OBSS-Science at the Bright end

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- It's hard to compete with GAIA at faint magnitudes
- Overall Mission accuracy is proportional to the number of photons collected per star:
- $\delta_{\rm X}$ ~ FWHM_{PSF} / $\sqrt{N_{\rm vM}}$ \longrightarrow Need LOTS of extra v's
- But GAIA saturates around G~12:
- It will measure G=2-12 via the diffraction wings that result from the rectangular aperture
- OR: will dump charge in drain such that final part of CCD transit is unsaturated
- Either way: Calibration is different from unsaturated stars which may cause systematic effects.
- Why do we care? <u>Precision Astrophysics can be</u> <u>achieved at bright magnitudes</u>

Astrometric Accuracy ~ Size FPA

(N. Zacharias)

- $\delta_{\rm X}$ ~ FWHM_{PSF} / $\sqrt{N_{\rm vM}}$
 - For a given Mission length:
 - Longer Integration times —> fewer observations
 - Same $N_{vM} \longrightarrow Same \delta_X$
 - 2x Larger Mirror &
 - Critical Sampling &
 - Same F-number &
 - Same (#) CCDs
 - \rightarrow 2² * Smaller Field of View ---> 1/2² * N_{Obs}
 - → If physical size of FPA is unlimited
 - → Same $N_{_{VM}}$ and FWHM/2 ---> $\delta_{_X}$ /2
 - → If physical size of FPA is identical --->

$$\rightarrow N_{OBS} = 2^{-3} \qquad \qquad ---> \delta_{X} *1$$

Science @ The Bright End

- Very Accurate π 's --> Luminosities
 - Double Stars
 - ◆ + Eclipsing Binaries --> Radii
 - ◆ + Spectroscopy --> Masses & [Fe/H]
 - Mass, Radius, Luminosity
 - For two components
 - \rightarrow M, R, L, log(g), He, Age
 - → For 10,000's Systems V<~ 14
 - Galactic Dynamics & Evolution with a sample of well-calibrated systems

Galaxy is NOT Static In the post- SIM&GAIA world

- ◆ Stellar mass has increased from 1% 10 Gyr ago to 100% now
- Dynamics of stars varies with time:
 - Potential changes adiabatically, Stars keep their initial energy
 - → Evolution of orbits
- In order to understand post-SIM/GAIA Galactic Dynamics, we need approximate ages of stars

Stellar Astrophysics

- ◆ Current <u>Astro</u>physics is limited by the fact that the Sun is the only star for which wee know the fundamental parameters to sufficient accuracy
- \bullet Differences among the resulting physical parameters amount to 0.1, 0.25 dex for [Fe/H] and log(g) and 1% for $T_{\rm eff}$
 - Even when using the most sophisticated models and high reolution data sets