place for this, but it is not clear how strong the requirements are
- Moon, Mercury, moons of Mars not great for life