



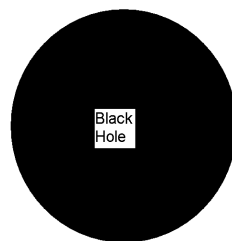
## Lecture 12: General Relativity II

- ★ Recap : gravitational redshift
- ★ Curved spacetime and geodesics
- ★ The General Theory of Relativity
- ★ Gravitational Lensing

3/7/11



## How to live for a 1000 years!



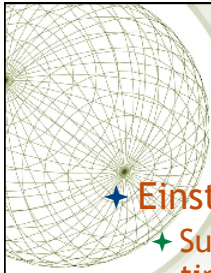
← Time runs very slowly at this point!



- ★ Go where gravity is very strong!
- ★ Observer on Earth would see astronaut's clock running **very** slowly when close to black hole - astronaut would age very slowly.
- ★ (In fact, there are other discomforts from of being near a black hole!)

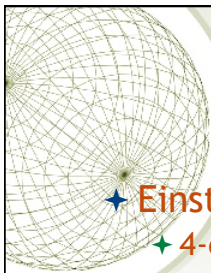
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## I: CURVED SPACE-TIME

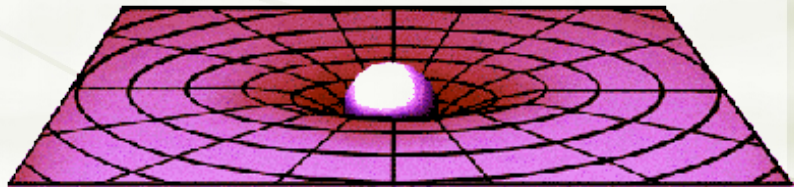
- ★ Einstein pondered several things...
  - ✦ Success of Special Relativity showed that space & time were closely linked
  - ✦ The “tower thought experiment” suggested that free-fall observers are (locally) free of effects of gravity
  - ✦ He wanted to say that gravity was an illusion caused by the fact that we live in an accelerating frame...
  - ✦ ... but there is no *single* accelerating frame that works! Somehow, you need to stick together frames of reference that are accelerating in different directions



- ★ Einstein’s suggestion
  - ✦ 4-dimensional space-time is curved
  - ✦ Free-falling objects move on “geodesics” (generalizations of straight lines) through curved space-time.
    - ✦ Matter and energy causes space-time to bend.
  - ★ What is a geodesic?
    - ✦ Shortest path between two points on a surface
    - ✦ E.g. path flown by aircraft
    - ✦ Geodesics that start parallel can converge or diverge (or even cross).



★ Curved space around the Earth looks something like this...



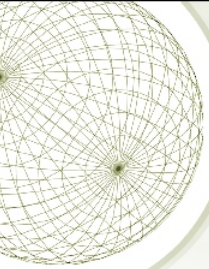
From web site of UCSD



## II: THE GENERAL THEORY OF RELATIVITY

- ★ Within a free-falling frame, the Special Theory of Relativity applies.
- ★ Free-falling particles/observers move on geodesics through curved space-time
- ★ The distribution of matter and energy determines how space-time is curved.

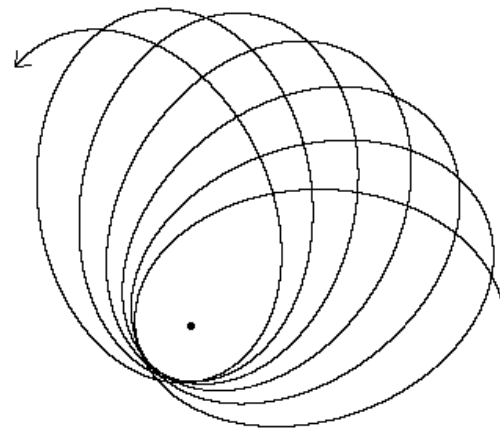
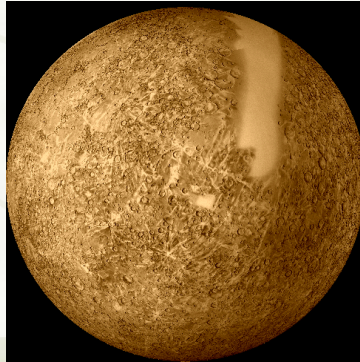
“Space-time curvature tells matter/energy how to move.  
Matter/energy tells space-time how to curve.”

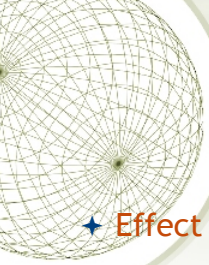

$$\underline{\underline{\mathbf{G}}} = \frac{8\pi G}{c^4} \underline{\underline{\mathbf{T}}}$$

- ★ Notes:
  - ★ The Einstein curvature tensor “ $\underline{\underline{\mathbf{G}}}$ ” is mathematical object describing curvature of 4-D space-time.
  - ★ The Stress-Energy tensor “ $\underline{\underline{\mathbf{T}}}$ ” is mathematical object describing distribution of mass/energy.
  - ★ Newton’s constant of gravitation “ $G$ ” and the speed of light “ $c$ ” appear as fundamental constants in this equation.
  - ★ This is actually a horrendous set of 10 coupled non-linear partial differential equations!!
- ★ For weak gravitational fields, this gives Newton’s law of gravitation.

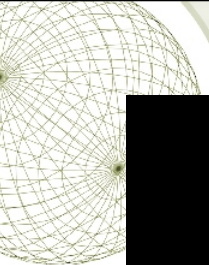
### III: GR EFFECTS IN THE SOLAR SYSTEM

- ★ Have already heard about bending of star light by the Sun (detected by Eddington).
- ★ Orbit of Mercury:

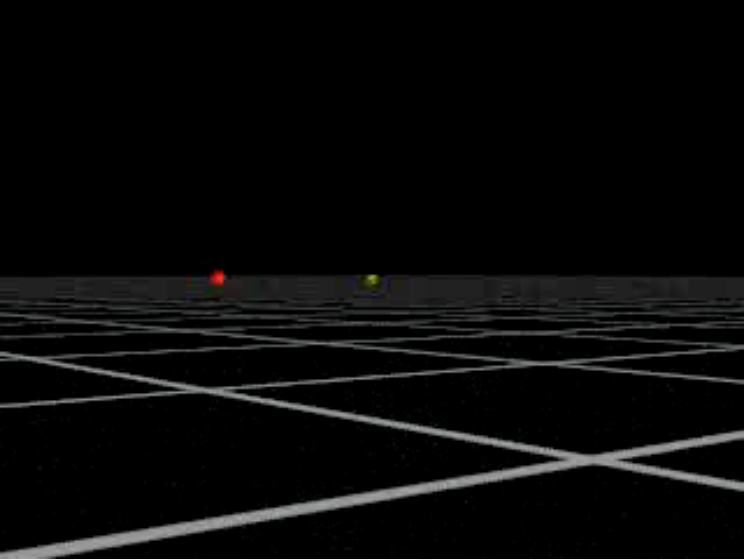




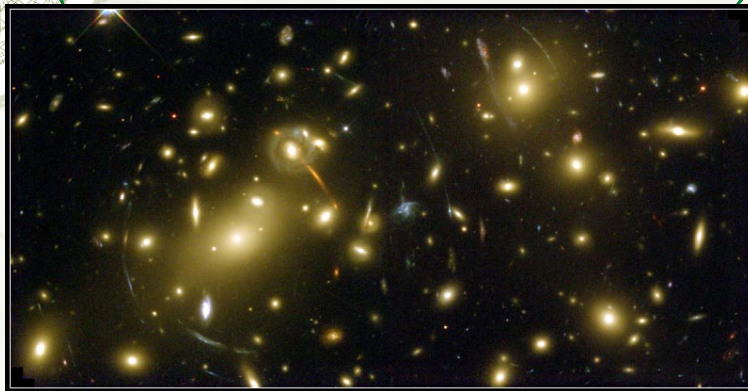
- ★ Effect called “precession of perihelion”.
- ★ Effect small - orbit twists by 5600 arc-seconds (1.56 degrees) per century
  - ★ With Newtonian gravity, can explain 5557 arc-seconds/century as due to
    - ★ Gravitational effect of other planets,
    - ★ deformation of the Sun,
    - ★ non-inertial nature of Earth’s frame
  - ★ But still leaves 43 arc-seconds per century unexplained...
- ★ Using GR, Einstein predicted (with no fiddling!) that Mercury should precess 43 arcseconds per century!



From the web site of  
The University of Oregon



## IV : THE BENDING OF LIGHT (GRAVITATIONAL LENSING)

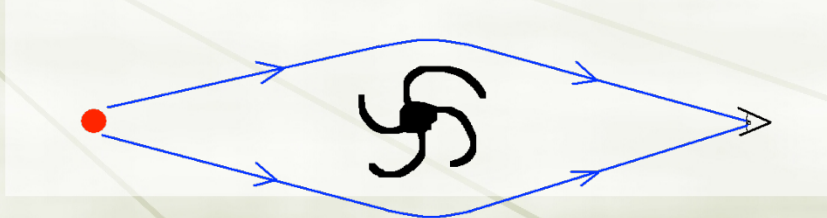
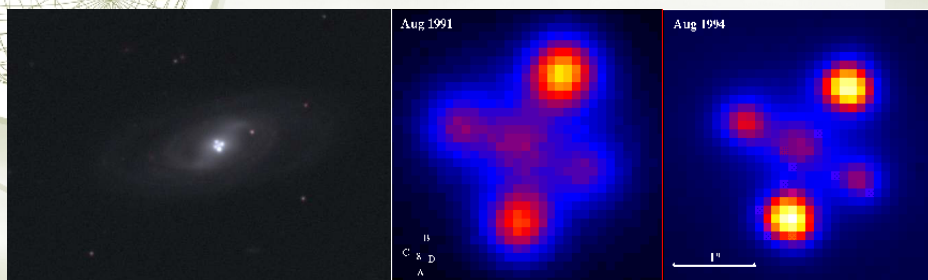


Galaxy Cluster Abell 2218

NASA, A. Fruchter and the ERO Team (STScI) • STScI-PRC00-08

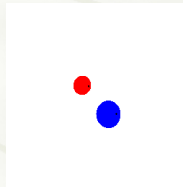
HST • WFPC2

## “The Einstein Cross”



## Gravitational micro-lensing

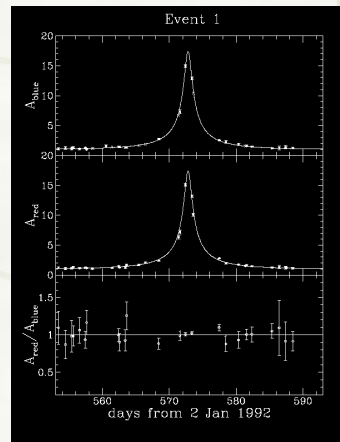
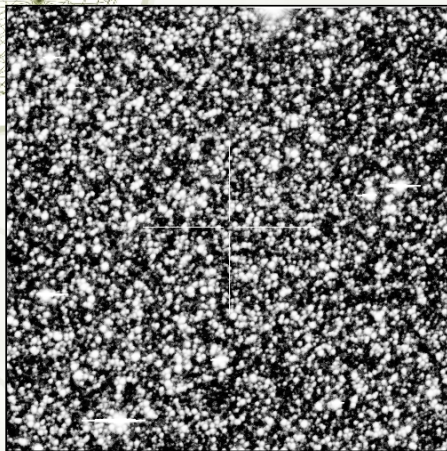
- ✦ Individual stars can also make a gravitational lens... microlensing.
- ✦ Suppose we...
  - ✦ Look at a distant star in our galaxy
  - ✦ Another massive (but dark) star passes in front...



From web site of  
Ned Wright (UCLA)

- ✦ Causes apparent increases in brightness of stellar image

## Really hard to do!



MACHO Project