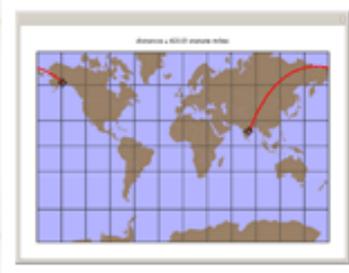
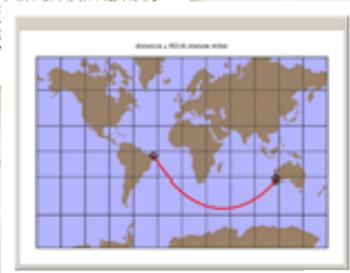




# Shortest flight paths are geodesics-geodesics are defined to be the shortest path between points in the space.

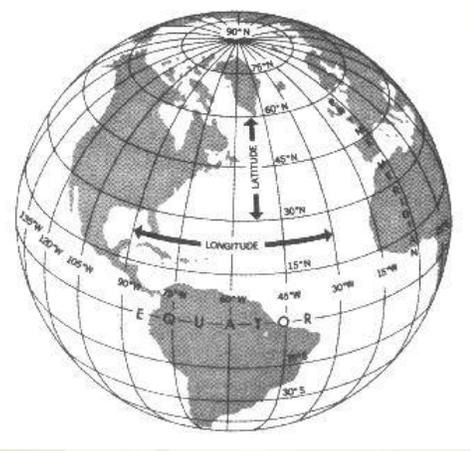


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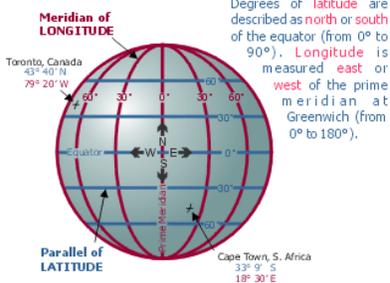


## On Globe...



**Meridian of LONGITUDE**

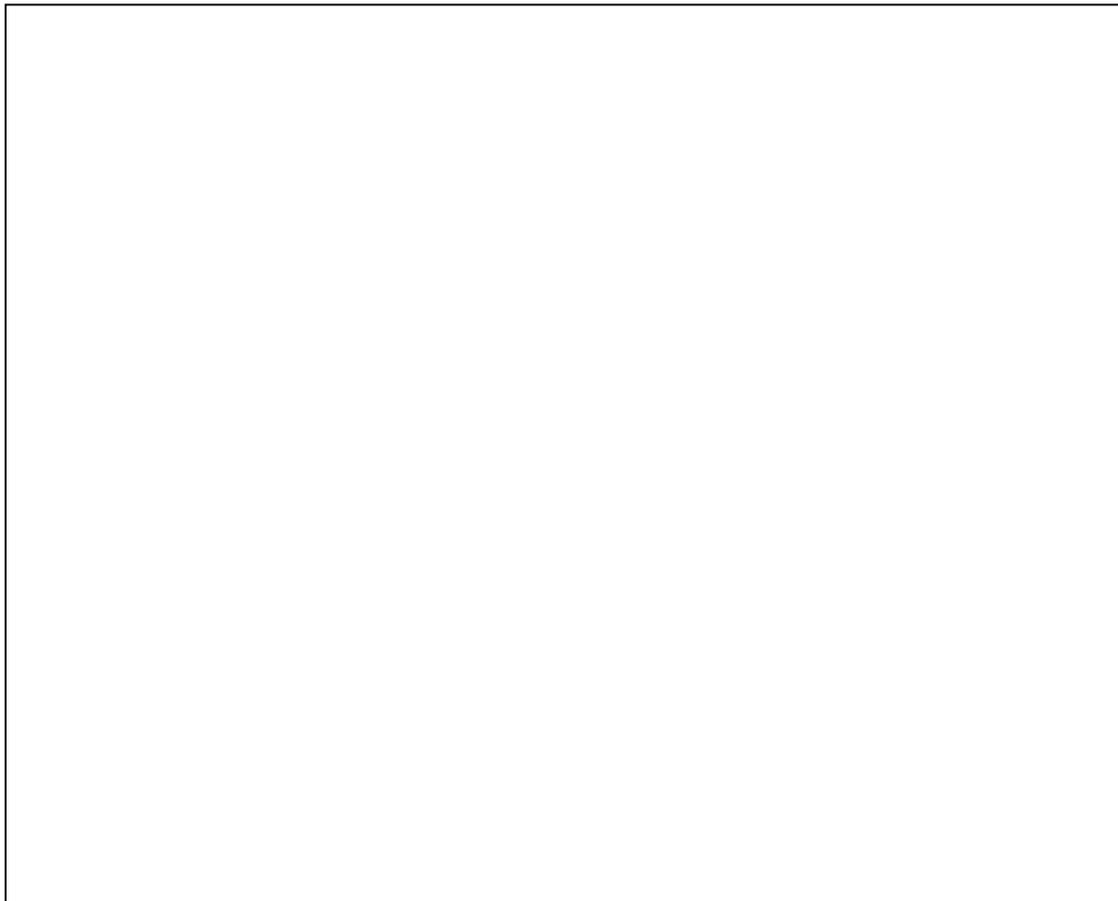
Degrees of latitude are described as north or south of the equator (from 0° to 90°). Longitude is measured east or west of the prime meridian at Greenwich (from 0° to 180°).

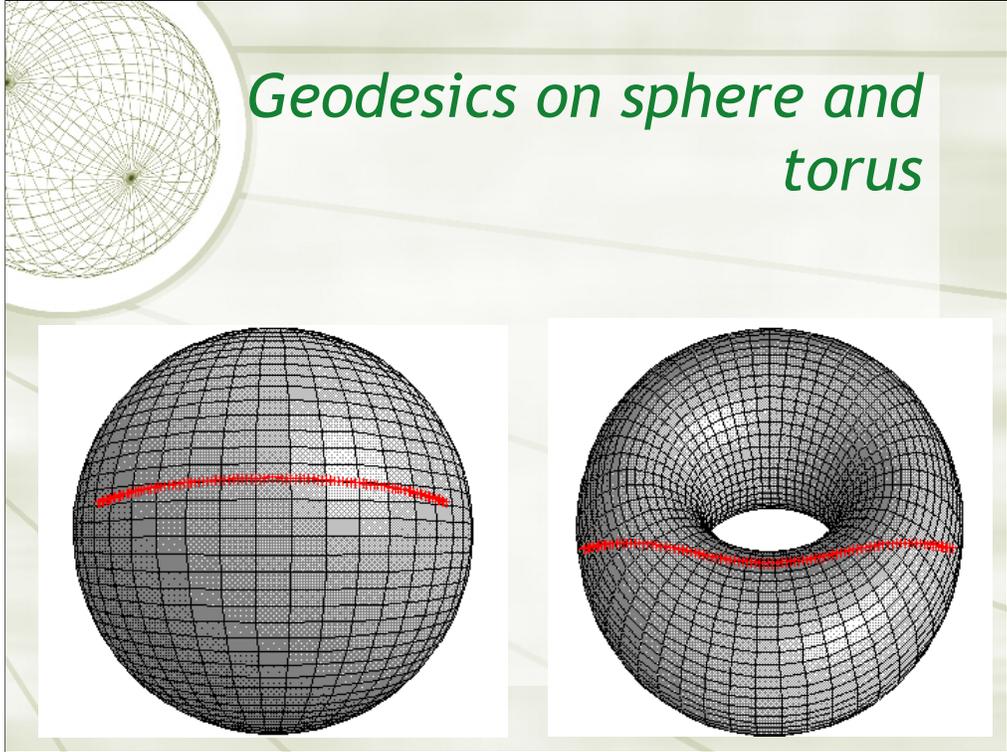


**Parallel of LATITUDE**

- ✦ Constant-longitude lines (meridians) are geodesics
- ✦ On the Earth, geodesics are Great Circles, the shortest distance between two points on the surface.
- ✦ Constant-latitude lines (parallels) are not geodesics

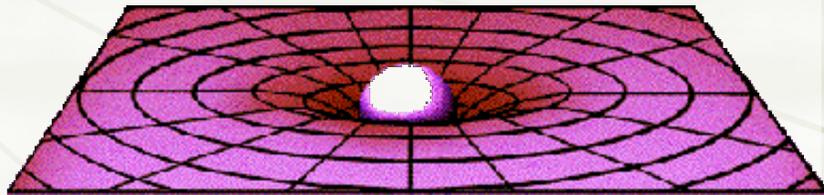
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## How does matter “warp” space?

- ★ Use two-dimensional space as an analogy: think of how rubber sheet is affected by weights
- ★ Any weight causes sheet to sag locally
- ★ Amount that sheet sags depends on how heavy weight is



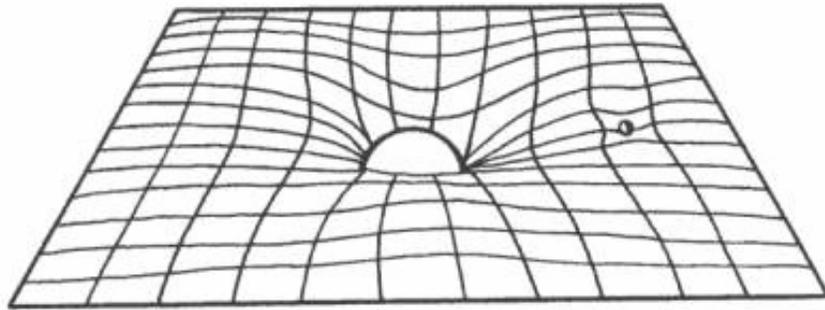
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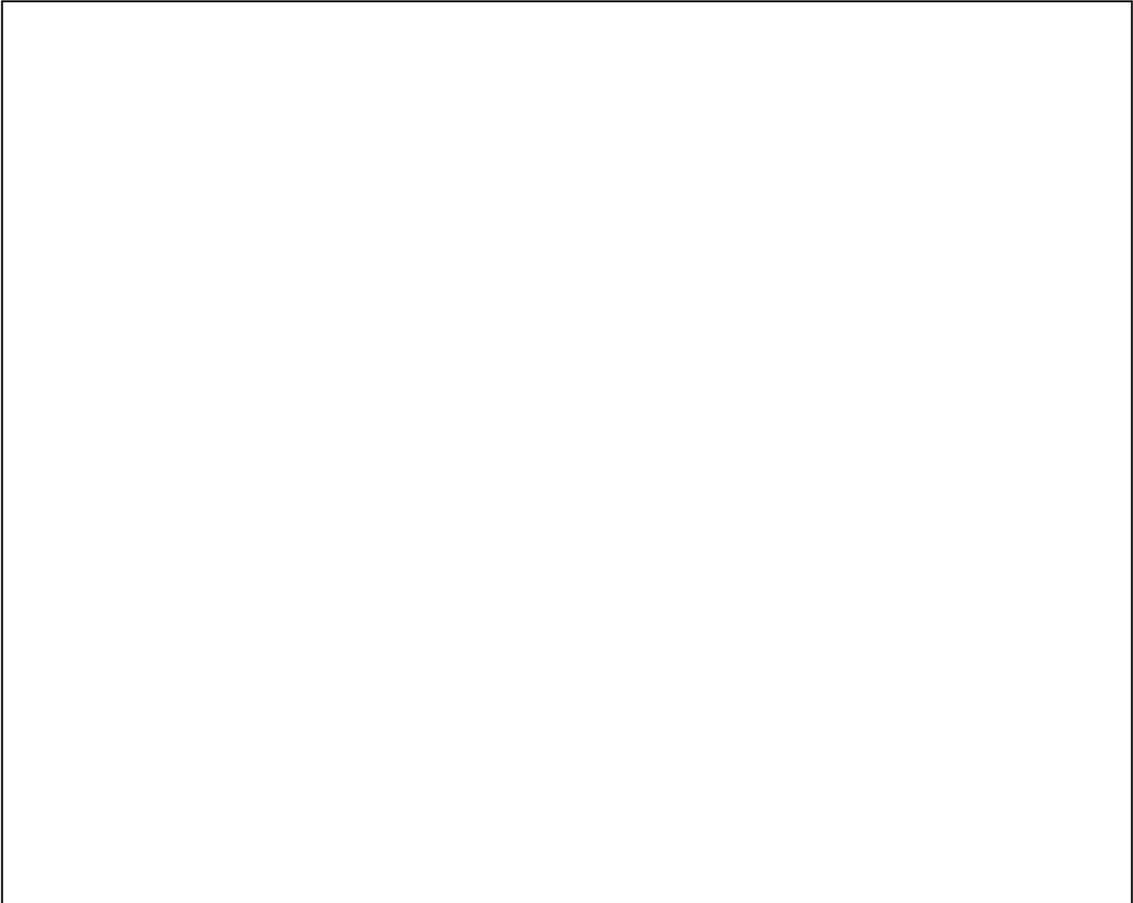
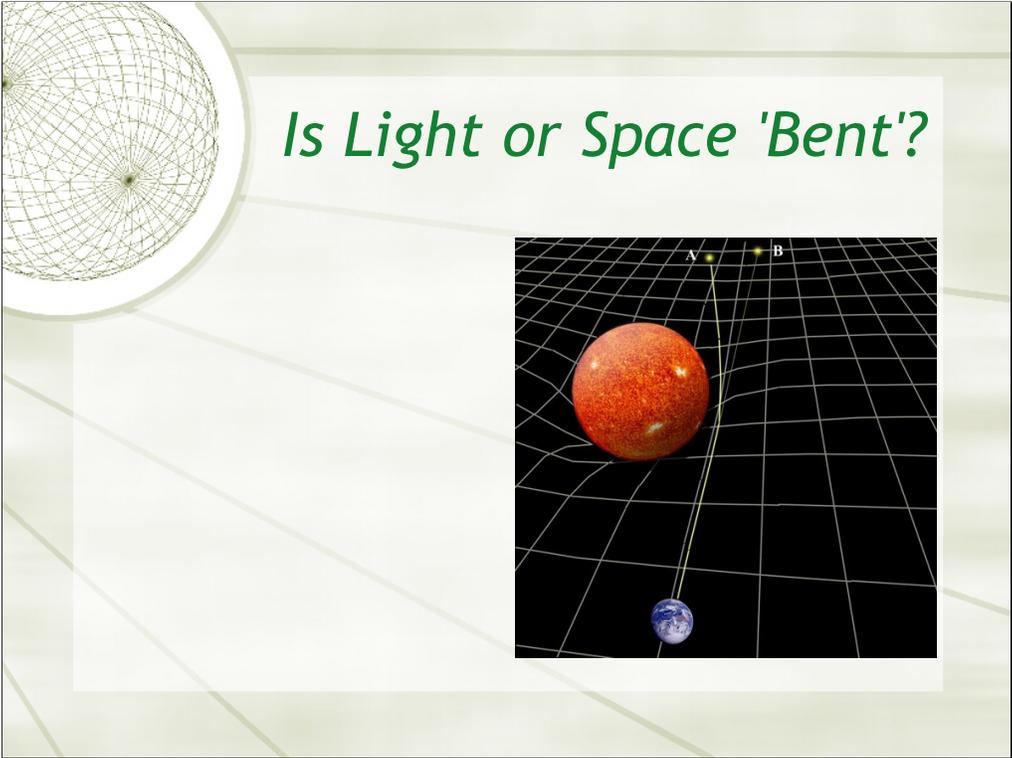
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## Effect of matter on coordinates

- ★ Lines that would be straight become curved (to external observer) when sheet is “weighted”



3.

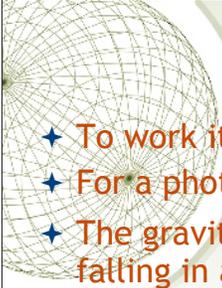




## *Lecture 11: General Relativity III*

- ★ Recap : gravitational redshift
- ★ Curved spacetime and geodesics
- ★ The General Theory of Relativity
- ★ Gravitational Lensing-again

3/8/14



## Gravitational Redshift

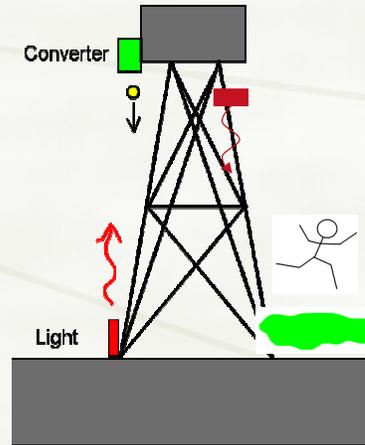
- ★ To work it out with equations
- ★ For a photon  $E=mc^2=h\nu$
- ★ The gravitational potential energy (the energy gained by falling in a gravitational field)  $PE=-GMm/r=(-GMh/rc^2)\nu_0$ . last lecture we use 'little' g;  $PE=mgh$ ; where 'little' g is the acceleration at the surface of the earth- these are the same
- ★ So as the photon rises up in the gravitational field
- ★  $h\nu=h\nu_0\{1-(GM/rc^2)\}$ ;  $\nu=\nu_0\{1-(GM/rc^2)\}$ ;  $\Delta\nu/\nu=GM/rc^2$
- ★ (where  $\Delta\nu$  is the change in frequency of the photon  $\Delta\nu=\nu-\nu_0$ )
- ★ Since the photon is reduced in frequency (and thus lengthened in wavelength) this is called the gravitational redshift

## Resolving the tower problem

- Now consider light ray aimed from top to bottom of tower
- Free-falling (FF) observer sees light ray travel **unaffected** by gravity, since freefall is an inertial frame
- From "Earth's" frame...

- Free-falling (FF) observer is traveling faster and faster
- Falling observer would see an increasing **redshift** of light source according to special relativity
- If FF observer is **supposed** to see a constant frequency light beam, then light must get relatively **blueshifted** as it falls in a gravitational field, to compensate
- Light beam aimed upward must conversely be increasingly **redshifted** with height
- Gravitational redshifting removes just the right amount of energy to solve the tower paradox!**

(remember that the energy of a photon is  $E=hf=hc/\lambda$ )



3/8/14

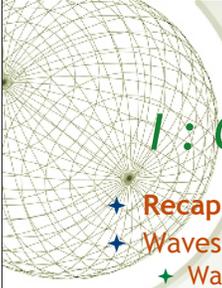
9

## Gravitational Redshift

- ✦ The conservation of energy-matter: not just the conservation of energy
  - ✦ 2. The equation for the potential energy in a gravitational field  $PE = m \cdot g \cdot h$  (or more generally  $PE = GM/r$ )\*
  - ✦ 3. The energy of a photon  $E = h\nu$
  - ✦ 4.  $E = mc^2$
  - ✦ So the 'trick' is: The energy received must equal the energy emitted minus the energy due to the potential energy
- \* In GR the potential energy is difficult to define quantity; for a spherical mass it is approximately  $PE = m_0 c^2 \sqrt{1 - 2GM/rc^2}$ : where  $m_0$  is the mass of the particle,  $G$  is the gravitational constant,  $M$  is the mass of the large body (e.g. earth or star and  $r$  is the distance from the center of the spherical large body) so this derivation is a little bit of a cheat

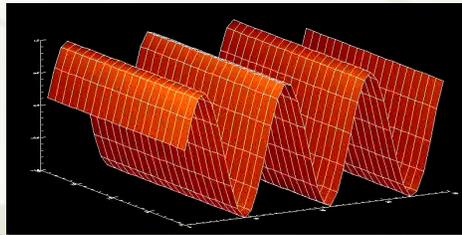
## Gravitational Redshift

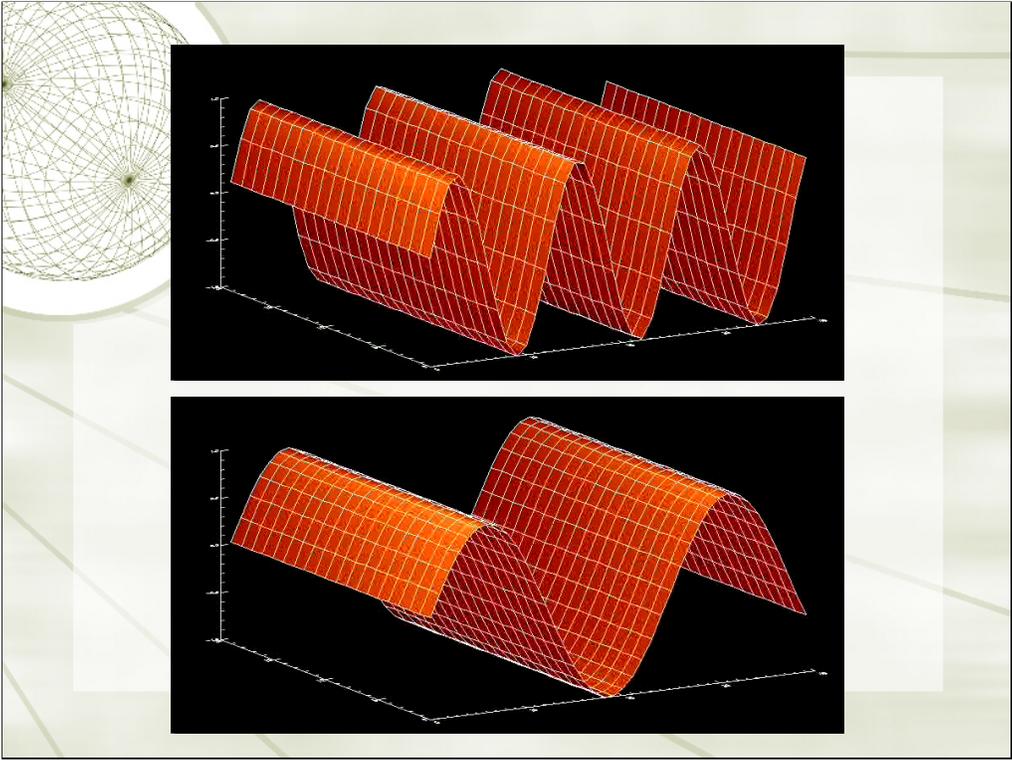
- ★ Finally another way of looking at this
- ★ From the principle of equivalence between gravity and force due to acceleration, the shift in frequency in a gravitational field can be related to the relativistic Doppler shift due to an accelerating light source.
- ★ Derivation:
- ★ for velocities  $v \ll c$  the Doppler shift formula is  $\nu = \nu_0 [1 + v/c]$
- ★ For an observer accelerated to velocity  $V$  in a time  $t = L/c$  ( $L$  is some length)  $V = at = aL/c$  and then  $\nu = \nu_0 [1 + aL/c^2]$  and replacing the arbitrary acceleration 'a' by gravitational acceleration 'g'  $\nu = \nu_0 [1 + gL/c^2]$



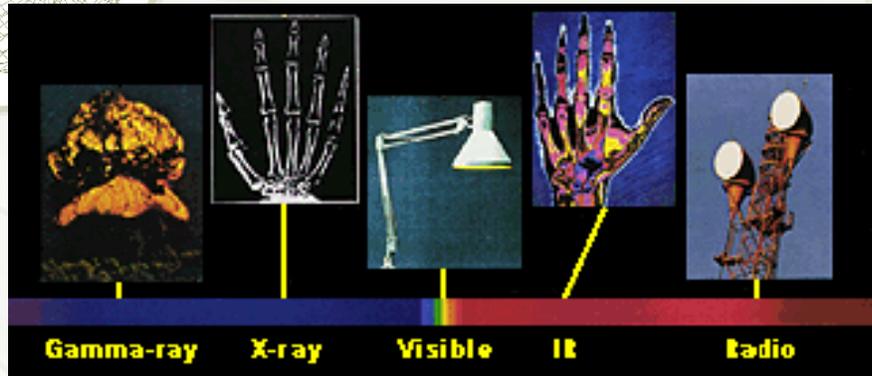
## I : GRAVITATIONAL TIME DILATION

- ★ **Recap of waves:**
- ★ Waves characterized by
  - ★ Wavelength ( $\lambda$ ) = distance between crests
  - ★ Frequency ( $f$  or  $\nu$ ) = number of crests passing a given point per second
- ★ Speed of a crest;  $c = \lambda \nu$
- ★ **Energy** of a wave is proportional to frequency  $\nu$ .



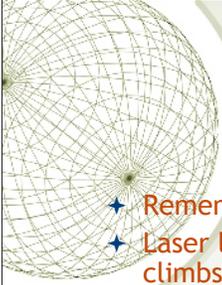


# The electromagnetic spectrum

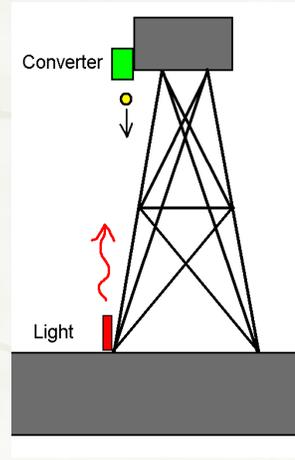


Small wavelength  
**High Frequency**  
**High energy**

Large wavelength  
**Low frequency**  
**Low energy**



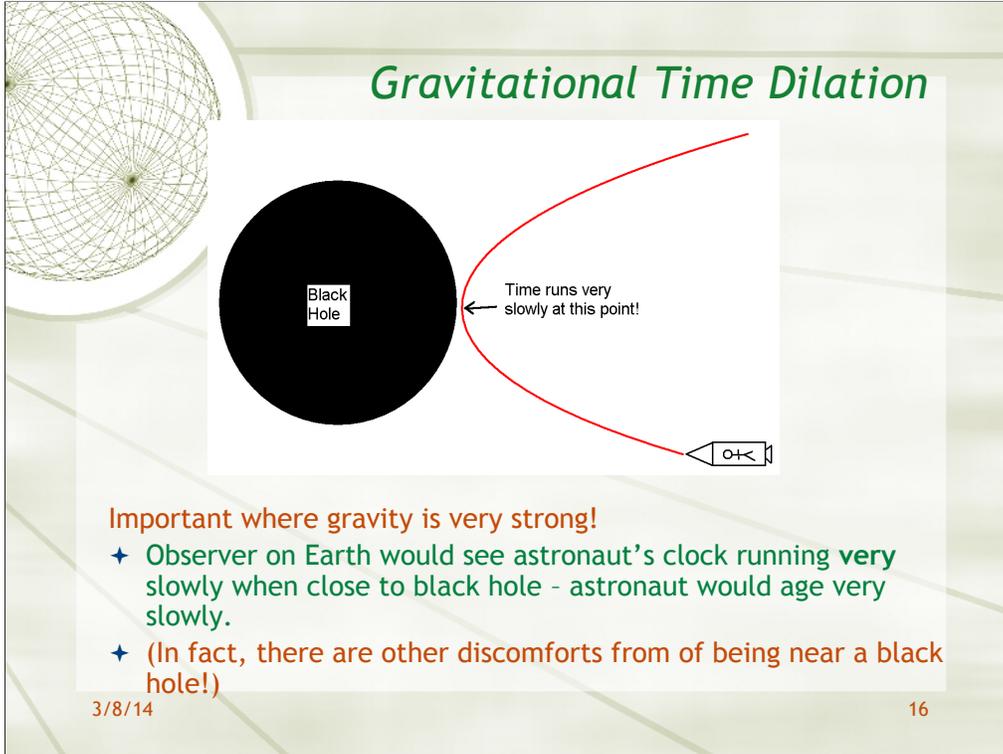
- ★ Remember the tower...
- ★ Laser light must lose energy as it climbs up
  - ★ So...frequency must decrease
  - ★ i.e., light is redshifted.
  - ★ Gravitational redshifting
- ★ Imagine a clock based on frequency of laser light...
  - ★ 1 "tick" = time taken for fixed number of crests to pass
  - ★ Gravitational redshifting slows down the clock.
  - ★ **Clocks in gravitational fields run slowly**



$$t_{grav} \approx \left(1 - \frac{GM}{c^2 r}\right) t_{space}$$

if gravitational field is "weak"

## Gravitational Time Dilation

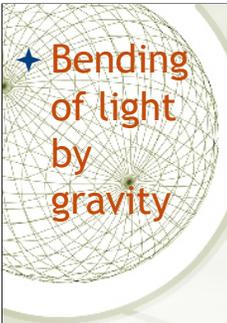


Important where gravity is very strong!

- ✦ Observer on Earth would see astronaut's clock running **very slowly** when close to black hole - astronaut would age very slowly.
- ✦ (In fact, there are other discomforts from of being near a black hole!)

3/8/14

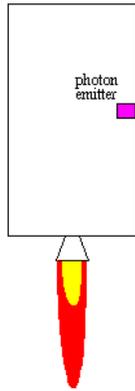
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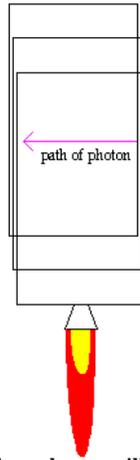
# Bending of light by gravity

## Gravity Bends Light

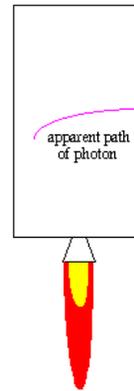
accelerating frame



view from outside



view from inside

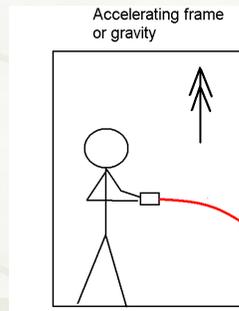
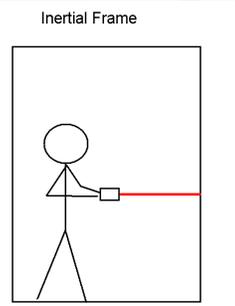


by the equivalence principle, a photon will also "fall" in a gravitational field

<http://abyss.uoregon.edu/~js/cosmo/lectures/lec06.html>

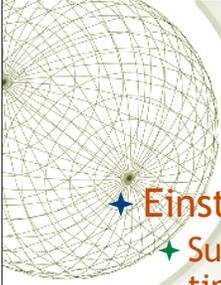
## What about light? It “falls”, too!

- ★ Astronaut in inertial frame with flashlight
  - ★ Inertial frame, so light goes in straight lines
  - ★ It doesn't matter whether this is free fall or far from masses
- ★ What if we now put flashlight in a gravitational field (accelerated frame)?
  - ★ Light beam will bend: it must accelerate at the same rate and direction as the elevator
  - ★ Strong equivalence principle  $\Rightarrow$  frame with gravity acts the same
  - ★ Important conclusion - light “falls” due to gravity!



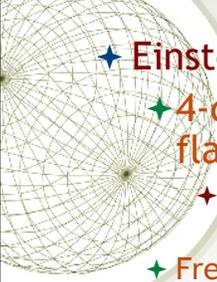
3/8/14

18



## I: CURVED SPACE-TIME

- ★ Einstein pondered several things...
  - ★ Success of Special Relativity showed that space & time were closely linked
  - ★ The “tower thought experiment” suggested that free-fall observers are (locally) free of effects of gravity
  - ★ He wanted to say that gravity was an illusion caused by the fact that we live in an accelerating frame...
  - ★ ... but there is no *single* accelerating frame that works! Somehow, you need to stick together frames of reference that are accelerating in different directions



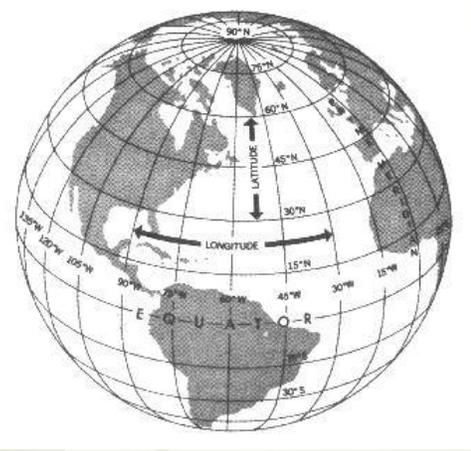
- ★ Einstein's proposal
  - ★ 4-dimensional space-time is “curved,” not flat
    - ★ Example: surface of sphere is curved 2D space; surface of football field is flat 2D space
    - ★ Free-falling objects move on **geodesics** through curved space-time
    - ★ The curvature (bending) of space-time is produced by matter and energy
  - ★ What is a geodesic?
    - ★ The generalization of a straight line in flat space to curved space
    - ★ It is the shortest path between two points on a surface; for instance, the path flown by an aircraft between cities on the globe
    - ★ Unlike straight lines in flat space, geodesics that start as parallel can converge or diverge (or even cross)

3/8/14 20



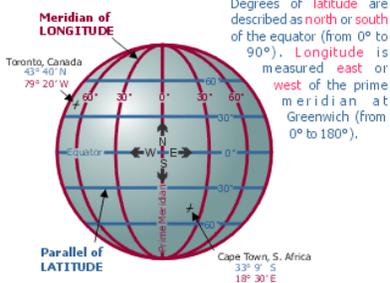


## On Globe...



**Meridian of LONGITUDE**

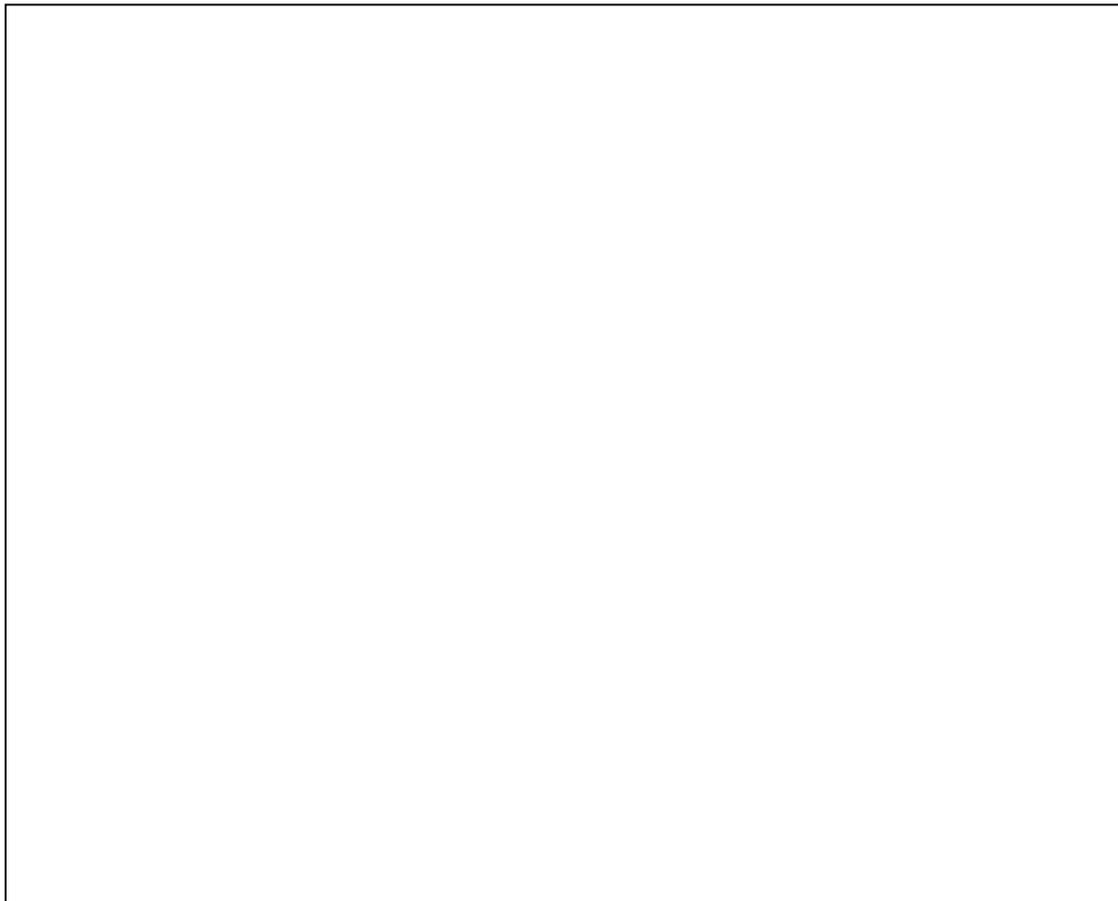
Degrees of latitude are described as north or south of the equator (from 0° to 90°). Longitude is measured east or west of the prime meridian at Greenwich (from 0° to 180°).



**Parallel of LATITUDE**

- ✦ Constant-longitude lines (meridians) are geodesics
- ✦ On the Earth, geodesics are Great Circles, the shortest distance between two points on the surface.
- ✦ Constant-latitude lines (parallels) are not geodesics

3/8/14
22

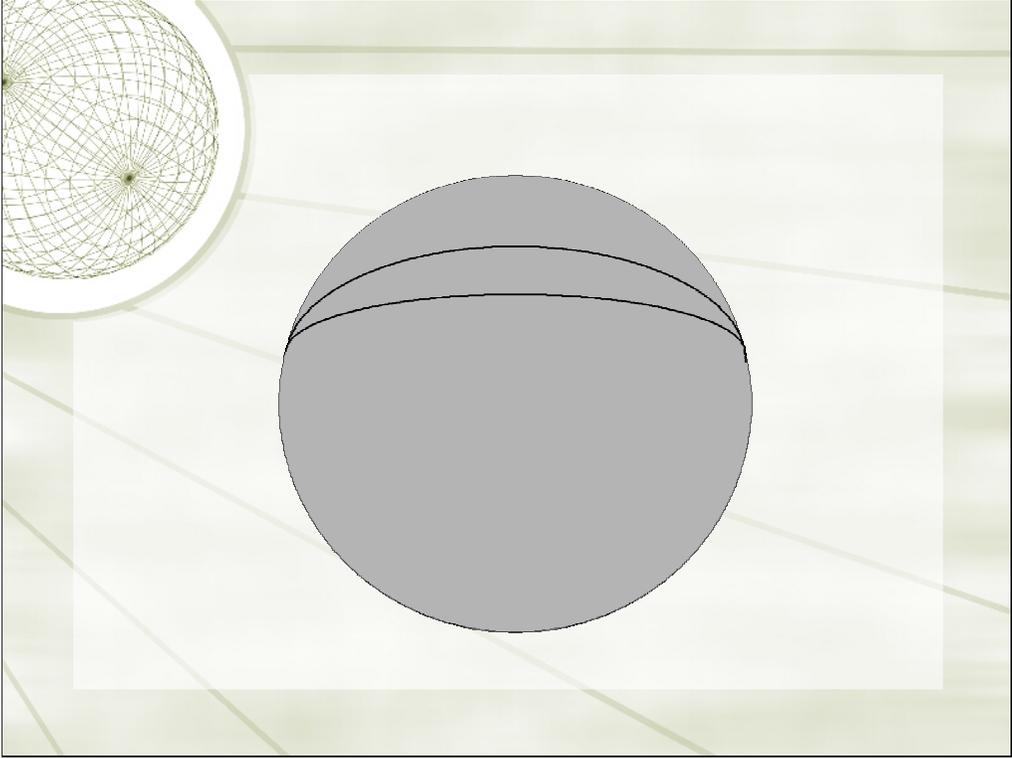


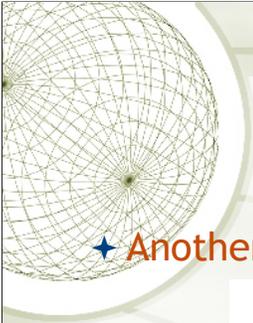
*Shortest flight paths are  
geodesics*



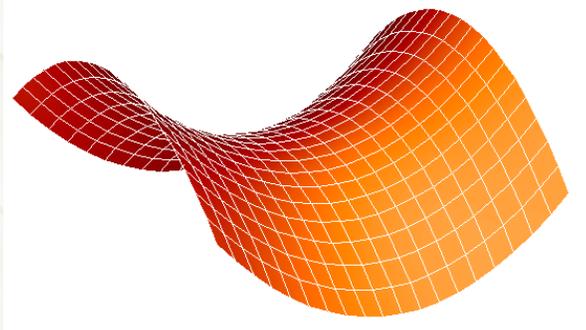
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23

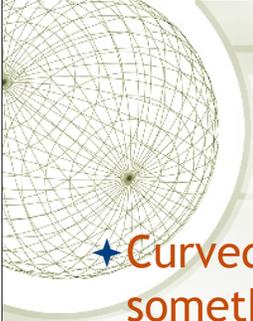




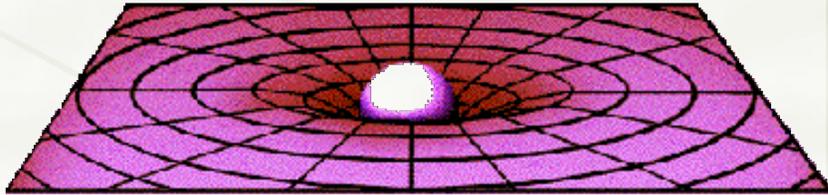
★ Another example - a “saddle”



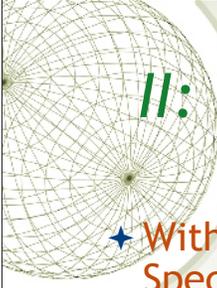
★ Geodesics diverge



★ Curved space around the Earth looks something like this...



From web site of UCSD

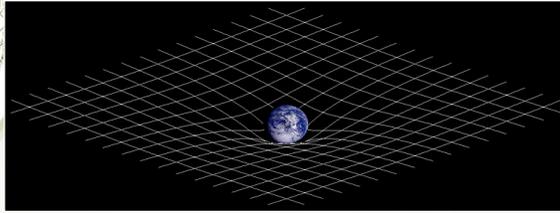
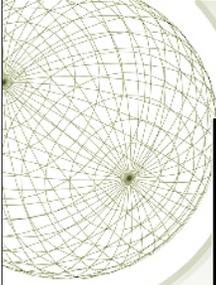


## II: THE GENERAL THEORY OF RELATIVITY

- ★ Within a free-falling (inertial) frame, the Special Theory of Relativity applies.
- ★ Free-falling particles/observers move on geodesics through curved space-time
- ★ The distribution of matter and energy determines how space-time is curved.

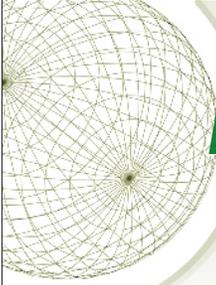
"Space-time curvature tells matter/energy how to move.  
Matter/energy tells space-time how to curve."

## Structure of GR



$$G_{\mu\nu} = 8\pi T_{\mu\nu}$$

- ★ General Relativity (Einstein 1915)
  - ★ Gravity is geometry, not a normal force!
  - ★ Free-fall particles follow “straight-lines” through the 4-d spacetime (Strong Equivalent Principle)...  
I.e., spacetime curvature tells matter how to move
  - ★ Distribution of matter and energy determines curvature of spacetime (Einstein Field Equations).  
I.e. matter/energy tells spacetime how to curve

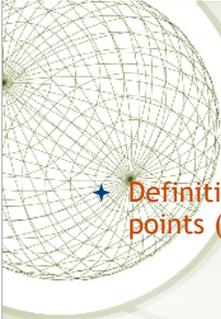


## Einsteins Master Equation

$$\underline{\underline{\mathbf{G}}} = \frac{8\pi G}{c^4} \underline{\underline{\mathbf{T}}}$$

### ★ Notes:

- ★ The Einstein curvature tensor “ $\underline{\underline{\mathbf{G}}}$ ” is mathematical object describing curvature of 4-D space-time.
- ★ The Stress-Energy tensor “ $\underline{\underline{\mathbf{T}}}$ ” is mathematical object describing distribution of mass/energy.
- ★ Newton’s constant of gravitation “ $G$ ” and the speed of light “ $c$ ” appear as fundamental constants in this equation.
- ★ This is actually a horrendous set of 10 coupled non-linear partial differential equations!!
- ★ For weak gravitational fields, this gives Newton’s law of gravitation.



## The Metric Tensor

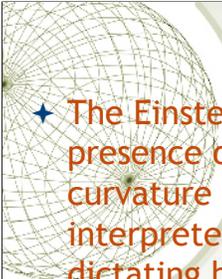
- ★ Definition of (3d) metric tensor  $g_{ij}$  ... the distance between points  $(x_1, x_2, x_3)$  and  $(x_1+\delta x_1, x_2+\delta x_2, x_3+\delta x_3)$  is

$$ds^2 = \sum_{i=1}^3 g_{ij} \delta x_i^2$$

- ★ Can simply generalize this to 4-d spacetime
- ★ Geodesics are those paths which minimize the distance between two points A and B

$$D = \int_A^B ds = \int_A^B \left( \sum_{i=0}^3 g_{ij} dx_i^2 \right)^{1/2}$$

- ★ For a given matter distribution, metric is determined by Einstein's equation (ten coupled partial differential equations)



★ The Einstein GR equations relate the presence of matter and the curvature of spacetime and are interpreted as a set of equations dictating how matter/energy determines the curvature of spacetime..

★ The Einstein field equations are a set of 10 simultaneous, non-linear, partial differential equations.

★ 
$$R_{\mu\nu} = g_{\mu\nu}R + g_{\mu\nu}\Lambda = 8\pi G/c^4 T_{\mu\nu}$$

★  $R_{\mu\nu}$  is the Ricci curvature tensor  $g_{\mu\nu}$  the metric tensor,  $\Lambda$  is the cosmological constant and  $T_{\mu\nu}$  is the stress energy tensor

## GR Metric

An alternate way of writing them is  $G_{\mu\nu} + g_{\mu\nu}\Lambda = 8\pi T_{\mu\nu}$  them where the stuff on the left represents the curvature of space-time and the tensor on the right is matter/energy density

The study of solutions of Einstein's field equations is one of the activities of cosmology. It leads to the prediction of black holes and to different models of evolution of the universe.

## Testing of GR

- ★ The metric of GR (the set of equations describing space-time) is mathematically complex.
- ★ However one can express it approximately as a algebraic series where the first time is 'Newtonian' and the higher order terms represent deviations from Newton's laws.
- ★  $ds^2 = (1 + 2\phi_N + 2\beta\phi_N^2 + \dots)c^2dt^2 - (1 - 2\gamma\phi_N + \dots)dx^2$ .
- ★ (Remember back in Special relativity we defined  $ds^2$  as the space time distance)
- ★ In GR  $\beta, \gamma = 1$  and different values would represent a deviation from GR.
- ★  $\beta, \gamma$  have been measured in the solar system and elsewhere that deviate by  $< 3 \times 10^{-5}$  from 1.