

## Astro 620 Homework 1

- 1) Explain why the intensity or surface brightness of a galaxy (usually measured in magnitudes per square arcsec or watts per square centimeter or equivalent) is nearly independent of its distance (ignore cosmological effects).
  
- 2) What is meant by the fundamental plane of elliptical galaxies? What parameters are involved and how are they correlated. With reference to problem 1 state why the observers used surface brightness as one of the parameters.
  
- 3) The luminosity function of galaxies is described by the Schechter function (S&G eqs 1.24,1.25 or B&T 1.19,1.19)
  - a)  $f(L) = (L/L^*)^{-a} \exp(-L/L^*)$ ; if  $a=1.25$  show that half the light comes from galaxies with  $L > 0.45L^*$  (incomplete gamma function integral... ) see MWB
  - b) What fraction of galaxies are brighter than  $L^*$ ,  $0.1 L^*$ ; what fraction of the light comes from galaxies between  $0.5-1.5 L^*$ ,.
  
- 4) What is the Tully-Fisher relation and what does it imply physically? How does one use it to measure distance. Estimate the distance to a galaxy with velocity width 200km/sec and an apparent magnitude of 14.
  
- 5) Since we are into integrals today describe the Sersic formula for the surface brightness of galaxies (B&T 1.17, MBW eq 2.22)
  - a) what has one to assume to fit this one dimensional formula
  - b) derive eq 2.23 (the Sersic law in magnitudes) for  $n=1$
  - c) Integrate the Sersic profile and determine the half light radius (that radius within which half the light is contained) for  $n=4$ . Explain how this is used in the fundamental plane.