

Nonlinear Equations

- Often (most of the time??) the relevant system of equations is not linear in the unknowns.
- Then, cannot decompose as $A\mathbf{x} = \mathbf{b}$. Oh well.
- Instead write as:
 - (1) $f(x) = 0$ function of one variable (1-D)
 - (2) $\mathbf{f}(\mathbf{x}) = \mathbf{0}$ $\mathbf{x} = (x_1, x_2, \dots, x_n)$, $\mathbf{f} = (f_1, f_2, \dots, f_n)$ (n -D)
- Not guaranteed to have any real solutions, but generally do for astrophysical problems.

Solutions in 1-D

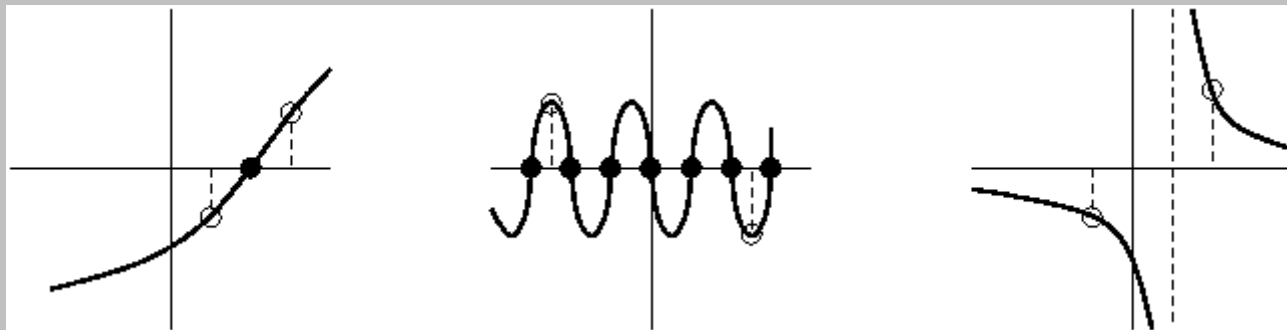
- Generally, solving multi-D equations is much harder, so we'll start with the 1-D case first...
- By writing $f(x) = 0$ we have reduced the problem to solving for the roots of f .
- In 1-D it is usually possible to trap or bracket the desired roots and hunt them down.
- Typically all root-finding methods proceed by iteration, improving a trial solution until some convergence criterion is satisfied.

Function Pathologies

- Before blindly applying a root-finding algorithm to a problem, it is essential that the form of the equation in question be understood: graph it!
- For smooth functions good algorithms will always converge, provided the initial guess is good enough.
- Pathologies include discontinuities, singularities, multiple or very close roots, or no roots at all!

Numerical Root Finding

- Suppose $f(a)$ and $f(b)$ have opposite sign.
- Then, if f is continuous on the interval (a,b) , there must be at least one root between a and b (this is the Intermediate Value Theorem).
- Such roots are said to be bracketed.



Bracketed root

Many roots

No roots

Example Application

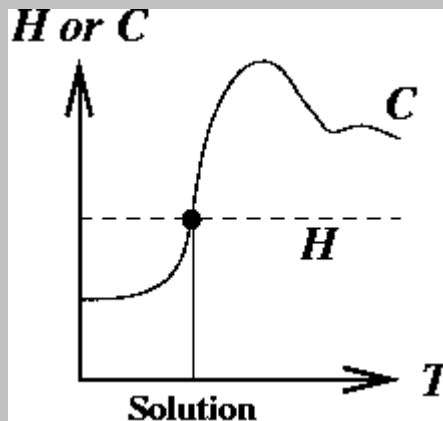
- Use root finding to calculate the equilibrium temperature of the ISM.
- The ISM is a very diffuse plasma.
 - Heated by nearby stars and cosmic rays.
 - Cooled by a variety of processes:
 - Bremsstrahlung: collisions between electrons and ions
 - Atom-electron collisions followed by radiative decay
 - Thermal radiation from dust grains

Example, Cont'd

- Equilibrium temperature given when:

Rate of Heating $H =$ Rate of Cooling C

- Often H is not a function of temperature T .
- Usually C is a complex, nonlinear function of T .



- To solve, find T such that $H - C(T) = 0$.

Bracketing and Bisection

- NRiC 9.1 lists some simple bracketing routines.
- Once bracketed, root is easy to find by bisection:
 - Evaluate f at interval midpoint $(a + b) / 2$.
 - Root must be bracketed by midpoint and whichever a or b gives f of opposite sign.
 - Bracketing interval decreases by 2 each iteration:

$$\epsilon_{n+1} = \epsilon_n / 2.$$

- Hence to achieve error tolerance of ϵ starting from interval of size ϵ_0 requires $n = \log_2(\epsilon_0/\epsilon)$ steps.

Convergence

- Bisection converges linearly (first power of ϵ).
- Methods in which

$$\epsilon_{n+1} = (\text{constant}) \times (\epsilon_n)^m \quad m > 1$$

are said to converge superlinearly.

- Note error actually decreases exponentially for bisection. It converges "linearly" because successive figures are won linearly with computational effort.

Tolerance

- What is a practical tolerance ε for convergence?
- Cannot be less than round-off error.
- For single-precision (`float`) accuracy, typically take $\varepsilon = 10^{-6}$ in fractional error.

i.e.
$$\frac{f(x) - f(x_r)}{f(x_r)} \sim 10^{-6}$$

where x = numerical solution, x_r = actual root.

- When $f(x_r) = 0$ this fails, so use $\varepsilon = 10^{-6}$ as absolute error (or perhaps use $\varepsilon(|a| + |b|)/2$).

Newton-Raphson Method

- Can one do better than linear convergence? Duh!
- Expand $f(x)$ in a Taylor series:

$$f(x + \delta) = f(x) + f'(x) \delta + f''(x) \delta^2/2 + \dots$$

- For $\delta \ll x$, drop higher order terms, so:

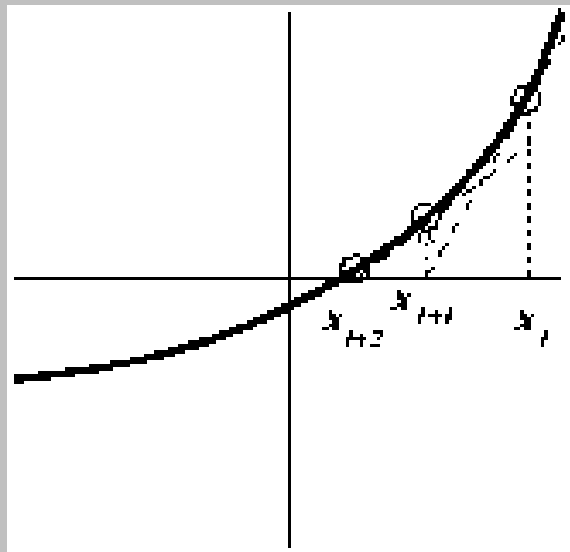
$$f(x + \delta) = 0 \Rightarrow \delta = -f(x) / f'(x)$$

- δ is correction added to current guess of root:

$$\text{i.e. } x_{i+1} = x_i + \delta$$

Newton-Raphson, Cont'd

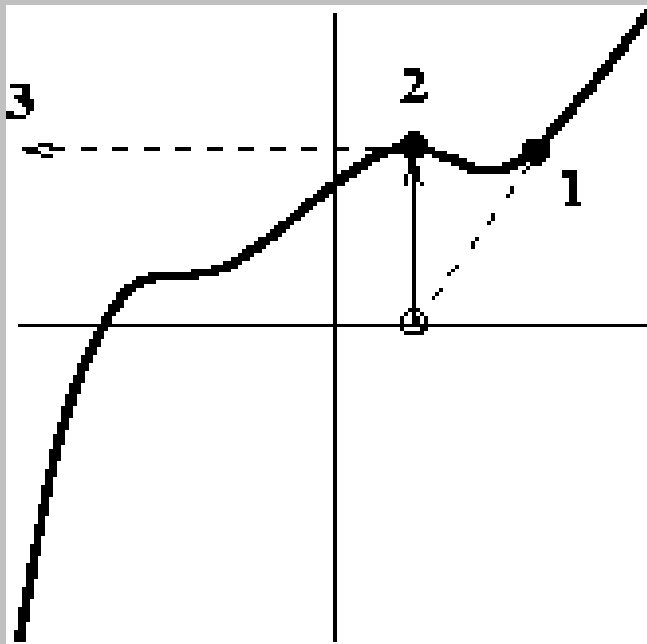
- Graphically, Newton-Raphson (NR) uses tangent line at x_i to find zero crossing, then uses x at zero crossing as next guess:



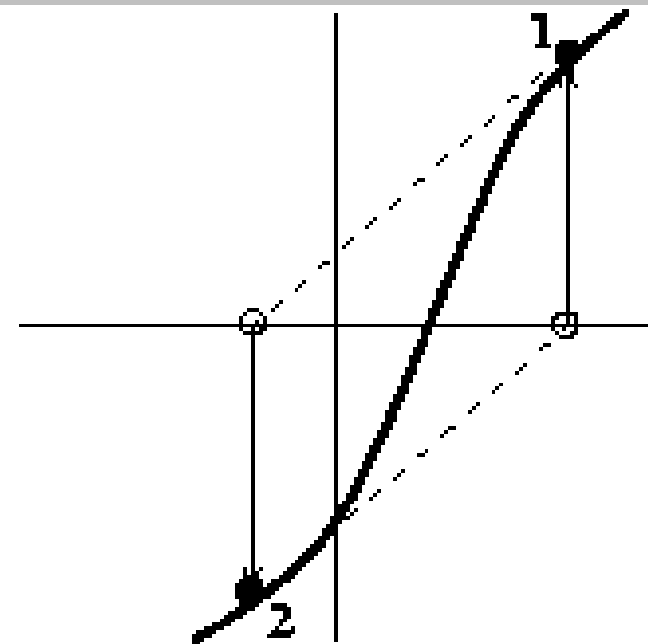
- Note: only works near root ($\delta \ll x$)...

Newton-Raphson, Cont'd

- When higher order terms important, NR fails spectacularly. Other pathologies exist too:



Shoots to infinity



Never converges

Newton-Raphson, Cont'd

- Why use NR if it fails so badly?
- Rate of convergence:

$$\epsilon_{i+1} = \epsilon_i - f(x_i) / f'(x_i)$$

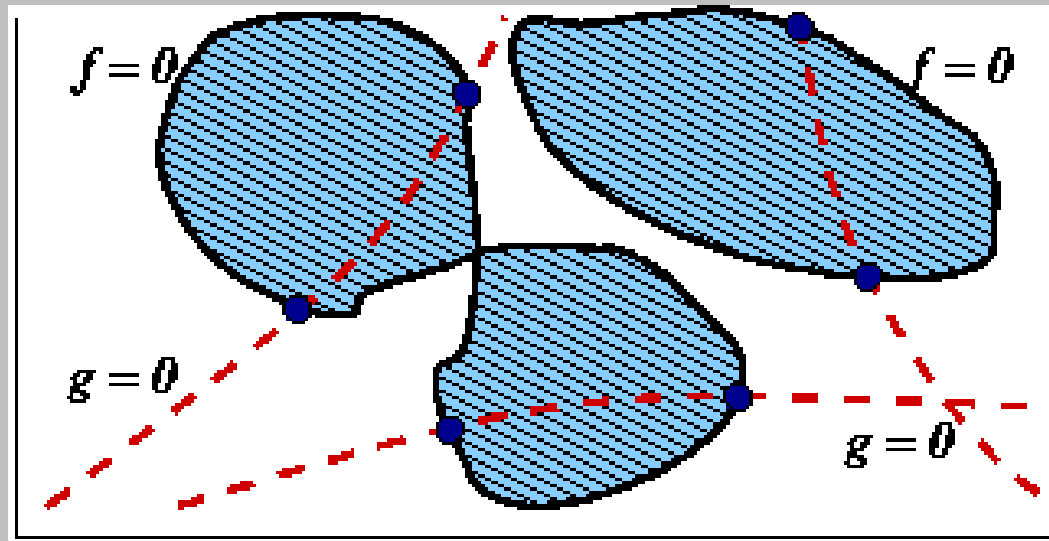
- Taylor expand $f(x_i)$ & $f'(x_i)$ to get:

$$\epsilon_{i+1} = - \epsilon_i^2 f''(x_i) / f'(x_i) \quad [\text{quadratic!}]$$

- Note both $f(x)$ and $f'(x)$ must be evaluated each iteration, plus both must be continuous near root.
- Best use of NR is to "polish-up" bisection root.

Nonlinear Systems of Equations

- Consider the system $f(x,y) = 0$, $g(x,y) = 0$. Plot zero contours of f & g :



- No information in f about g , and vice versa.
 - In general, no good method for finding roots.

Nonlinear Systems, Cont'd

- If you are near root, best bet is NR.

e.g. For $\mathbf{F}(\mathbf{x}) = \mathbf{0}$, choose $\mathbf{x}_{i+1} = \mathbf{x}_i + \delta$, where

$$\mathbf{F}'(\mathbf{x}) \delta = -\mathbf{F}(\mathbf{x})$$

- This is a matrix equation: $\mathbf{F}'(\mathbf{x})$ is a matrix with elements $\partial F_i / \partial x_j$. The matrix is called the Jacobian.
- Written out:

$$\frac{\partial f}{\partial x} \delta_x + \frac{\partial f}{\partial y} \delta_y = -f(x, y)$$

$$\frac{\partial g}{\partial x} \delta_x + \frac{\partial g}{\partial y} \delta_y = -g(x, y)$$

Nonlinear Systems, Cont'd

- Given initial guess, must evaluate matrix elements and RHS, solve system for δ , and compute next iteration \mathbf{x}_{i+1} . Then repeat (must solve 2×2 linear system each time).
- Essentially the non-linear system has been linearized to make it easier to work with.
- NriC 9.7 discusses a global convergence strategy that combines multi-D NR with "backtracking" to improve chances of finding solutions.

Example: Interstellar Chemistry

- ISM is multiphase plasma consisting of electrons, ions, atoms, and molecules.
- Originally, the ISM was thought to be too hostile for molecules.
- But in 1968-69, radio observations discovered absorption/emission lines of NH_3 , H_2CO , H_2O , ...
- Lots of organic molecules, e.g. $\text{CH}_3\text{CH}_2\text{OH}$ (ethanol).

Example, Cont'd

- In some places, all atoms have been incorporated into molecules.
- For example, molecular clouds: dense, cold clouds of gas composed primarily of molecules.
($T \sim 30$ K, $n \sim 10^6$ cm⁻³, $M \sim 10^{5-6} M_{\odot}$, $R \sim 10$ -100 pc).
- How do you predict what abundances of different molecules should be, given n and T ?
- Need to solve a chemical reaction network.

Example, Cont'd

- Consider reaction between two species A and B:



- Reverse also possible:



- In equilibrium:

$$(1) \quad n_A n_B R_{AB} = n_{AB} R'_{AB}$$

$$(2) \quad n_A + n_{AB} = n_A^0 \quad \backslash \text{Normalizations: \#}$$

$$(3) \quad n_B + n_{AB} = n_B^0 \quad / \text{A \& B conserved}$$

Example, Cont'd

- Substitute normalization equations into reaction equations to get quadratic in n_{AB} , easily solved.
- However, many more possible reactions:
 - $AC + B \leftrightarrow AB + C$ (exchange reaction)
 - $ABC \leftrightarrow AB + C$ (dissociation reaction)
- Wind up with large nonlinear system describing all forward/reverse reactions, involving known reaction rates R . Must solve given fixed n^0 & T .

Numerical Derivatives

- For NR and function minimization, often need derivatives of functions. It's always better to use an analytical derivative if it's available.
- If you're stuck, could try:

$$f'(x) \approx \frac{f(x+h) - f(x)}{h}$$

- However, this is very susceptible to RE. Better:

$$f'(x) \approx \frac{f(x+h) - f(x-h)}{2h}$$

- Read NriC 5.7 before trying this!